





## Photon identification in ATLAS: a revolution

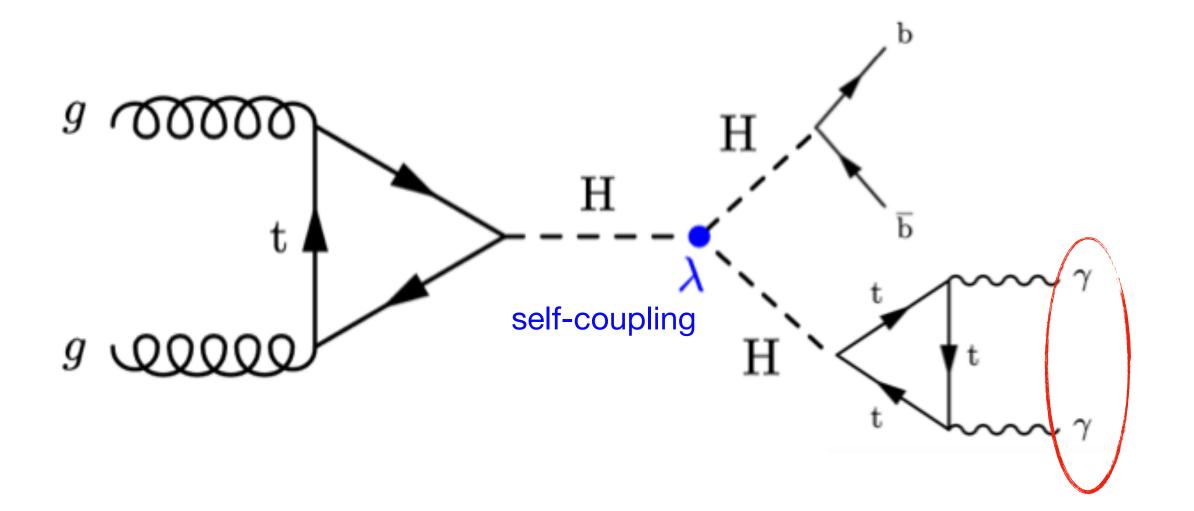
Luca Franco

#### A few remarks

- This project is done in the ATLAS combined performance EGamma group
- I am collaborating with the current EGamma photon identification (ID) convener (Mohamed Belfkir, UAEU)
- Much of what I will show today is very WIP, with mostly ideas and plans
- For some things I will follow a "pedagogical" order, rather than a chronological order

#### Motivation

- HH→bbγγ analysis
  - Current Photon ID efficiency ~85%
  - 100% Ph ID → 15% more events → ~7% larger HH significance
- H→Zγ analysis: ~80%
- All analyses with photons which are statistically limited

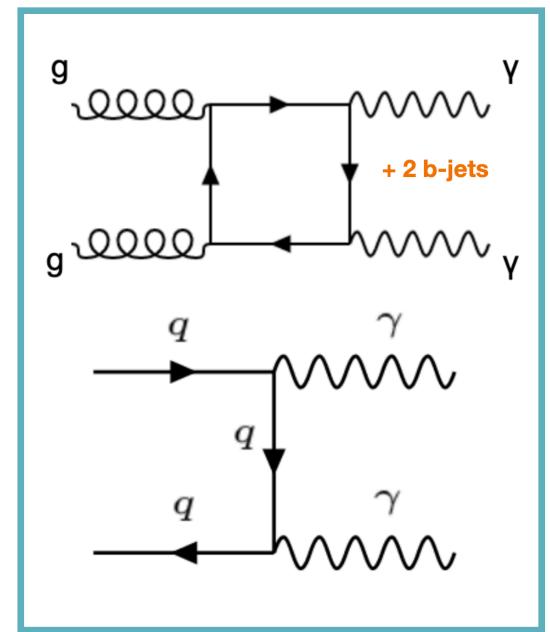


#### Cutflow of the Run2 HH→bbγγ analysis

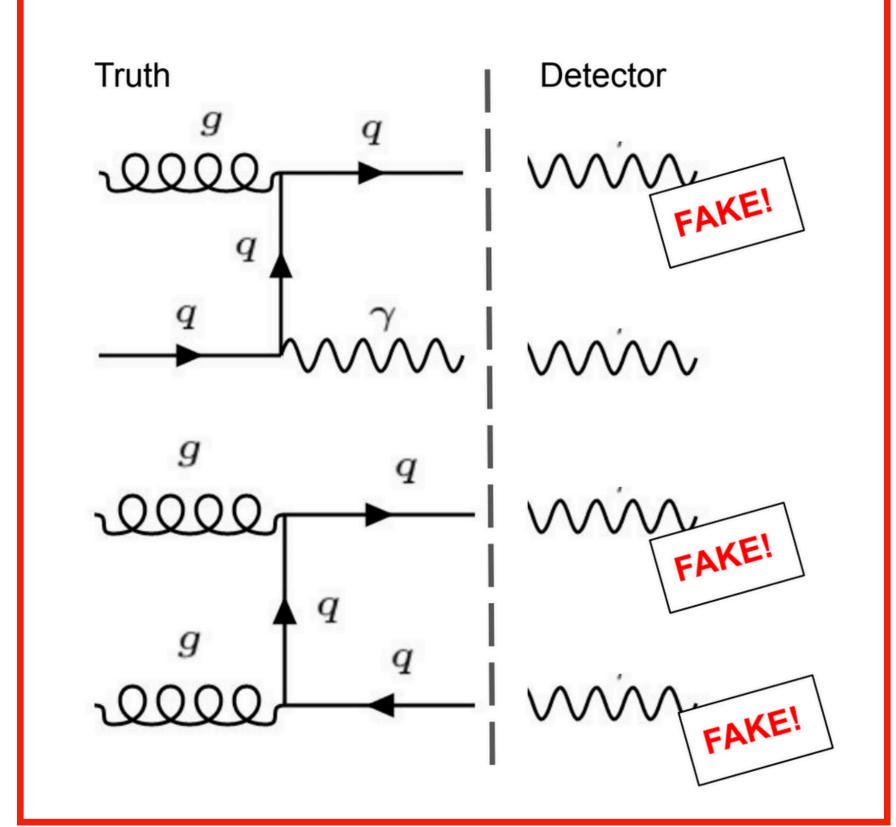
Cuts	SM ggF HH		
Cuts	Yield	Efficiency [%]	
All events	12.600	100.00	
Pass trigger	9.368	74.35	
Has primary vertex	9.368	74.35	
2 loose photons	6.875	54.57	
$e - \gamma$ ambiguity	6.872	54.54	
Trigger match	6.833	54.23	
Photons tight ID cut	5.974	47.41	
Photons isolation cut	5.332	42.31	
Rel. $p_T$ cuts	4.788	38.00	
$m_{\gamma\gamma} \in [105, 160] \text{ GeV}$	4.782	37.95	
$N_{\text{lep}} = 0$	4.759	37.77	
$N_{\rm jets} \geq 2$	4.221	33.50	
$N_{\rm central\ jets} < 6$	4.118	32.68	
2 b-jets with 77% WP	1.628	12.92	

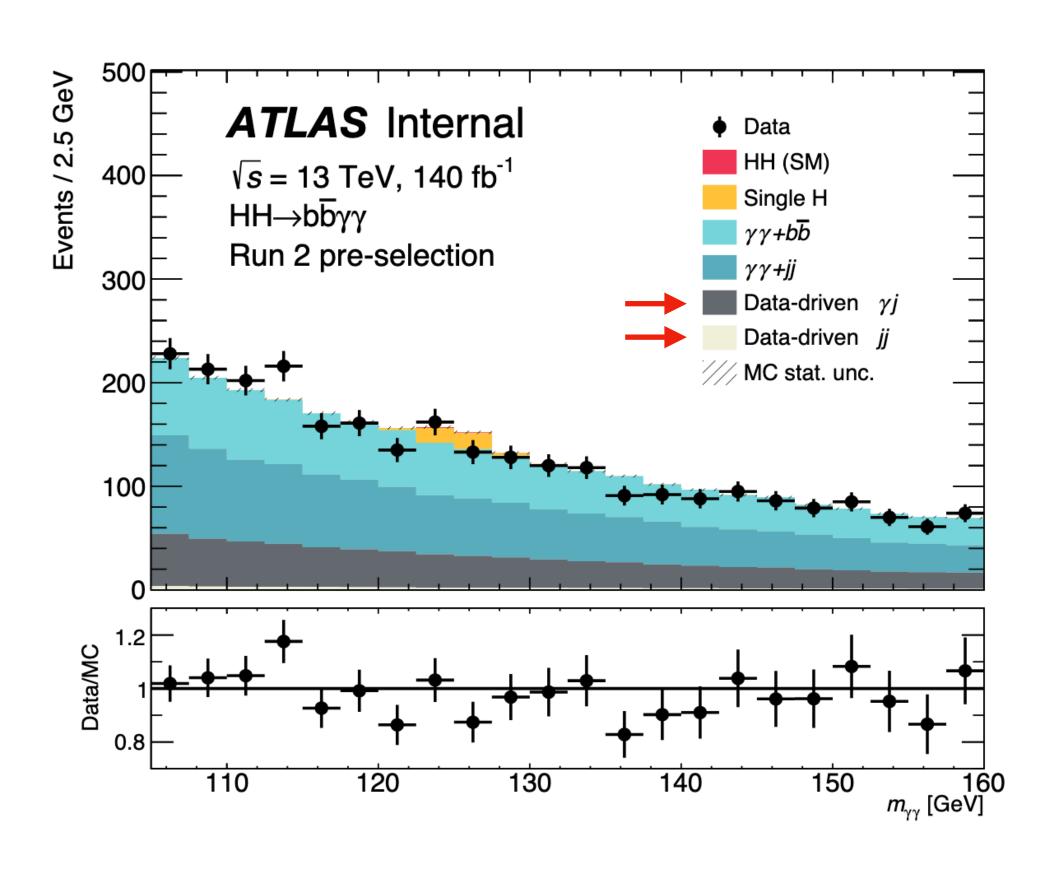
We want to improve this selection cut

#### Jets faking photons



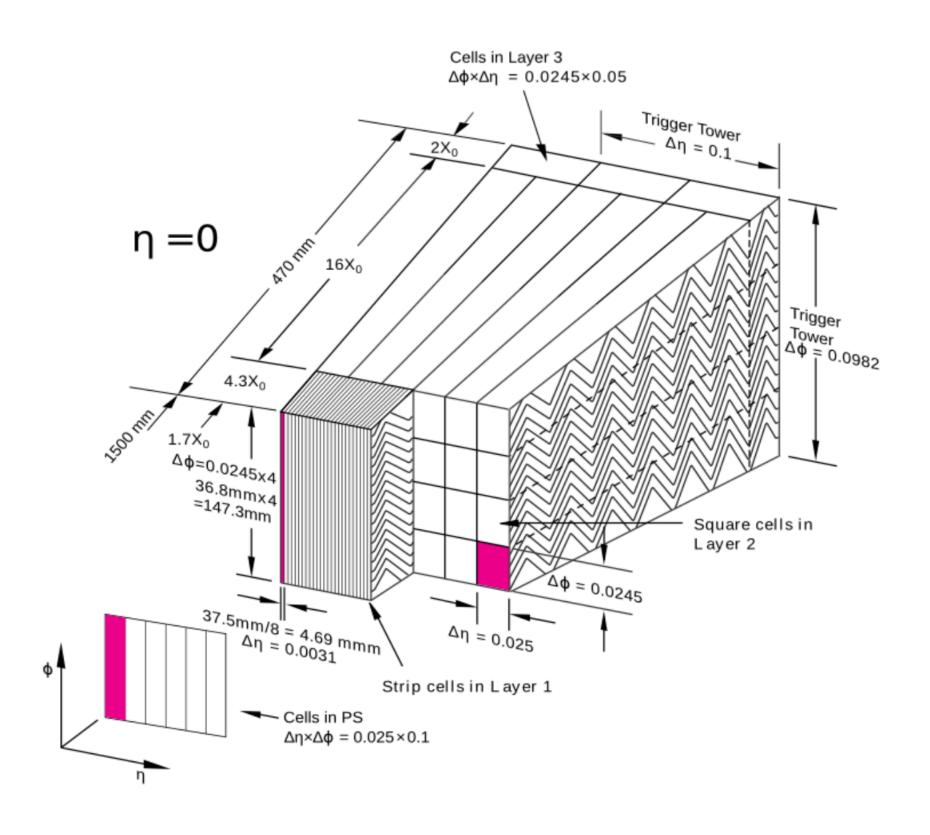
Irreducible γγ BKG





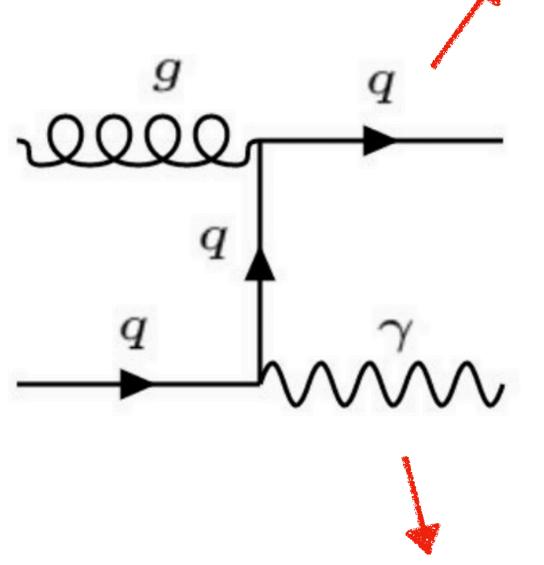
Reducible γj - jj BKG

#### Photons in the detector

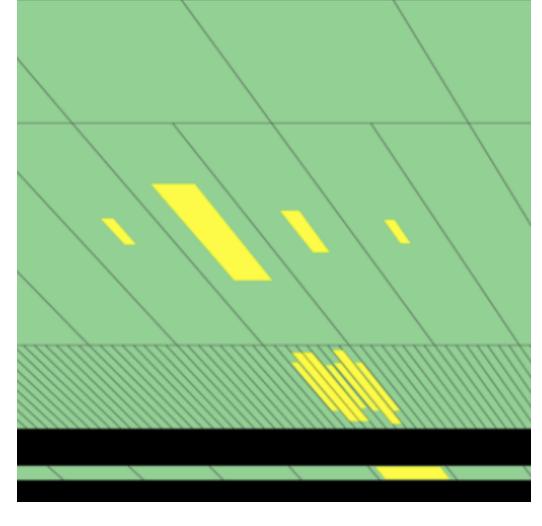


Section of the liquid Argon (LAr) electromagnetic calorimeter

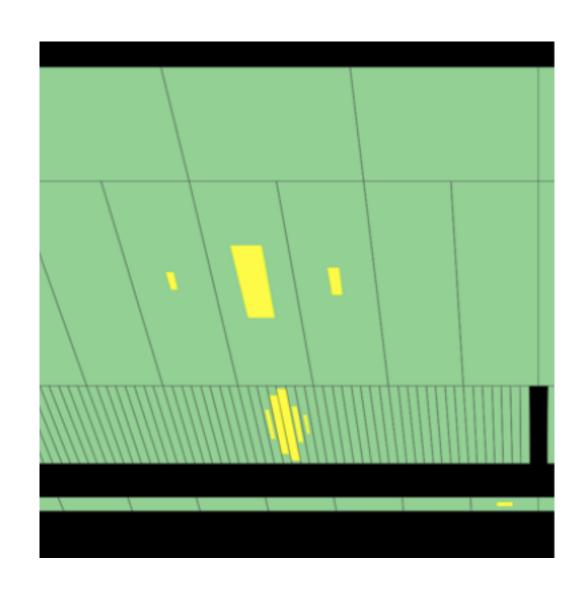




Photon from the **hard-scattering** will reach the calorimeter



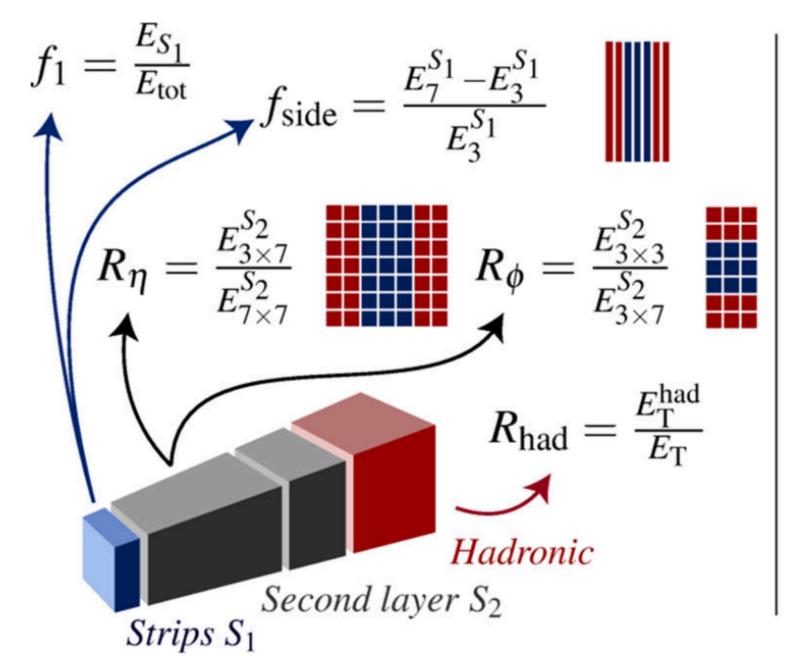
Non-prompt (fake) photon coming from  $\pi 0 \rightarrow \gamma \gamma$ 

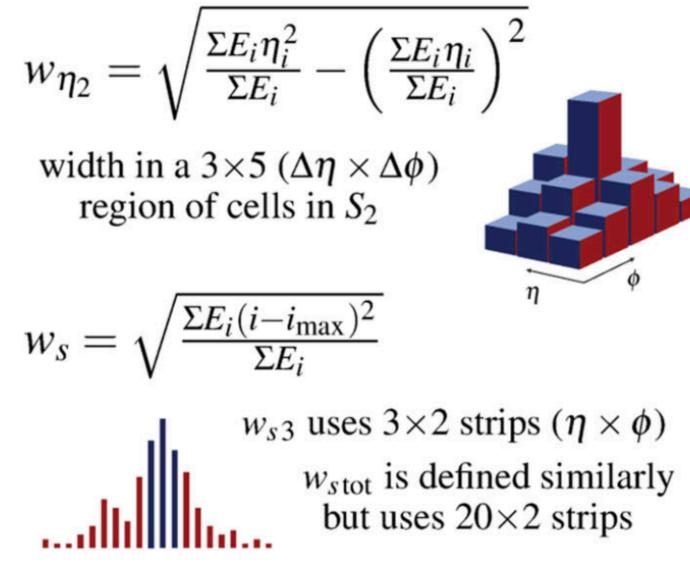


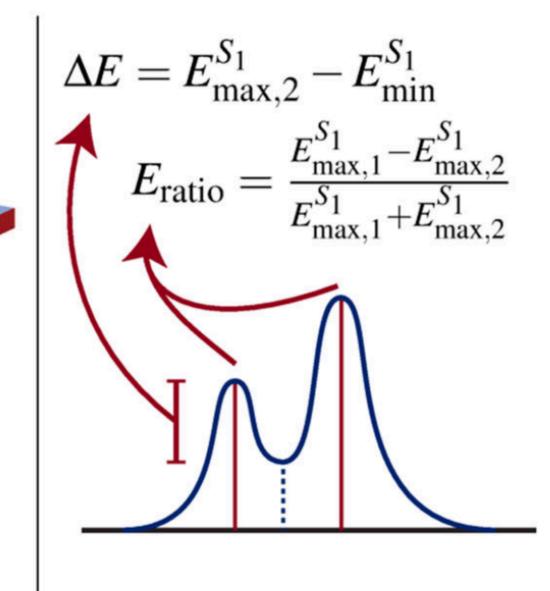
Prompt (real) photon in the EM calorimeter

