# SEARCHES FOR HEAVY NEUTRAL LEPTONS IN THE ATLAS EXPERIMENT

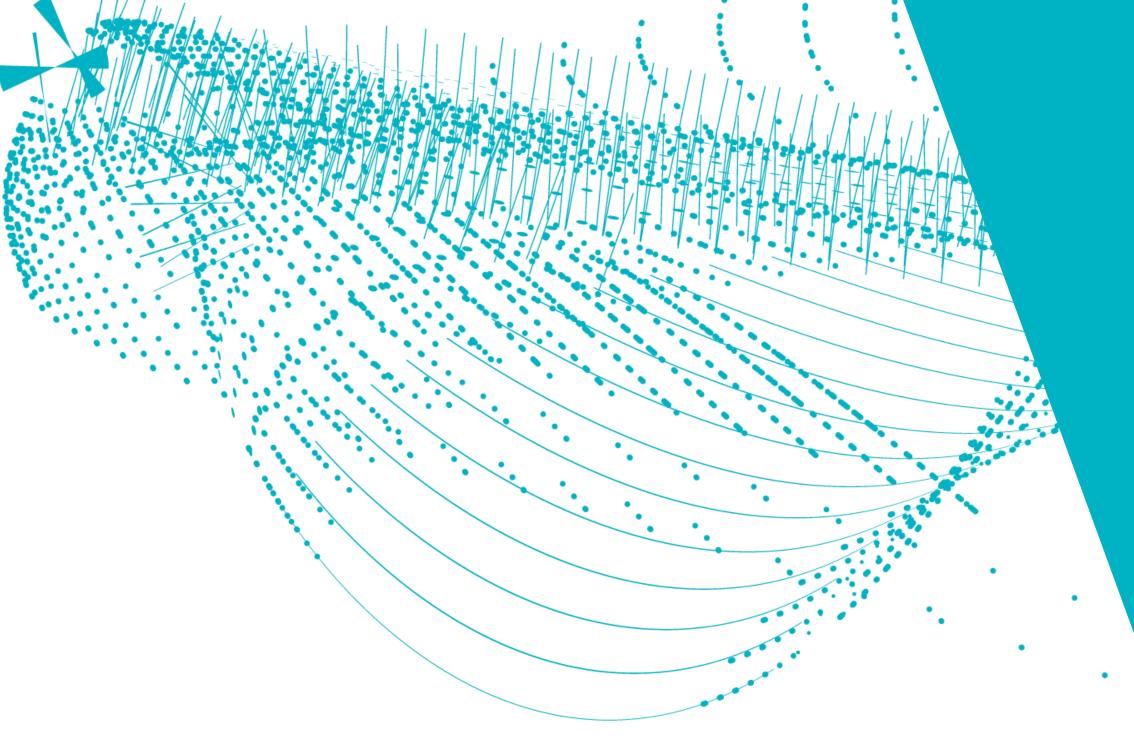
# 7<sup>TH</sup> NOVEMBER 2024

NNV FALL CONFERENCE



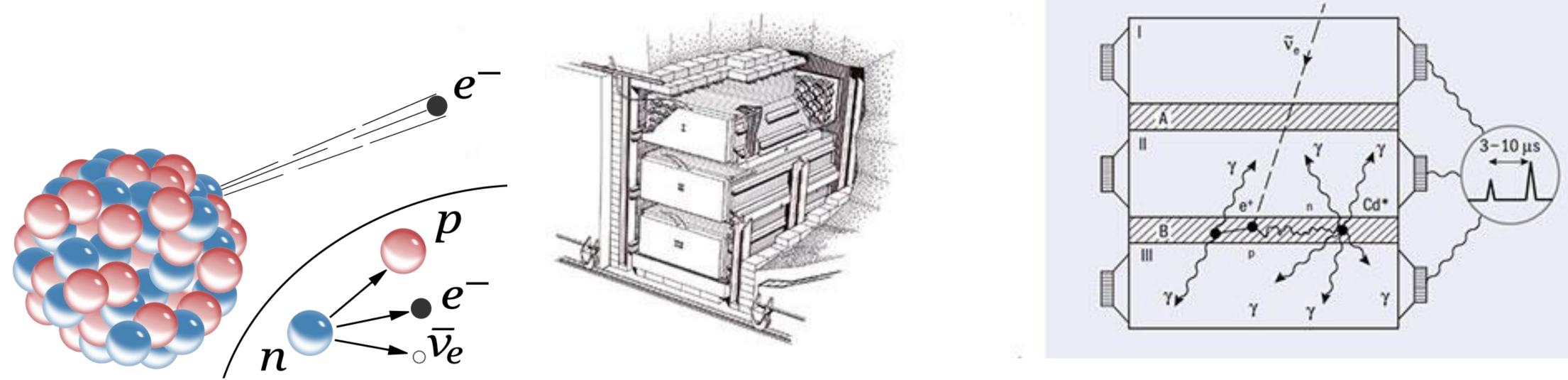


Nikhef





## Early neutrino physics



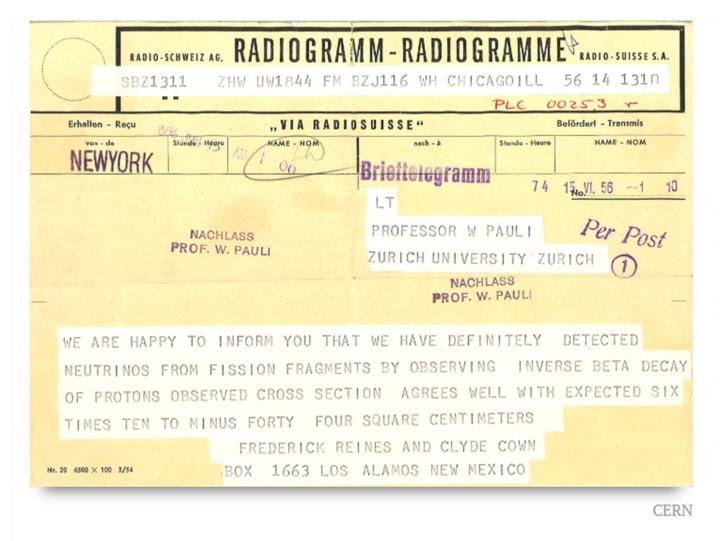
#### $(A, Z) \rightarrow (A, Z \pm 1) + e^{\mp} + \text{nothing else visible}$

Beta decay

### **NEUTRINO NAMED**

Enrico Fermi, father of the world's first nuclear reactor, popularized the term "neutrino." It is Italian for "little neutral one," summing up key properties of the neutrino: it's lack of charge and it's incredibly tiny size. Fermi <u>proposed a theory</u> that included Pauli's hypothesized particle; this was the first theory of one of the four fundamental forces, the weak force.

#### **NVV Fall Conference**



# FIRST EXPERIMENTAL EVIDENCE

A team of scientists led by Los Alamos National Laboratory physicists Frederick Reines and Clyde Cowan observed the first evidence of neutrinos as part of <u>Project Poltergeist</u>. The neutrino was then, as it still is now, considered a "ghostly" particle, flitting through matter without leaving much of a trace. The scientists used a nuclear reactor, an incredibly dense source of neutrinos, to finally catch the ghost. For five months, they collected data with a 10-ton detector placed next to a fission reactor at the Savannah River Plant. They announced their success in a 1956 telegram to Wolfgang Pauli: "We are happy to inform you that we have definitively detected neutrinos."



