

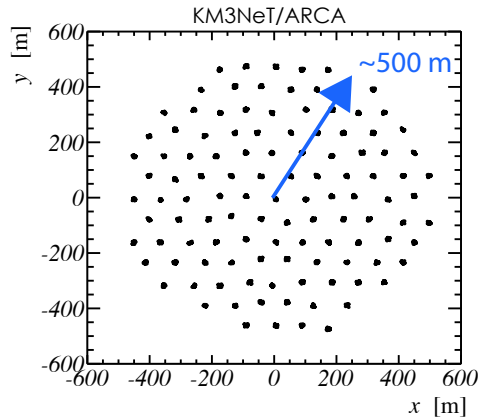
# KM3NeT/ARCA Reconstruction Algorithms

Karel Melis on Behalf of the KM3NeT Collaboration

35<sup>th</sup> ICRC (2017) – Busan - South Korea



# The KM3NeT Detectors

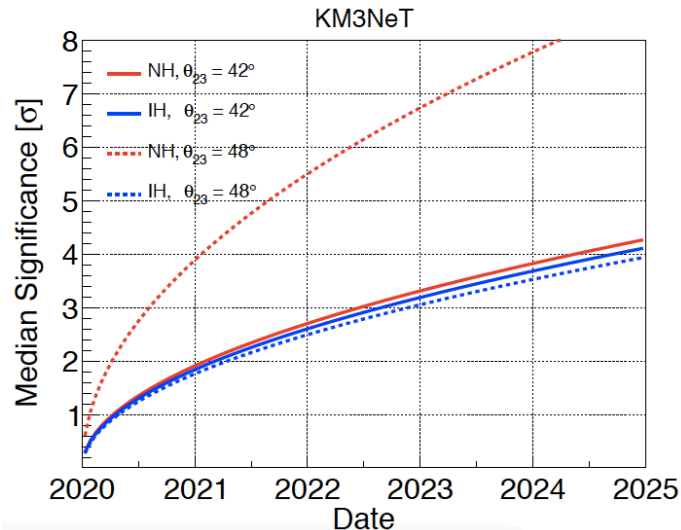


- Digital Optical Module (DOM)
  - 31 x 3-inch PMTs
  - DAQ + Calibration devices
- Detection Unit (DU)
  - 18 DOMs
  - ARCA: DOMs  $\sim 36\text{m}$  apart
  - ORCA: DOMs  $\sim 9\text{m}$  apart
- Building Block
  - 115 DUs
  - ARCA: DUs  $\sim 95\text{m}$  apart
  - ORCA: DUs  $\sim 23\text{m}$  apart

# KM3NeT Science Objectives

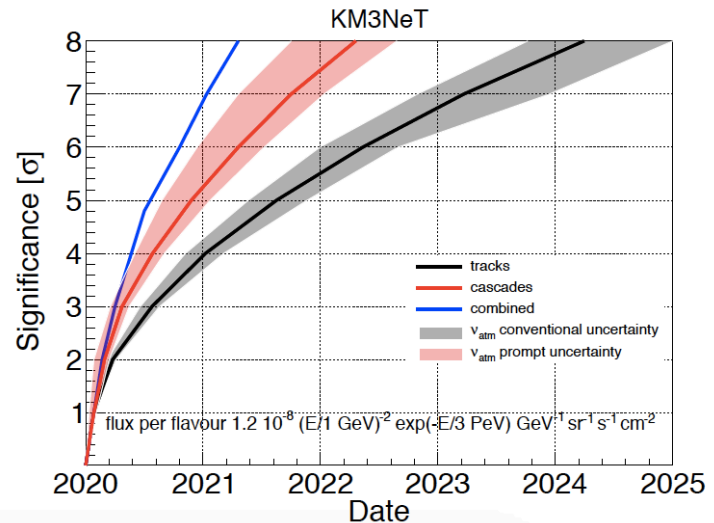
## KM3NeT/ORCA

- Neutrino mass hierarchy
- Low-energy atmospheric neutrinos
- Sensitivity: **3 sigma in 3 years**

