

# Robust parameter estimation within minutes on gravitational wave signals from binary neutron stars

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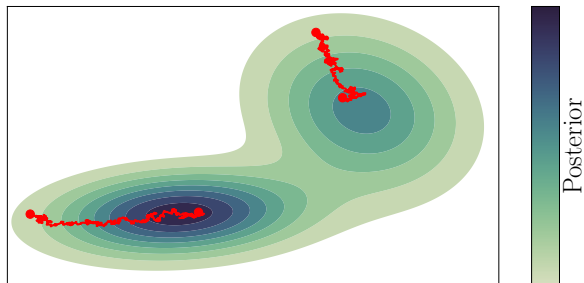
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# Parameter estimation

Parameter estimation (PE): get **posterior** of GW parameters  $\theta$

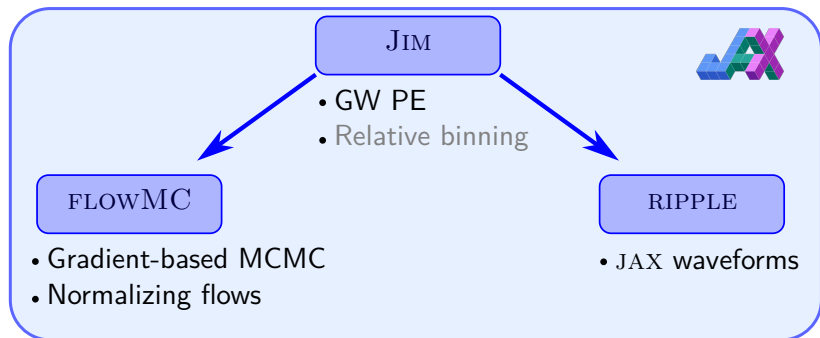
$$p(\theta|d) = \frac{p(d|\theta)p(\theta)}{p(d)} = \frac{\text{likelihood} \times \text{prior}}{\text{evidence}}$$

**Problem:** Markov Chain Monte Carlo (MCMC): computationally expensive for binary neutron stars (BNS)



JIM: fast parameter estimation of GW signals with JAX

- MCMC sampler: FLOWMC
- Waveforms: RIPPLE



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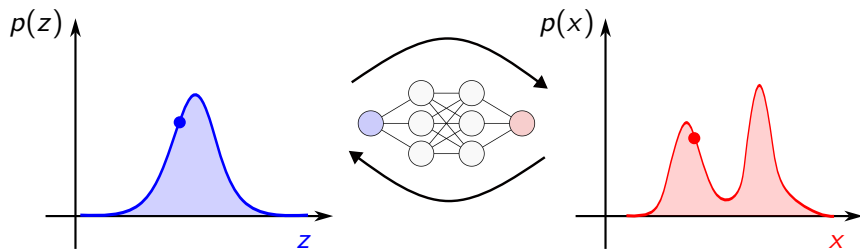
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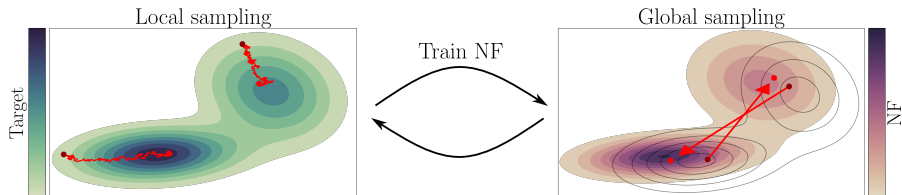
# Normalizing flows

- Generative machine learning model
- Learn mapping between **latent** and **parameter** space
- Enable approximate sampling from complicated distributions



FLOWMC: normalizing-flow (NF) enhanced MCMC sampling

- 1 Gradient-based sampler (local sampler)
- 2 Train NF with samples from local sampler
- 3 Normalizing flows (global sampler)



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# Results

- Waveforms: TaylorF2 (TF2), IMRPhenomD\_NRTidalv2 (NRTv2)
- JIM wall time: (i) computing reference parameters for relative binning, (ii) training NF, (iii) sampling

Event	Waveform	JIM (1 GPU)	pBILBY (480 cores)	RB-BILBY (24 cores)	ROQ-BILBY (24 cores)
GW170817	TF2	(9.70 + 17.00) min	9.64 h	3.18 h	–
	NRTv2	(5.69 + 28.02) min	10.99 h	4.68 h	1.65 h
GW190425	TF2	(5.13 + 16.49) min	4.08 h	2.30 h	–
	NRTv2	(6.15 + 15.37) min	4.69 h	4.68 h	0.97 h
Injection	TF2	24.76 min	–	–	–
	NRTv2	18.02 min	–	–	–

(pBILBY = PARALLEL BILBY, RB = relative binning, ROQ = reduced order quadrature)

# Environmental impact

JIM is **more environmentally friendly** than existing pipelines

- Energy consumption for all 204 runs of paper
- Convert to number of trees to capture the emitted CO<sub>2</sub> in a year.