# Next decades – physicists in caves and mountains



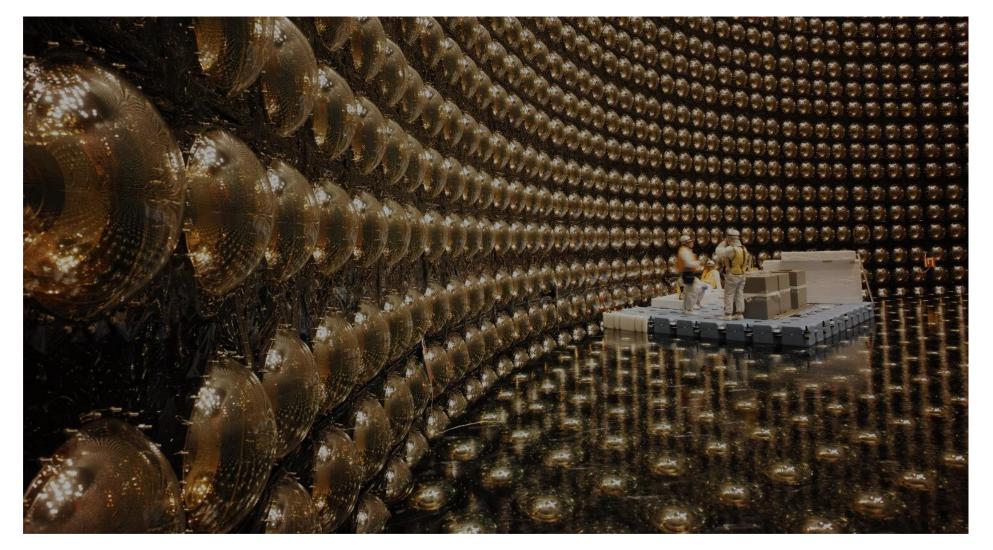
#### Homestake



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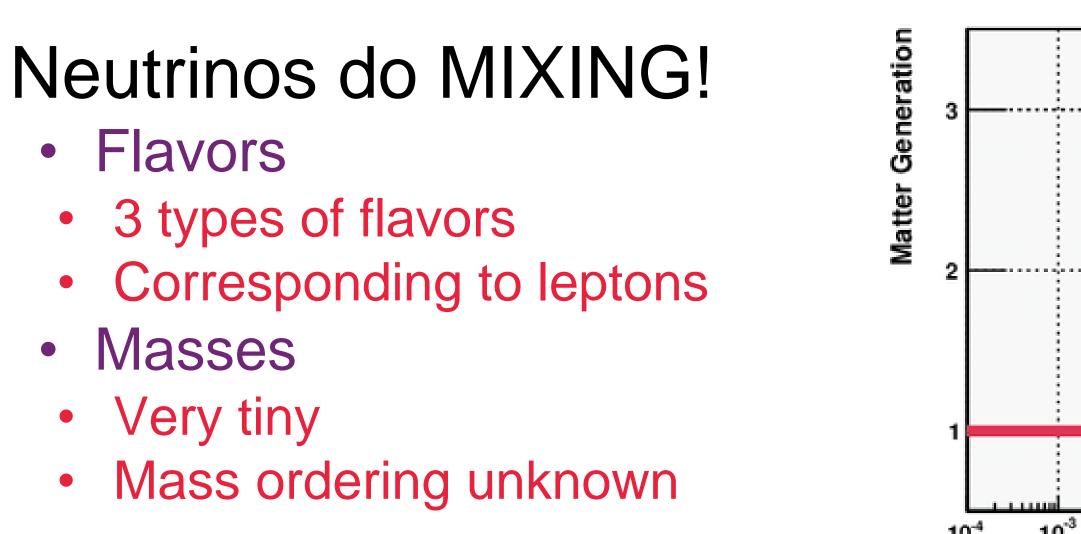
