

# Quantum chemistry for fundamental physics

## From the Standard Model to molecular structure theory

Konstantin Gaul

Helmholtz Institut Mainz  
Johannes Gutenberg-Universität Mainz, Germany

**HIM**  
**HELMHOLTZ**  
Helmholtz-Institut Mainz



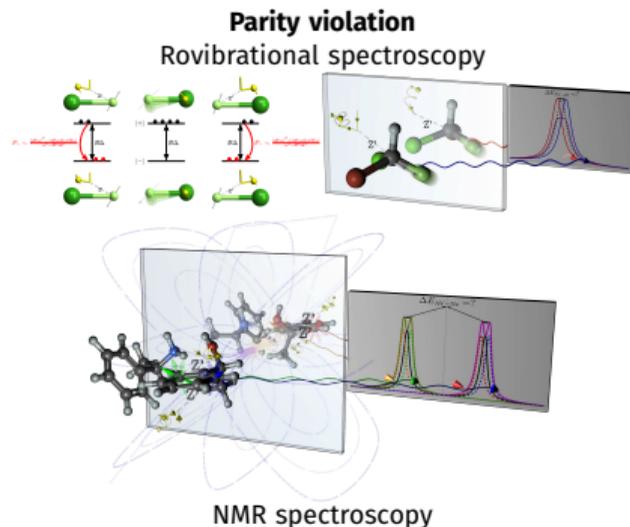
Search for new physics with low-energy precision tests, UG summer school, June 29 to July 4 2025

## The plan

## Today:

### From the Standard Model to molecular structure theory

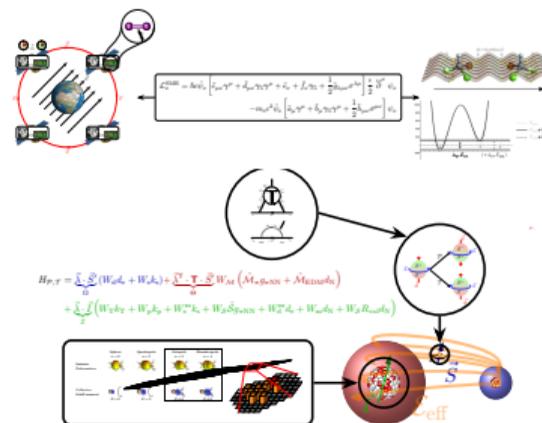
- Standard Model of particle physics, cosmology and beyond
- A crash course in relativistic molecular structure



## Thursday:

### Molecular probes of New Physics

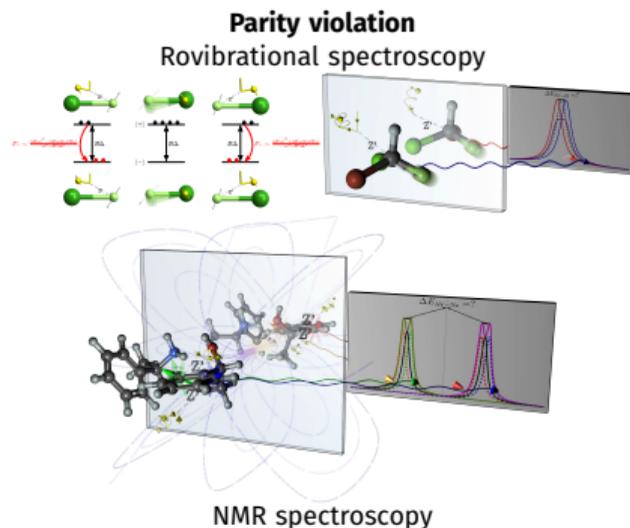
- Molecular parity violation within and beyond the Standard Model
- Lorentz invariance violation
- New bosons/fifth force
- Electric dipole moments in atoms and molecules



## Today:

### From the Standard Model to molecular structure theory

- Standard Model of particle physics, cosmology and beyond
- A crash course in relativistic molecular structure



## Thursday:

### Molecular probes of New Physics

- Molecular parity violation within and beyond the Standard Model
- Lorentz invariance violation
- New bosons/fifth force
- Electric dipole moments in atoms and molecules



## From the Standard Model to molecular structure theory

# The Standard Models of particle physics, cosmology and beyond

# The Standard Models of particle physics, cosmology and beyond

The building blocks of nature: bosons and fermions



Force	rel. Strength	Range/m	Mediator
Gravity	$1 \times 10^{-43}$	$\infty$	graviton?
Weak	$1 \times 10^{-14}$	$1 \times 10^{-18}$	W and Z
Electromagnetic	$1 \times 10^{-3}$	$\infty$	photon
Strong	1	$1 \times 10^{-15}$	gluon

# The Standard Models of particle physics, cosmology and beyond

The building blocks of nature: bosons and fermions



Force	rel. Strength	Range/m	Mediator
Gravity	$1 \times 10^{-43}$	$\infty$	graviton?
Weak	$1 \times 10^{-14}$	$1 \times 10^{-18}$	W and Z
Electromagnetic	$1 \times 10^{-3}$	$\infty$	photon
Strong	1	$1 \times 10^{-15}$	gluon

# The Standard Models of particle physics, cosmology and beyond

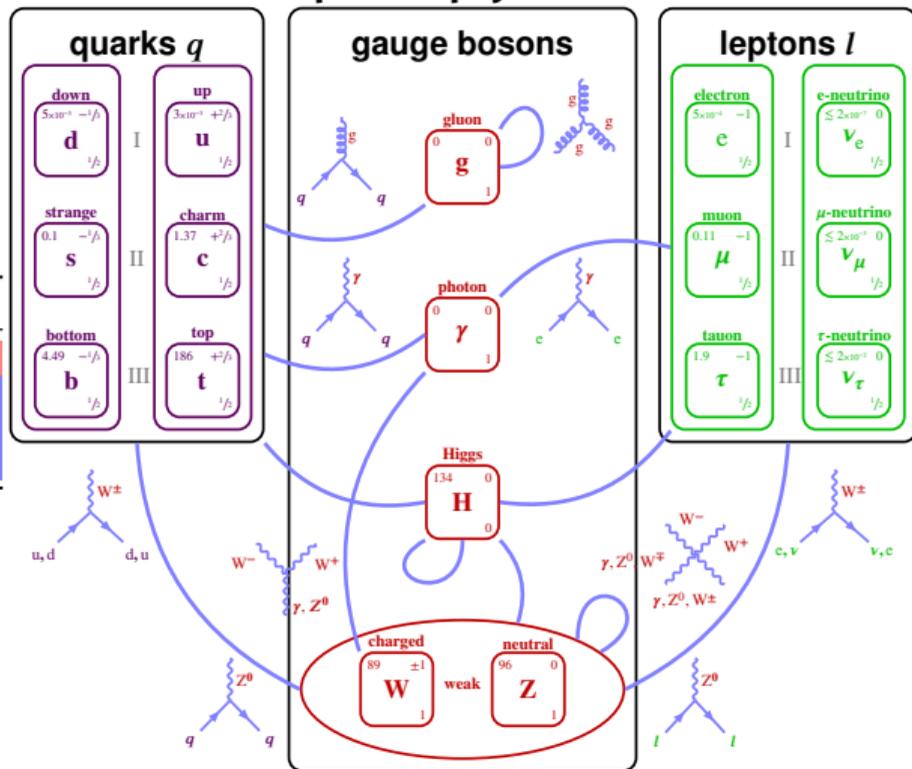
The building blocks of nature: bosons and fermions



Force	rel. Strength	Range/m	Mediator
Gravity	$1 \times 10^{-43}$	$\infty$	graviton?
Weak	$1 \times 10^{-14}$	$1 \times 10^{-18}$	W and Z
Electromagnetic	$1 \times 10^{-3}$	$\infty$	photon
Strong	1	$1 \times 10^{-15}$	gluon

- Standard Model (Quantum field theory)
  - Standard electroweak theory
    - ★ Quantum electrodynamics
    - ★ Fermi's theory of weak interactions
  - Quantum chromodynamics
- General relativity

## Standard model of particle physics



# The Standard Models of particle physics, cosmology and beyond

## Symmetries in the Standard Model

Symmetry	Conserved quantity	Non-observable	SM
<b>Continuous symmetries</b>			
translation in time	energy	absolute time	✓
translation in space	linear momentum	absolute spatial position	✓
rotation	angular momentum	absolute spatial direction	✓
Lorentz symmetry	$CPT$	absolute velocity	✓
<b>Discrete symmetries</b>			
permutation symmetry	Bose-Einstein/Fermi-Dirac statistics	difference between identical particles	✓
inversion in space	parity ( $\mathcal{P}$ )	absolute left/right	✗
inversion in time	time-reversal ( $\mathcal{T}$ )	absolute direction of time	✗
inversion of electric charge	charge conjugation ( $C$ )	absolute sign of electric charge	✗
<b>Unitary symmetries</b>			
$U(1)$ gauge invariance	electric charge $Q$	phase shifts between states of different $Q$	✓
$SU_L(2)$ gauge invariance	weak charge $Q_W$	phase shifts between states of different $Q_W$	✓
$SU(3)$ gauge invariance	color charge $Q_c$	phase shifts between states of different $Q_c$	✓

# The Standard Models of particle physics, cosmology and beyond

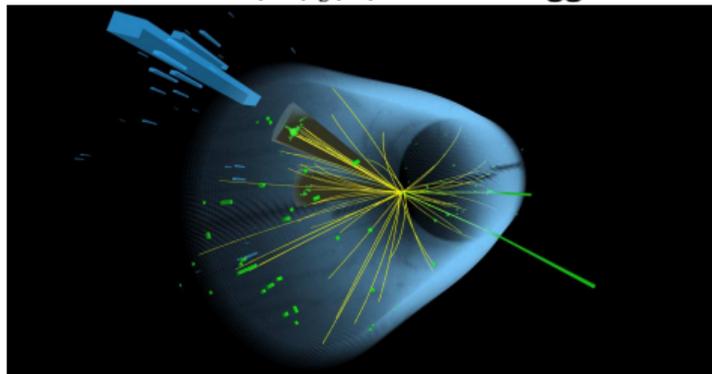
## Symmetries in the Standard Model

Symmetry	Conserved quantity	Non-observable	SM
<b>Continuous symmetries</b>			
translation in time	energy	absolute time	✓
translation in space	linear momentum	absolute spatial position	✓
rotation	angular momentum	absolute spatial direction	✓
<b>Lorentz symmetry</b>	<b><math>CPT</math></b>	<b>absolute velocity</b>	✓
<b>Discrete symmetries</b>			
permutation symmetry	Bose-Einstein/Fermi-Dirac statistics	difference between identical particles	✓
<b>inversion in space</b>	<b>parity (<math>\mathcal{P}</math>)</b>	<b>absolute left/right</b>	✗
<b>inversion in time</b>	<b>time-reversal (<math>\mathcal{T}</math>)</b>	<b>absolute direction of time</b>	✗
<b>inversion of electric charge</b>	<b>charge conjugation (<math>\mathcal{C}</math>)</b>	<b>absolute sign of electric charge</b>	✗
<b>Unitary symmetries</b>			
$U(1)$ gauge invariance	electric charge $Q$	phase shifts between states of different $Q$	✓
$SU_L(2)$ gauge invariance	weak charge $Q_W$	phase shifts between states of different $Q_W$	✓
$SU(3)$ gauge invariance	color charge $Q_c$	phase shifts between states of different $Q_c$	✓

# The Standard Models of particle physics, cosmology and beyond

The SM: An impressive success story

- The Standard Model agrees with experimental data to highest precision in history of physics!
- Prediction of  $W$ ,  $Z$ ,  $g$ ,  $t$ ,  $c$ ... and Higgs boson!