Pipeline	Trees
JIM	0.55
pBILBY	59.02
RB-BILBY	1.49
ROQ-BILBY	sampling 0.52
	precompute <sup>†</sup> 92.40

<sup>†</sup>Estimated cost to build ROQ bases.

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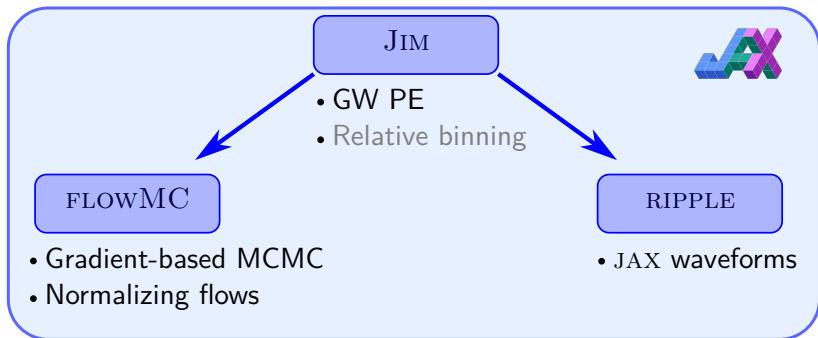
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# Conclusion

JIM: a fast and environmentally friendly PE pipeline for GW signals

- TaylorF2 and IMRPhenomD\_NRTidalv2 in RIPPLE
- Parameter estimation of BNS in 15 – 30 minutes sampling time without pretraining



# References

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- [3] Marylou Gabrié, Grant M Rotskoff, and Eric Vanden-Eijnden. “Efficient bayesian sampling using normalizing flows to assist markov chain monte carlo methods”. In: *arXiv preprint arXiv:2107.08001* (2021).
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- [6] Barak Zackay, Liang Dai, and Tejaswi Venumadhav. “Relative Binning and Fast Likelihood Evaluation for Gravitational Wave Parameter Estimation”. In: (June 2018). arXiv: 1806.08792 [astro-ph.IM].
- [7] Patrick Kidger and Cristian Garcia. EQUINOX: *neural networks in JAX via callable PyTrees and filtered transformations*. Available at: <https://github.com/patrick-kidger/equinox>. 2021. arXiv: 2111.00254 [cs.LG].

## **APPENDIX**

# Normalizing flow

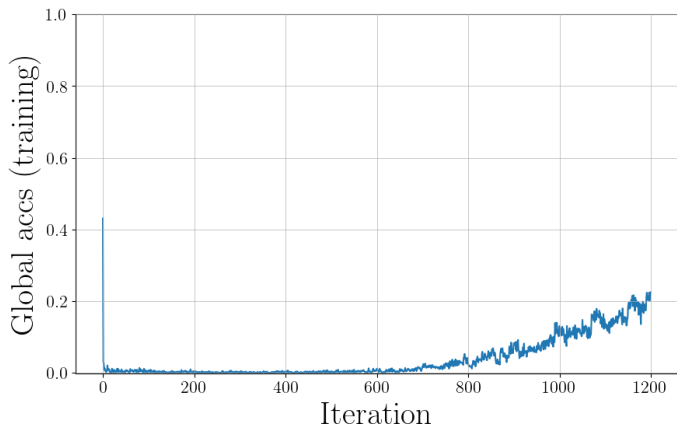
Normalizing flows details:

- Rational-quadratic neural spline flows
- 10 layers, 8 bins
- 128 neurons in hidden layers
- Adam optimizer, learning rate decayed (polynomial schedule)
- Deep learning library: EQUINOX

# Stopping criterion

We stop training the NF if we achieve a mean Metropolis-Hastings acceptance rate of 10% (20%) for real events (injections).

