Axions at Nikhef?

Niels van Bakel



Used slides from Axel Lindner

Motivation - Physics

• The QCD axion:

- Solves the strong CP problem in SM, and couples to two photons
- People interested at Nikhef. What about ALPs?

• Axion Like Particles - ALPs:

- Does not solve open issue in SM but motivated by: astrophysical hints, cold DM candidate, and appears in string theory
- Small(er) group interested at Nikhef

Why WIMPs <u>and</u> WISPs

- ALPs parameter space should be explored, other experiments (LHC, DM) show no BSM physics so far
- WIMPs and WISPs can both exist two leading DM candidates
- No WIMPs and no WISPs
 - Our understanding of gravity is not complete, (e.g. Verlinde) and no need for DM

Experimental Landscape

- Purely laboratory
 experiments
 - "light-shining-throughwalls" \Rightarrow optical photons

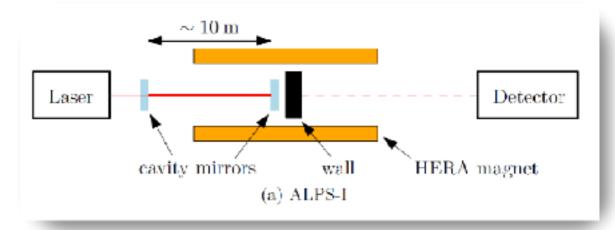
Helioscopes

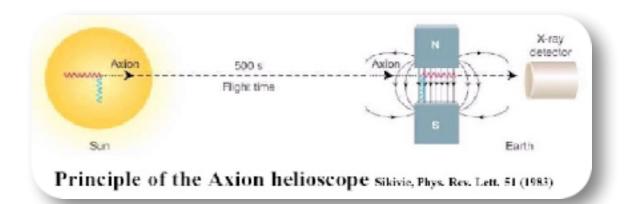