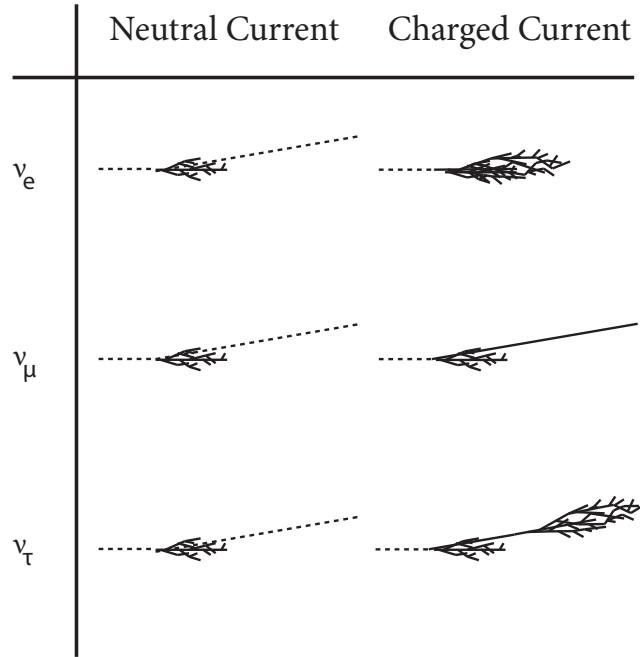


## KM3NeT/ARCA

- Cosmic high-energy neutrino sources
- Diffuse flux
- Sensitivity: **5 sigma in 0.5 year**

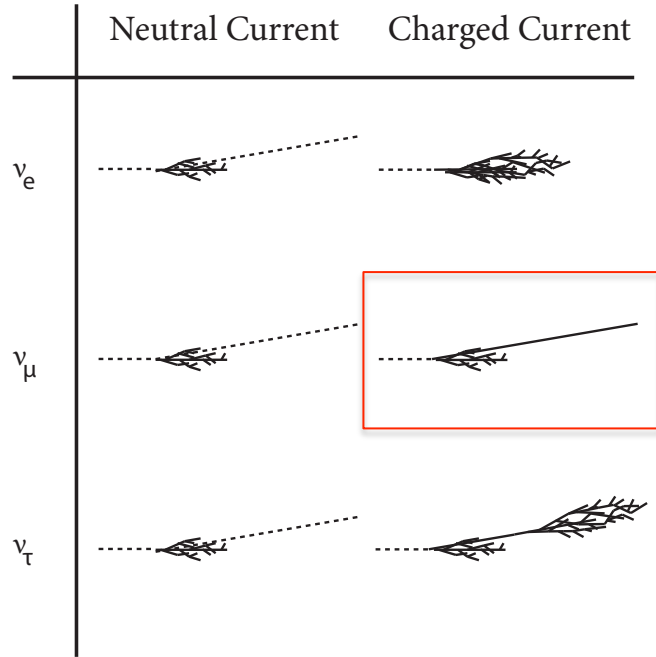


# Event Reconstruction



- Three event topologies:
  - Shower-like
  - Track-like
  - Tau double-bang
- Reconstruction of:
  - Position and time
  - Direction and energy