Example: GW170817, TaylorF2 with 20%:

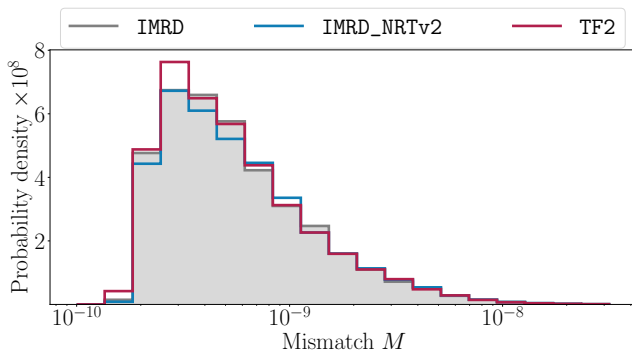




# Validation – Mismatch waveforms

Cross-check against LALSUITE: mismatch histogram based on 10 000 waveforms, from uniform samples with following ranges:

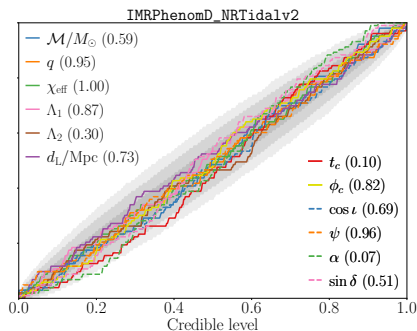
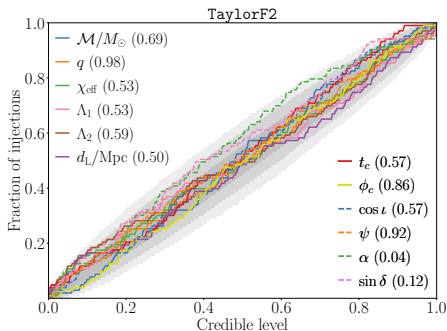
Parameter	Range
Component masses	$[0.5M_{\odot}, 3M_{\odot}]$
Component aligned spins	$[-0.05, 0.05]$
Dimensionless tidal deformabilities	$[0, 5000]$
Inclination angle	$[0, \pi]$



# Validation – p-p plot

We demonstrate the robustness of JIM:

- 100 GW events with HLV at design sensitivity and  $T = 128$  s,
- NRTv2: reference waveform relative binning without taper,
- Priors: Table 1.



*Table 1:* Prior ranges used in our analyses. All priors are uniform priors with the specified range.

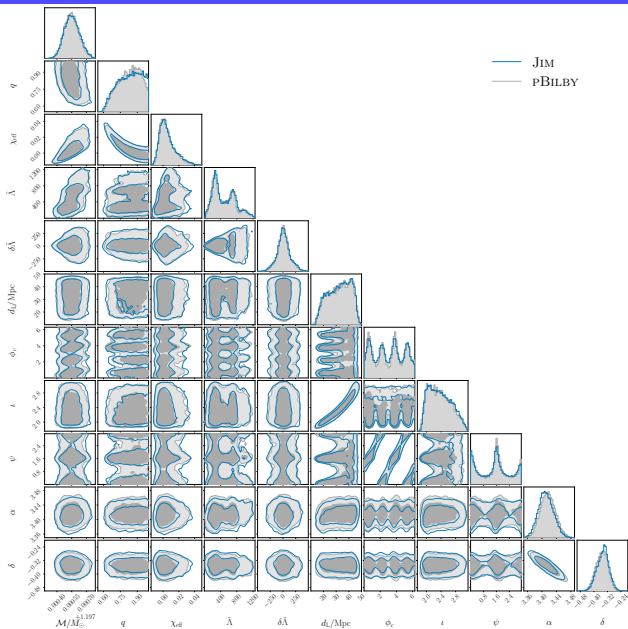
Parameter	Injection	GW170817	GW190425
$\mathcal{M}$ [ $M_{\odot}$ ]	[0.88, 2.61]	[1.18, 1.21]	[1.485, 1.490]
$q$	[0.5, 1]	[0.125, 1]	[0.125, 1]
$\chi_i$	[-0.05, 0.05]	[-0.05, 0.05]	[-0.05, 0.05]
$\Lambda_i$	[0, 5000]	[0, 5000]	[0, 5000]
$d_L$ [Mpc]	[30, 300]	[1, 75]	[1, 500]
$t_c$ [s]	[-0.1, 0.1]	[-0.1, 0.1]	[-0.1, 0.1]
$\phi_c$	[0, $2\pi$ ]	[0, $2\pi$ ]	[0, $2\pi$ ]
$\cos \iota$	[-1, 1]	[-1, 1]	[-1, 1]
$\psi$	[0, $\pi$ ]	[0, $\pi$ ]	[0, $\pi$ ]
$\alpha$	[0, $2\pi$ ]	[0, $2\pi$ ]	[0, $2\pi$ ]
$\sin \delta$	[-1, 1]	[-1, 1]	[-1, 1]

# GW170817 & GW190425: Jensen-Shannon divergences

*Table 2:* Jensen-Shannon divergences (in bits) between the marginal posterior obtained for GW170817 and GW190425 using TaylorF2 and IMRPhenomD\_NRTidalv2 with JIM and PBILBY, with the highest value of each comparison in bold. The divergences are bound between  $[0, 1]$ .