#### Photon ID in ATLAS today

- Electromagnetic showers from real and fake photons have a different development in the calorimeter
- ~10 variables describing the properties of the <u>shower shape</u>

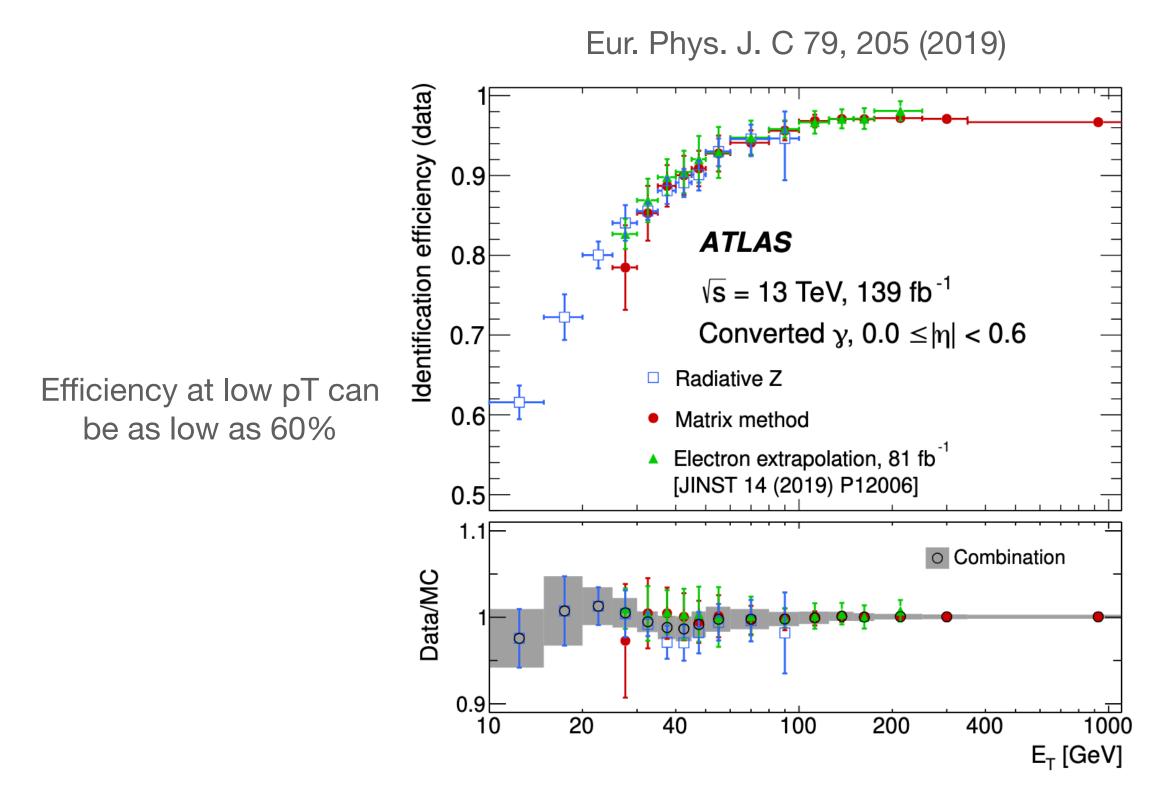


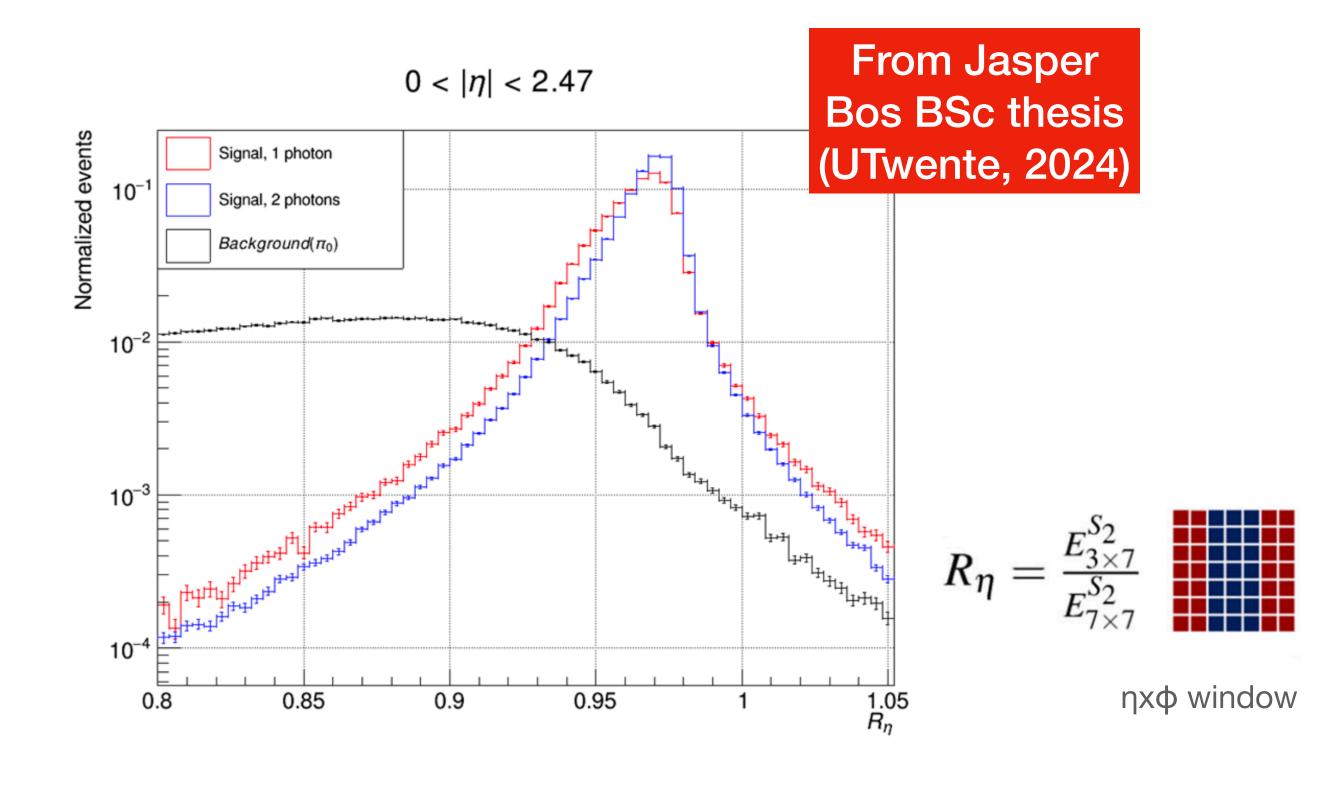




#### Photon ID in ATLAS today

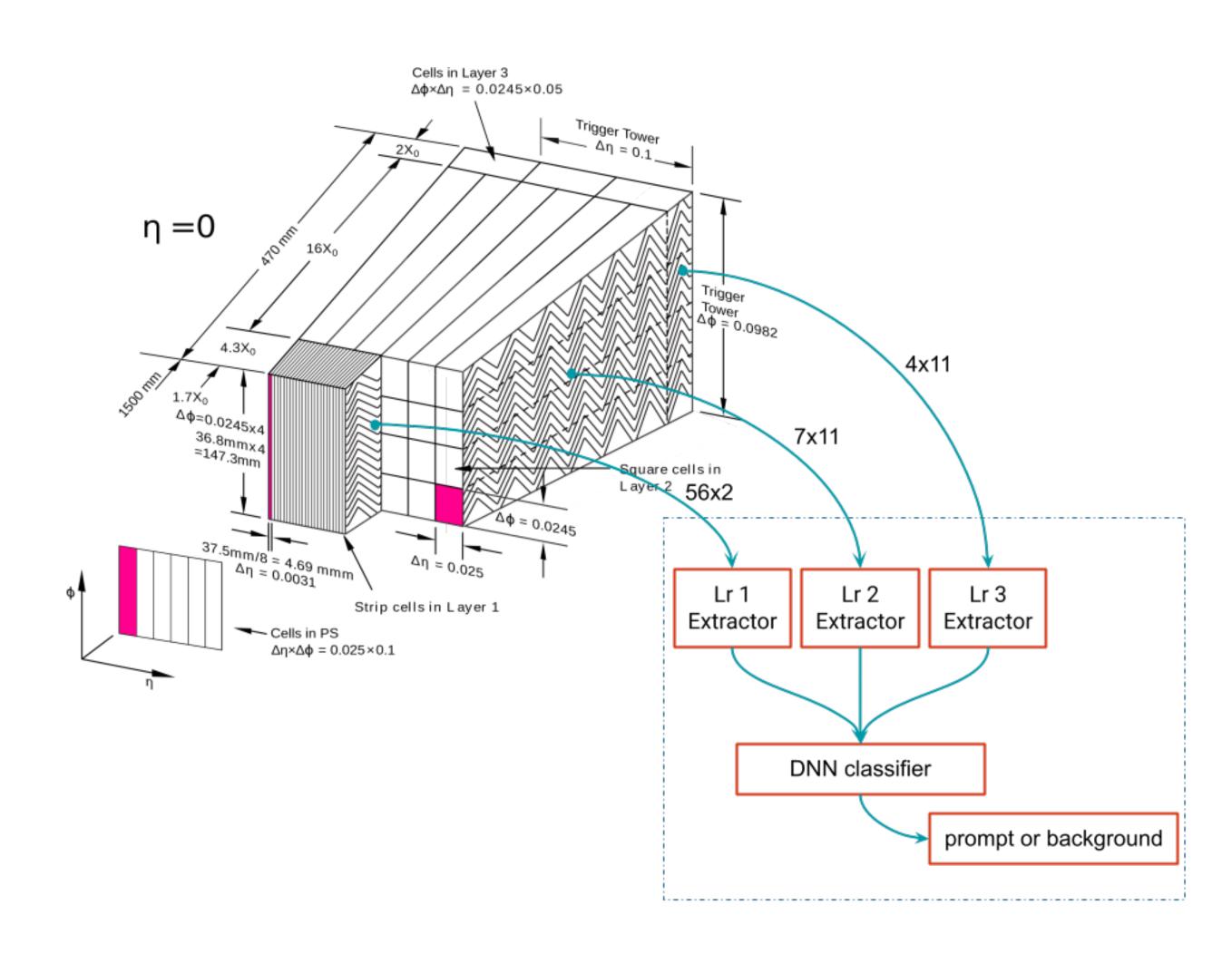
- These shower shape variables are used in a series of rectangular cuts (i.e. no Machine Learning)
- Optimization performed in different η-pT regions, separately for converted-unconverted photons



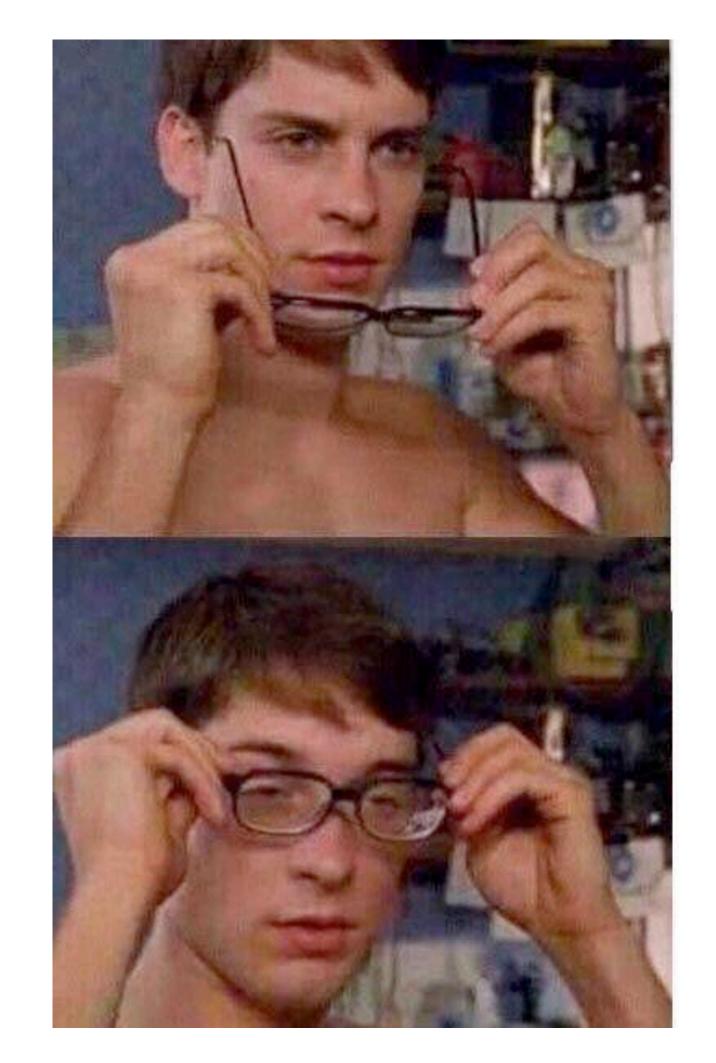


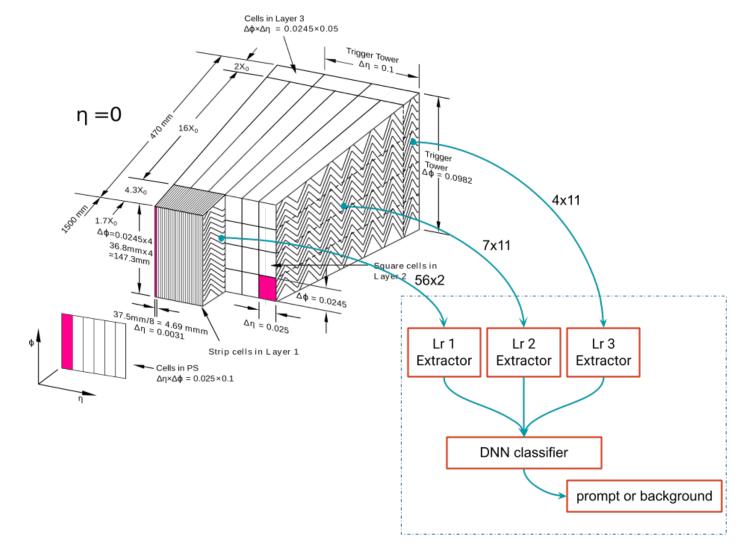
#### Photon ID in ATLAS tomorrow?

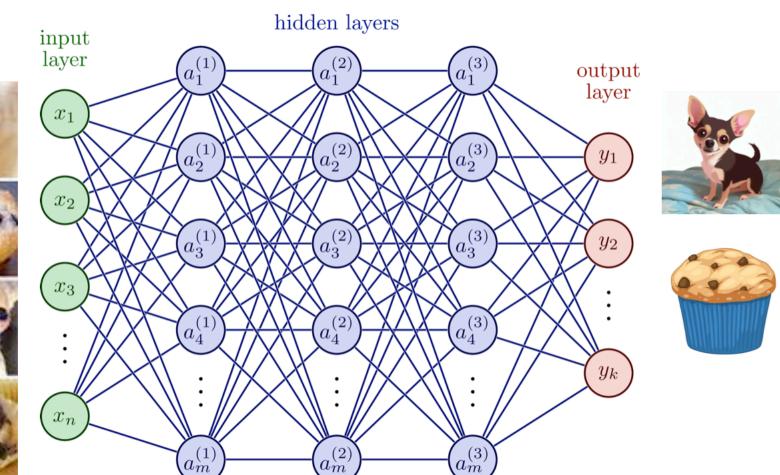
- Exploit Machine Learning
- Drop shower shape (high-level) variables
- Feed 2D (ηxφ) energy clusters to a convolutional neural network (CNN)
- Low-level input: LAr calorimeter cells (3 layers)
- Classification task with prompt photon as signal and non-prompt photons as background



#### Photon ID in ATLAS tomorrow?







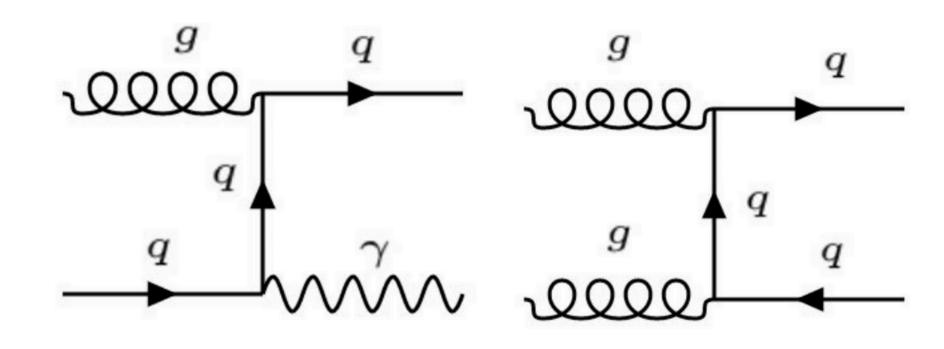
Photon ID becomes an image recognition problem with machine learning