CERN

- Most precise predictions in QED

- ▶ Anomalous magnetic dipole moment of the electron

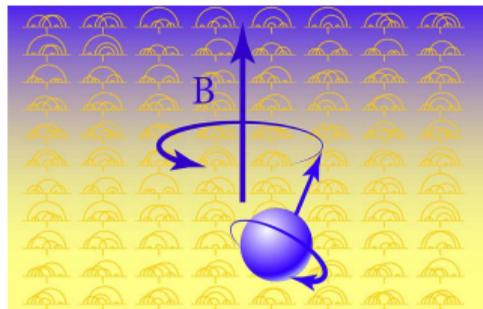
$$a_e = (g - 2)/2:$$

$$a_e(\text{exp}) = 0.00115965218059(13)$$

$$a_e(\text{theo}) = 0.00115965218178(77)$$

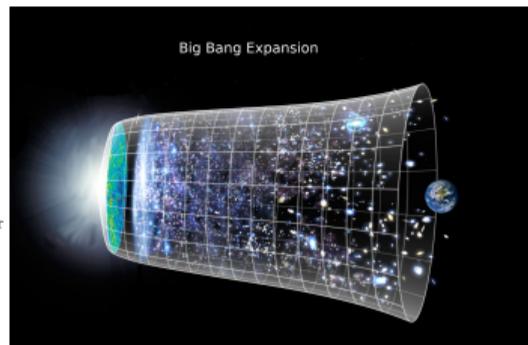
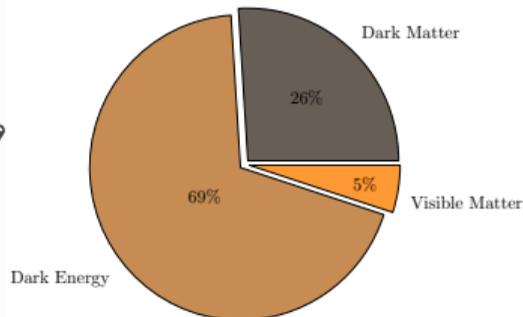
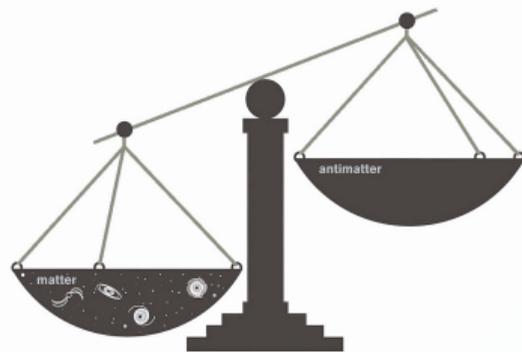
1 part in  $10^{12}$  agreement!

- ▶ Other magnetic dipole moments (muon, tauon, etc.)
- ▶ Lifetime of muon
- ▶ Many more...



# The Standard Models of particle physics, cosmology and beyond

## Limitations of the Standard Model



### Unexplained phenomena:

- Gravity
- Dark matter and dark energy (Standard Model of cosmology)
- Inflation (flatness and isotropy of the universe)
- Baryon asymmetry (imbalance of matter and antimatter in the observed universe)
- Neutrino oscillations (extension of the Standard Model)

### “Discrepancies”:

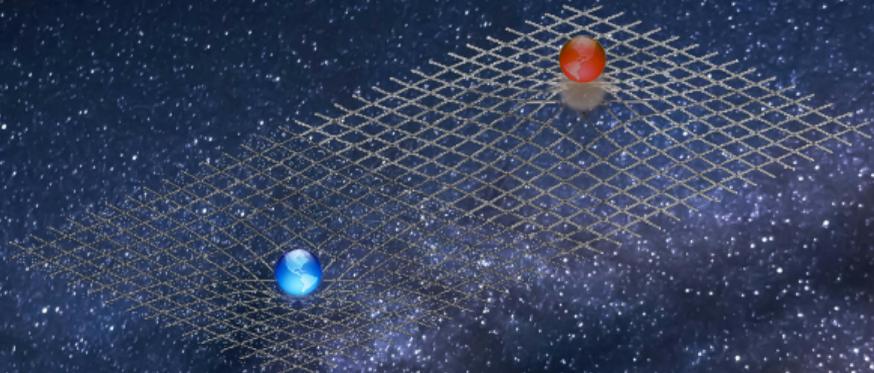
“Neutron lifetime puzzle”, B meson decay, missing hadrons and “glueballs”, sterile neutrinos, muon- $\mu$ , proton radius

### Theoretical/conceptual “problems”:

Empirical formulas (Koide, CKM mixing matrix, masses and Higgs vacuum expectation value), solving QCD, hierarchy problem, number of parameters (19/28), quantum triviality, strong  $CP$  “problem”, etc.

# The Standard Models of particle physics, cosmology and beyond

Einstein's General Relativity



"Size of curvature"

Measure of deformation  
along shortest path

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = 8\pi GT_{\mu\nu}$$

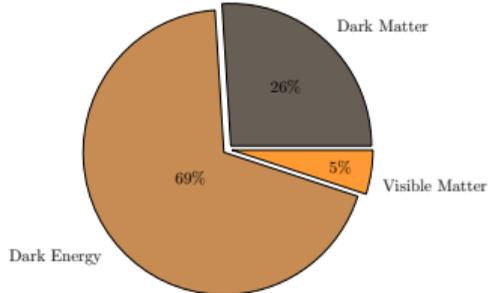
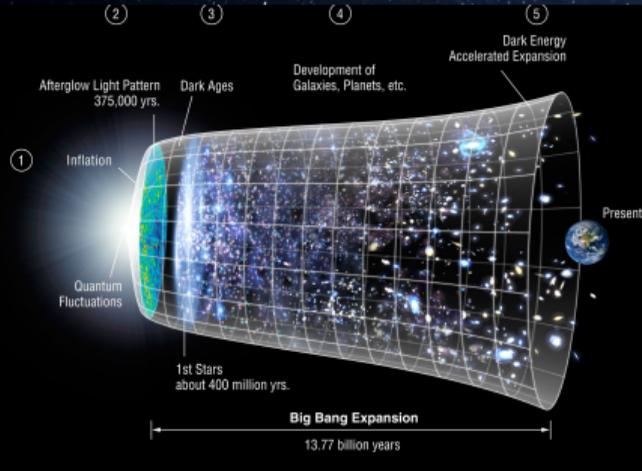
energy/momentum  
density and flux

Metric tensor

⇒ Gravity is a **pseudo-force**

# The Standard Models of particle physics, cosmology and beyond

## $\Lambda$ CDM: The standard model of cosmology

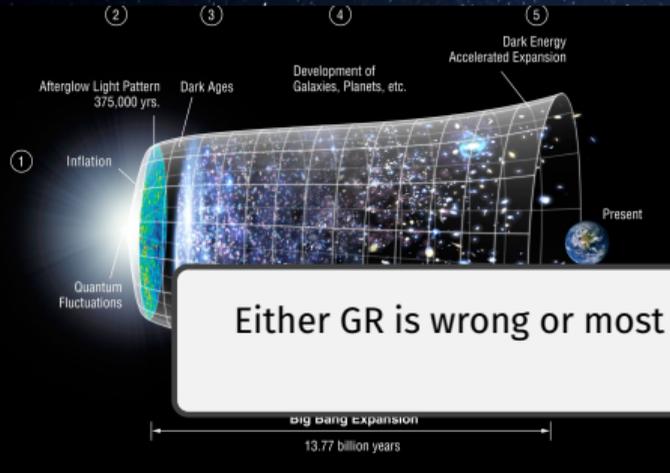


- In GR **absolute energy** matters!
- Constant vacuum energy density  $\rho_{\text{vac}}$  adds to  $T_{\mu\nu}$
- $\Lambda = \frac{\rho_{\text{vac}}}{8\pi G}$  is the cosmological constant
- Our universe is fairly flat and matter dominated:  
 $1 - \Omega_c \approx \Omega_\Lambda + \Omega_M$
- From the Friedmann equation<sup>a</sup>
- Matter- (M), radiation- (R) and vacuum ( $\Lambda$ ) density parameters today:  
 $\Omega_\Lambda \propto \Omega_c a^2 \propto \Omega_M a^3 \propto \Omega_M a^4$
- Observation  $\Omega_c \approx 0$ ,  $\Omega_M = 0.3 \pm 0.1$ ,  $\Omega_\Lambda = 0.7 \pm 0.01$   
 $\leftrightarrow$  Best estimates of baryon density  $\Omega_b = 0.04 \pm 0.02!$

<sup>a</sup>S. M. Carroll, Spacetime and Geometry: An Introduction to General Relativity, Cambridge University Press, 2019.

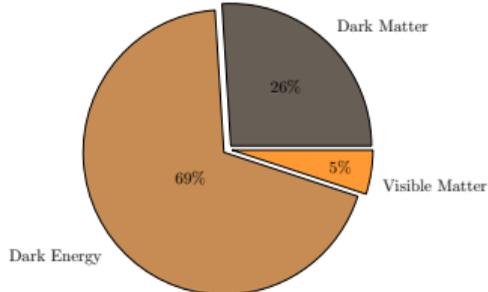
# The Standard Models of particle physics, cosmology and beyond

$\Lambda$ CDM: The standard model of cosmology



- In GR **absolute energy** matters!
- Constant vacuum energy density  $\rho_{\text{vac}}$  adds to  $T_{\mu\nu}$
- $\Lambda = \frac{\rho_{\text{vac}}}{8\pi G}$  is the cosmological constant
- Our universe is fairly flat and matter dominated:

Either GR is wrong or most of the matter content of the universe is dark!



- Matter- ( $\Omega_M$ ), radiation- ( $\Omega_R$ ) and vacuum ( $\Omega_\Lambda$ ) density parameters today:  
 $\Omega_\Lambda \propto \Omega_c a^2 \propto \Omega_M a^3 \propto \Omega_M a^4$
- Observation  $\Omega_c \approx 0$ ,  $\Omega_M = 0.3 \pm 0.1$ ,  $\Omega_\Lambda = 0.7 \pm 0.01$   
 $\leftrightarrow$  Best estimates of baryon density  $\Omega_b = 0.04 \pm 0.02!$

<sup>a</sup>S. M. Carroll, Spacetime and Geometry: An Introduction to General Relativity, Cambridge University Press, 2019.

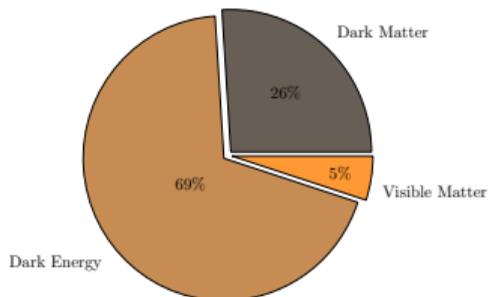
# The Standard Models of particle physics, cosmology and beyond

$\Lambda$ CDM: The standard model of cosmology



- In GR **absolute energy** matters!
- Constant vacuum energy density  $\rho_{\text{vac}}$  adds to  $T_{\mu\nu}$
- $\Lambda = \frac{\rho_{\text{vac}}}{8\pi G}$  is the cosmological constant
- Our universe is fairly flat and matter dominated:

Either GR is wrong or most of the matter content of the universe is dark!  
Modified Newtonian Dynamics (MOND)? →  
Not able to explain gravitational lensing, waves...