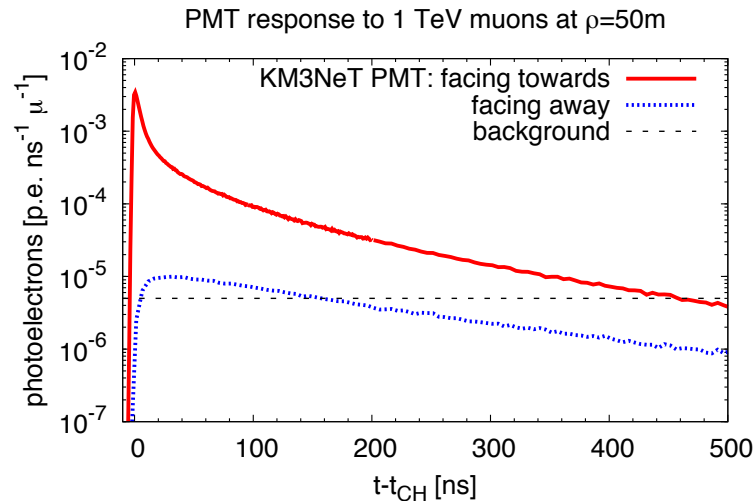
# Event Reconstruction



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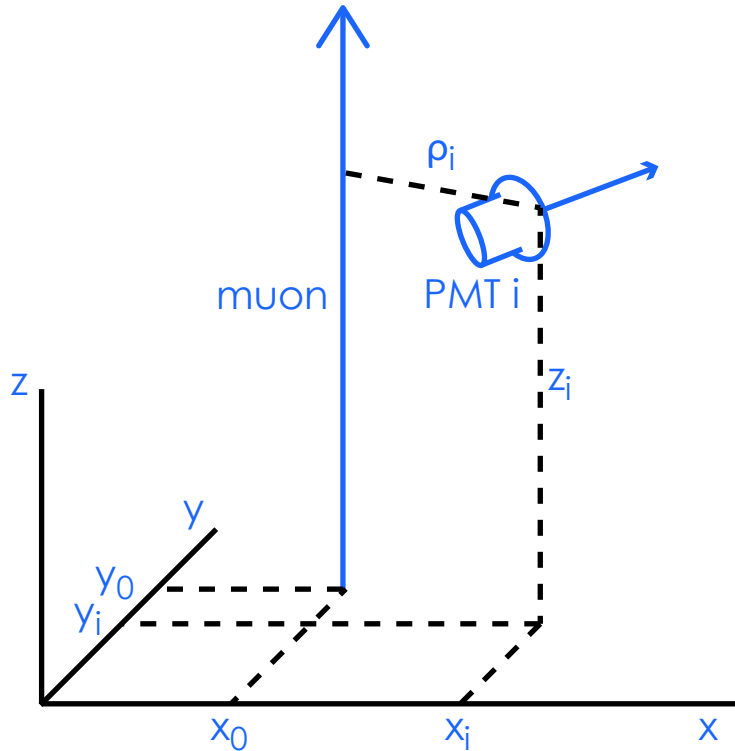
# Track Direction and Energy

$$\mathcal{L} = \prod_{\text{hit PMTs}} \left[ \frac{\partial P}{\partial t}(\rho_i, \theta_i, \phi_i, \Delta t) \right]$$



