### Axions at Nikhef?

Niels van Bakel



Used slides from Axel Lindner

# Motivation - Physics

#### • The QCD axion:

- Solves the strong CP problem in SM, and couples to two photons
- People interested at Nikhef. What about ALPs?

#### • Axion Like Particles - ALPs:

- Does not solve open issue in SM but motivated by: astrophysical hints, cold DM candidate, and appears in string theory
- Small(er) group interested at Nikhef ......

#### Why WIMPs <u>and</u> WISPs

- ALPs parameter space should be explored, other experiments (LHC, DM) show no BSM physics so far
- WIMPs and WISPs can both exist ..... two leading DM candidates
- No WIMPs and no WISPs
  - Our understanding of gravity is not complete, (e.g. Verlinde) and no need for DM

# Experimental Landscape

- Purely laboratory
  experiments
  - "light-shining-throughwalls"  $\Rightarrow$  optical photons

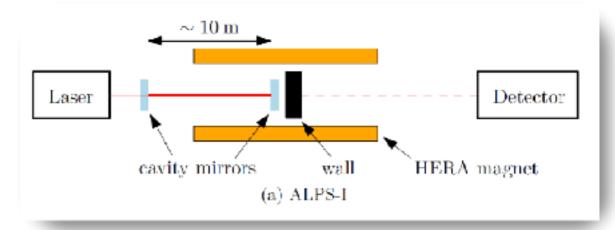
### Helioscopes

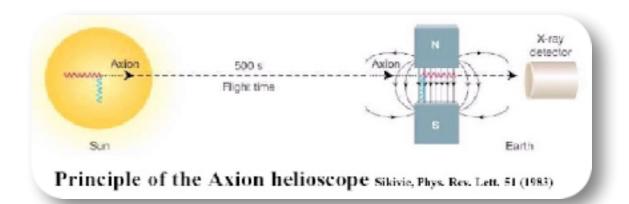
WISPs emitted by the sun
 ⇒ X-rays

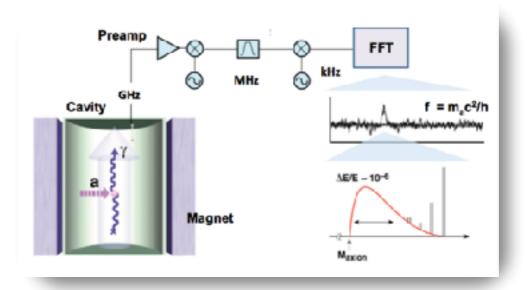
### Haloscopes

- Looking for axions in Milky Way halo (DM)  $\Rightarrow \mu$ -waves

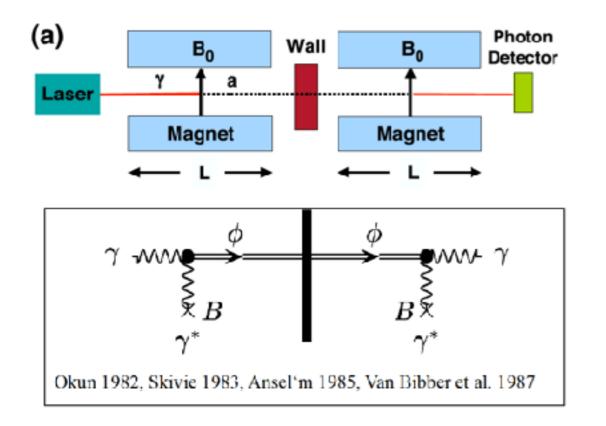








### ALPS-2



Photon regeneration rate:

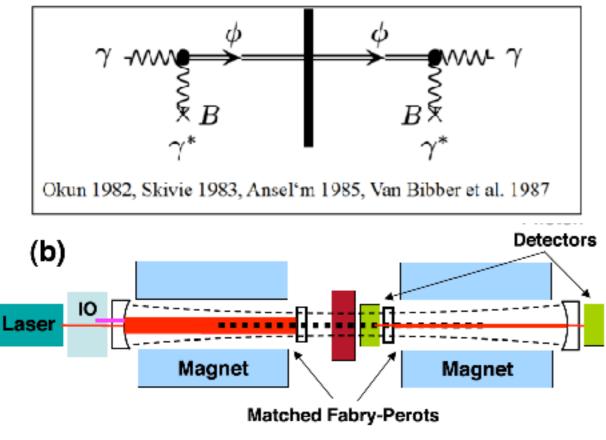
$$P_{\gamma \to \phi \to \gamma} = \frac{1}{16} \cdot \left(g_{a\gamma} B l\right)^4$$

$$= 6 \cdot 10^{-38} \cdot \left(\frac{g_{a\gamma}}{10^{-10} GeV^{-1}} \frac{B}{1 T} \frac{l}{10m}\right)^4$$