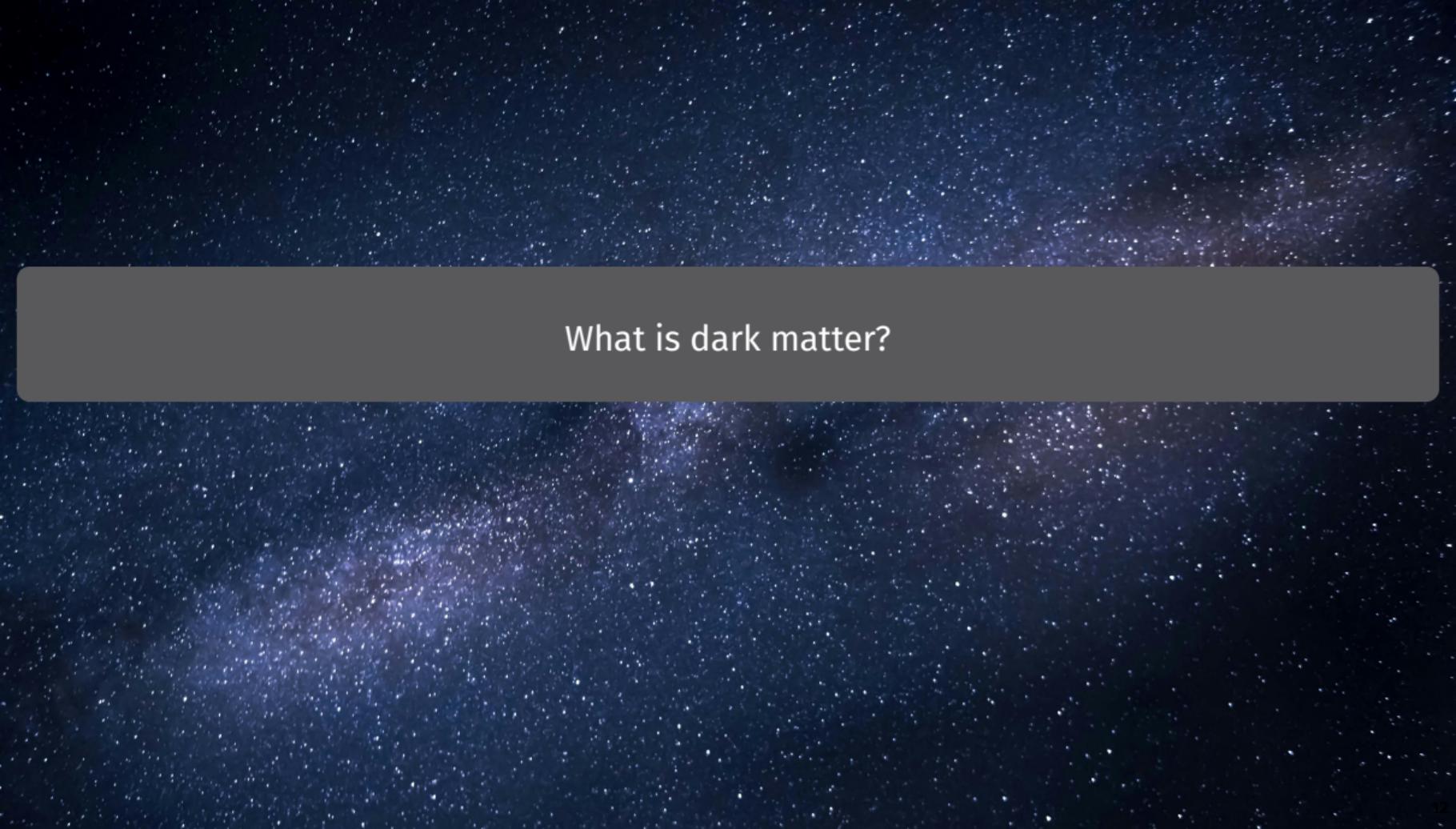


parameters today.

$$\Omega_{\Lambda} \propto \Omega_c a^2 \propto \Omega_M a^3 \propto \Omega_M a^4$$

- Observation  $\Omega_c \approx 0$ ,  $\Omega_M = 0.3 \pm 0.1$ ,  $\Omega_{\Lambda} = 0.7 \pm 0.01$   
↔ Best estimates of baryon density  $\Omega_b = 0.04 \pm 0.02!$

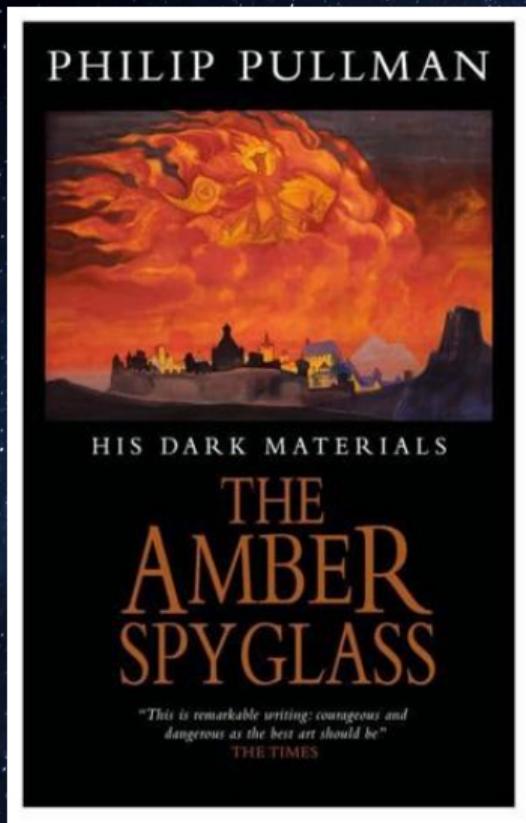
<sup>a</sup>S. M. Carroll, Spacetime and Geometry: An Introduction to General Relativity, Cambridge University Press, 2019.



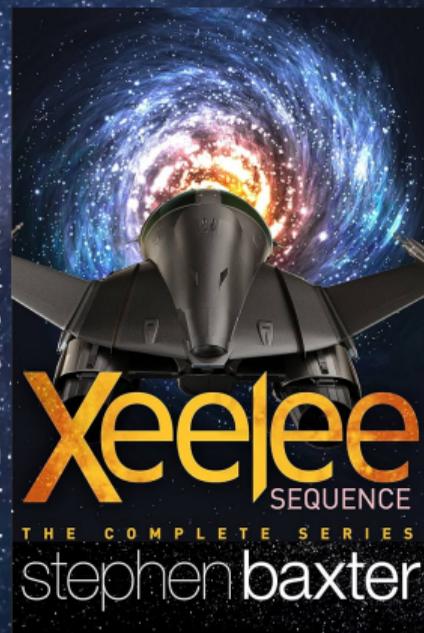
What is dark matter?

# The Standard Models of particle physics, cosmology and beyond

Dark matter in science-fiction



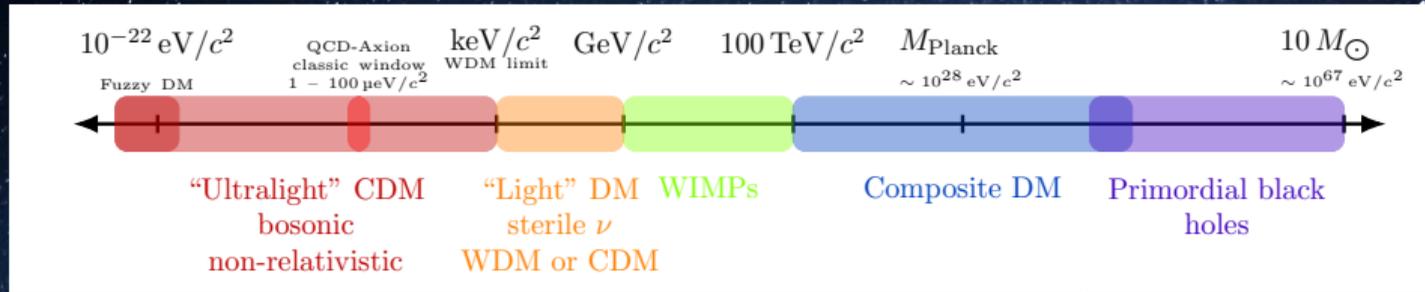
"Shadows are particles of consciousness. You ever heard anything so stupid? No wonder we can't get our grant renewed... And here goes the crazy part: you can't see them unless you expect to. Unless you put your mind in a certain state." —Mary Melone (Dark matter research unit, Oxford) in "The amber spyglass" by Philip Pullman.



Dark matter creatures...  
They might tell us what they are made of?

# The Standard Models of particle physics, cosmology and beyond

Dark matter in science ~~fiction~~



Pseudoscalar bosons for solving strong  $CP$  problem

$$\theta \frac{g_s^2}{32\pi^2} \tilde{G}_a^{\mu\nu} G_{\mu\nu}^a$$

(Weinberg and Wilczek)  
 Many low-energy searches:  
**axion-electric effect** (light through wall) or the **axion-wind** (magnetic field sensing)

Primordial black holes (MACHOs)

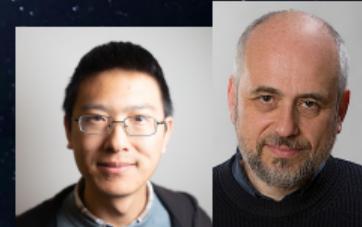
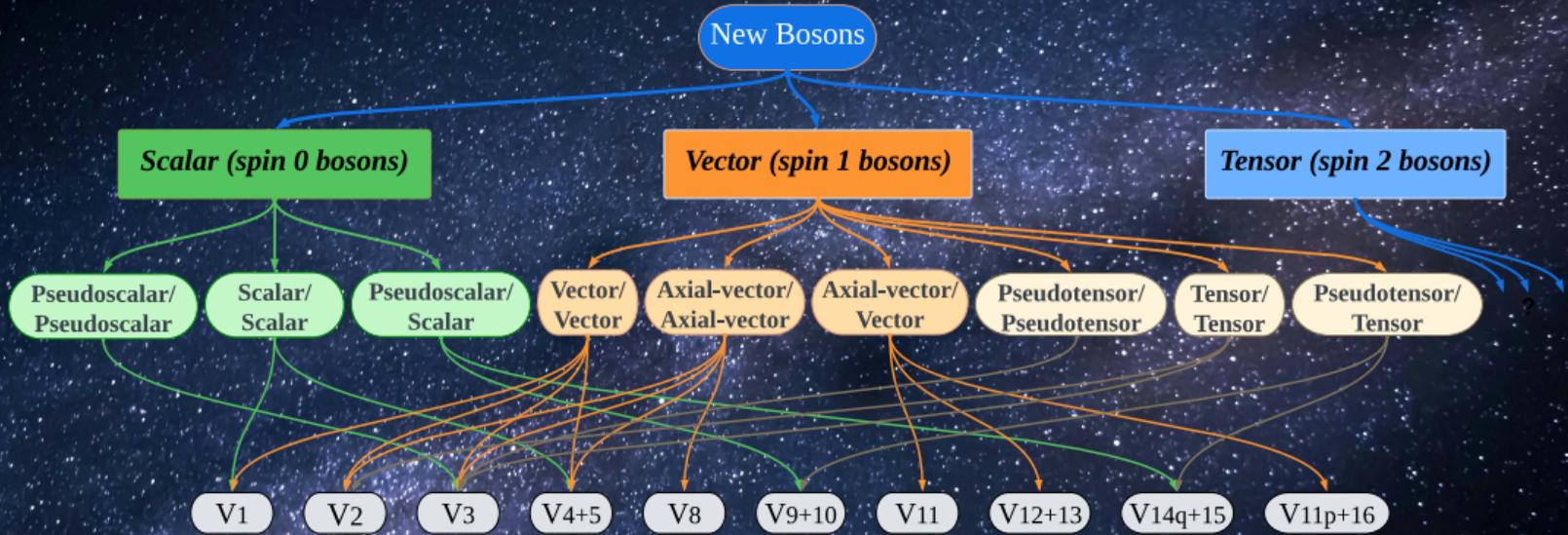


Zel'dovich, Nowikow and later Hawking  
 Formed shortly after big bang (before stars!)  
 Masses down to  $10^{12} \text{ m} \Leftrightarrow$  size of a fm!