WISPs emitted by the sun
 ⇒ X-rays

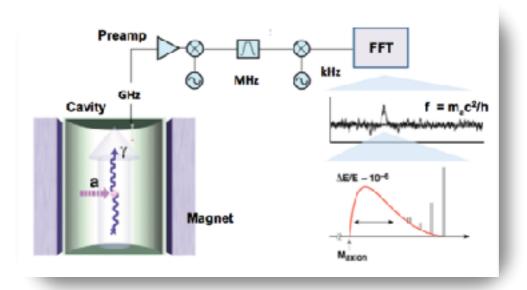
Haloscopes

- Looking for axions in Milky Way halo (DM) $\Rightarrow \mu$ -waves

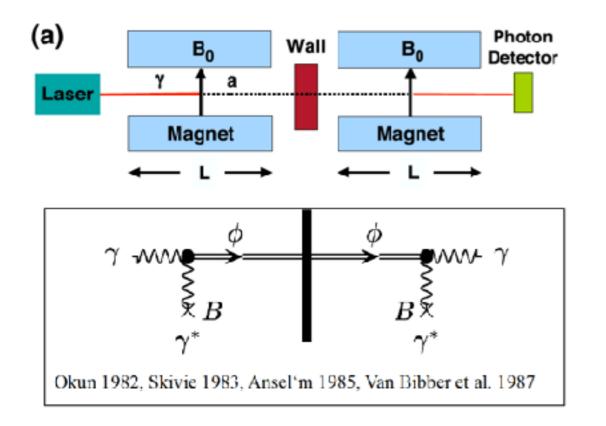








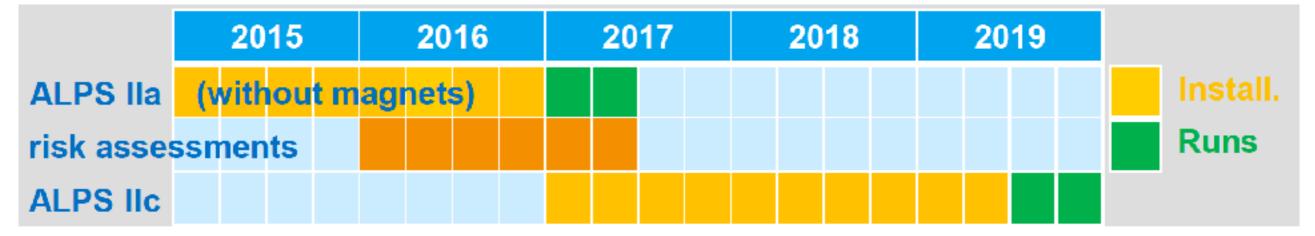
ALPS-2



Photon regeneration rate:

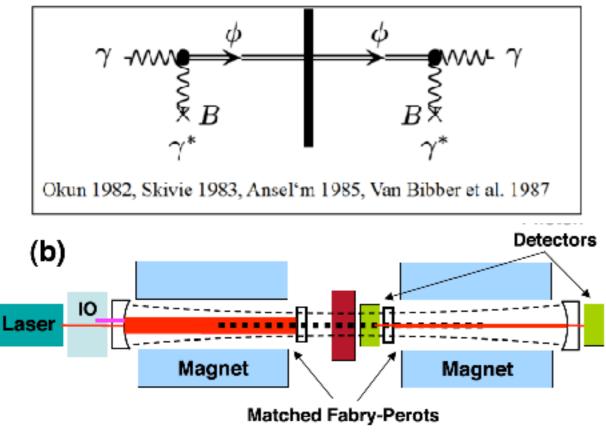
$$P_{\gamma \to \phi \to \gamma} = \frac{1}{16} \cdot \left(g_{a\gamma} B l\right)^4$$

$$= 6 \cdot 10^{-38} \cdot \left(\frac{g_{a\gamma}}{10^{-10} GeV^{-1}} \frac{B}{1 T} \frac{l}{10m}\right)^4$$



Axel Lindner Oct 2016

ALPS-2

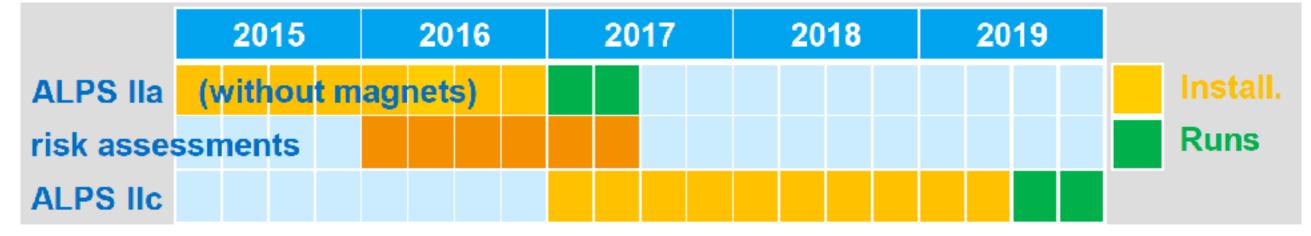


Sikivie PRL 98, 172002 (2007)

Photon regeneration rate:

$$P_{\gamma \to \phi \to \gamma} = \frac{1}{16} \cdot \mathcal{F}_{PC} \mathcal{F}_{RC} \cdot (g_{a\gamma\gamma} Bl)^4$$

$$= 6 \cdot 10^{-38} \cdot \mathcal{F}_{PC} \mathcal{F}_{RC} \cdot \left(\frac{g_{a\gamma\gamma}}{10^{-10} GeV^{-1}} \frac{B}{1T} \frac{l}{10m}\right)^4$$