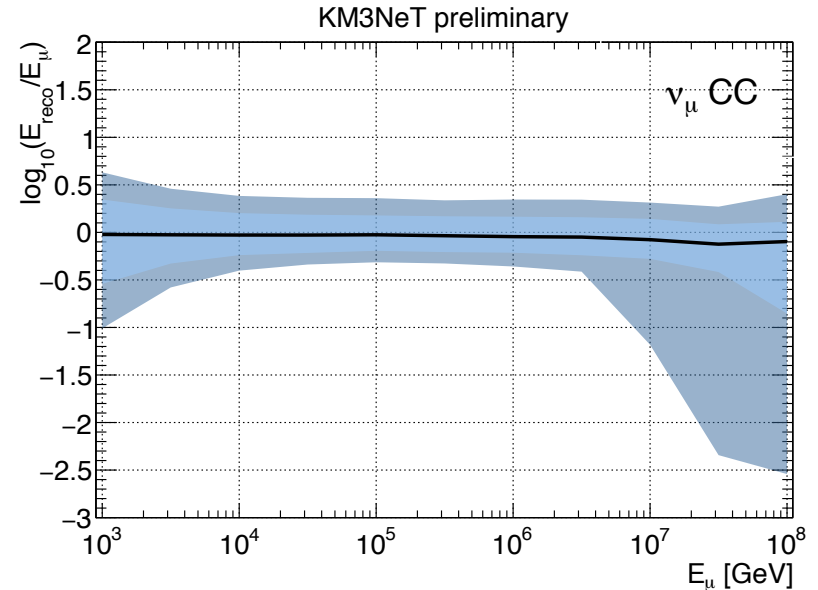
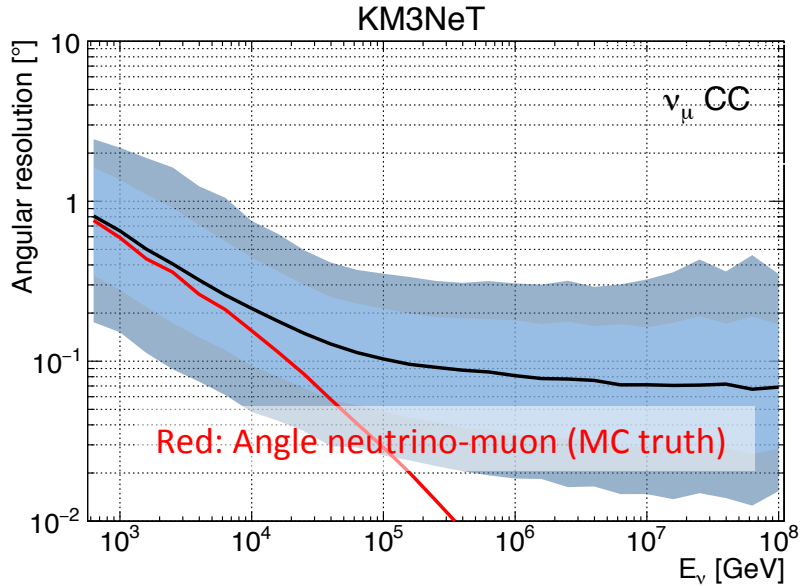
- Data: time of 1<sup>st</sup> hit on PMT
- PDF:
  - Interpolated tables in 4D
  - Includes ionization losses, light propagation, detector response (TTS), optical backgrounds etc.
- **Water ideal medium**
  - Very sharply peaked PDF
  - Very narrow maximum

# Track Prefit



- Provide set of (12) best start values
  - Many sharp local maxima in likelihood landscape
- Scan over assumed track directions
  - Linear  $\chi^2$  fit of track trajectory
- Outlier removal
  - Clustering algorithm
  - One less hit must lead to  $>3$  sigma  $\chi^2$  improvement
- Good start value in 90% of events