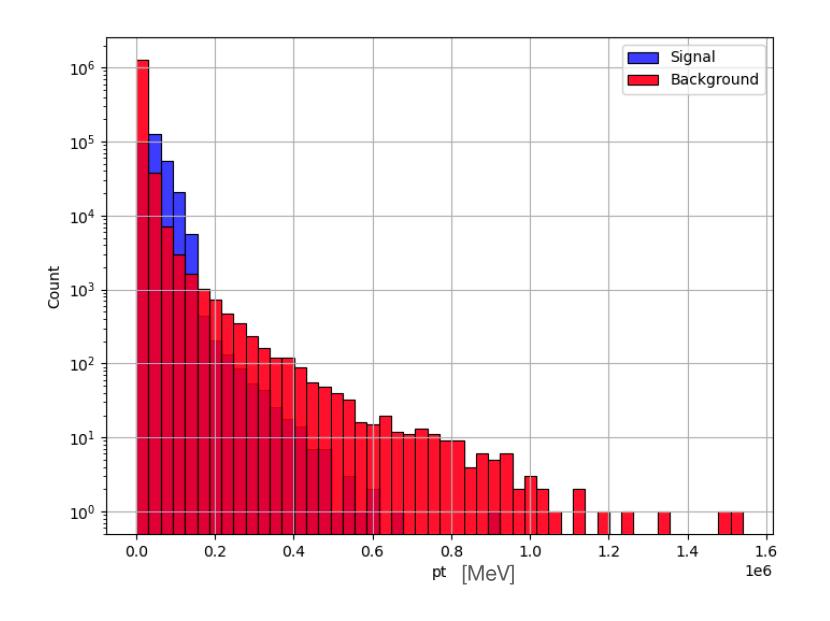


A revolution

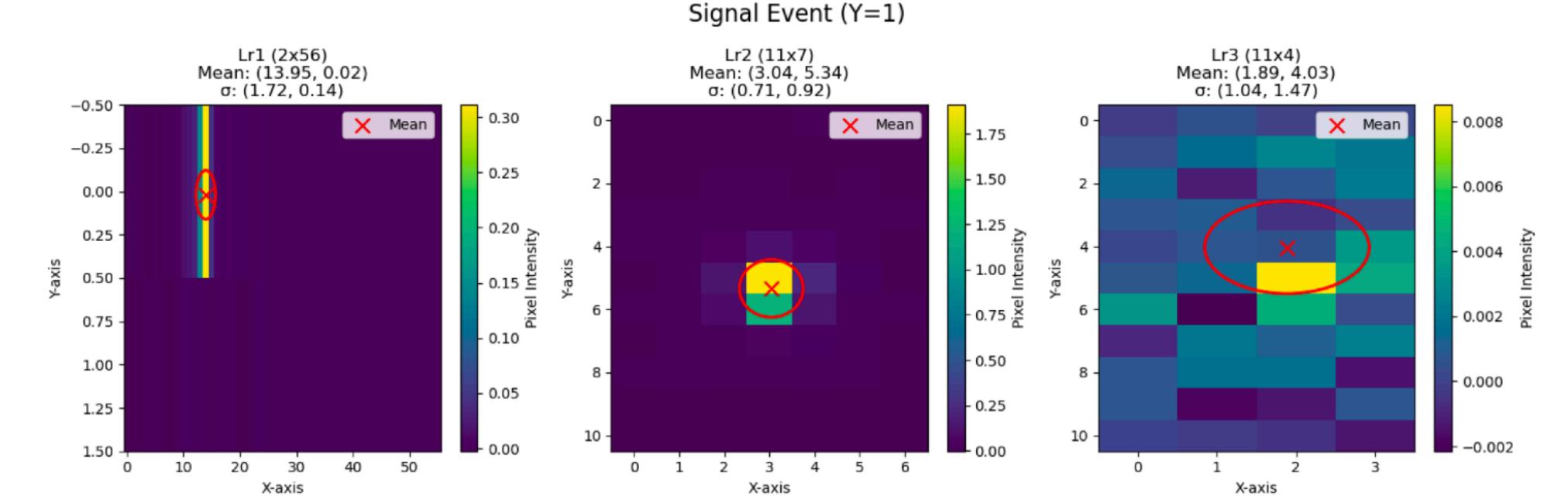
#### Design of a CNN for Photon ID

- Training with Monte Carlo (MC) simulations of γj and jj processes
- Truth MC label (prompt/non-prompt) used as class label
- Photon pT between ~1 GeV and ~1 TeV
- pT-reweighting in the training to compensate for class imbalance
- 1.5M entries: 80% Training (of which 33% validation), 20% Test

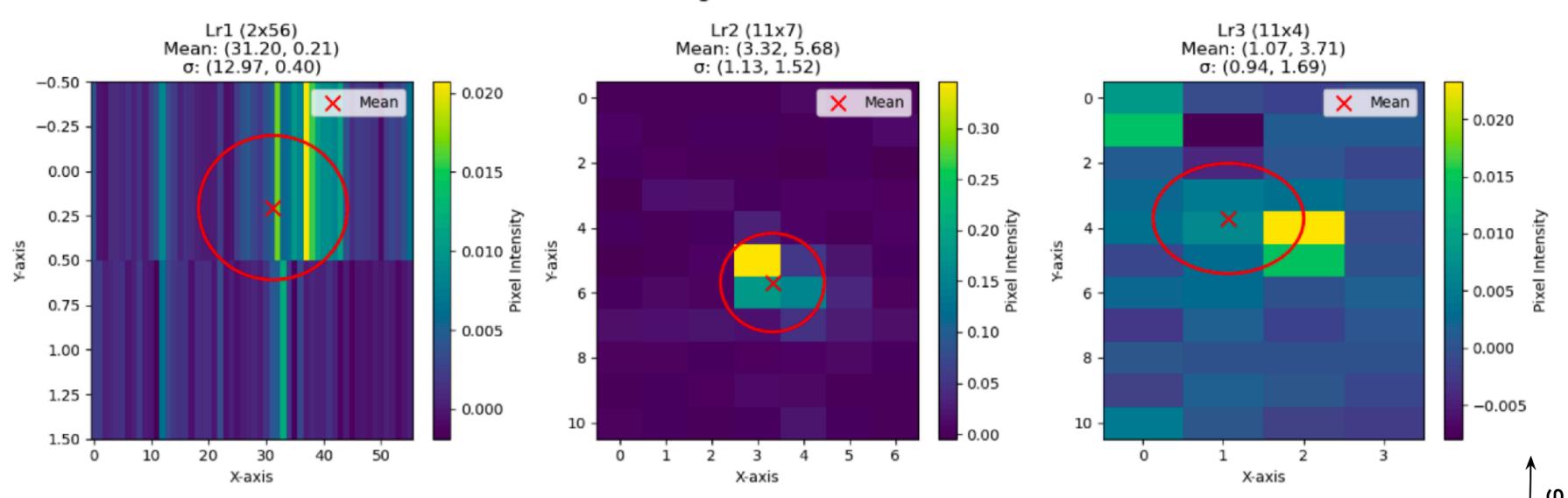




## Input features

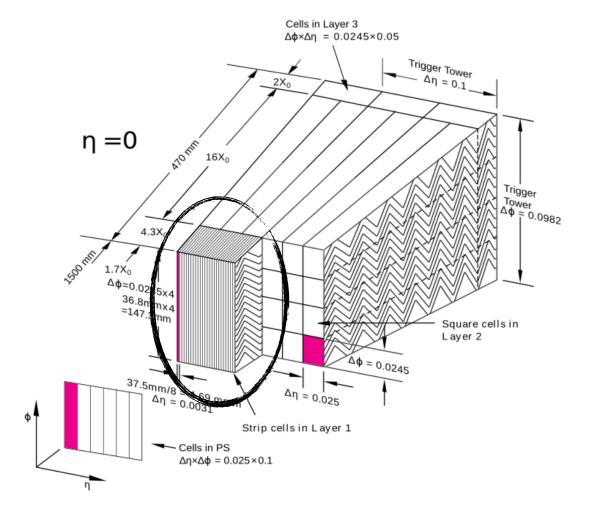




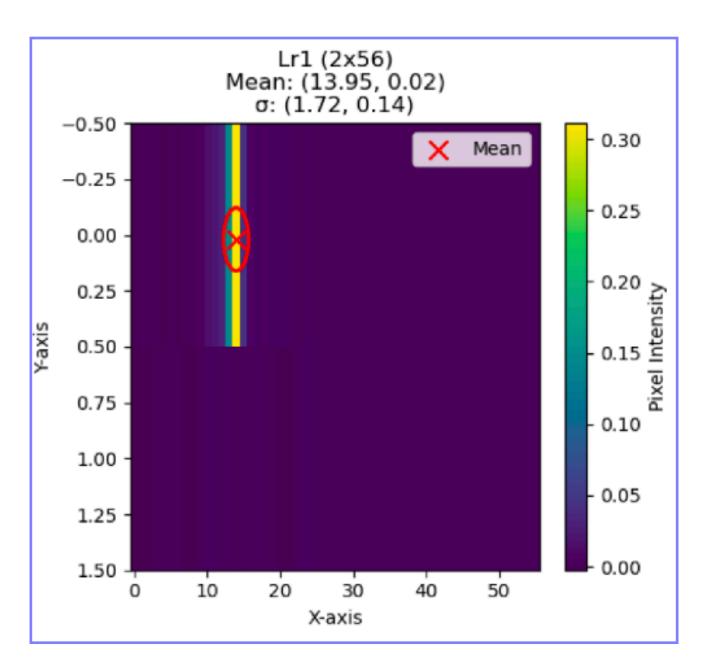


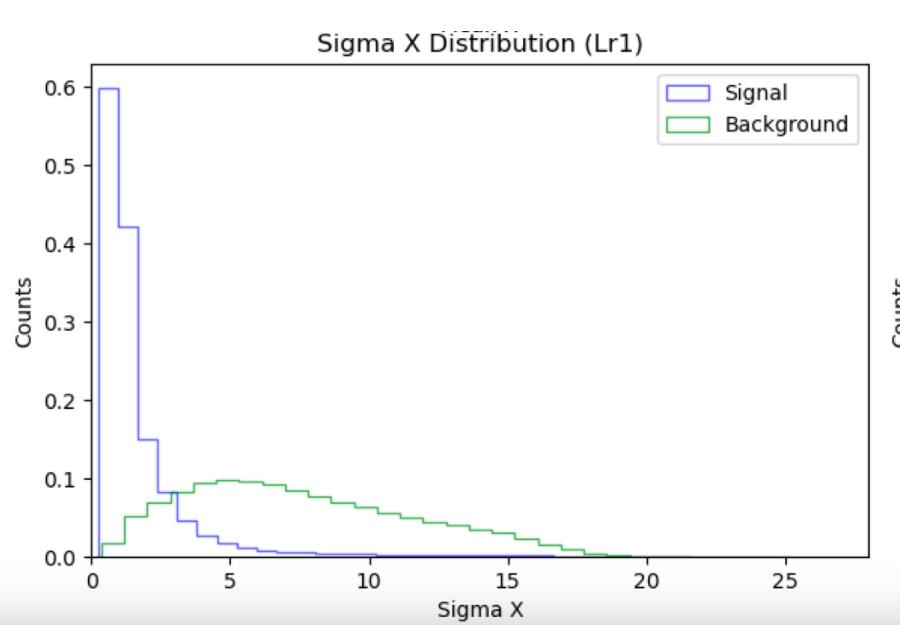
## Input features

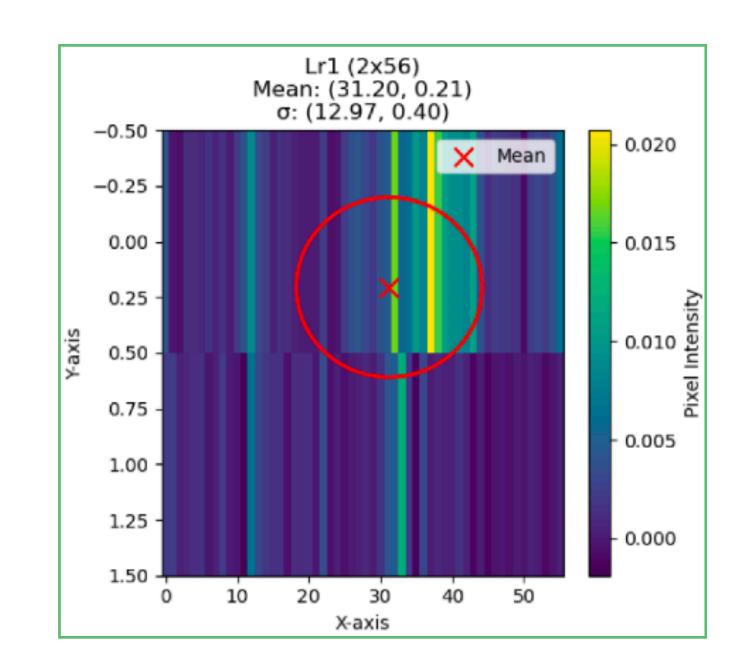
For one entry →

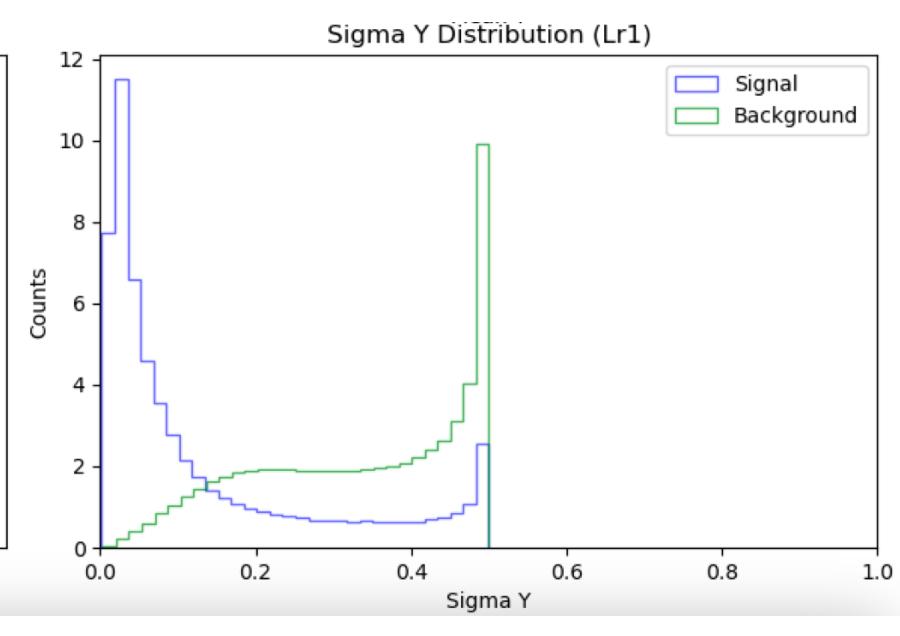


For all entries →

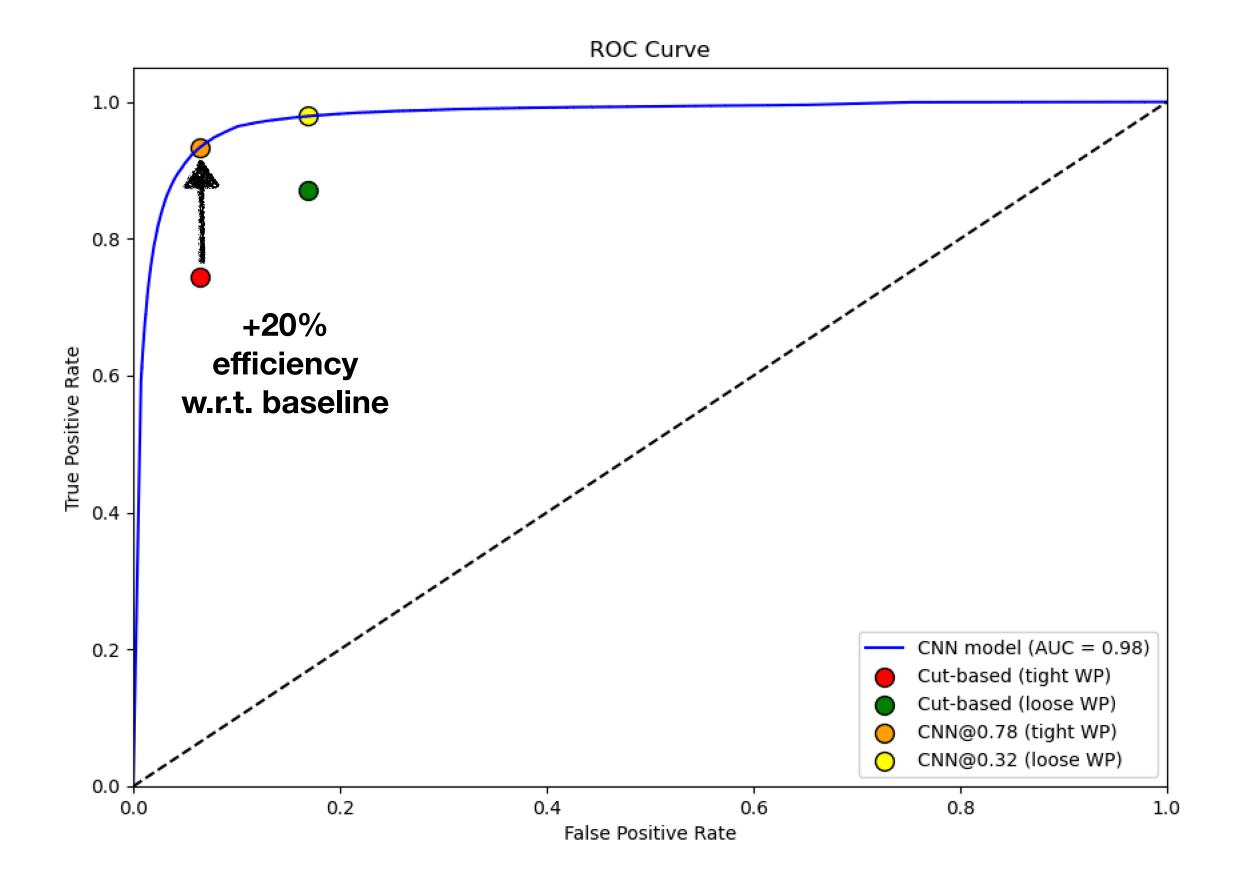




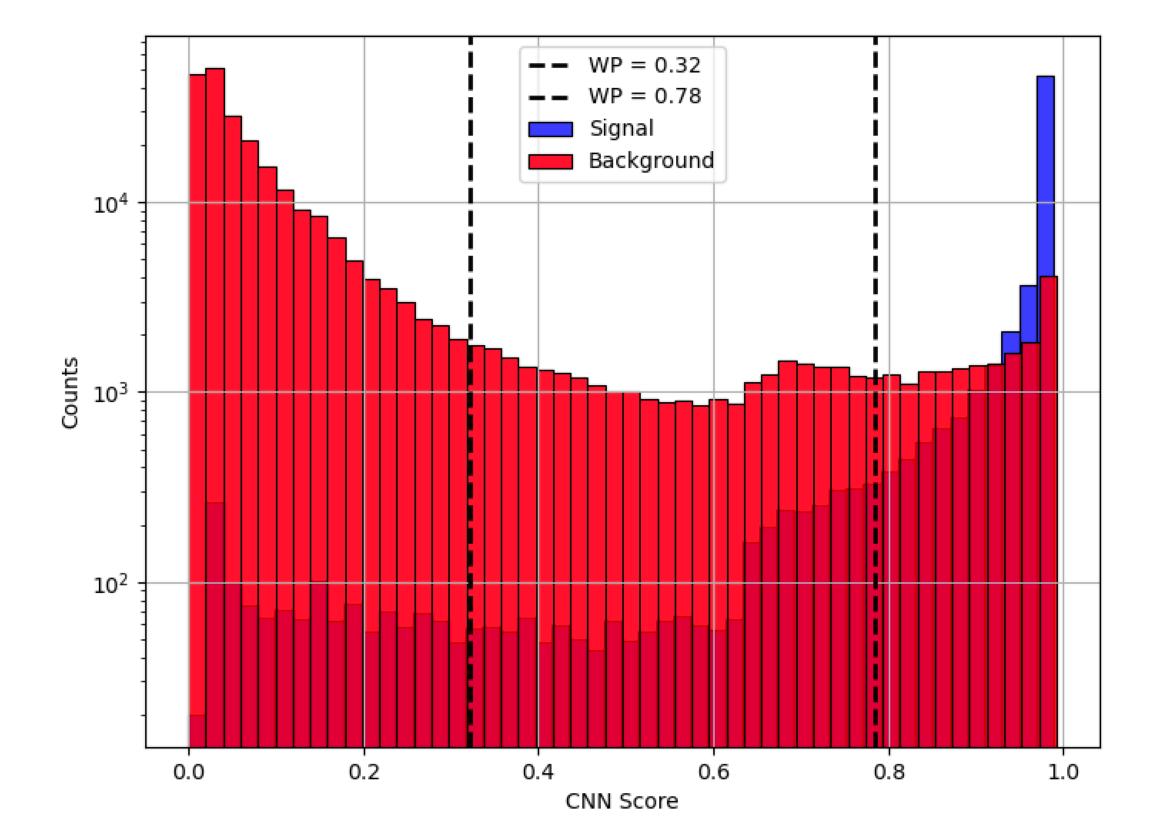




#### CNN performance



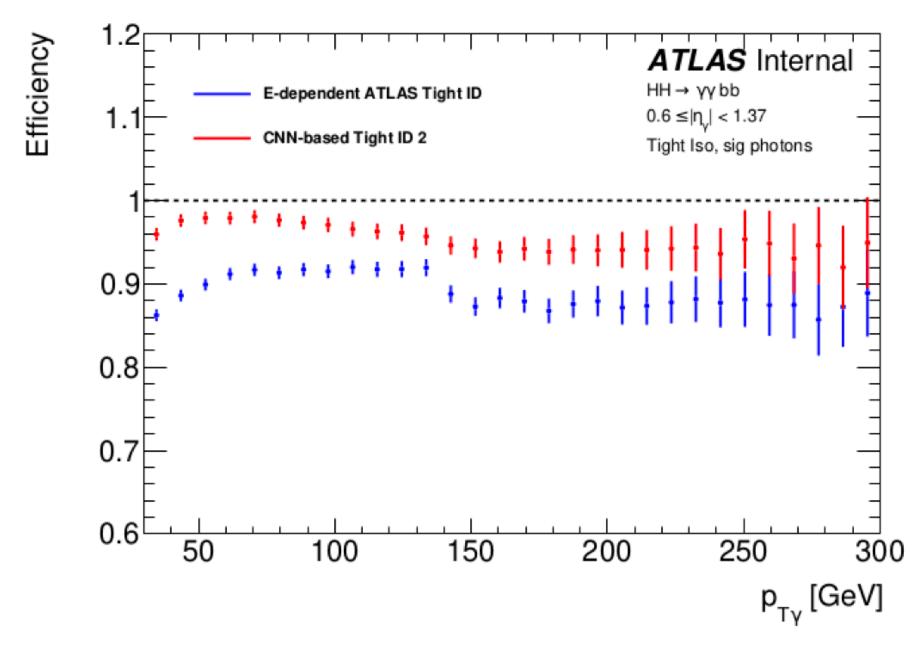
<u>DISCLAIMER</u>: for a better comparison, WPs will have to be defined in regions of pT,  $\eta$ , conversion status