# The Standard Models of particle physics, cosmology and beyond

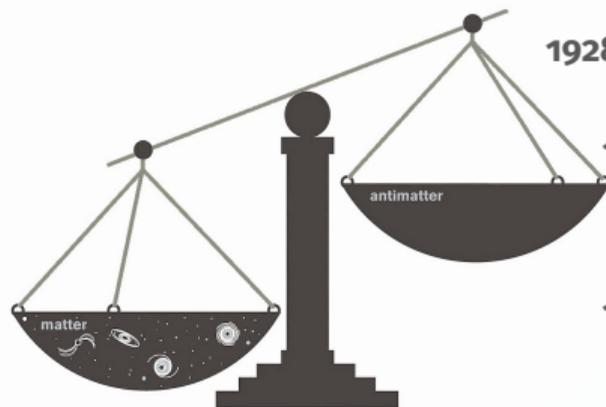
New bosons?



Why is there matter at all?

# The Standard Models of particle physics, cosmology and beyond

## Baryon asymmetry



<https://www.symmetrymagazine.org/article/october-2005/explain-it-in-60-seconds>

**1928/32** Feynman–Stückelberg interpretation of Dirac equation  
→ Antimatter

**1933** Discovery of the positron by Anderson.

### Where is the antimatter?

Hidden anti-matter clusters? →  $N_{\text{anti}}/N_{\text{mat}} < 10^{-6}$ !

**1967** Sakharov conditions

- 1 Baryon number ( $B$ ) violation  $\Delta B$ ,
- 2  $C$ - and  $CP$ -violation,
- 3 Deviation from thermal equilibrium.

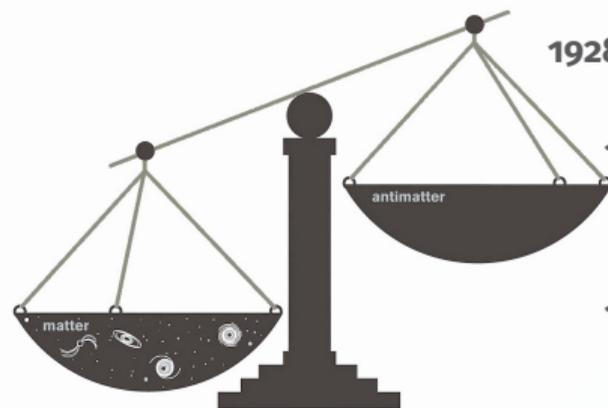
A. D. Sakharov, *JETP Lett.* **1967**, 5, 24, A. D. Sakharov, *Sov. Phys. Usp.* **1991**, 34, 392–393.

L. Canetti et al., *New J. Phys.* **2012**, 14, 095012.

G. 't Hooft, *Phys. Rev. Lett.* **1976**, 37, 8–11.

# The Standard Models of particle physics, cosmology and beyond

## Baryon asymmetry



<https://www.symmetrymagazine.org/article/october-2005/explain-it-in-60-seconds>

**1928/32** Feynman–Stückelberg interpretation of Dirac equation  
→ Antimatter

**1933** Discovery of the positron by Anderson.

### Where is the antimatter?

Hidden anti-matter clusters? →  $N_{\text{anti}}/N_{\text{mat}} < 10^{-6}!$

**1967** Sakharov conditions

- 1 Baryon number ( $B$ ) violation  $\Delta B$ ,
- 2  $C$ - and  $CP$ -violation,
- 3 Deviation from thermal equilibrium.

✓  $C$  is maximally violated in the SM (electroweak)

✓  $CP$  is violated in the SM (CKM, neutrino oscillations)

✓  $B$  is violated in the SM (non-perturbative, electroweak)

✓ Primordial plasma was out of equilibrium in the **expanding universe**

A. D. Sakharov, *JETP Lett.* **1967**, 5, 24, A. D. Sakharov, *Sov. Phys. Usp.* **1991**, 34, 392–393.

L. Canetti et al., *New J. Phys.* **2012**, 14, 095012.

G. 't Hooft, *Phys. Rev. Lett.* **1976**, 37, 8–11.

# The Standard Models of particle physics, cosmology and beyond

## Baryon asymmetry



1928/32 Feynman–Stückelberg interpretation of Dirac equation  
→ Antimatter

1933 Discovery of the positron by Anderson.

**Where is the antimatter?**

**X**In the SM only  $C$ -violation is sufficient!  
⇒ We need other mechanisms for  $CP$  and  $B$  violation!

<https://www.symmetrymagazine.org/article/october-2005/explain-it-in-60-seconds>

8 Deviation from thermal equilibrium.

- ✓  $C$  is maximally violated in the SM (electroweak)
- ✓  $CP$  is violated in the SM (CKM, neutrino oscillations)
- ✓  $B$  is violated in the SM (non-perturbative, electroweak)
- ✓ Primordial plasma was out of equilibrium in the **expanding universe**
- X**  $CP$ -violation very small in the SM (Jarlskog invariant  $\sim 10^{-5}$ )
- X**  $B$ -violation minuscule (sphalerons, at  $T = 0$ :  $\Gamma/V \sim m_W^4 \exp(-8\pi^2/g^2) \sim m_W^4 10^{-180}$ )

A. D. Sakharov, *JETP Lett.* **1967**, 5, 24, A. D. Sakharov, *Sov. Phys. Usp.* **1991**, 34, 392–393.

L. Canetti et al., *New J. Phys.* **2012**, 14, 095012.

G. 't Hooft, *Phys. Rev. Lett.* **1976**, 37, 8–11.

## Searching for New Physics at different scales

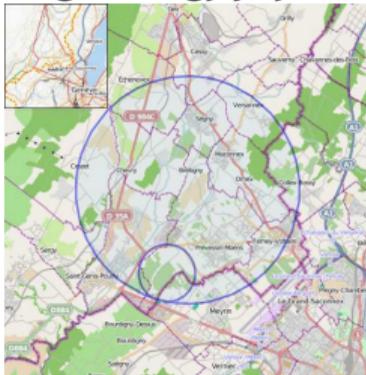
# Searching for New Physics at different scales

## Astrophysics



<https://www.eso.org/public/germany/images/potw1604a/>

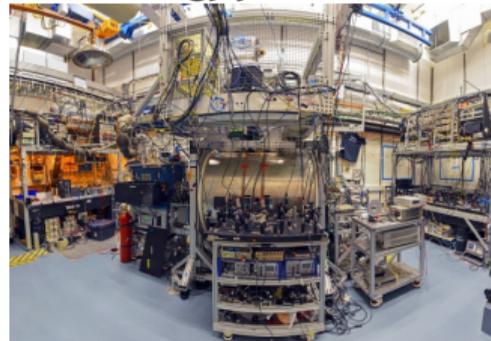
## High-energy physics



CERN: <https://home.cern/resources/image/accelerators/lhc-images-gallery>

<https://home.cern/resources/image/accelerators/lhc-images-gallery>

## Low-energy precision tests



ACME Collaboration: <http://www.electroedm.org/>



<https://first-tf.fr/grand-public-scolaires/ressources-grand-public/>

# Searching for New Physics at different scales

How to do a low-energy precision test: Two strategies

## 1. Precise comparison of theory and experiment

- ✓ Small well controllable systems:  
H,D,He,HD,H<sub>2</sub> etc.
- ✓ Accuracy of calculations  $\ll 1\%$
- ✗ System selection usually limited by nuclear structure

## 2. Looking for non-observables of the SM

- ✓ Symmetry violation beyond the Standard Model
- ✓ Forbidden transitions in atoms and molecules
- ✓ Measuring very precisely zero
- ✓ Weakly dependent on theory uncertainty
- ✗ Very important to control experimental systematics

REVIEWS OF MODERN PHYSICS, VOLUME 90, APRIL–JUNE 2018

### Search for new physics with atoms and molecules

M.S. Safronova<sup>1,2</sup>, D. Budker<sup>3,4,5</sup>, D. DeMille<sup>6</sup>, Derek F. Jackson Kimball<sup>7</sup>, A. Derevianko<sup>8</sup> and Charles W. Clark<sup>2</sup>

# Searching for New Physics at different scales

How to do a low-energy precision test: Two strategies

## 1. Precise comparison of theory and experiment

- ✓ Small well controllable systems: H,D,He,HD,H<sub>2</sub> etc.
- ✓ Accuracy of calculations  $\ll 1\%$
- ✗ System selection usually limited by nuclear structure

## 2. Looking for non-observables of the SM

- ✓ Symmetry violation beyond the Standard Model
- ✓ Forbidden transitions in atoms and molecules
- ✓ Measuring very precisely zero
- ✓ Weakly dependent on theory uncertainty
- ✗ Very important to control experimental systematics

REVIEWS OF MODERN PHYSICS, VOLUME 90, APRIL–JUNE 2018

### Search for new physics with atoms and molecules

M.S. Safronova<sup>1,2</sup>, D. Budker<sup>3,4,5</sup>, D. DeMille<sup>6</sup>, Derek F. Jackson Kimball<sup>7</sup>, A. Derevianko<sup>8</sup> and Charles W. Clark<sup>2</sup>

***We will follow the second route!***

Searches for  $\mathcal{P}$ -violation,  $CP/\mathcal{P}$ ,  $\mathcal{T}$ -violation,  $\mathcal{P}$ -violating dark matter, local Lorentz invariance violations, variations of fundamental constants, etc...

# Searching for New Physics at different scales

why molecules?

## Pro

- ✓ Complex many body system
- ✓ Internally broken symmetries
- ✓ Nucleus-nucleus interactions
- ✓ Can have simpler electronic structure than atoms
- ✓ Recent breakthroughs in molecular spectroscopy

## Contra

- ✗ Complex many body system
- ✗ Lower resolution than in atoms
- ✗ Ab initio description can be limited

1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	
1	H																18	
2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	18	
3	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	
4	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36
5	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54
6	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72
7	87	88	89	90	91	92	93	94	95	96	97	98	99	100	101	102	103	104
	105	106	107	108	109	110	111	112	113	114	115	116	117	118	119	120	121	122
	123	124	125	126	127	128	129	130	131	132	133	134	135	136	137	138	139	140



- Computational limitations come most often from the nucleus.
  - Experimental limitations can be outrivaled by internal enhancement mechanisms.
- ⇒ Tailoring a tuned quantum sensor from the periodic table!

## A crash course in relativistic molecular structure

# A crash course in relativistic molecular structure

From the SM to atoms and molecules

Molecules are composed of electrons and nuclei

✓ In good approximation only electromagnetic force matters!