Axel Lindner Oct 2016

# ALPS-2



Sikivie PRL 98, 172002 (2007)

Photon regeneration rate:

$$P_{\gamma \to \phi \to \gamma} = \frac{1}{16} \cdot \mathcal{F}_{PC} \mathcal{F}_{RC} \cdot (g_{a\gamma\gamma} Bl)^4$$

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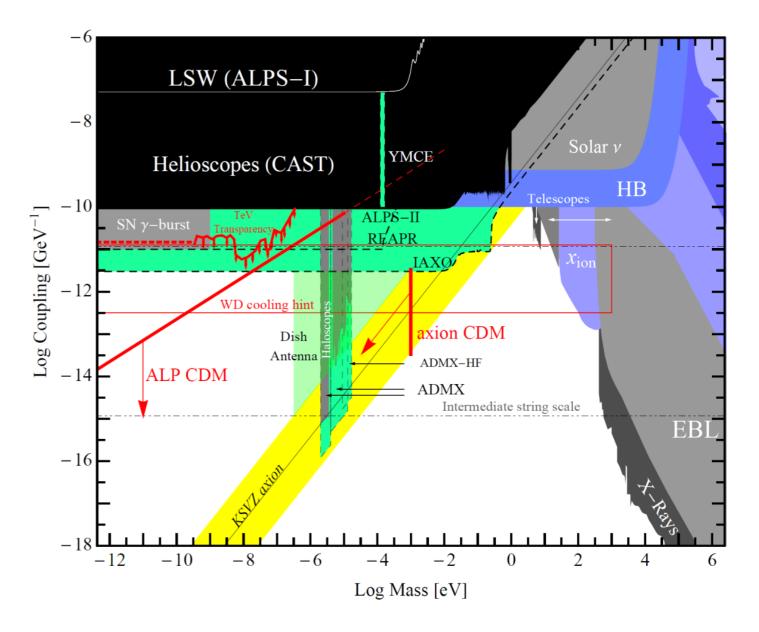
### Parameter space

### Motivation

- QCD axion, Axion CDM, ALP CDM
- WD cooling hint , TeV transparency
- Limits from astro-particle
  physics:
  - FermiLAT result?

Exclude: coupling >  $5 \cdot 10^{-12}$  GeV<sup>-1</sup> for ALP masses 0.5 < m<sub>a</sub> < 5 neV

- SN γ-burst?

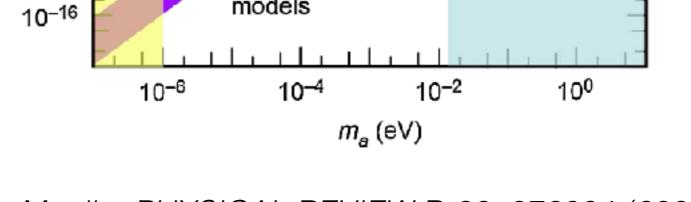


arXiv:1306.2841v1 [hep-ph] 12 Jun 2013

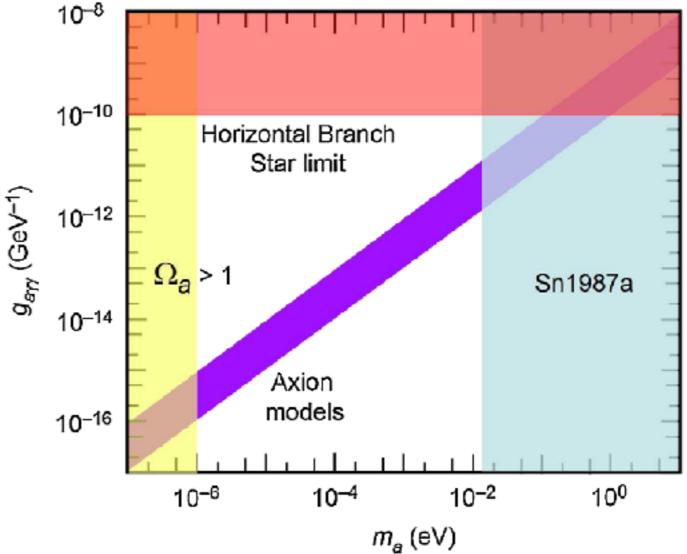
### Parameter space

#### **Constraints on axion mass** and coupling

- Cosmological constraint ullet $(\rho_a \sim 1/m_a)$ : many uncertainties - yellow
- Laboratory experiments & stellar evolution (Type 2 supernovae) - blue
- Stellar evolution (energy loss stars) - red
- Compare with previous slide?