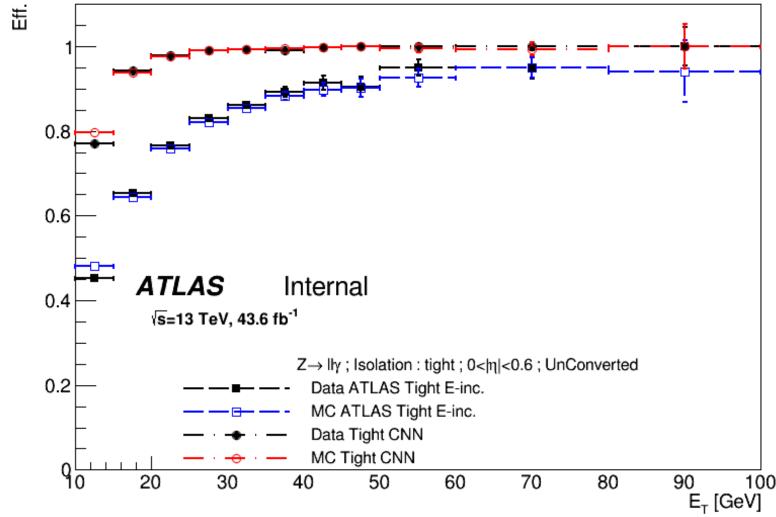


- Two working points (WP) defined on the CNN score:
  - Loose for higher signal efficiency
  - Tight for higher background rejection
- Both match the bkg rej. of cut-based method

#### Application to physics analyses

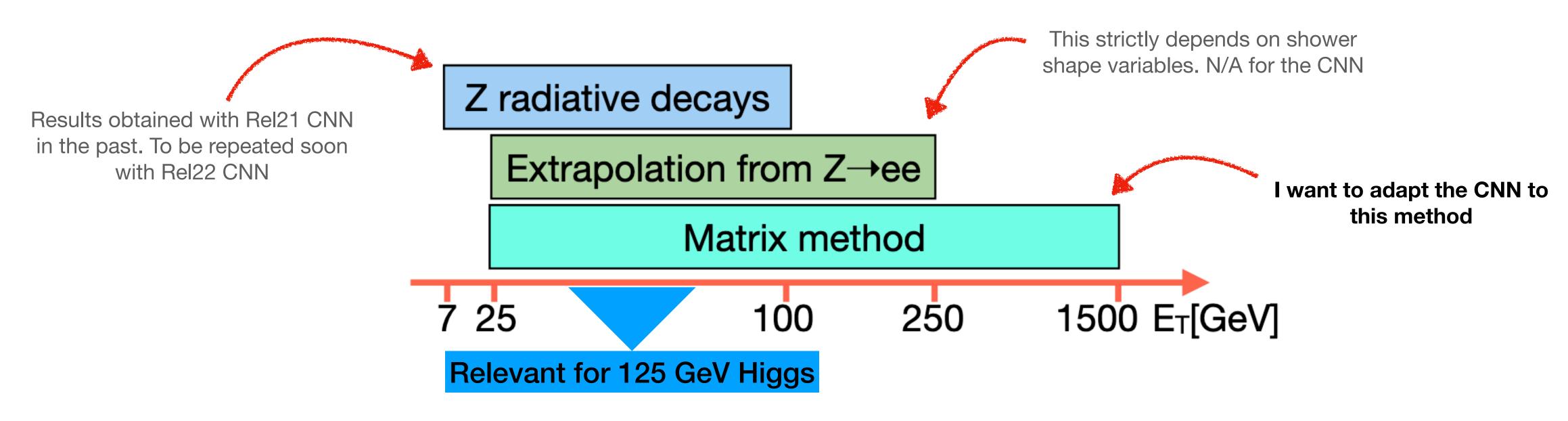
- CNN in HH→bbγγ
  - up to 8% more HH signal in selection
  - final results (HH limit) improved by 5% (GN2 brought ~8%)
- CNN in Z→IIγ (Z radiative decays used for calibration)
  - x2 improvement in efficiency at low pT w.r.t. baseline
  - Out-of-sample validation: CNN trained on inclusive photon production and applied to Z→IIγ





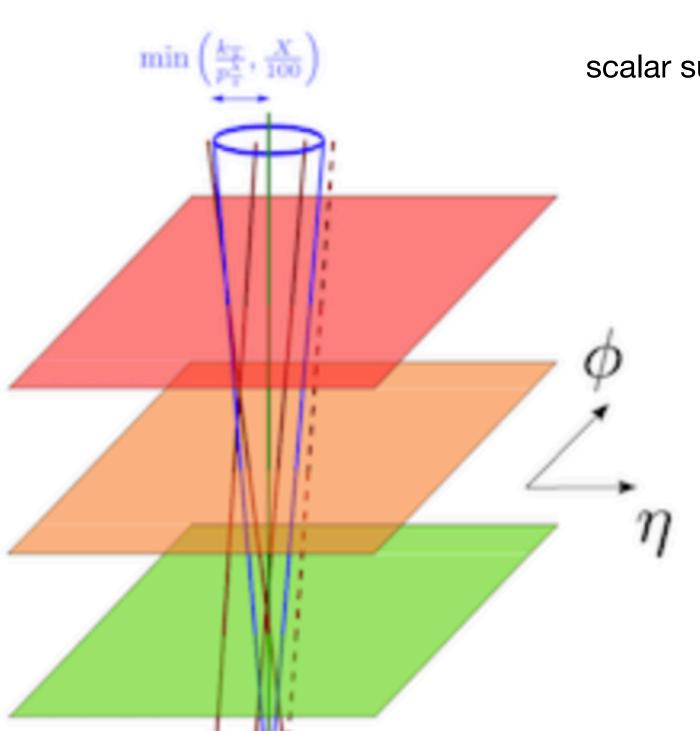
#### The real challenge: data/MC calibration

- The CNN shows promising results when applied to MC
- What about ID efficiencies in <u>data</u>?
- In EGamma there are 3 methods for the measurement of the ID efficiencies in data: all of them were designed for a non-ML-based ID



# Existing calibration methods rely on the assumption that ID and isolation are weakly correlated

#### Isolation variables



#### **Track-based isolation**

scalar sum of pT of all the tracks found inside the isolation cone of size XX

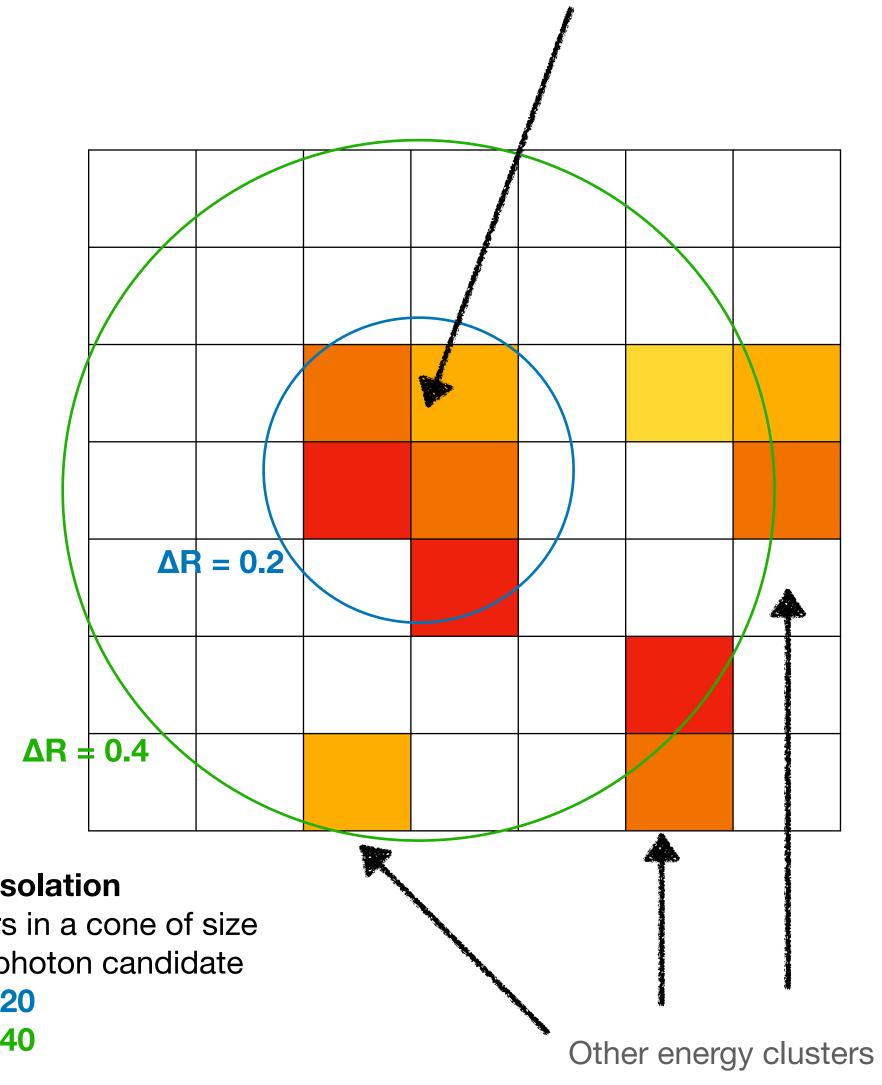
> Ptcone20 Ptcone40

Recommended (loose) WP: used in the bbyy analysis and others...

$$E_{\mathrm{T}}^{\mathrm{iso}}\Big|_{\Delta R < 0.2} < 0.065 \cdot E_{\mathrm{T}} \quad \mathrm{and} \quad p_{\mathrm{T}}^{\mathrm{iso}}\Big|_{\Delta R < 0.2} < 0.05 \cdot E_{\mathrm{T}}$$

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Energy cluster associated to photon candidate

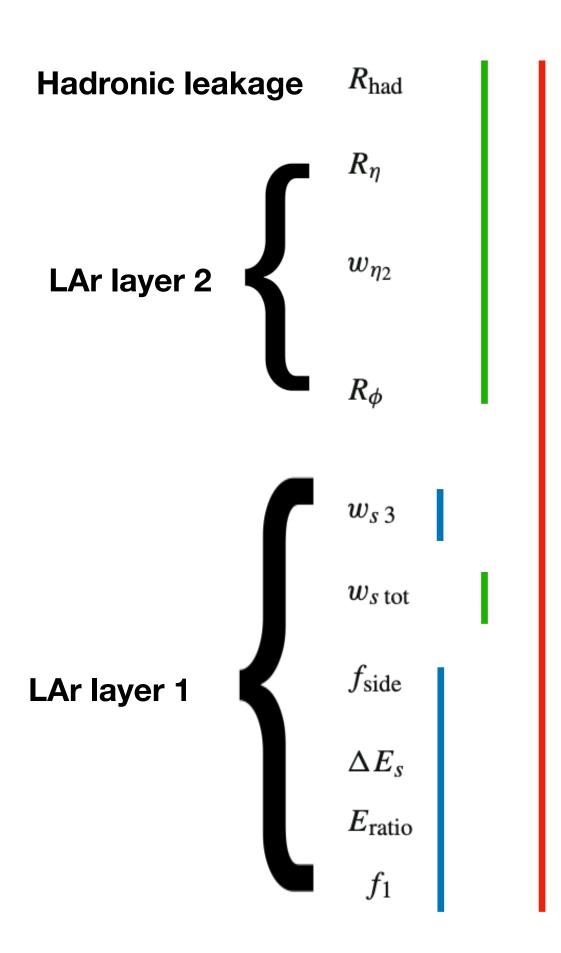


#### **Calo-based isolation**

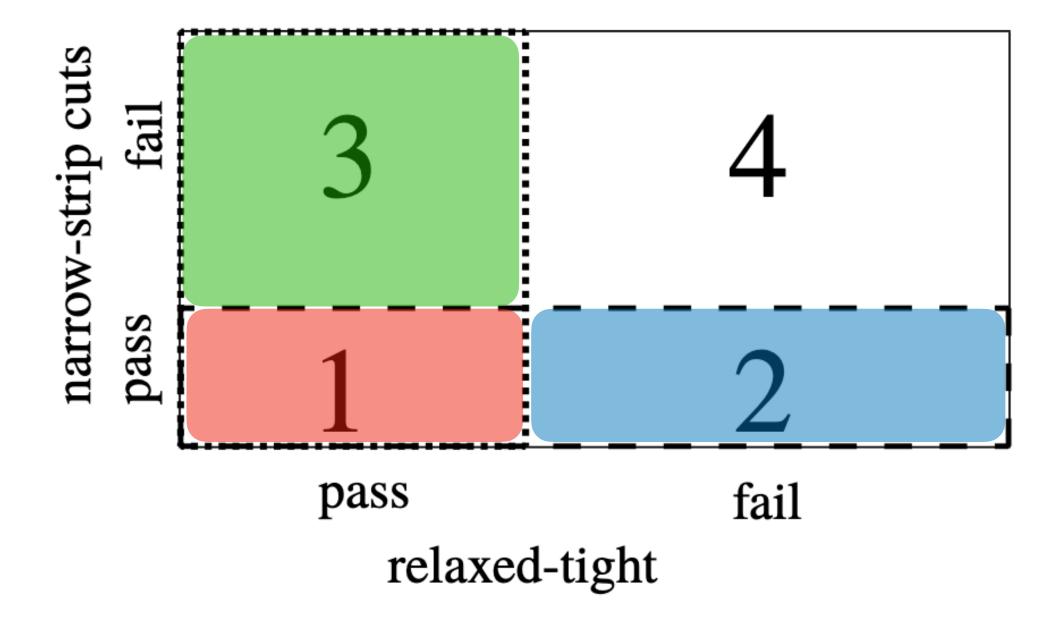
Sum of ET of all clusters in a cone of size XX, after removig the photon candidate

Etcone20 Etcone40

#### Efficiency measurement with Matrix Method



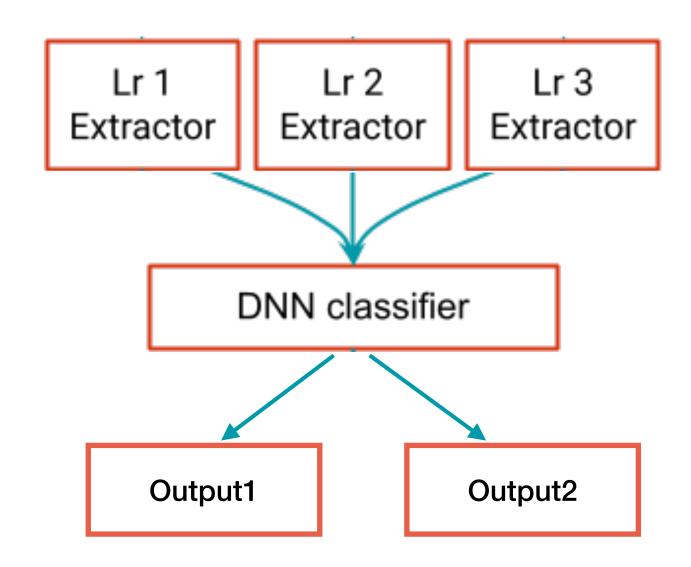
- Shower shapes are split into two subsets: narrow-strip and relaxed-tight
- These are used to define 4 regions
- Track-Isolation used to estimate efficiency in each region



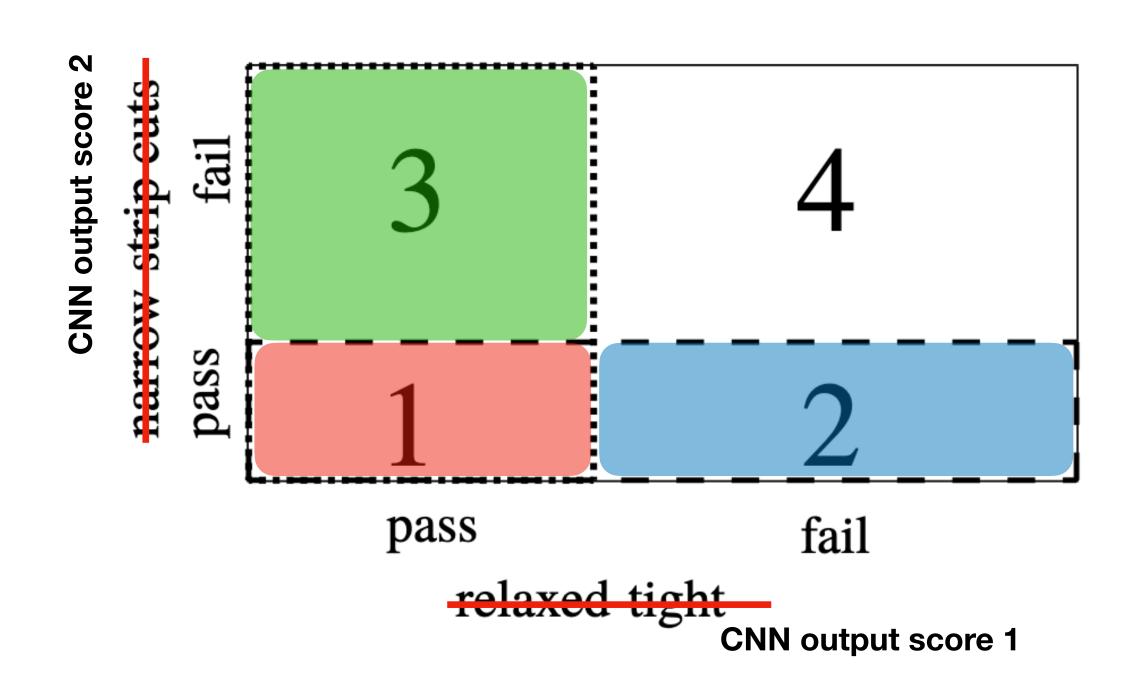
	cut-based	feature
1 must pass	all shower shapes	best performance
2 must pass	narrow-strip variables	weak correlation with Iso
3 must pass	relaxed-tight variables	orthogonal (?) to narrow-strip