# Track Resolutions

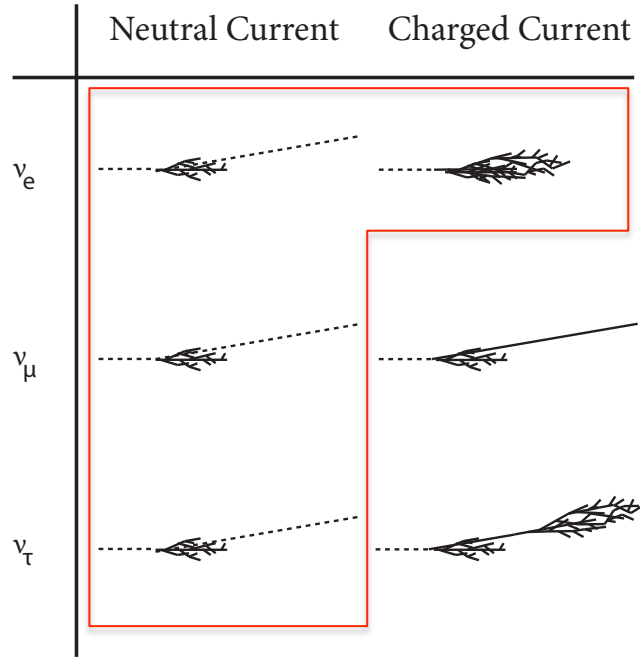


**Direction resolution: < 0.1 degree**

**Energy resolution:  $\sigma < 0.3$  in  $\log_{10}(E_{\text{reco}}/E_{\text{mu}})$**

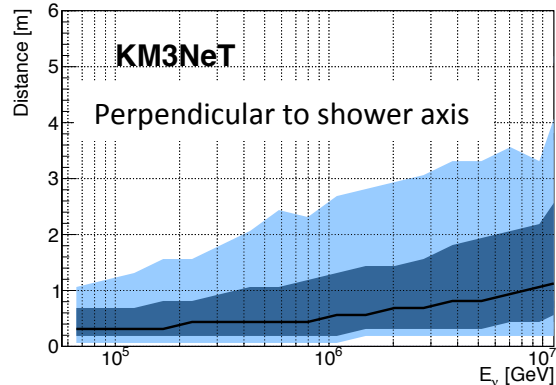
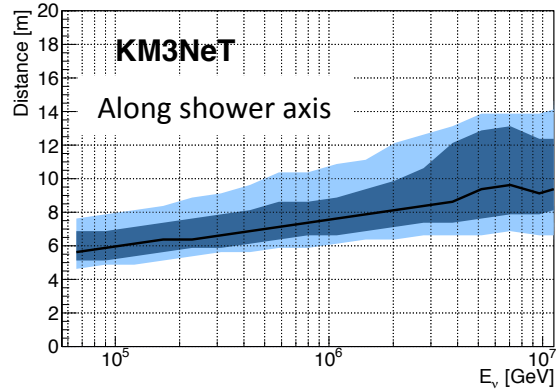


# Event Reconstruction



- Three event topologies:
  - Shower-like
  - Track-like
  - Tau double-bang
- Reconstruction of:
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# Shower Position

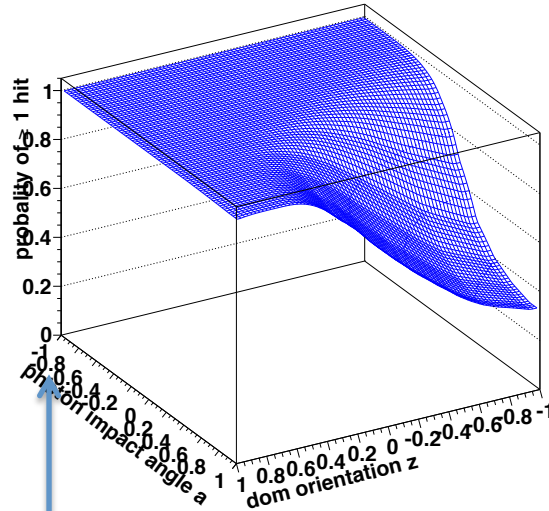


- Data: Hit times of local coincidences in DOM
  - High-purity sample
- Model: Expanding sphere from shower position
- M-estimator score function
  - Background hits further suppressed
- **Vertex Fit Resolution: <1 meter**
  - Shower maximum fitted rather than interaction vertex

# Shower Direction and Energy Fit

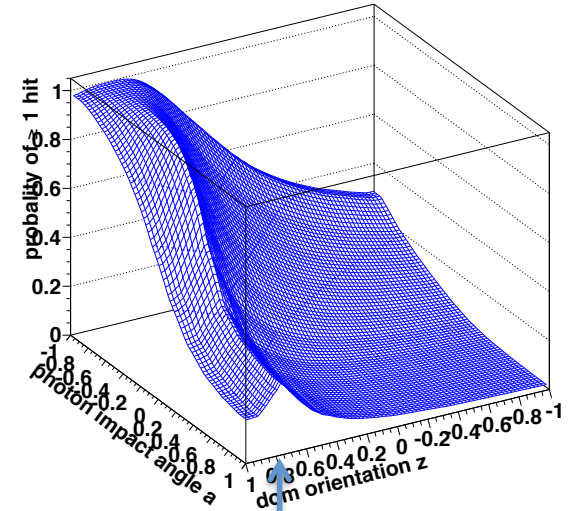
- Fitted vertex as pivot point
- Spatial distribution of hit/not hit PMTs
- Number of photons proportional to shower energy