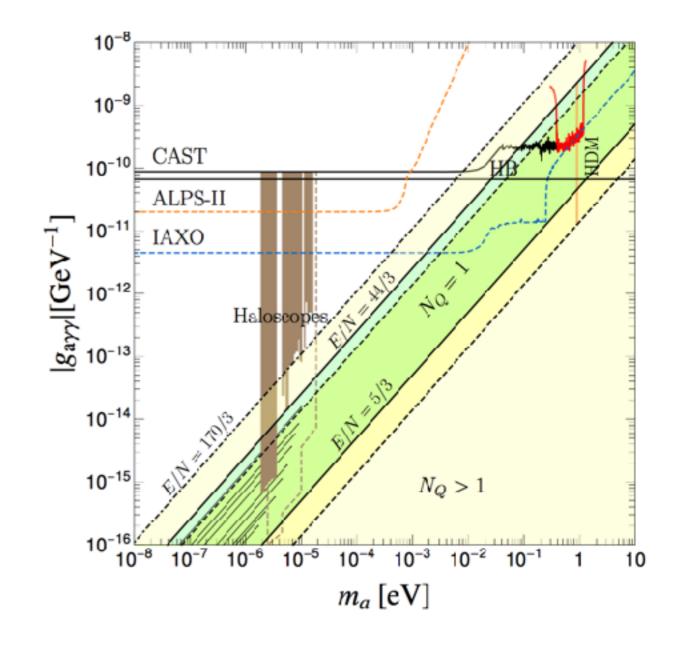


Mueller PHYSICAL REVIEW D 80, 072004 (2009)



# Parameter space?

- QCD axion band changes
- Freedom in parameter space of QCD axion?

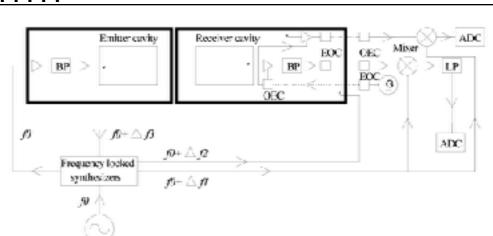


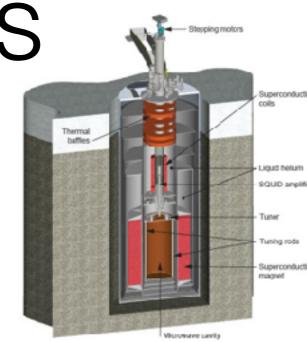
arXiv:1610.07593

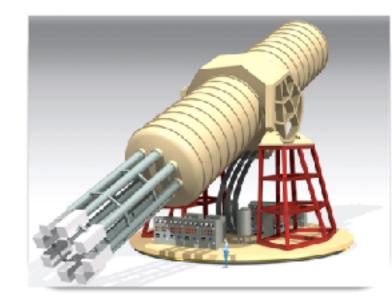
# Many experiments

- ADMX (US)- world leading
- Many experiments with different techniques......
- What to choose:
  - Impact of ALPS-2?
  - Room to improve (LSW) experiments sensitive for QCD axion?
  - Future experiments? ALPS2 IAXO.....

*Microwave cavity LSW experiment: arXiv 0908.0759* 







# ALPS-2 @ Nikhef

### Motivation

- Many uncertainties on ALPs parameter space and on nature of DM
- Overlap with instrumentation expertise GW & DR&D
- Use ALPS-2 as start to enter new & interesting field @ low cost
- Limited **effort** few staff part-time (25%)
- Collaborate on **new technologies**: e.g. SRON
  Kinetic Inductance Detector .... (CMB)
- Future experiments: e.g. x-ray detector with GridPix and magnets (Ten Cate - CERN, Twente University) for IAXO

### Contribution

- **Start helping** ALPS2 on instrumentation: manpower
- Instrumentation development (future) and Theory
- Link with other programs GW, LHC, eEDM, Xenon1T
- **Small activity** not a major new program part of a few smaller experiments Vista25
- Funding Start with projectruimte