# A crash course in relativistic molecular structure

From the SM to atoms and molecules

Molecules are composed of electrons and nuclei

✓ In good approximation only electromagnetic force matters!

$$\mathcal{L}_{\text{QED}} = \underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu}}_{\mathcal{L}_{\text{photon}}} + \sum_i \bar{\psi}_i \left[ \underbrace{i\hbar c \gamma^\mu \partial_\mu \psi_i}_{\mathcal{L}_{\text{kin}}} \underbrace{- m_i c^2}_{\mathcal{L}_{\text{mass}}} \underbrace{- q_i c \gamma^\mu A^\mu}_{\mathcal{L}_{\text{int}}} \right] \psi_i$$

$$\mu = 0, 1, 2, 3; \quad x_\mu = (ct, x, y, z)_\mu; \quad \partial^\mu = \frac{\partial}{\partial x_\mu};$$

$$A^\mu = [(\phi, -A_x, -A_y, -A_z)^\top]_\mu; \quad F_{\mu\nu} = \begin{pmatrix} 0 & \mathcal{E}_x/c & \mathcal{E}_y/c & \mathcal{E}_z/c \\ -\mathcal{E}_x/c & 0 & -\mathcal{B}_z & \mathcal{B}_y \\ -\mathcal{E}_y/c & \mathcal{B}_z & 0 & -\mathcal{B}_x \\ -\mathcal{E}_z/c & -\mathcal{B}_y & -\mathcal{B}_x & 0 \end{pmatrix}_{\mu\nu}$$

$$\gamma^0 = \begin{pmatrix} 1 & \mathbf{0} \\ \mathbf{0} & -1 \end{pmatrix}; \quad \gamma^k = \begin{pmatrix} 0 & \sigma^k \\ -\sigma^k & 0 \end{pmatrix}; \quad \gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3; \quad \beta = \gamma^0; \quad \vec{\alpha} = \gamma^0\vec{\gamma}; \quad \vec{\Sigma} = \gamma^0\gamma^5\vec{\gamma}$$

# A crash course in relativistic molecular structure

From the SM to atoms and molecules

Molecules are composed of electrons and nuclei

✓ In good approximation only electromagnetic force matters!

$$\mathcal{L}_{\text{QED}} = \underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu}}_{\mathcal{L}_{\text{photon}}} + \sum_i \bar{\psi}_i \left[ \underbrace{i\hbar c \gamma^\mu \partial_\mu \psi_i}_{\mathcal{L}_{\text{kin}}} \underbrace{-m_i c^2}_{\mathcal{L}_{\text{mass}}} \underbrace{-q_i c \gamma^\mu A^\mu}_{\mathcal{L}_{\text{int}}} \right] \psi_i$$

$\mu = 0, 1, 2, 3; x_\mu = (ct, x, y, z)_\mu; \partial^\mu = \frac{\partial}{\partial x_\mu};$

$A^\mu = [(\phi, -A_x, -A_y, -A_z)^\top]_\mu; F_{\mu\nu} = \begin{pmatrix} 0 & \mathcal{E}_x/c & \mathcal{E}_y/c & \mathcal{E}_z/c \\ -\mathcal{E}_x/c & 0 & -\mathcal{B}_z & \mathcal{B}_y \\ -\mathcal{E}_y/c & \mathcal{B}_z & 0 & -\mathcal{B}_x \\ -\mathcal{E}_z/c & -\mathcal{B}_y & -\mathcal{B}_x & 0 \end{pmatrix}_{\mu\nu}$

$\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}; \gamma^k = \begin{pmatrix} 0 & \sigma^k \\ -\sigma^k & 0 \end{pmatrix}; \gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3; \beta = \gamma^0; \vec{\alpha} = \gamma^0\vec{\gamma}; \vec{\Sigma} = \gamma^0\gamma^5\vec{\gamma}$

- ✗ Interaction term is **particle field dependent!** → Many-body problem is tough in QED (For **two** particles: **Bethe-Salpeter equation**)
- ✗ Combination of electron correlation and QED an open problem (how to treat virtual positronic states?)!

# A crash course in relativistic molecular structure

From the SM to atoms and molecules

Molecules are composed of electrons and nuclei

✓ In good approximation only electromagnetic force matters!

$$\mathcal{L}_{\text{QED}} = \underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu}}_{\mathcal{L}_{\text{photon}}} + \sum_i \bar{\psi}_i \left[ \underbrace{i\hbar c \gamma^\mu \partial_\mu \psi_i}_{\mathcal{L}_{\text{kin}}} \underbrace{- m_i c^2}_{\mathcal{L}_{\text{mass}}} \underbrace{- q_i c \gamma^\mu A^\mu}_{\mathcal{L}_{\text{int}}} \right] \psi_i$$

$\mu = 0, 1, 2, 3; x_\mu = (ct, x, y, z)_\mu; \partial^\mu = \frac{\partial}{\partial x_\mu};$

$A^\mu = [(\phi, -A_x, -A_y, -A_z)^\top]_\mu; F_{\mu\nu} = \begin{pmatrix} 0 & \mathcal{E}_x/c & \mathcal{E}_y/c & \mathcal{E}_z/c \\ -\mathcal{E}_x/c & 0 & -\mathcal{B}_z & \mathcal{B}_y \\ -\mathcal{E}_y/c & \mathcal{B}_z & 0 & -\mathcal{B}_x \\ -\mathcal{E}_z/c & -\mathcal{B}_y & -\mathcal{B}_x & 0 \end{pmatrix}_{\mu\nu}$

$\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}; \gamma^k = \begin{pmatrix} 0 & \sigma^k \\ -\sigma^k & 0 \end{pmatrix}; \gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3; \beta = \gamma^0; \vec{\alpha} = \gamma^0\vec{\gamma}; \vec{\Sigma} = \gamma^0\gamma^5\vec{\gamma}$

- ✗ Interaction term is **particle field dependent!** → Many-body problem is tough in QED (For **two** particles: **Bethe-Salpeter equation**)
- ✗ Combination of electron correlation and QED an open problem (how to treat virtual positronic states?)
- ✓ Fortunately: QED is perturbative!

# A crash course in relativistic molecular structure

From the SM to atoms and molecules

Molecules are composed of electrons and nuclei

✓ In good approximation only electromagnetic force matters!

$$\mathcal{L}_{\text{QED}} = \underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu}}_{\mathcal{L}_{\text{photon}}} + \sum_i \bar{\psi}_i \left[ \underbrace{i\hbar c \gamma^\mu \partial_\mu \psi_i}_{\mathcal{L}_{\text{kin}}} \underbrace{- m_i c^2}_{\mathcal{L}_{\text{mass}}} \underbrace{- q_i c \gamma^\mu A^\mu}_{\mathcal{L}_{\text{int}}} \right] \psi_i$$

$\mu = 0, 1, 2, 3; x_\mu = (ct, x, y, z)_\mu; \partial^\mu = \frac{\partial}{\partial x_\mu};$   
 $A^\mu = [(\phi, -A_x, -A_y, -A_z)^\top]_\mu; F_{\mu\nu} = \begin{pmatrix} 0 & \mathcal{E}_x/c & \mathcal{E}_y/c & \mathcal{E}_z/c \\ -\mathcal{E}_x/c & 0 & -\mathcal{B}_z & \mathcal{B}_y \\ -\mathcal{E}_y/c & \mathcal{B}_z & 0 & -\mathcal{B}_x \\ -\mathcal{E}_z/c & -\mathcal{B}_y & -\mathcal{B}_x & 0 \end{pmatrix}_{\mu\nu}$   
 $\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}; \gamma^k = \begin{pmatrix} 0 & \sigma^k \\ -\sigma^k & 0 \end{pmatrix}; \gamma^5 = i\gamma^0\gamma^1\gamma^2\gamma^3; \beta = \gamma^0; \vec{\alpha} = \gamma^0\vec{\gamma}; \vec{\Sigma} = \gamma^0\vec{\gamma}^5\vec{\gamma}$

- ✗ Interaction term is **particle field dependent!** → Many-body problem is tough in QED (For **two** particles: **Bethe-Salpeter equation**)
- ✗ Combination of electron correlation and QED an open problem (how to treat virtual positronic states?)
- ✓ Fortunately: QED is perturbative!
- ✓ Common approximations:
  - ▶ “No-pair” approximation (Electron-positron pair creation at ~ 1 MeV!)
  - ▶ Instantaneous interactions (classical Coulomb)
    - Perturbative QED correction from the photon field
  - ▶ Nuclei are hard spherical charges  $eZ\rho(r)$
  - ▶ Electronic and nuclear motion is separated (Born–Oppenheimer approximation)
  - ▶  $\alpha \rightarrow \infty$  (non-relativistic limit) (Perturbation theory on the Schrödinger equation is predictive for  $E \ll m_e c^2 \sim 511 \text{ keV}$ )
  - ▶ Bound-state atomic/molecular Hamiltonians:

$$\hat{H} = c\vec{\alpha} \cdot \hat{\vec{p}} + \beta m_e c^2 + \hat{V}_{eN} + \hat{V}_{ee} + \hat{V}_{NN}, \quad \lim_{\alpha \rightarrow 0} \hat{H} - m_e c^2 = \hat{H}_{\text{nr}} = \frac{\hat{\vec{p}}^2}{2m_e} + \hat{V}_{eN} + \hat{V}_{ee} + \hat{V}_{NN}$$

# A crash course in relativistic molecular structure

One-particle Dirac equation and Schrödinger equation

Time-independent Dirac/Schrödinger equation for a central Coulomb potential ( $V_{eN} = \frac{-eZ}{4\pi\epsilon_0 r}$  for  $\rho(r) = 4\pi\delta(r)$ ):

$$\hat{H}\Psi_n = E_n\Psi_n$$

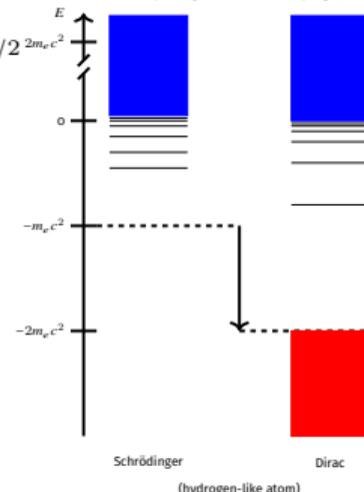
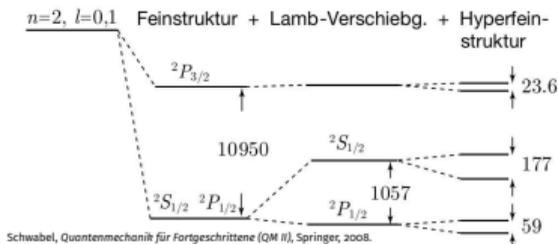
For Dirac  $\Psi_n = \begin{pmatrix} \text{blue wave} \\ \text{red wave} \end{pmatrix} \otimes \begin{pmatrix} |\uparrow\rangle \\ |\downarrow\rangle \end{pmatrix}$  is a bi-spinor (four-vector)

“Large component” for **electronic states**  
and “small component” for **positronic states**

$$E_{n,j} = mc^2 \left[ 1 + \frac{Z\alpha}{n - j - \frac{1}{2} + \underbrace{\sqrt{(j + \frac{1}{2})^2 - (Z\alpha)^2}}_{\text{Electronic Lorentz factor}}} \right]^{1/2} - \frac{1}{2} 2m_e c^2$$

$$\lim_{\alpha \rightarrow 0} E_{n,j} - mc^2 = -\frac{mc^2(Z\alpha)^2}{2n^2}$$

Shifting energies by  $mc^2$  to obtain a non-relativistic analogue!  
**Dirac sea? → Negative energy states are empty!**



# A crash course in relativistic molecular structure

Relativistic effects? Paul Dirac's assessment

## § 1. *Introduction.*

The general theory of quantum mechanics is now almost complete, the imperfections that still remain being in connection with the exact fitting in of the theory with relativity ideas. These give rise to difficulties only when high-speed particles are involved, and are therefore of no importance in the consideration of atomic and molecular structure and ordinary chemical reactions, in which it is, indeed, usually sufficiently accurate if one neglects relativity variation of mass with velocity and assumes only Coulomb forces between the various electrons and atomic nuclei. The underlying physical laws necessary for the mathematical theory of a large part of physics and the whole of chemistry are thus completely known, and the difficulty is only that the exact application of these laws leads to equations much too complicated to be soluble. It therefore becomes desirable that approximate practical methods of applying quantum

# A crash course in relativistic molecular structure

Do relativistic effects really not matter for physics and chemistry?