#### Matrix Method with the CNN

- Idea: design a multi-output CNN to "mimic" narrow-strip and relaxed-tight cuts
- Define a separate CNN output which is explicitly decorrelated w.r.t. track-isolation in the loss



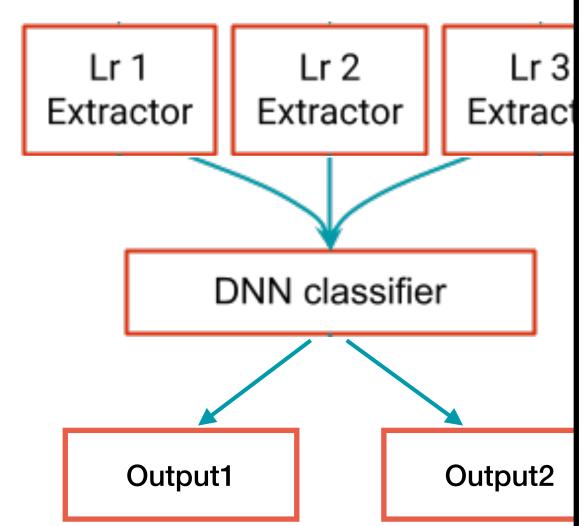
$$\mathcal{L} = \sum_{i=1}^{2} BCE(Out_i, y) + \lambda Decorr(Out_2, ptcone40)$$

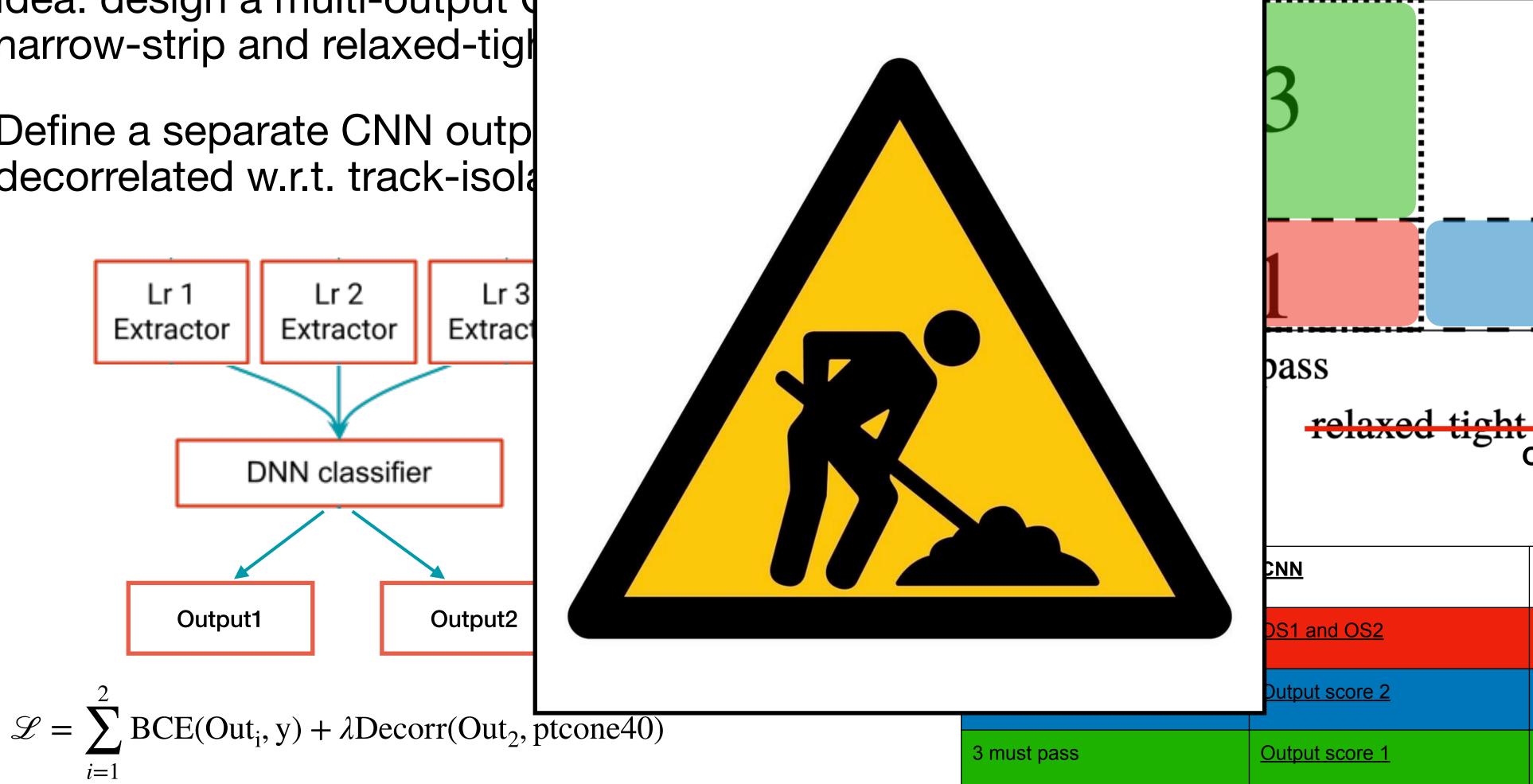


	CNN	feature
1 must pass	OS1 and OS2	best performance
2 must pass	Output score 2	weak correlation with Iso
3 must pass	Output score 1	orthogonal (?) to OS2

#### Matrix Method with the CNN

- Idea: design a multi-output ( narrow-strip and relaxed-tight
- Define a separate CNN outpl decorrelated w.r.t. track-isola





3 must pass

fail

feature

Output score 1

**CNN** output score 1

best performance

weak correlation with Iso

orthogonal (?) to OS2

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#### Conclusions

- My mission for the next ~2 years (hopefully in time for full Run3 EGamma recommendations): revolution of photon ID in ATLAS
- Progress is slow (CNN requires dedicated MC sample production with EM cells) but we are getting there!
- Potential gain for a wide range of physics analyses, mainly searches with photons
- Is it a low-hanging fruit? Not really, data/MC calibration is difficult
- Multi-dimensional project: ML developments, design of calibration method, interplay with isolation, application to analyses
- I will present the status at the next ATLAS EGamma workshop in a few weeks

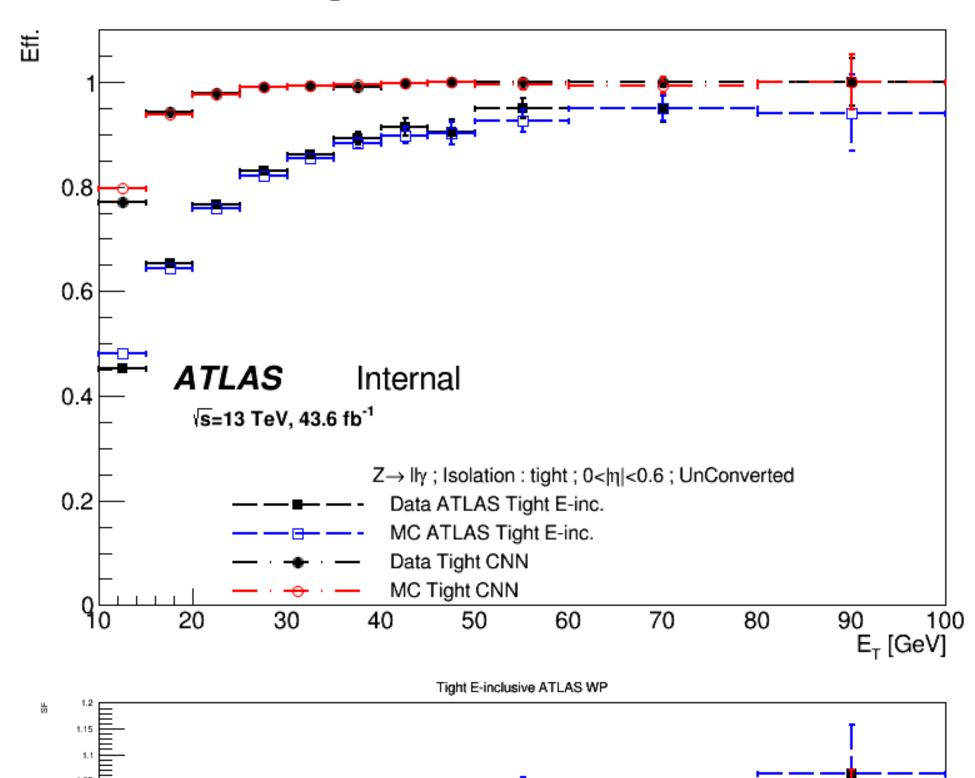
### Back-up

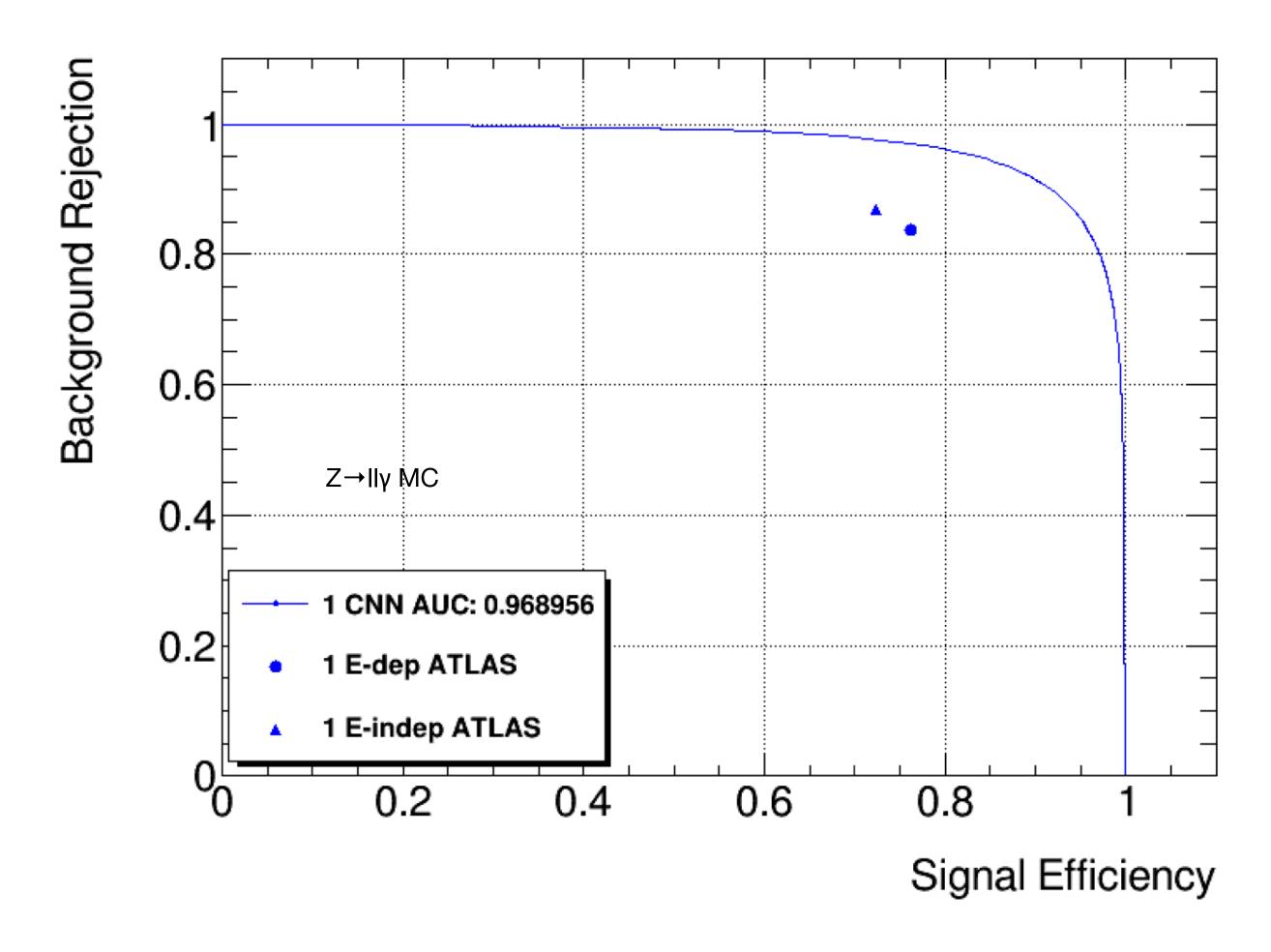
 Table 1 Discriminating variables used for loose and tight photon identification

Category	Description	Name	loose	tight
Acceptance	$ \eta  < 2.37$ , with $1.37 \le  \eta  < 1.52$ excluded	_	✓	<b>√</b>
Hadronic leakage	Ratio of $E_{\rm T}$ in the first sampling layer of the hadronic calorimeter to $E_{\rm T}$ of the EM cluster (used over the range $ \eta  < 0.8$ or $ \eta  > 1.52$ )		✓	✓
	Ratio of $E_{\rm T}$ in the hadronic calorimeter to $E_{\rm T}$ of the EM cluster (used over the range $0.8 <  \eta  < 1.37$ )	$R_{\rm had}$	✓	$\checkmark$
EM middle layer	Ratio of the energy in $3 \times 7 \eta \times \phi$ cells over the energy in $7 \times 7$ cells centered around the photon cluster position	$R_{\eta}$	✓	$\checkmark$
Lateral shower width, $\sqrt{(\Sigma E_i \eta_i^2)/(\Sigma E_i) - ((\Sigma E_i \eta_i)/(\Sigma E_i))^2}$ , where $E_i$ the energy and $\eta_i$ is the pseudorapidity of cell $i$ and the sum is calculated within a window of $3 \times 5$ cells		$w_{\eta_2}$	✓	✓
	Ratio of the energy in $3 \times 3 \eta \times \phi$ cells over the energy of $3 \times 7$ cells centered around the photon cluster position	$R_{oldsymbol{\phi}}$		$\checkmark$
EM strip layer	Lateral shower width, $\sqrt{(\Sigma E_i(i-i_{\max})^2)/(\Sigma E_i)}$ , where <i>i</i> runs over all strips in a window of $3\times 2$ $\eta\times \phi$ strips, and $i_{\max}$ is the index of the highest-energy strip calculated from three strips around the strip with maximum energy deposit	$w_{s3}$		✓
	Total lateral shower width $\sqrt{(\Sigma E_i(i-i_{\max})^2)/(\Sigma E_i)}$ , where <i>i</i> runs over all strips in a window of $20 \times 2 \eta \times \phi$ strips, and $i_{\max}$ is the index of the highest-energy strip measured in the strip layer	$w_{s  ext{ tot}}$		✓
	Energy outside the core of the three central strips but within seven strips divided by energy within the three central strips	$f_{ m side}$		$\checkmark$
	Difference between the energy associated with the second maximum in the strip layer and the energy reconstructed in the strip with the minimum value found between the first and second maxima	$\Delta E_s$		✓
	Ratio of the energy difference between the maximum energy deposit and the energy deposit in the secondary maximum in the cluster to the sum of these energies	$E_{ m ratio}$		✓
	Ratio of the energy in the first layer to the to the total energy of the EM cluster	$f_1$		$\checkmark$