Axel Lindner Oct 2016

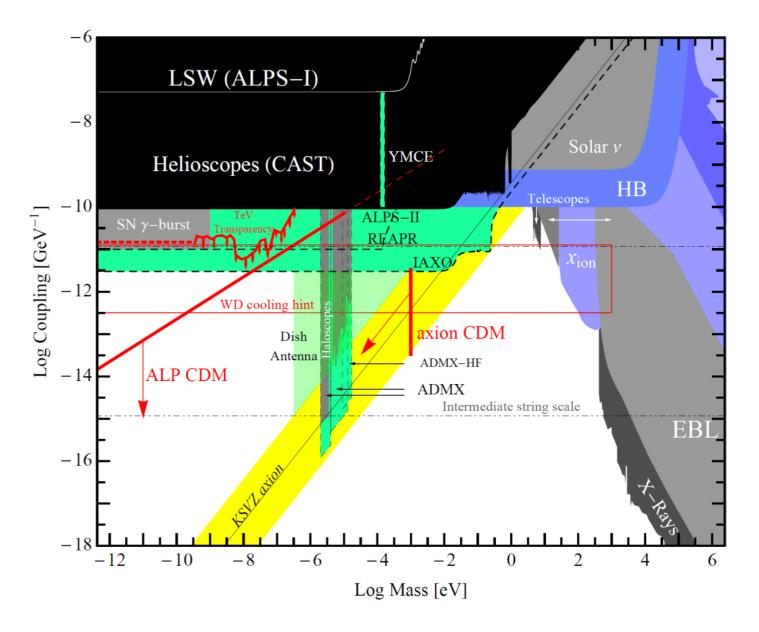
Parameter space

Motivation

- QCD axion, Axion CDM, ALP CDM
- WD cooling hint , TeV transparency
- Limits from astro-particle
 physics:
 - FermiLAT result?

Exclude: coupling > $5 \cdot 10^{-12}$ GeV⁻¹ for ALP masses 0.5 < m_a < 5 neV

- SN γ-burst?

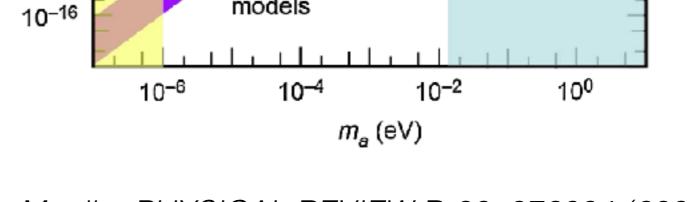


arXiv:1306.2841v1 [hep-ph] 12 Jun 2013

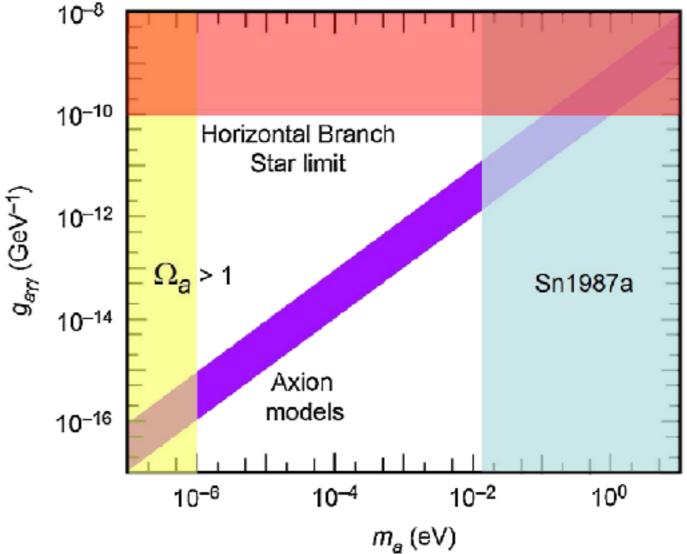
Parameter space

Constraints on axion mass and coupling

- Cosmological constraint ullet $(\rho_a \sim 1/m_a)$: many uncertainties - yellow
- Laboratory experiments & stellar evolution (Type 2 supernovae) - blue
- Stellar evolution (energy loss stars) - red
- Compare with previous slide?