Classics in chemistry:

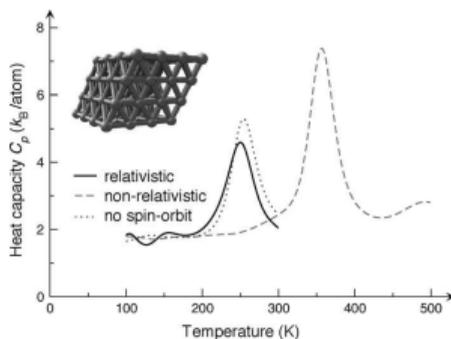
- Why is gold so shiny-yellow?



- Why does your (old) car start?



- Why is mercury liquid at room temperature?



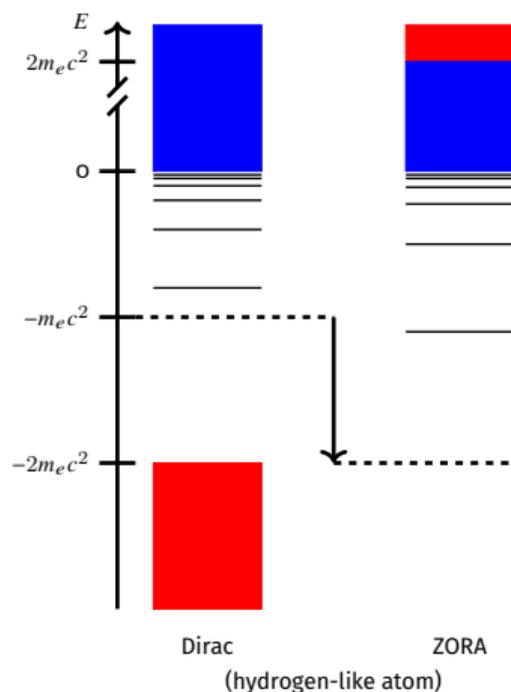
# A crash course in relativistic molecular structure

## Two-component approximations/Quasi-relativistic calculations

- Positronic states not populated in the molecule  $\Rightarrow$  implicit treatment of the small component!
- Elimination of the small component at the one-particle level:
  - ▶ Two-step approaches (elimination on the matrix level, RI): DKH, X2C
  - ▶ One-step approaches (elimination on the operator level): (Beit-)Pauli, ZORA

$$\begin{pmatrix} \hat{V} - \epsilon & c\vec{\sigma} \cdot \hat{p} \\ c\vec{\sigma} \cdot \hat{p} & \hat{V} - \epsilon - 2m_e c^2 \end{pmatrix} \begin{pmatrix} \psi_L \\ \psi_S \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$$
$$\Rightarrow \psi_S = \left( 2m_e c^2 - \hat{V} + \epsilon \right)^{-1} c\vec{\sigma} \cdot \hat{p} \psi_L$$

- Expansion of  $\left( 2m_e c^2 - \hat{V} + \epsilon \right)^{-1}$ 
  - ▶ Pauli:  $\frac{1}{2m_e c^2} \sum_{k=0}^{\infty} \left[ \frac{\hat{V} - \epsilon}{2m_e c^2} \right]^k$
  - ▶ Regular:  $\frac{1}{2m_e c^2 - \hat{V}} \sum_{k=0}^{\infty} \left[ \frac{-\epsilon}{2m_e c^2 - \hat{V}} \right]^k \Rightarrow$  ZORA:  $\frac{1}{2m_e c^2 - \hat{V}}$
- ZORA usually captures all important relativistic effects for valence states!
- Core states can be renormalized (IORA)
- Missing two-electron effects usually negligible  $\ll 1\%$  for heavy elements!



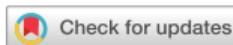
# A crash course in relativistic molecular structure

Two-component approximations / Quasi-relativistic calculations

- Positron of the s
- Elimina
  - ▶ Two-ste
  - ▶ One-ste

## Quantum electrodynamic corrections for molecules: Vacuum polarization and electron self-energy in a two-component relativistic framework

Kjell Janke ; Andrés Emilio Wedenig ; Peter Schwerdtfeger ; Konstantin Gaul ; Robert Berger  



*J. Chem. Phys.* 162, 104111 (2025)

<https://doi.org/10.1063/5.0252409>

**TABLE III.** SE (Flambaum–Ginges) and VP (Uehling) contributions in eV to the  ${}^2P_{1/2} \leftarrow {}^2S_{1/2}$  and  ${}^2P_{3/2} \leftarrow {}^2S_{1/2}$  transitions calculated as expectation values based on ZORA-HF/dyall.cv3z calculations, using prefactor (6). A comparison<sup>31</sup> with four-component numerical DHF calculations with perturbative treatment of the QED contributions.  $Z$  is the nuclear charge and  $N$  is the number of electrons.

$Z$	$N$	$V_{SE,2P1/2}$	Dev. (%)	$V_{SE,2P3/2}$	Dev. (%)	$V_{VP,2P1/2}$	Dev. (%)	$V_{VP,2P3/2}$	Dev. (%)
10	3	$-1.498 \times 10^{-2}$	0.0	$-1.436 \times 10^{-2}$	0.0	$7.643 \times 10^{-4}$	0.0	$7.643 \times 10^{-4}$	0.0
20	3	$-2.065 \times 10^{-1}$	0.5	$-1.936 \times 10^{-1}$	0.0	$1.414 \times 10^{-2}$	1.4 <sup>a</sup>	$1.417 \times 10^{-2}$	0.7 <sup>b</sup>
30	3	$-8.864 \times 10^{-1}$	0.1	$-8.179 \times 10^{-1}$	0.1	$7.562 \times 10^{-2}$	0.7	$7.616 \times 10^{-2}$	0.7
40	3	-2.448	0.3	-2.238	0.4	$2.526 \times 10^{-1}$	0.4	$2.568 \times 10^{-1}$	0.4
50	3	-5.379	0.5	-4.906	0.6	$6.631 \times 10^{-1}$	0.3	$6.838 \times 10^{-1}$	0.2
60	3	$-1.032 \times 10^1$	1.1	-9.467	1.0	1.513	0.1	1.591	0.3
70	3	$-1.810 \times 10^1$	1.1	$-1.691 \times 10^1$	1.3	3.169	0.4	3.420	0.6
80	3	$-2.988 \times 10^1$	1.2	$-2.885 \times 10^1$	1.6	6.293	0.5	7.040	0.9

- Expans
  - ▶ Pauli: 2
  - ▶ Regular

- ZORA usually captures all important relativistic effects for valence states.
- Core states can be renormalized (IORA)
- Missing two-electron effects usually negligible  $\ll 1\%$  for heavy elements!



ZORA  
(atom)

# A crash course in relativistic molecular structure

Many electrons

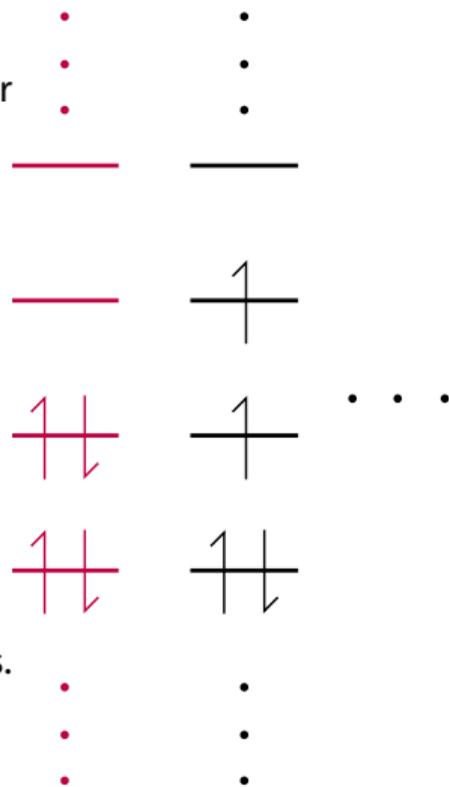
Many-electron wave function  $\Psi(\vec{r}_1, \dots, \vec{r}_N)$  must be totally anti-symmetric under exchange of two electrons (Spin-statistics theorem):

$$\Phi_j \sim \hat{\mathcal{A}} \prod_i \psi_{ij}(\vec{r}_i) \text{ Slater determinant}$$

$\Psi_j$  does not need to be of structure  $\Phi_j$  but it can be shown that for any  $\Psi_j$ :

$$\Psi_j = \sum_I c_{Ij} \Phi_I \text{ (Configuration Interaction),}$$

where the sum runs over all configurations  $\Phi_I$  in the full single-particle space...  
Number of determinants grows  $\binom{M}{N}$  for  $N$  electrons and  $M$  one-particle functions.  
⇒ Intractable problem for most atoms and molecules!



# A crash course in relativistic molecular structure

Quantum Chemistry

Quantum chemistry  $\Leftrightarrow$  Obtaining the most efficient and accurate approximations to  $\Psi_j$ .

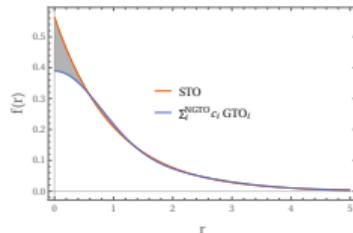
## Approximating the one-particle space

Expand single-particle function  $\phi_i$  in set of known functions  $\{\chi_\mu\}$  inspired by H-atom solutions:

$$\chi_\mu(\vec{r}) = (x - X_\mu)^{l_{x\mu}} (y - Y_\mu)^{l_{y\mu}} (z - Z_\mu)^{l_{z\mu}} \times \exp\left(-\zeta_\mu \left|\vec{r} - \vec{R}_\mu\right|^k\right)$$

$$\phi_i(\vec{r}) = \sum_{\mu=1}^M C_{\mu i} \chi_\mu(\vec{r})$$

In most cases we choose  $k = 2$  (Gaussian) for numerical simplicity. Closer to H-atom solution would be  $k = 1$ .



## Approximating the $N$ -particle space

$$\Psi_j \approx \Phi_{j0}, \text{ scales } M^3$$

Hartree-Fock  
 $\min_{\phi_i} E[\Phi_{j0}(\{\phi_i\})]$

$$\Psi_j \approx \sum_{I \in M_K} \Phi_{jI}, \text{ scales } \binom{N+K}{N}$$

Multi-reference methods  
MCSCF, CASSCF, RASSCF, DMRG, ...  
 $\min_{\phi_i, \Phi_{jI}} E[\Phi_j(\{\phi_i\})]$

$$\Psi_j \approx \sum_k^{N_{\text{exc}}} \hat{T}_k \Phi_{j0}, \text{ scales } M^5 - M^7, \dots$$

Single-reference correlated methods  
CI, CC, MBPT, MPn, ...  
Not necessarily variational

$$\Psi_j \approx \sum_k^{N_{\text{exc}}} \hat{T}_k \tilde{\Psi}_j; \tilde{\Psi}_j \approx \sum_{I \in M_K} \Phi_{jI}$$

Multi-reference correlated methods  
MRCI, MRCC, CASPTn, GASCI, FSCC, ...