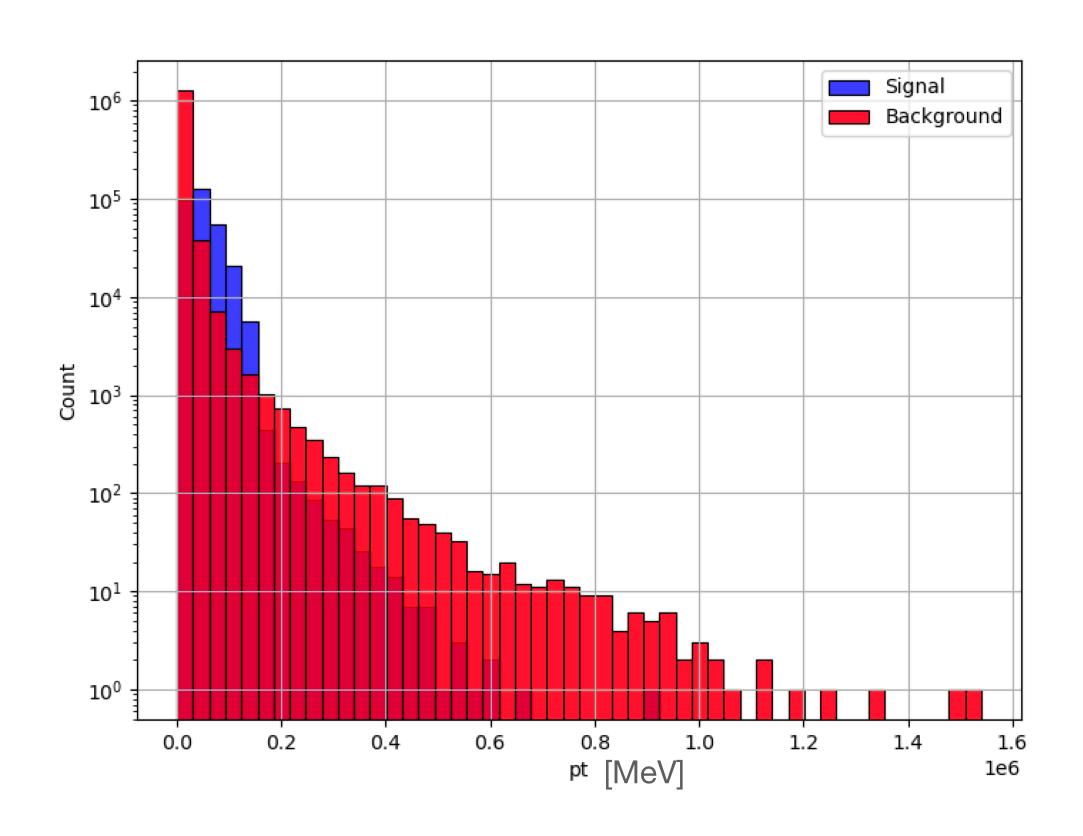
#### **CNN** with radiative Z

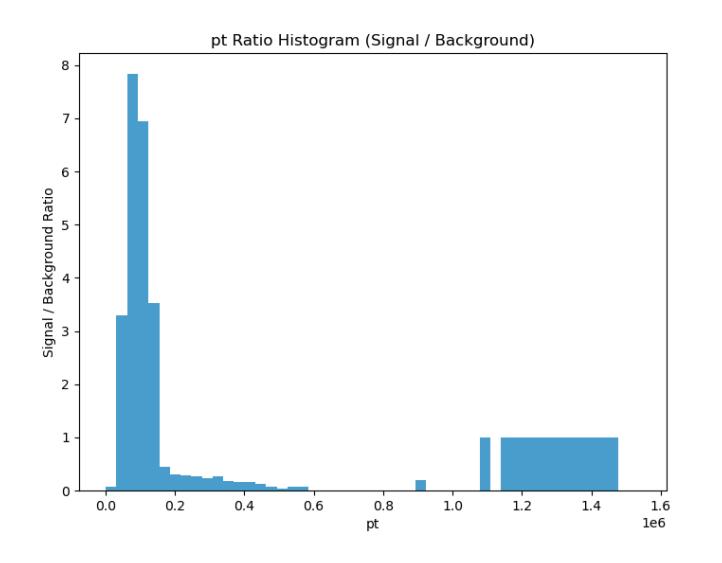
Tight E-inclusive ATLAS WP

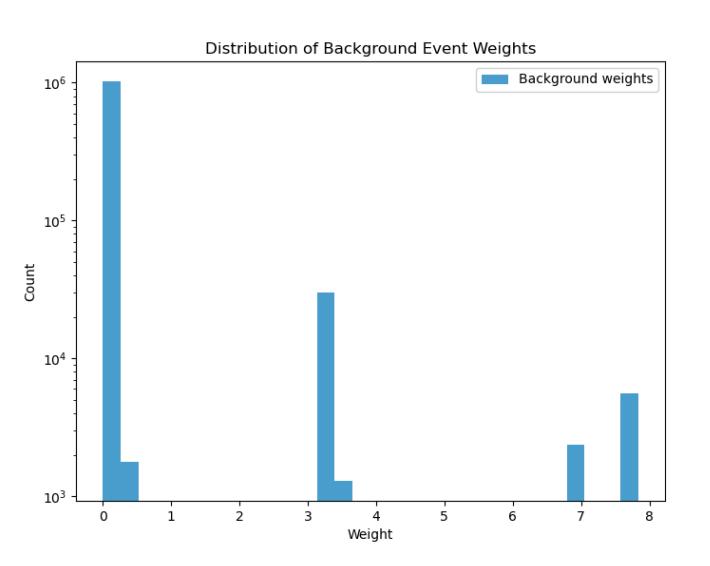




#### pT reweighting



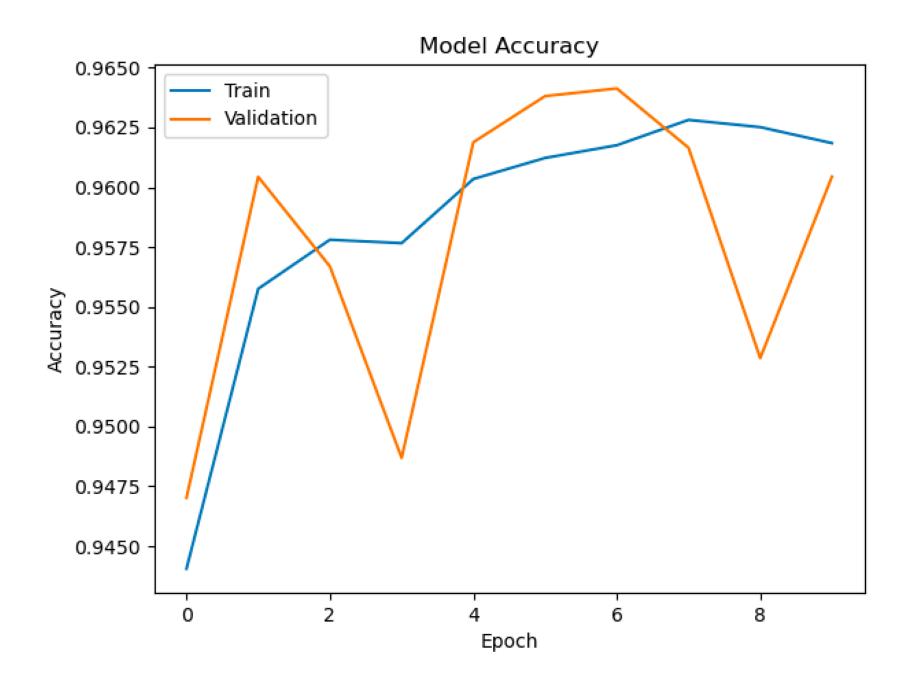


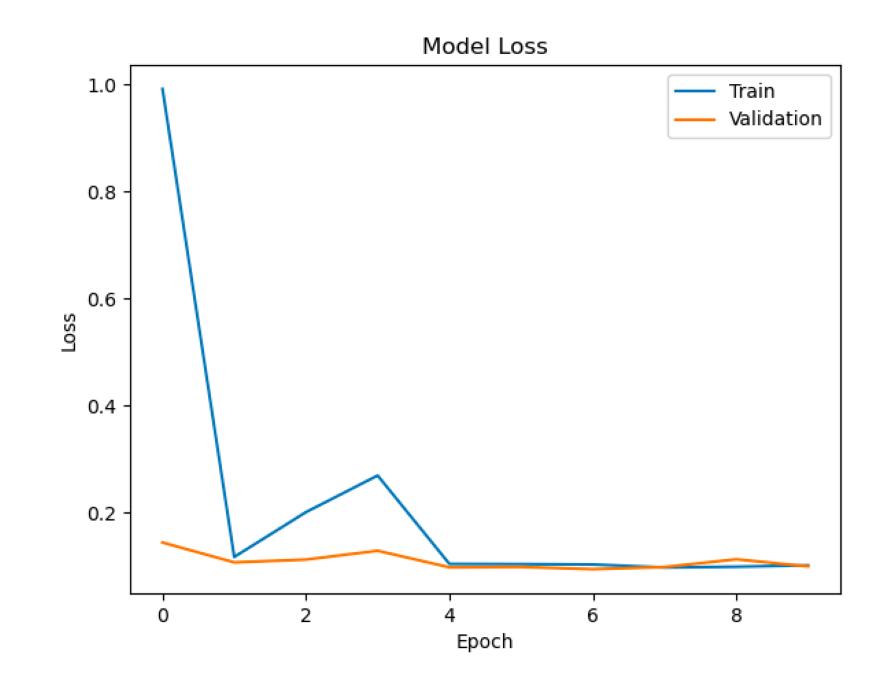


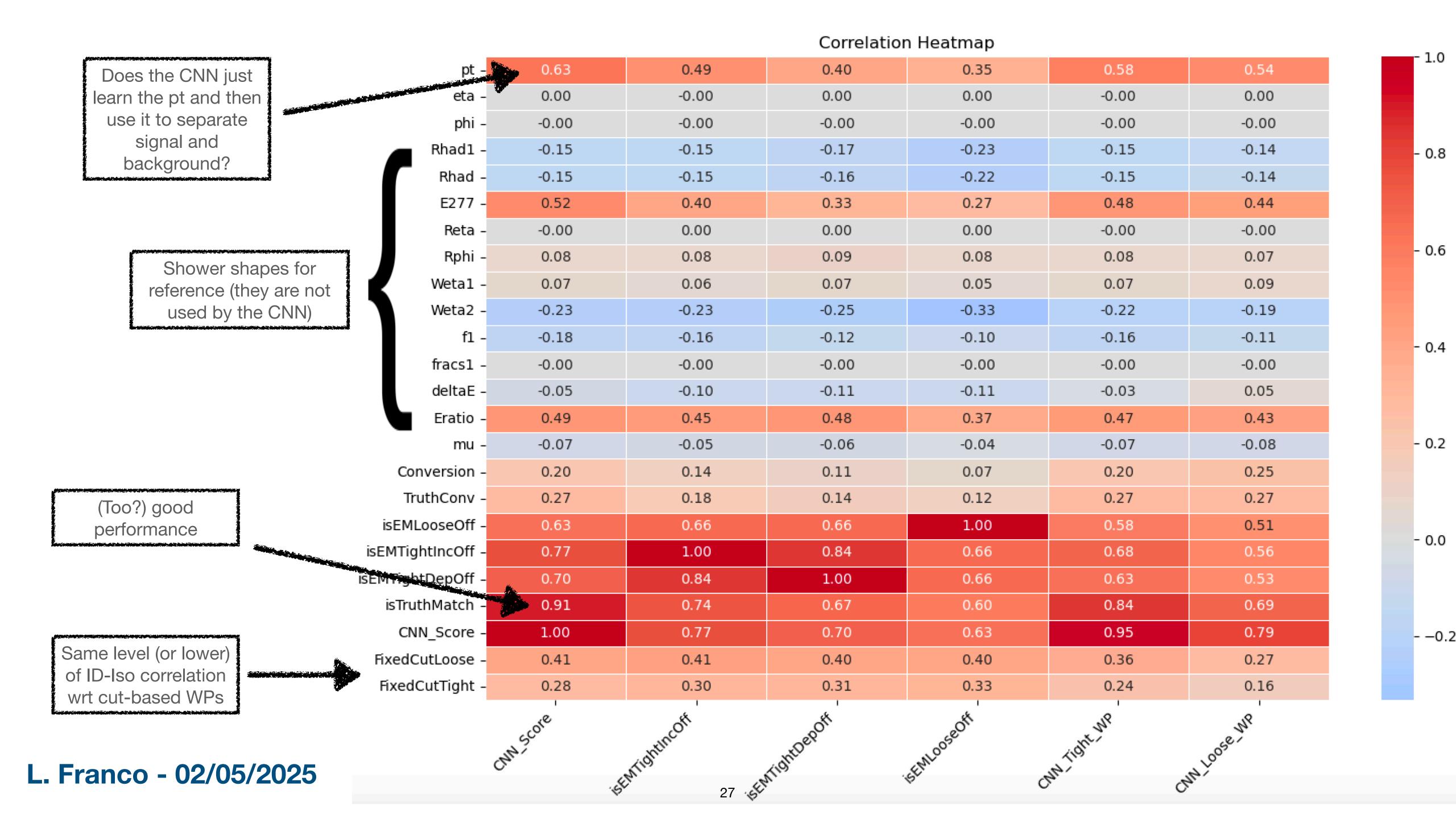
- compute pt ratio of sig/bkg
- reweight each bkg event with the ratio (sig events are weighted with 1)
- provide weights to the CNN during training to over/undercompensate for lack/abundance of signal events in given regions

#### Training

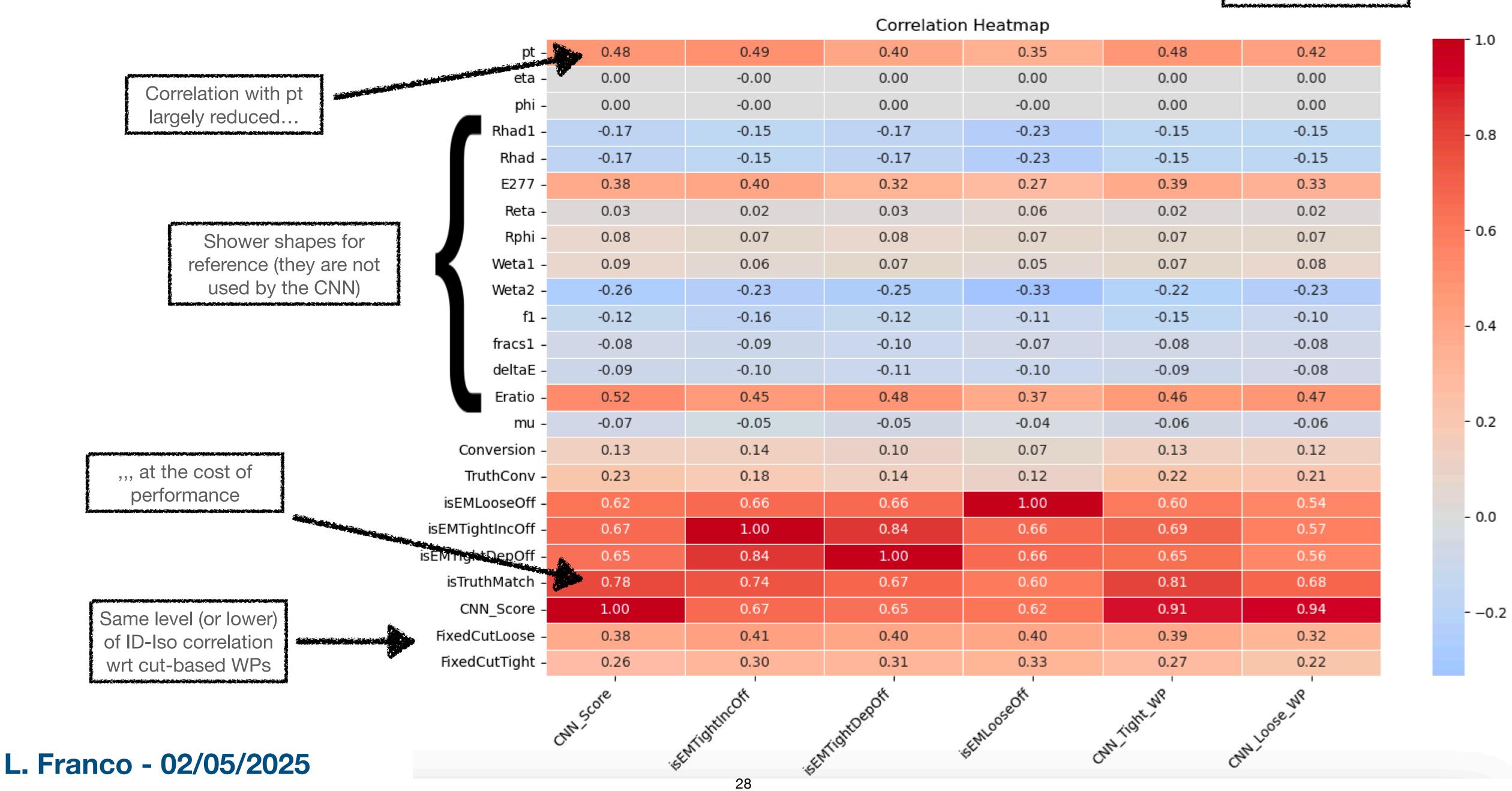
- Loss function: binary\_crossentropy
- Optimizer: Adam(learning\_rate=0.001)
- batch\_size=128
- Total params: 1248513 (4.76 MB)
- epochs=10 (~2-3 mins each on a 16-core CPU, i.e. on lxplus)





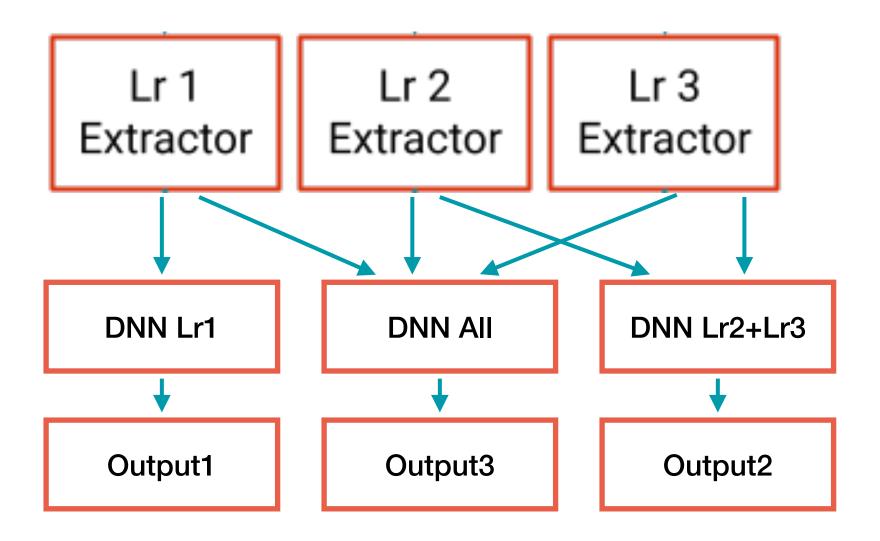






#### Approach #1: split per layer

- Idea: design a multi-output CNN to "mimic" narrow-strip and relaxedtight cuts
- Narrow-strip variables are computed using energy in Lr1: define a separate CNN output based only on Lr1 input



From Matrix Method INT note (link)

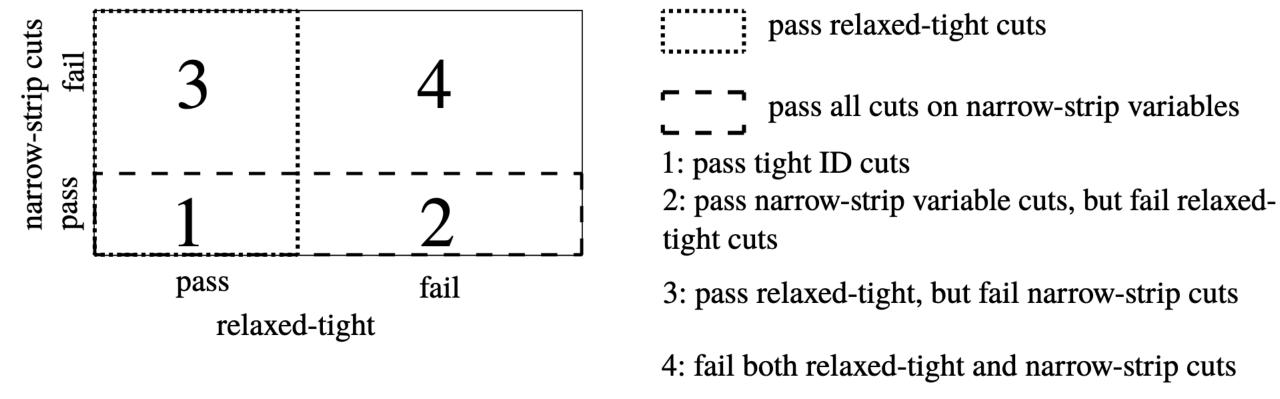
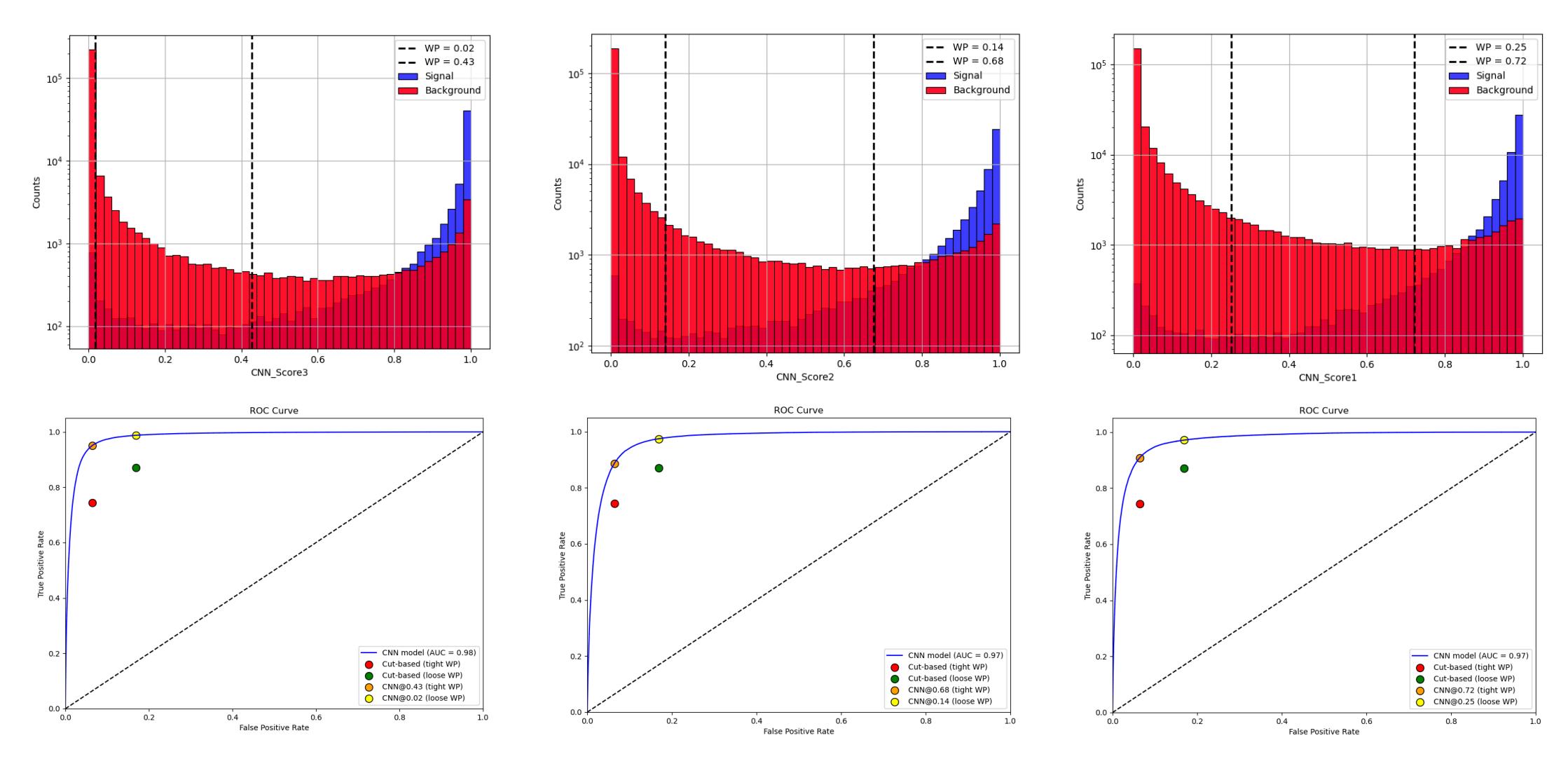


Figure 3: An illustration of the photon classification.