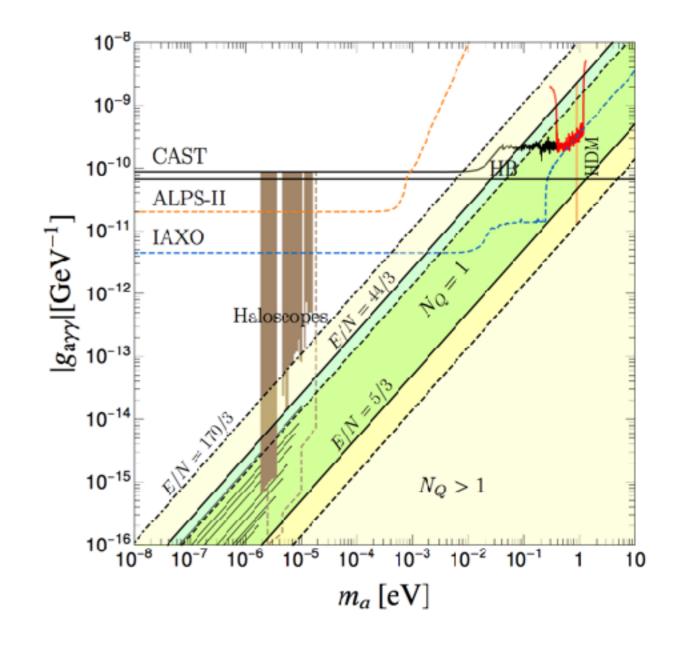


Mueller PHYSICAL REVIEW D 80, 072004 (2009)



Parameter space?

- QCD axion band changes
- Freedom in parameter space of QCD axion?

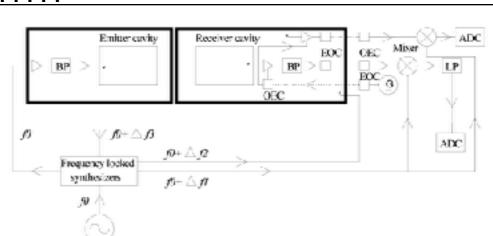


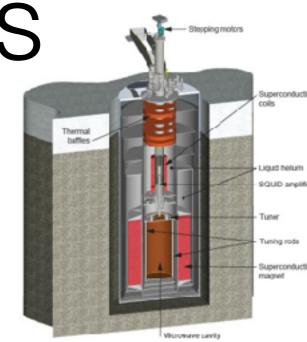
arXiv:1610.07593

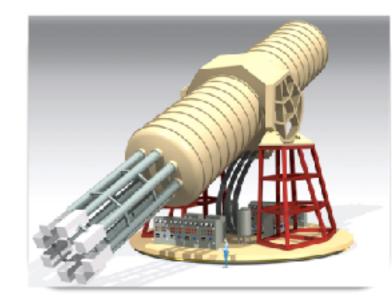
Many experiments

- ADMX (US)- world leading
- Many experiments with different techniques......
- What to choose:
 - Impact of ALPS-2?
 - Room to improve (LSW) experiments sensitive for QCD axion?
 - Future experiments? ALPS2 IAXO.....

Microwave cavity LSW experiment: arXiv 0908.0759







ALPS-2 @ Nikhef

Motivation

- Many uncertainties on ALPs parameter space and on nature of DM
- Overlap with instrumentation expertise GW & DR&D
- Use ALPS-2 as start to enter new & interesting field @ low cost
- Limited **effort** few staff part-time (25%)
- Collaborate on **new technologies**: e.g. SRON
 Kinetic Inductance Detector (CMB)
- Future experiments: e.g. x-ray detector with GridPix and magnets (Ten Cate - CERN, Twente University) for IAXO

Contribution

- **Start helping** ALPS2 on instrumentation: manpower
- Instrumentation development (future) and Theory
- Link with other programs GW, LHC, eEDM, Xenon1T
- **Small activity** not a major new program part of a few smaller experiments Vista25
- Funding Start with projectruimte