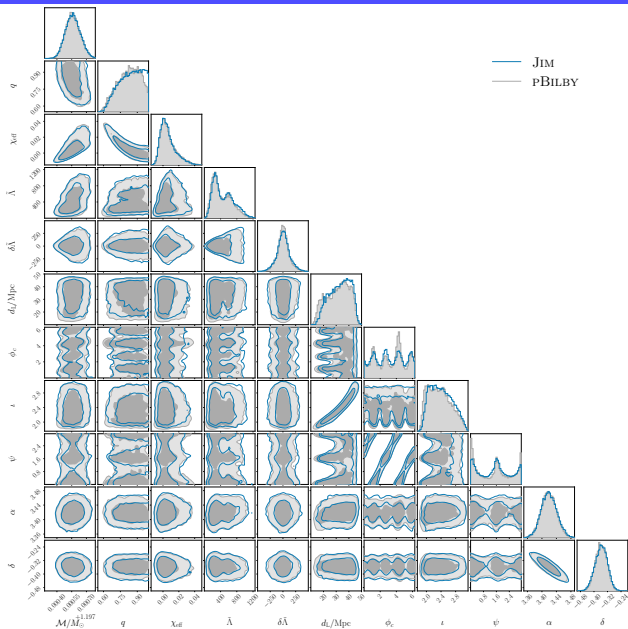
Parameter	GW170817		GW190425	
	TF2	NRTv2	TF2	NRTv2
$\mathcal{M}$	0.001725	0.000516	0.003557	0.002461
$q$	0.005212	0.007894	0.004837	0.002960
$\chi_1$	0.005633	0.004301	0.002794	0.004825
$\chi_2$	0.003030	0.002671	0.002416	0.003041
$\Lambda_1$	0.001062	0.002208	0.008556	0.000783
$\Lambda_2$	0.000559	0.002186	0.005808	0.003576
$d_L$	0.001544	<b>0.01847</b>	0.001273	0.002878
$\phi_c$	0.003500	0.010714	0.003338	0.006126
$\cos \iota$	0.001615	0.012851	0.006400	0.005279
$\psi$	0.004048	0.011036	0.001516	0.003730
$\alpha$	<b>0.014008</b>	0.001258	<b>0.009822</b>	<b>0.012291</b>
$\sin \delta$	0.009570	0.001761	0.008934	0.009228

Figure 1



# GW170817 with IMRPhenomD\_NRTidalv2

Figure 2



# GW190425 with TaylorF2

Figure 3

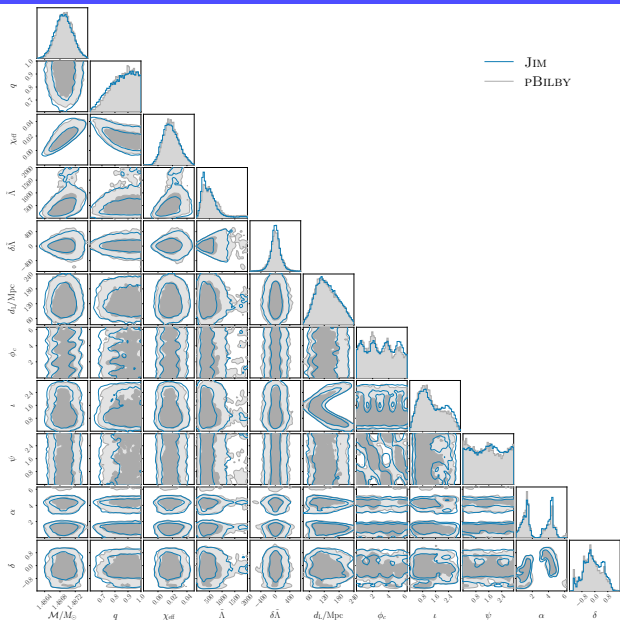


Figure 4