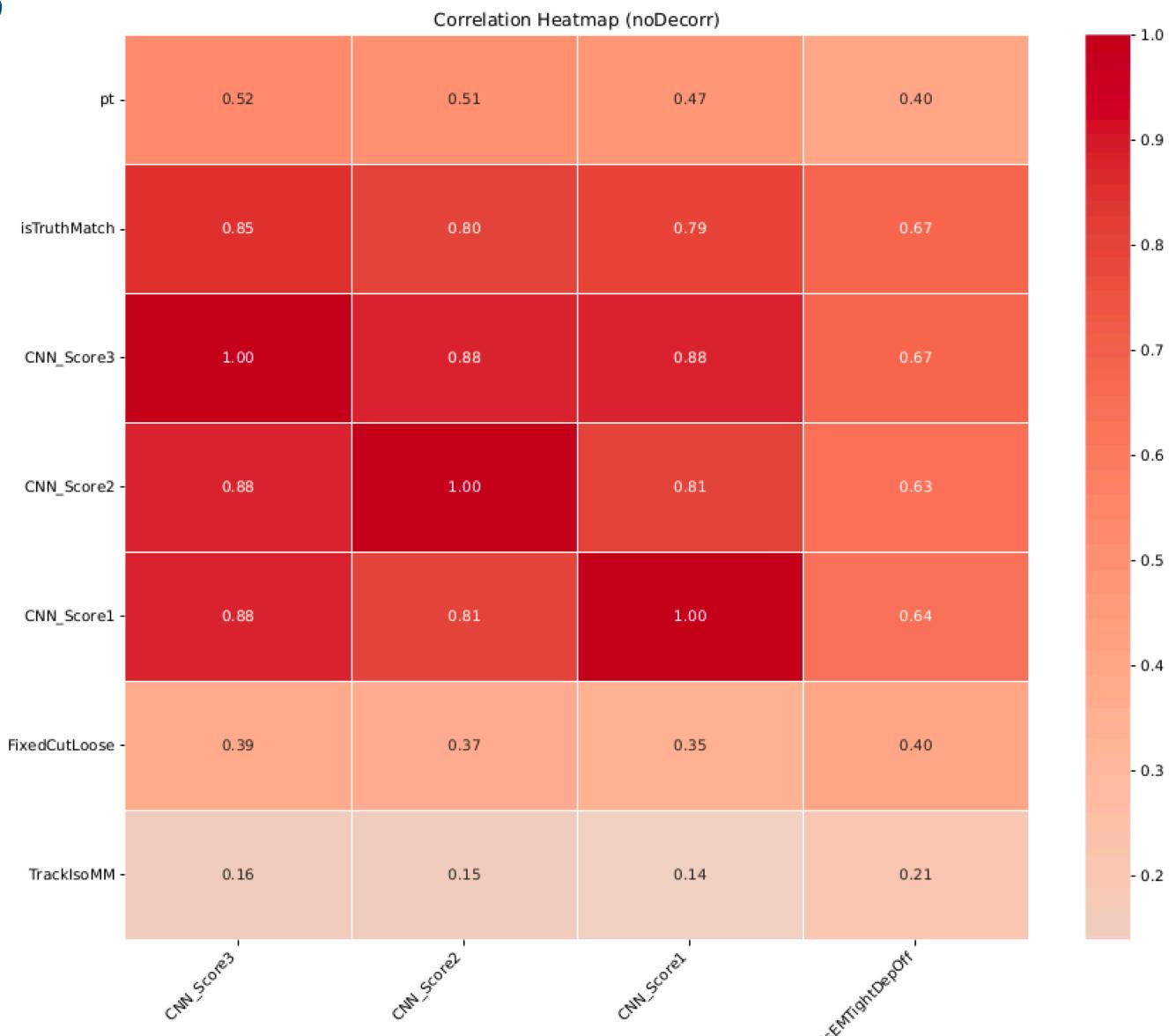
	cut-based	feature	CNN
all shower shapes best perform		best performance	final combined output
2 narrow-strip va	narrow-strip variables	weak correlation with Iso	only Lr1 input
relaxed-tight variables orthogonal (?) to narrow-strip		rest: Lr2, Lr3, HCal (?)	

#### Approach #1: CNN outputs



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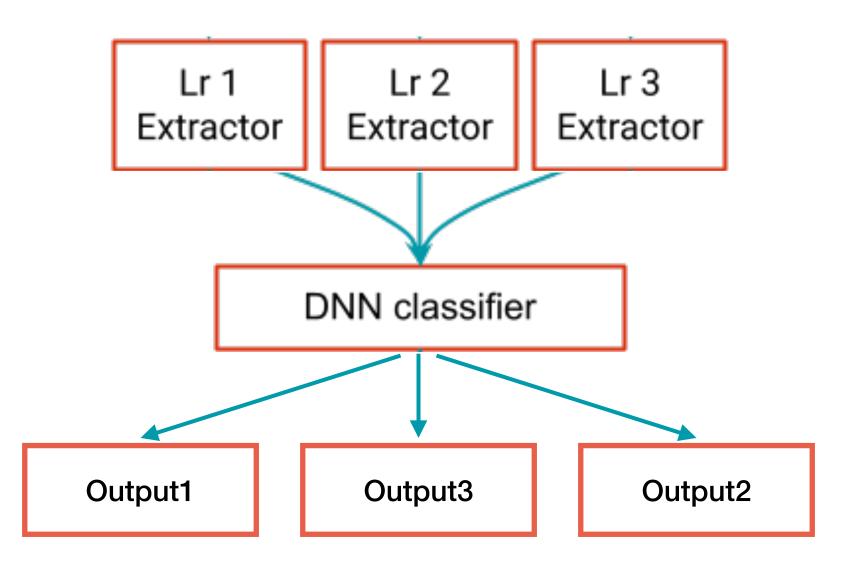
#### Approach #1: results



#### Approach #2: decorrelation in the loss

- Idea: design a multi-output CNN to "mimic" narrow-strip and relaxedtight cuts
- Define a separate CNN output which is explicitly decorrelated w.r.t. track-isolation in the loss

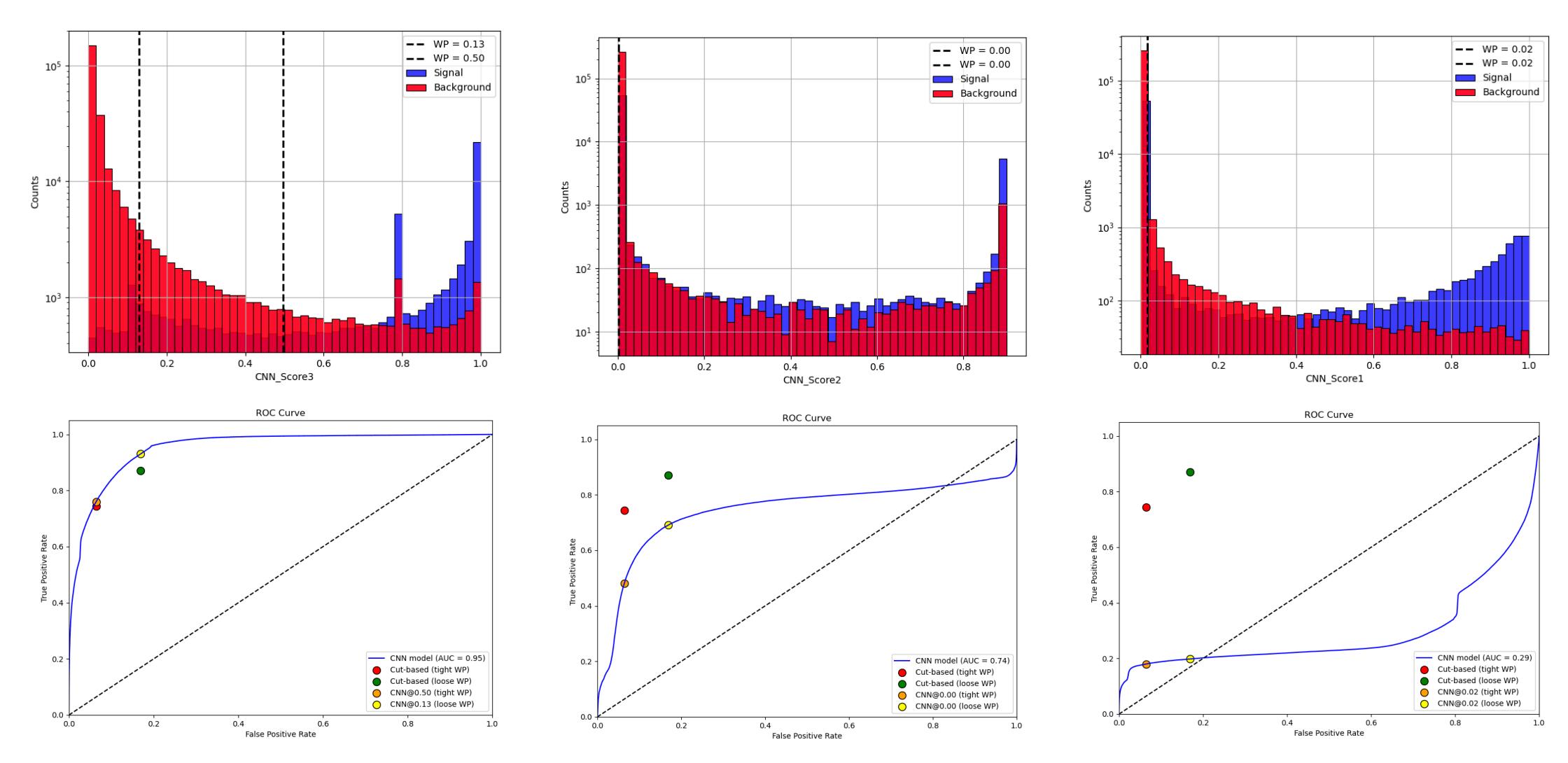
	cut-based	feature	CNN
1	all shower shapes	best performance	no penalization in loss
2	narrow-strip variables	weak correlation with Iso	regularization term w.r.t. TrackIsoMM
3	relaxed-tight variables	orthogonal (?) to narrow-strip	regularization term w.r.t. output1



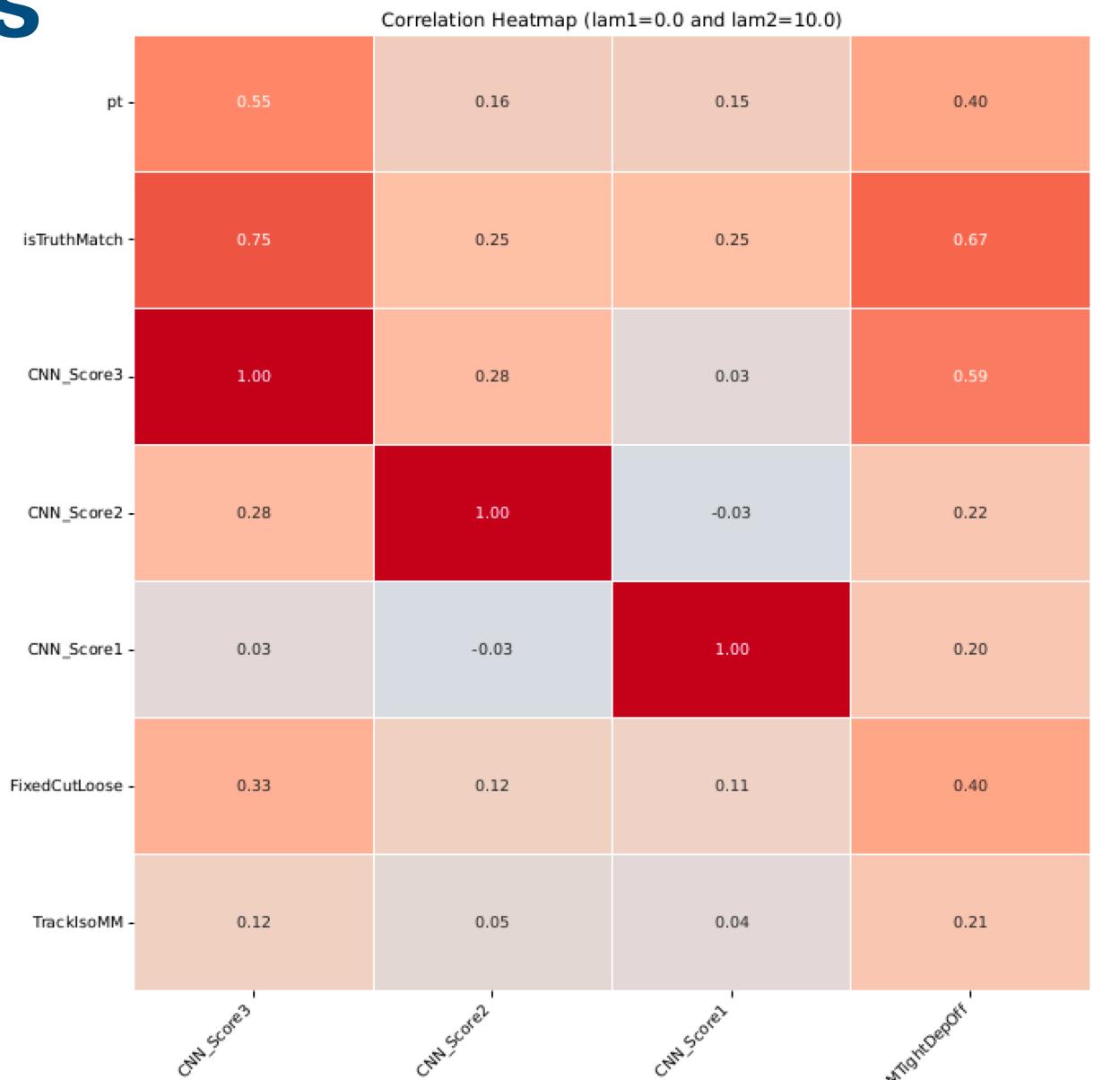
$$\mathcal{L} = \sum_{i=1}^{3} BCE(Out_i, y) + \lambda_1 DisCo(Out_1, ptcone40) + \lambda_2 DisCo(Out_2, Out_1)$$

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#### Approach #2: CNN outputs