$r=100\text{ m}$ ,  $E_s=1\text{ PeV}$



PMT facing shower

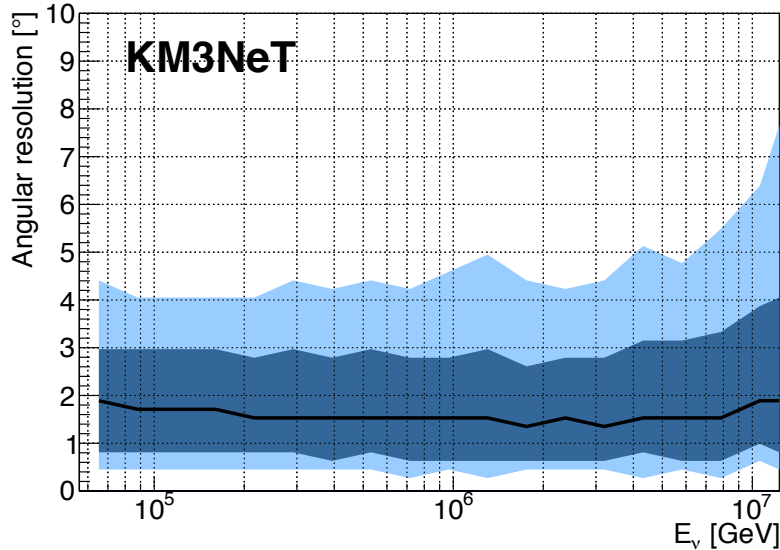
$r=200\text{ m}$ ,  $E_s=1\text{ PeV}$



PMT at Cherenkov angle

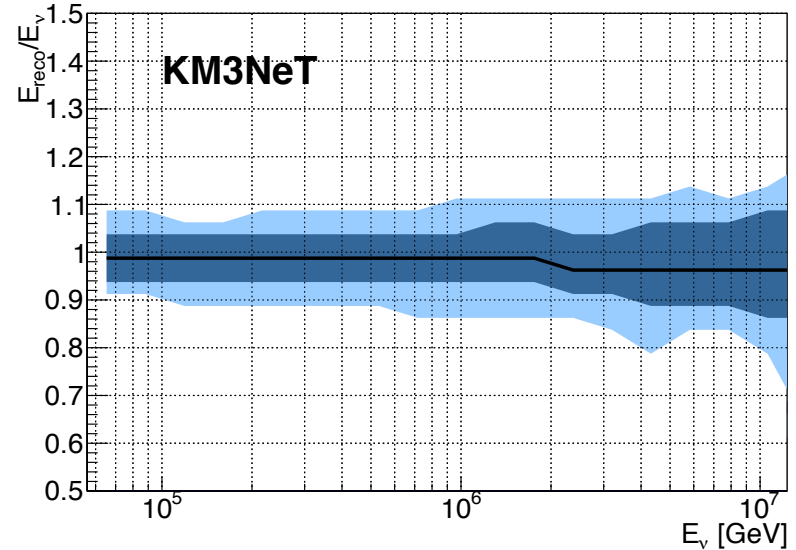
# Shower Resolutions

Ang. resolution vs  $E_\nu$



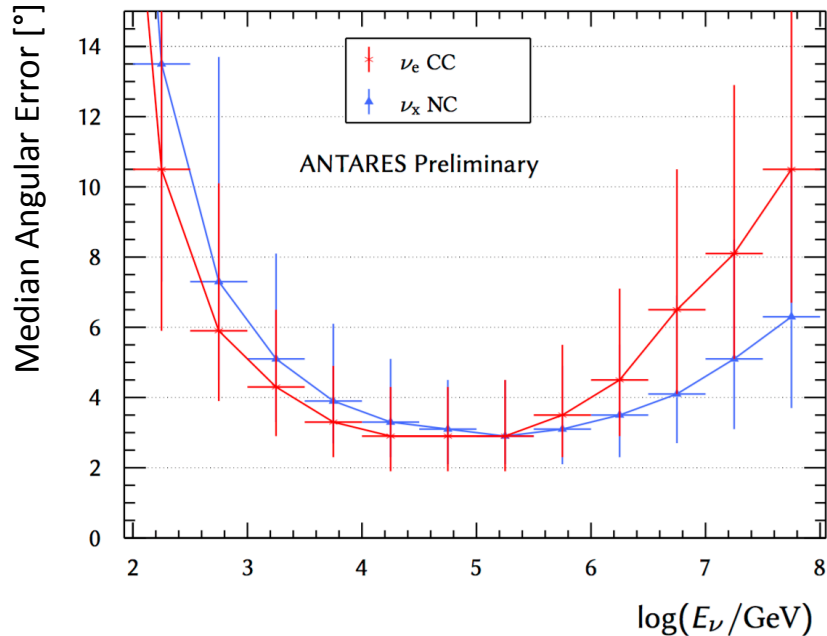
**Direction resolution: ~1.5 degree**

$E_{\text{reco}}/E_\nu$  vs  $E_\nu$