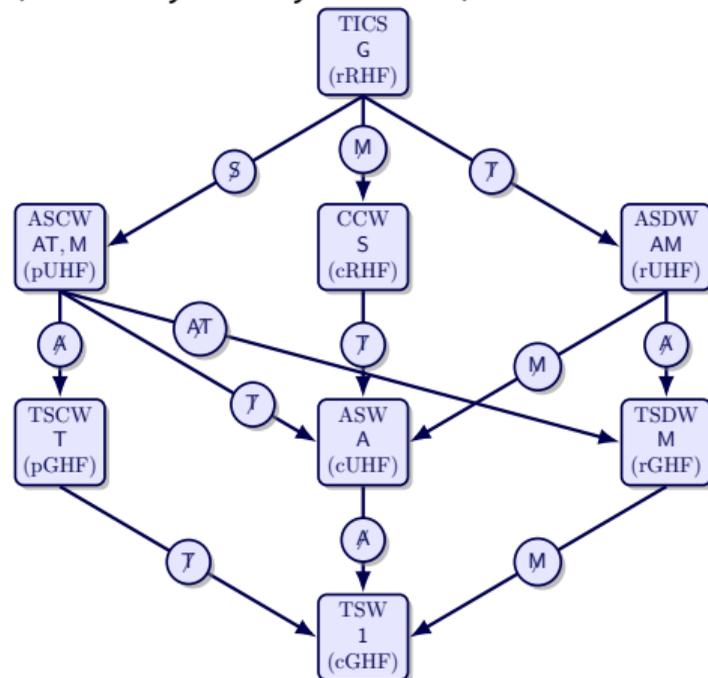
$$\Psi_j \approx \sum_I \Phi_{jI}$$

Full CI  
exact diagonalization

# A crash course in relativistic molecular structure

“Cheated” correlation: Broken symmetry Hartree–Fock

We can obtain mixed electronic configuration by breaking the symmetry of the Hamiltonian (Löwdin symmetry dilemma):



Hamiltonian invariant under  $T = \{-1, 1, -\hat{\Theta}, \hat{\Theta}\}$

Non-relativistic Hamiltonian additionally invariant under  $\hat{K}$  &  $s(\vec{n}, \theta) = \exp\left(i\frac{\theta}{2}n_i\sigma^i\right)$   $S = \{s(\vec{n}, \theta) : \vec{n} = \mathbb{R}^3, \|\vec{n}\| = 1, \theta \in \{0, 4\pi\}\}$

$G = S \otimes T$ , has eight subgroups:

$G, S, T, A(\vec{n}) = \{s(\vec{n}, \theta) : \theta \in \{0, 4\pi\}\}, M(\vec{n}) = \{1, \Theta s(\vec{n}, \pi)\}, A(\vec{n})M(\vec{n}') \text{ with } \vec{n} \perp \vec{n}', A \otimes T \text{ and } E = \{-1, 1\}$

- ✓ Multi-reference wave-function at the cost of HF ( $\propto N^3$ )!
- ✓ Simple application of perturbation theory (property calculations)
- ✗ Possibility for unphysical mixtures of configurations
- ✗ Wave-function has wrong symmetry
- ✗ Difficult to describe low-spin states

# A crash course in relativistic molecular structure

## Density functional theory (DFT)

The ground state energy is fully determined by the one-electron four-current  $j^\mu(\vec{r})$ :

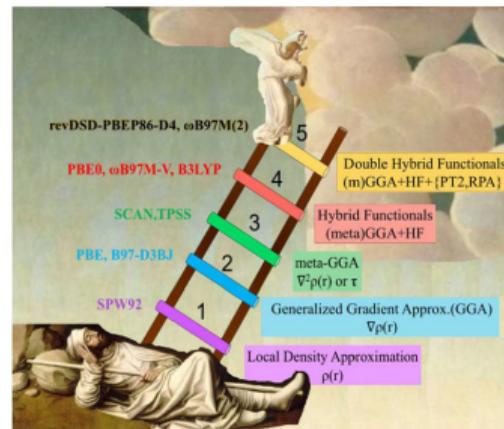
$$j^\mu(\vec{r}_1) = N \int \cdots \int d\vec{r}_2 \cdots d\vec{r}_N \bar{\Psi}_{4c} \gamma_1^\mu \otimes \gamma_2^0 \otimes \cdots \otimes \gamma_N^0 \Psi_{4c},$$

Idea: instead of using the complex  $3N$  dimensional wave functions minimize energy with respect to  $j^\mu(\vec{r})$ :

$$E[j^\mu(\vec{r})] = \underbrace{T[j^\mu(\vec{r})]}_{\text{kinetic energy}} + \underbrace{V_{\text{ext}}[j^\mu(\vec{r})]}_{\text{external potential (incl e-N interaction)}} + \underbrace{V_{\text{H}}[j^\mu(\vec{r})]}_{\text{Hartree potential (classical Coulomb repulsion)}} + \underbrace{E_{\text{XC}}[j^\mu(\vec{r})]}_{\text{exchange correlation functional}}$$

**Problem: Exact closed expression for  $E_{\text{XC}}$  unknown.**

- ✓ DFT is formally exact
- ✓ DFT can be formulated consistently with QFT
- ✓ Pure Kohn-Sham DFT is of lower cost than HF ( $N(N+1)/2N_{\text{grid}}$ )
- ✗  $E_{\text{XC}}$  is unknown
- ✗ Semi-empirical estimates of  $E_{\text{XC}}$  introduce systematic errors
- ✗ Relativistic calculations rely currently on non-relativistic  $E_{\text{XC}}$



# A crash course in relativistic molecular structure

Adiabatic approximation/Born–Oppenheimer approximation

Nuclei are assumed to be non-relativistic!

Table 1. Three frequently discussed adiabatic approximations

	I Crude adiabatic approximation	II Born–Oppenheimer adiabatic approximation	III Born–Huang adiabatic approximation
Adiabatic wavefunction	$\Psi_j^{CA}(r, Q) = \psi_j^i(r, Q_0) \chi_j^{CA}(Q)$	$\Psi_j^{BO}(r, Q) = \psi_j^i(r, Q) \chi_j^{BO}(Q)$	$\Psi_j^{BH}(r, Q) = \psi_j^i(r, Q) \chi_j^{BH}(Q)$
Electronic equation	$[T_e(r) + U(r, Q_0) - \epsilon_j^i(Q_0)] \times \psi_j^i(r, Q_0) = 0$	$[T_e(r) + U(r, Q) - \epsilon_j(Q)] \times \psi_j^i(r, Q) = 0$	$[T_e(r) + U(r, Q) - \epsilon_j(Q)] \times \psi_j^i(r, Q) = 0$
Vibrational equation	$[T_N(Q) + V(Q) + \epsilon_j^i(Q_0) + \langle \psi_j^i(r, Q_0)   \Delta U(r, Q)   \psi_j^i(r, Q_0) \rangle - E_j^{CA}] \chi_j^{CA}(Q) = 0$	$[T_N(Q) + V(Q) + \epsilon_j(Q) - E_j^{BO}] \chi_j^{BO}(Q) = 0$	$[T_N(Q) + V(Q) + \epsilon_j(Q) + \langle \psi_j^i(r, Q)   T_N(Q)   \psi_j^i(r, Q) \rangle - E_j^{BH}] \chi_j^{BH}(Q) = 0$
Adopted approximation	$\langle \psi_j^i(r, Q_0)   \Delta U(r, Q)   \psi_k^i(r, Q_0) \rangle = 0$ for $k \neq j$	$\langle \psi_j^i(r, Q)   T_N(Q)   \psi_k^i(r, Q) \rangle = 0$ and $\langle \psi_j^i(r, Q)   \frac{\partial}{\partial Q_m}   \psi_k^i(r, Q) \rangle = 0$	$\langle \psi_j^i(r, Q)   T_N(Q)   \psi_k^i(r, Q) \rangle = 0$ for $k \neq j$ and $\langle \psi_j^i(r, Q)   \frac{\partial}{\partial Q_m}   \psi_k^i(r, Q) \rangle = 0$

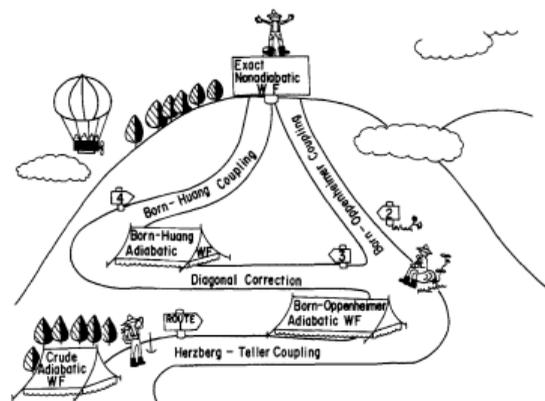


Figure 1. Various routes to approach the exact non-adiabatic wavefunction. As to the three types of adiabatic wavefunctions, see Table 1.

BO is an effective field theory. Perturbation series in the electron-to-proton-mass-ratio:

$$E_{\text{elec}} \propto \sqrt{\frac{m_e}{m_p}} E_{\text{vib}} \propto \frac{m_e}{m_p} E_{\text{rot}}$$

# A crash course in relativistic molecular structure

## Computation of molecular properties

- For small perturbations (new physics effects are definitely small!)  $\hat{H}_i$  we can use perturbation theory!
- First order for variational  $\Psi_j$

$$\frac{\partial E}{\partial \lambda_i} = \langle \Psi_j | \hat{H}_i | \Psi_j \rangle$$

- Not true for CC or MPn!  
→ numeric differentiation (finite field) or perturbation of Lagrangian within response theory!
- For second order:

$$\frac{\partial^2 E}{\partial \lambda_i \partial \lambda_l} = \langle \Psi_j | \hat{H}_{il} | \Psi_j \rangle + \left\langle \frac{\partial \Psi_j}{\partial \lambda_l} \left| \hat{H}_i \right| \Psi_j \right\rangle + \text{hc}$$

$$\left\langle \frac{\partial \Psi_j}{\partial \lambda_l} \left| \hat{H}_i \right| \Psi_j \right\rangle = \sum_{aj} \frac{\langle \Psi_j | \hat{H}_i | \Psi_a \rangle \langle \Psi_a | \hat{H}_l | \Psi_j \rangle}{E_j - E_a}$$

**Depends on excited states!**

⇒ HF/DFT level: Linear response theory (random phase approximation, RPA)

# A crash course in relativistic molecular structure

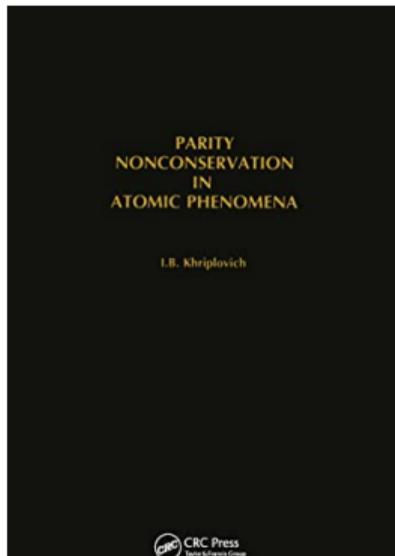
Do we really need all those ab initio calculations? Fermi–Segré model

1933 Fermi and Segré: simple effective one-electron approach to describe hyperfine coupling constants within quantum defect theory.

1958 Improvement of the model by Foldy.

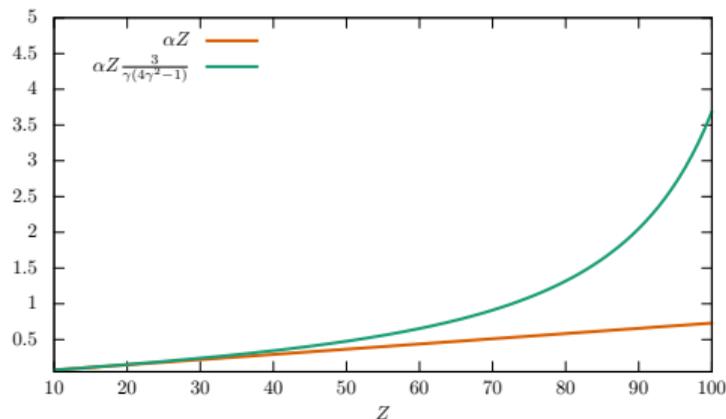
1972 Generalization of the model by Fröman and Fröman.

