KM3NeT

The Cubic Kilometre Neutrino Telescope Anna Hurhina

Introduction

Cubic Kilometre Neutrino Telescope:

- Water Chereknov detector.
- Detects products of neutrino water interactions.
- At the bottom of the Mediterranean Sea.
- Several km³ of instrumented volume.





Google Maps

Astroparticle Research with Cosmics in the Abyss:

- Study high-energy neutrino sources in the Universe.
- 2 building blocks.
- At 3500 m depth, 100 km offshore Italy.

- Horizontal spacing 95 m.
- TeV neutrino energies.





[1] [2]

STADE

Oscillation Research with Cosmics in the Abyss:

- Study of neutrino flavour oscillations and the neutrinos mass hierarchy.
- At 2450 m depth, 40 km offshore France.
- 1 building block.
- Horizontal spacing 20 m.
- GeV neutrino energies.





Detection String

- Each block has 115 strings.
- Each string has 18 digital optical modules (DOMs).
- ARCA: spread out configuration; height: 700 m with DOMs vertical spacing 36 m.
- ORCA: dense configuration; height: 200 m with DOMs vertical spacing 9 m.
- Cables for power and data transmission.
- Float at the top.
- Calibration equipment in the base.









[7]

The DOM



- 44 cm in diameter.
- Pressure-resistant glass sphere.
- 31 PMTs.

[8]

- Optical gel to fill all cavities.
- Calibration sensors: LED nano beacon, compass, tiltmeter and acoustic sensor.

[1] [2] [5]





[7]

[7]



PMT Characteristics

- Minimum gain 10^6 .
- Photocathode diameter 72 mm + reflector ring.
- Length 122 mm.
- Small dark count rate.
- High quantum efficiency.
- Time resolution <5 ns.
- Individual HV base with integrated amplification and tuneable signal discrimination.

[1] [3] [4]





Interactions

[2]

Charged current events:

 $v_e + N \rightarrow e + X$ EM + hadronic shower

 $\nu_{\mu} + N \rightarrow \mu + X$ Muon track + hadronic shower

 $\nu_{\tau} + N \rightarrow \tau + X$ Hadronic shower + 2nd decay signature

 $\tau \rightarrow \nu_{\tau} \overline{\nu}_{\mu} \mu$ / $\tau \rightarrow \nu_{\tau} \overline{\nu}_{e} e$ / $\tau \rightarrow \nu_{\tau} + hadrons$

Neutral current:	
$\nu + N \rightarrow \nu + X$	Hadronic shower
Scattering:	
$\overline{\nu}_e + e \rightarrow \overline{\nu}_e + e$	EM shower
	[2]

Cherenkov Radiation

Emission angle of Cherenkov radiation:

$$\cos(\theta) = \frac{1}{n\beta}$$

 β is the ratio of particle speed over the speed of light in vacuum.

Frequency spectrum of Cherenkov radiation given by Frank-Tamm equation:

$$\frac{d^{2}E}{dxd\omega} = \frac{q^{2}\mu(\omega)\omega}{4\pi} \left(1 - \frac{1}{\beta^{2}n^{2}(\omega)}\right)$$

Amount of energy E emitted from Cherenkov radiation, per unit length travelled x and per frequency ω . $\mu(\omega)$ is permeability and $n(\omega)$ is frequency dependent index of refraction.





- Measurement: time of arrival and TOT.
- Each PMT individual time-todigital converter.
- Each measurement: 6 B from PMT + position and orientation from calibration equipment.
- Information sent to shore: 1B for PMT address, 4B for time and 1B for TOT.

[1]



<u>Track</u>

Good direction measurment, not good energy measurement.

Cascade

Good energy measurment, bad direction measurement.





Background

- 40 K β -decay.
- Bioluminescence.
- Atmospheric muons and neutrinos. [1] [2]





The Science

- Point-like neutrino sources.
- Diffuse galactic neutrino flux.
- Multi-messenger astronomy.





- Neutrino flavour states oscillations.
- Order of neutrino mass states.



Exotic Science

- SUSY Neutralinos (neutrinos flux)
- Magnetic monopoles (Cherenkov radiation)
- Nuclearites strange quark stable matter from cosmic rays (black-body radiation)
- Violation of Lorentz Invariance (unexpected interference in neutrino oscillation patters)



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Thank you for attention!

KMaNe