**Energy resolution: ~5%**

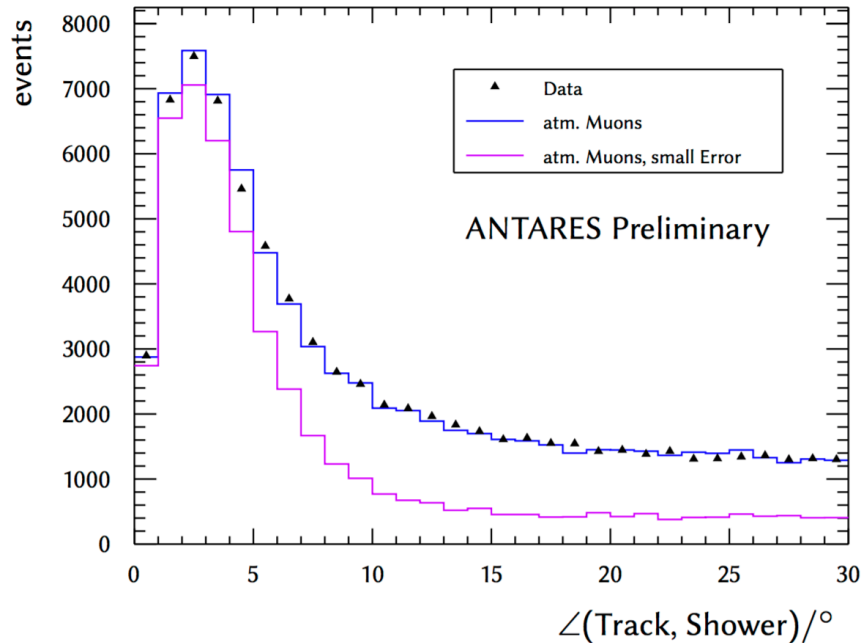
# Showers in Antares



- More or less same algorithm used in Antares
  - Direction resolution:  $\sim 4$  degrees
- Antares
  - Smaller detector
  - Fewer PMTs

Neutrino point source search including cascade events with the ANTARES neutrino telescope, T. Michael, ICRC (2015)