1975 Bouchiat and Bouchiat: extend it to parity violation.



- Simple effective one-electron wave functions at distances  $r \ll a_0 Z^{-1/3}$ :

$$f_{\kappa}(r) = \frac{\kappa}{|\kappa|r\sqrt{Za_0\nu_{\kappa}^3}} \left( (\kappa + \gamma_{\kappa}) J_{2\gamma_{\kappa}} \left( \sqrt{\frac{8Zr}{a_0}} \right) - \sqrt{\frac{2Zr}{a_0}} J_{2\gamma_{\kappa}-1} \left( \sqrt{\frac{8Zr}{a_0}} \right) \right),$$
$$g_{\kappa}(r) = \frac{\kappa Z \alpha}{|\kappa|r\sqrt{Za_0\nu_{\kappa}^3}} J_{2\gamma_{\kappa}} \left( \sqrt{\frac{8Zr}{a_0}} \right),$$



- Effective quantum numbers  $\nu$  contain all many-body information ( $O(\nu_{\kappa}) \sim 1!$ )  
⇒ Estimate the expectable order of magnitude of a property!

**⇒ Fundamental symmetry violating properties are often relativistically enhanced!**

E. Fermi, E. Segrè, *Z. Phys.* **1933**, 82, 729–749, L. L. Foldy, *Phys. Rev.* **1958**, 111, 1093–1098, N. Fröman, P. O. Fröman, *Phys. Rev. A* **1972**, 6, 2064–2067, C. Bouchiat, M. A. Bouchiat, *J. Phys. (Paris)* **1975**, 36, 493–509.

# A crash course in relativistic molecular structure

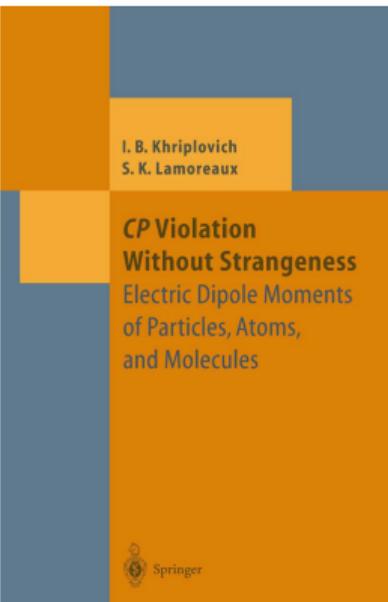
Do we really need all those ab initio calculations? Fermi–Segré model

1933 Fermi and Segré: simple effective one-electron approach to describe hyperfine coupling constants within quantum defect theory.

1958 Improvement of the model by Foldy.

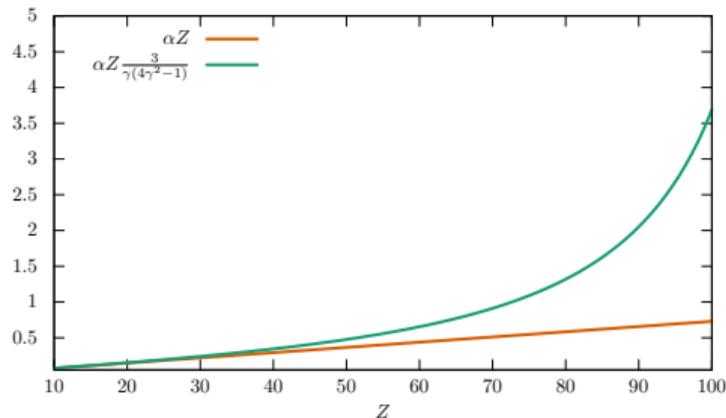
1972 Generalization of the model by Fröman and Fröman.

1975 Bouchiat and Bouchiat: extend it to parity violation.



- Simple effective one-electron wave functions at distances  $r \ll a_0 Z^{-1/3}$ :

$$f_{\kappa}(r) = \frac{\kappa}{|\kappa|r\sqrt{Za_0\nu_{\kappa}^3}} \left( (\kappa + \gamma_{\kappa}) J_{2\gamma_{\kappa}} \left( \sqrt{\frac{8Zr}{a_0}} \right) - \sqrt{\frac{2Zr}{a_0}} J_{2\gamma_{\kappa}-1} \left( \sqrt{\frac{8Zr}{a_0}} \right) \right),$$
$$g_{\kappa}(r) = \frac{\kappa Z \alpha}{|\kappa|r\sqrt{Za_0\nu_{\kappa}^3}} J_{2\gamma_{\kappa}} \left( \sqrt{\frac{8Zr}{a_0}} \right),$$



- Effective quantum numbers  $\nu$  contain all many-body information ( $O(\nu_{\kappa}) \sim 1!$ )  
⇒ Estimate the expectable order of magnitude of a property!

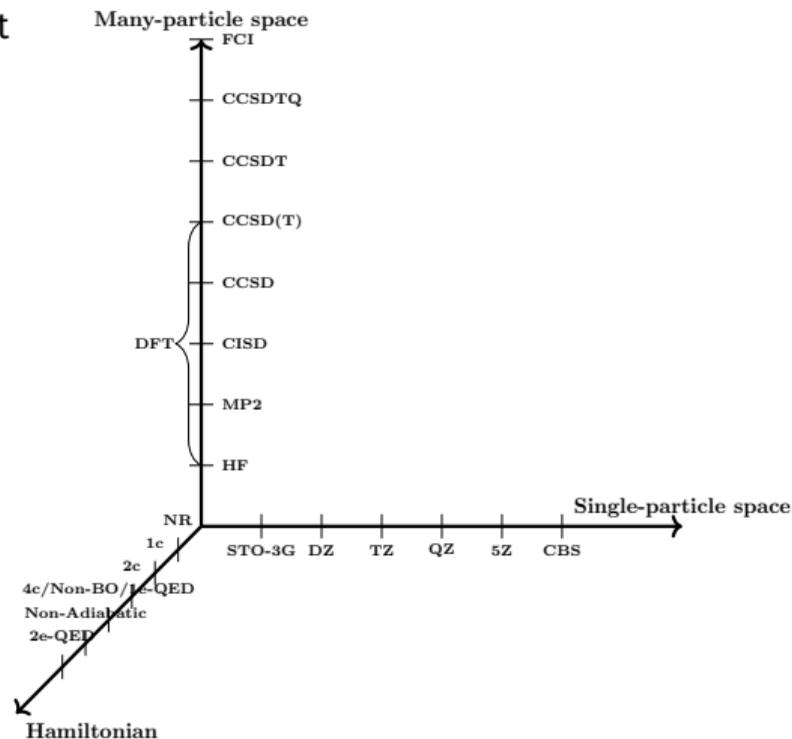
**⇒ Fundamental symmetry violating properties are often relativistically enhanced!**

E. Fermi, E. Segré, *Z. Phys.* **1933**, 82, 729–749, L. L. Foldy, *Phys. Rev.* **1958**, 111, 1093–1098, N. Fröman, P. O. Fröman, *Phys. Rev. A* **1972**, 6, 2064–2067, C. Bouchiat, M. A. Bouchiat, *J. Phys. (Paris)* **1975**, 36, 493–509.

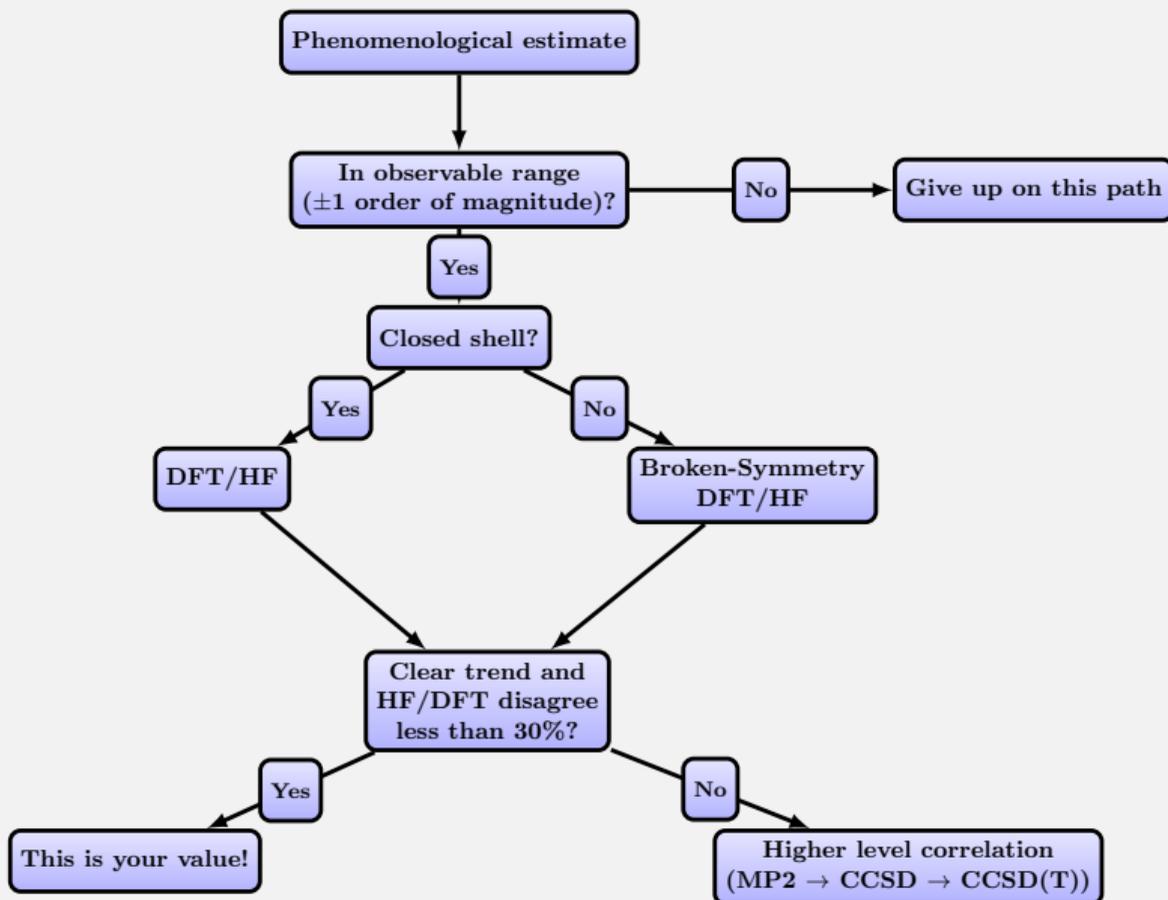
# A crash course in relativistic molecular structure

Quantum chemistry for new physics searches: a few remarks

- For identifying molecular probes **DFT** usually sufficient (if **open-shell broken-symmetry HF/DFT!**)
- Broken symmetries appear in the valence: **Two-component** methods are quantitative  $< 1\%$
- Standard basis sets are not designed for physics beyond the Standard Model properties!  
→ **Be very careful!**
- Assisting spectroscopy requires usually accurate predictions with relative errors  $< 10\%$
- Extraction of limits on new physics is dependent on the theory uncertainty and requires
  - ▶ Highly correlated methods (CC, MRCI, etc.) if the uncertainty is limited by electronic structure
  - ▶ HF/DFT is usually OK if the uncertainty is limited by nuclear structure
- Second order properties can be tough!



- For identical systems (if open shell)
- Broken symmetry
- Two-component
- Standard beyond → Be v
- Assistance predict
- Extract the theoretical value
  - High accuracy if the method is
  - HF/DFT if the method is
- Second order



particle space →