Approach #2: results



- 0.8

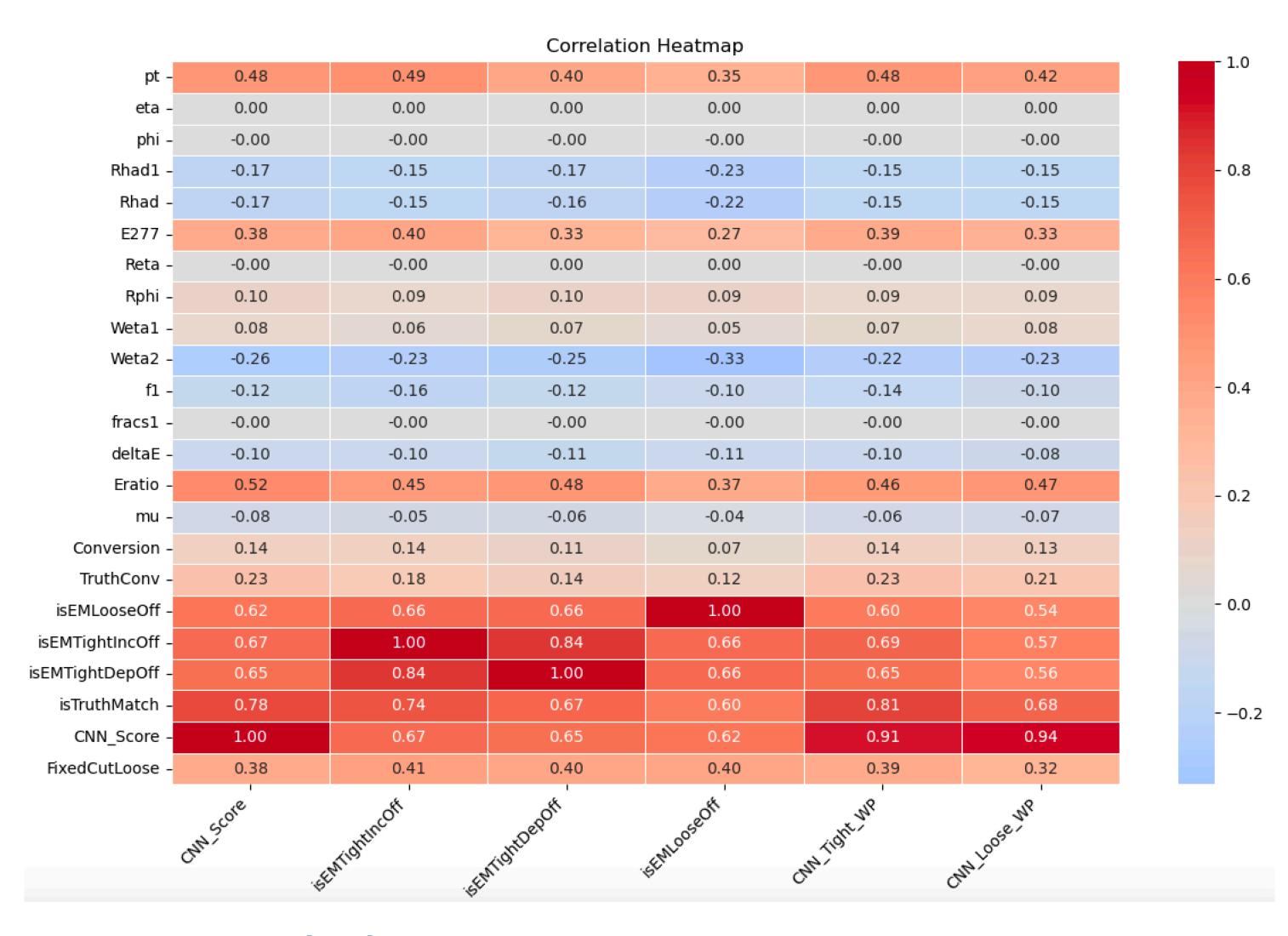
- 0.6

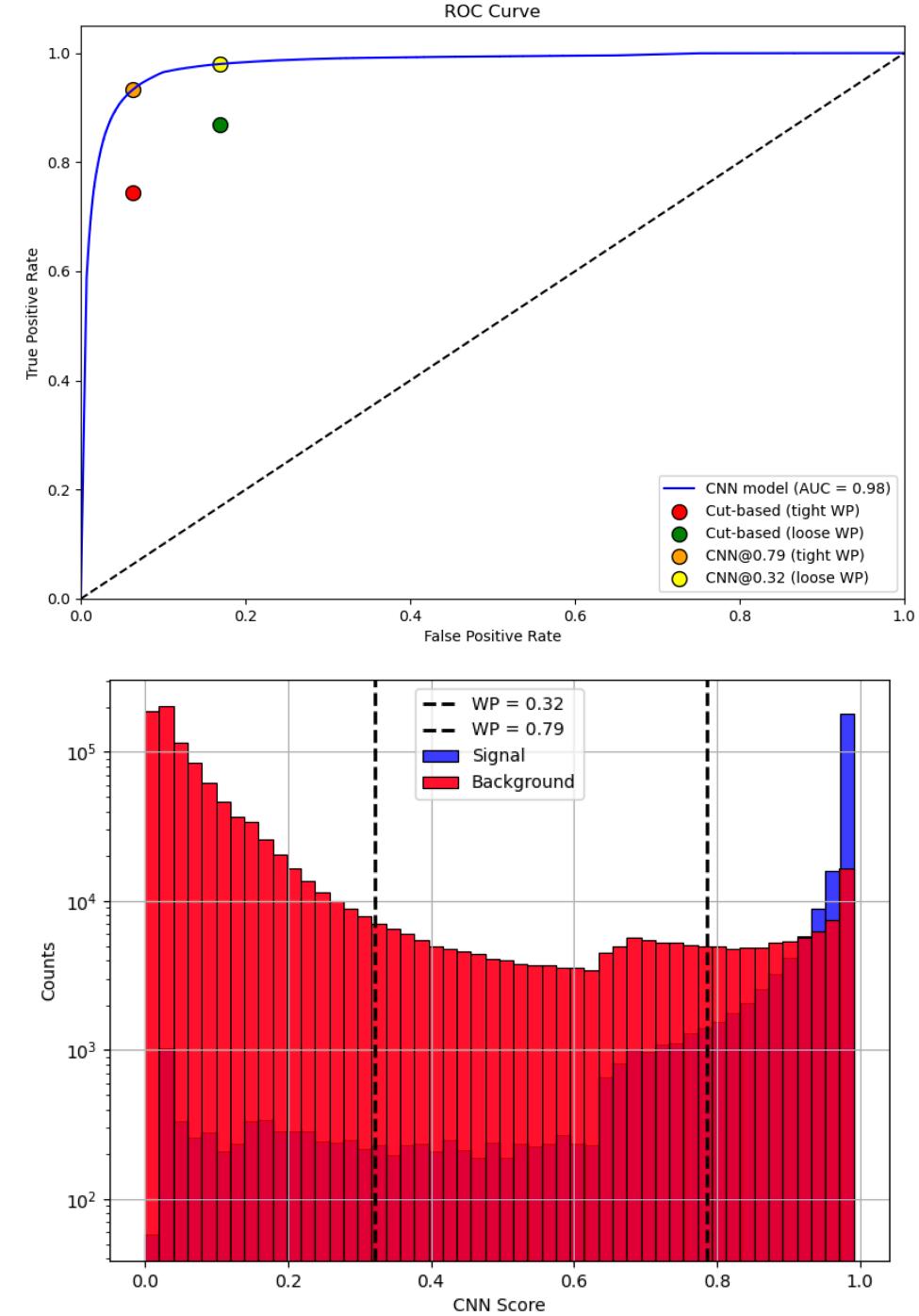
- 0.4

- 0.2

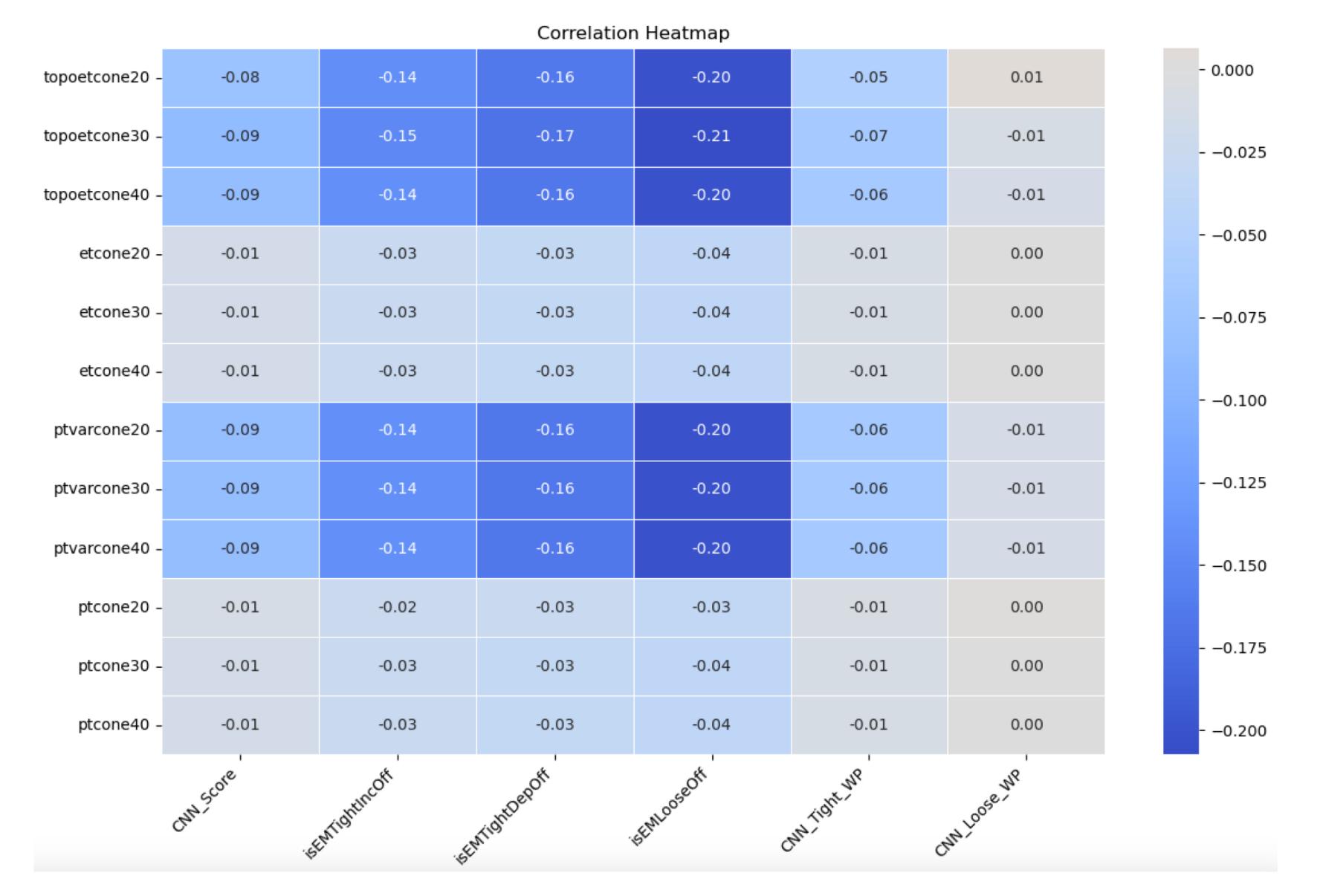
- 0.0

#### Train set



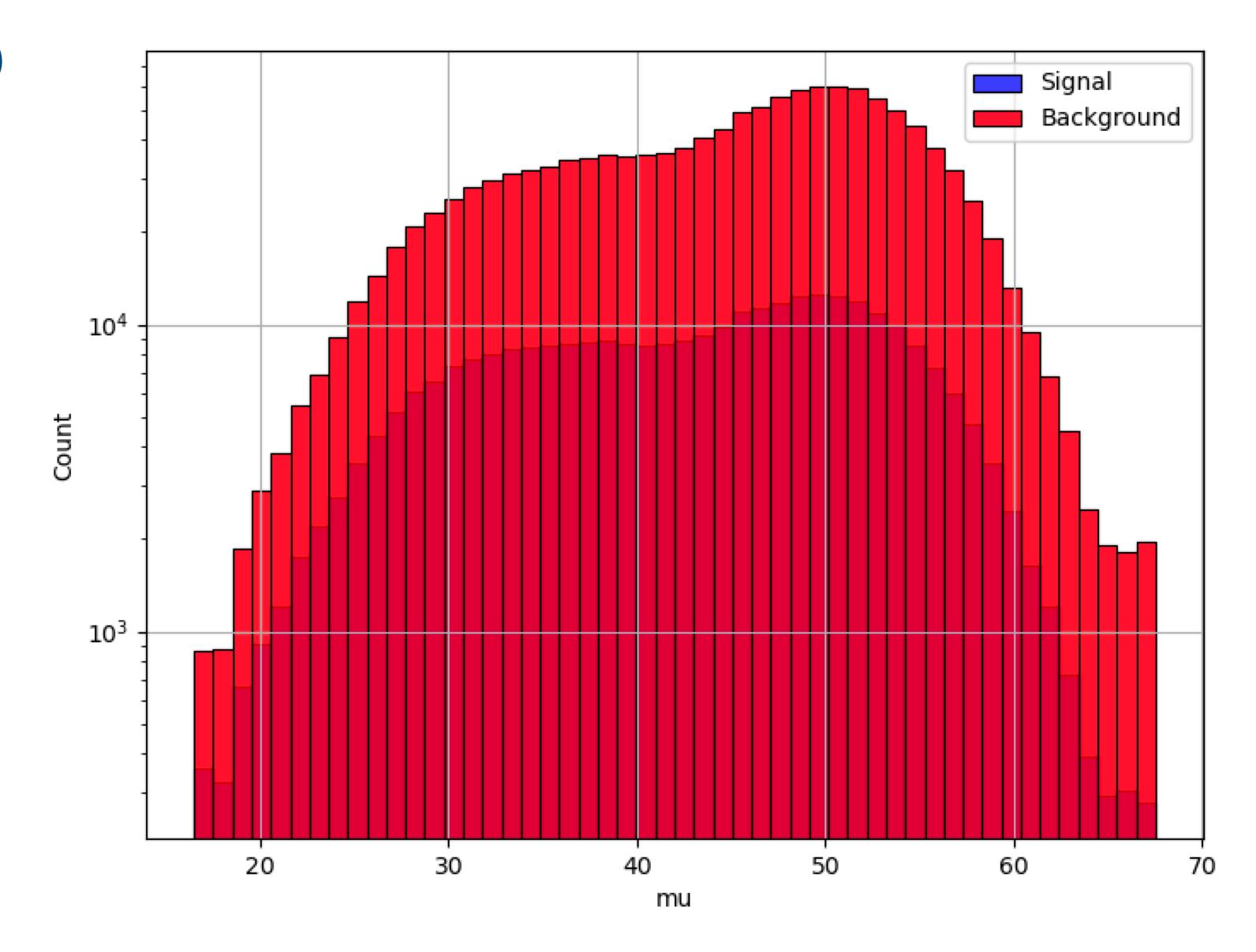


# Correlation wrt Isolation variables





#### Pile-up



#### Model definition

L. Franco - 02/05/2025

Model: "model"

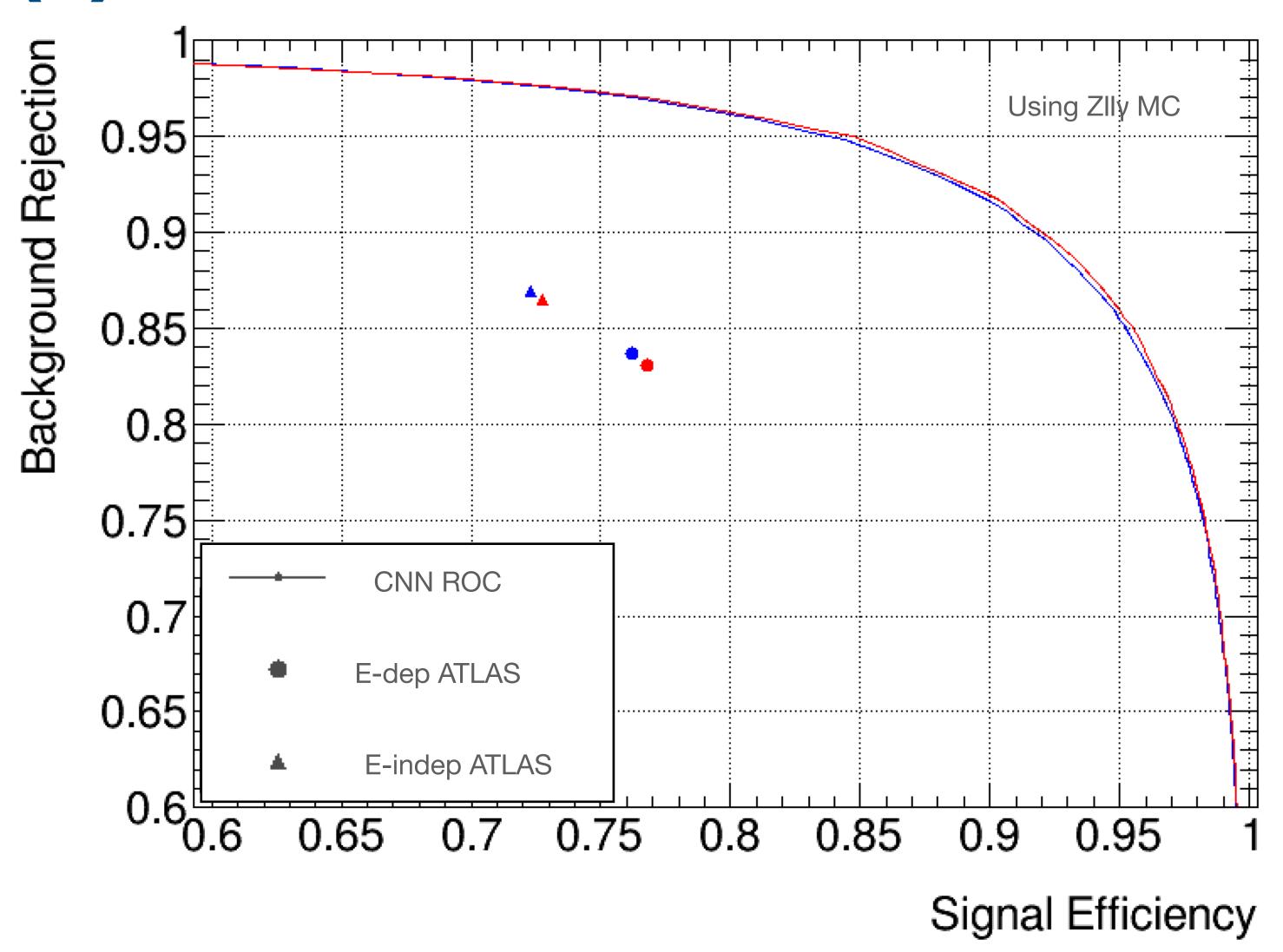
Layer (type)	Output Shape	Param #	Connected to
<pre>Input1 (InputLayer)</pre>	======================================	0 0	[]
Input2 (InputLayer)	[(None, 11, 7, 1)]	0	[]
Input3 (InputLayer)	[(None, 11, 4, 1)]	0	[]
Cov11 (Conv2D)	(None, 1, 55, 128)	640	['Input1[0][0]']
Cov21 (Conv2D)	(None, 10, 6, 128)	640	['Input2[0][0]']
Cov31 (Conv2D)	(None, 10, 3, 128)	640	['Input3[0][0]']
MaxP11 (MaxPooling2D)	(None, 1, 27, 128)	0	['Cov11[0][0]']
MaxP21 (MaxPooling2D)	(None, 5, 6, 128)	0	['Cov21[0][0]']
MaxP31 (MaxPooling2D)	(None, 5, 3, 128)	0	['Cov31[0][0]']
Cov12 (Conv2D)	(None, 1, 26, 128)	32896	['MaxP11[0][0]']
Cov22 (Conv2D)	(None, 4, 6, 128)	32896	['MaxP21[0][0]']
Cov32 (Conv2D)	(None, 4, 3, 128)	32896	['MaxP31[0][0]']
MaxP12 (MaxPooling2D)	(None, 1, 13, 128)	0	['Cov12[0][0]']
MaxP22 (MaxPooling2D)	(None, 2, 6, 128)	0	['Cov22[0][0]']
MaxP32 (MaxPooling2D)	(None, 2, 3, 128)	0	['Cov32[0][0]']
Flat1 (Flatten)	(None, 1664)	0	['MaxP12[0][0]']
Flat2 (Flatten)	(None, 1536)	0	['MaxP22[0][0]']
Flat3 (Flatten)	(None, 768)	0	['MaxP32[0][0]']
Concatenate (Concatenate)	(None, 3968)	0	['Flat1[0][0]', 'Flat2[0][0]', 'Flat3[0][0]']
D1 (Dense)	(None, 256)	1016064	['Concatenate[0][0]']
<pre>gaussian_noise (GaussianNo ise)</pre>	(None, 256)	0	['D1[0][0]']
D2 (Dense)	(None, 256)	65792	['gaussian_noise[0][0]']
<pre>gaussian_noise_1 (Gaussian Noise)</pre>	(None, 256)	0	['D2[0][0]']
D3 (Dense)	(None, 256)	65792	['gaussian_noise_1[0][0]']
Output (Dense)	(None, 1)	257	['D3[0][0]']

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Total params: 1248513 (4.76 MB)

#### Iso-ID correlation (1)

- CNN trained with Tight Isolation
- Applied to
  - Loose Iso
  - Tight Iso
- No change in performance of the algorithm



#### Iso-ID correlation (2)

- CNN trained with Tight Isolation
- Applied to
  - Topoetcone20/pt<-0.1</li>
  - Topoetcone20/pt>=-0.1
- CNN has same trend as ATLAS tight WPs: more isolated worse performance, because you remove info around the center of the cluster