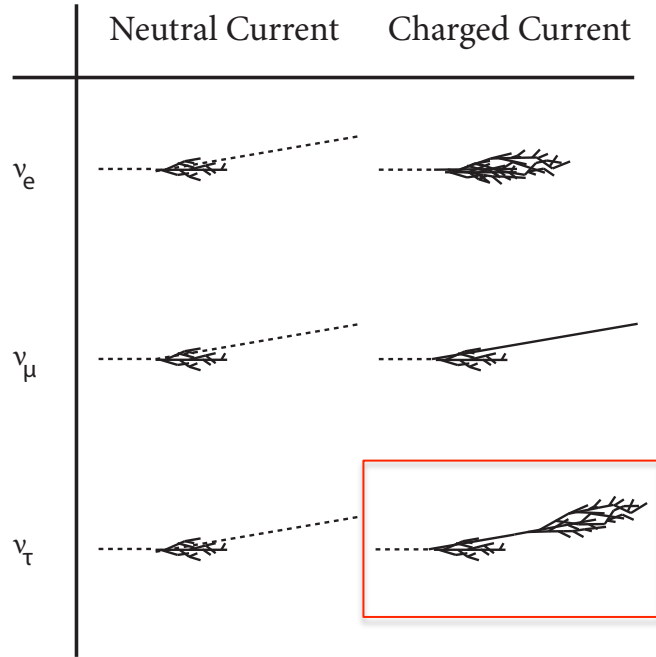
# Showers in Antares



- More or less same algorithm in Antares
  - Direction resolution:  $\sim 4$  degrees
- Reconstructed muon tracks with luminous (Bremsstrahlung) shower along axis
  - Both approximately same direction
- Confirmation of angular resolution

Neutrino point source search including cascade events with the ANTARES neutrino telescope, T. Michael, ICRC (2015)

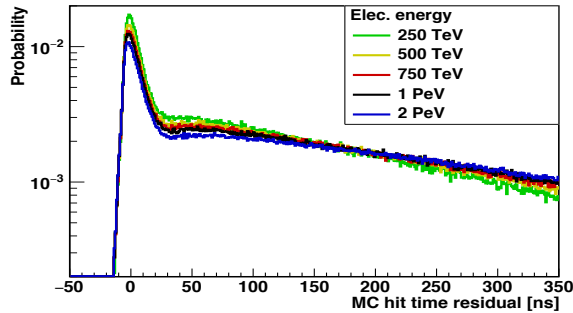
# Event Reconstruction



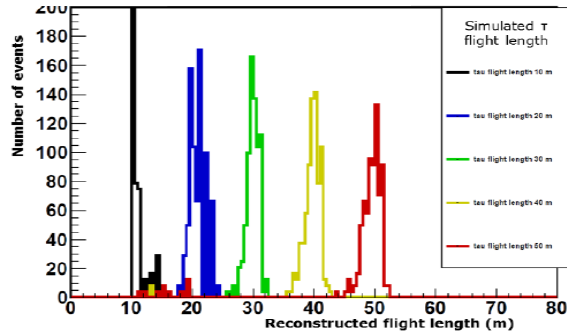
- Three event topologies:
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  - Track-like
  - **Tau double-bang**
- Reconstruction of:
  - Position and time
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# Tau Double Bang Events

PDF:



Tau flight length resolution:



- Study by R. Bormuth with preliminary results
- Clustering hit selection algorithm
  - High purity sample related to single shower
- Single shower fit
  - Vertex + Direction
- Full likelihood fit
  - Data: Hit times of first hits on all PMTs
  - Model: Very simple PDF based on hit times
- Good resolving power of shower positions
  - Expected angular resolution:  $\sim 1$  degree



# Conclusions

- All-flavour reconstruction algorithms tested on Monte Carlo & Antares data
- $5\sigma$  confirmation of IC high-energy diffuse flux in 0.5 year
- Pinpoint cosmic neutrino sources with unprecedented resolutions:
- Muon tracks:
  - Direction:  $< 0.1$  degree,
  - Energy:  $< 0.3$  in  $\log_{10}(E_{\text{reco}}/E_{\text{mu}})$
- Showers:
  - Direction:  $\sim 1.5$  degree
  - Energy:  $\sim 5\%$
- Tau double bang:
  - Preliminary results very promising