#### Kamiokande





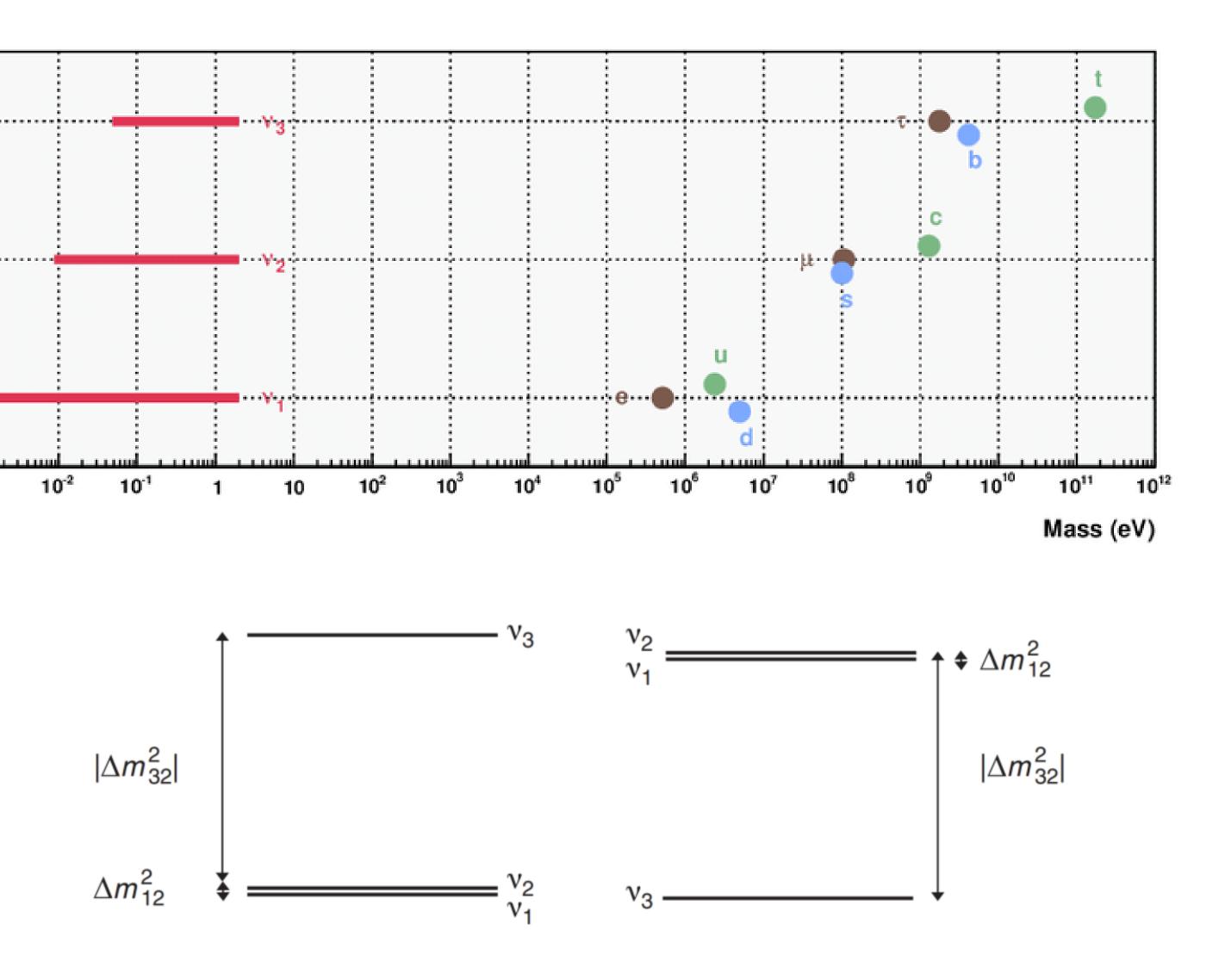


# What have they found?



#### Extra parameters in the SM! How can we explain the masses?

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## Type-I seesaw mechanism

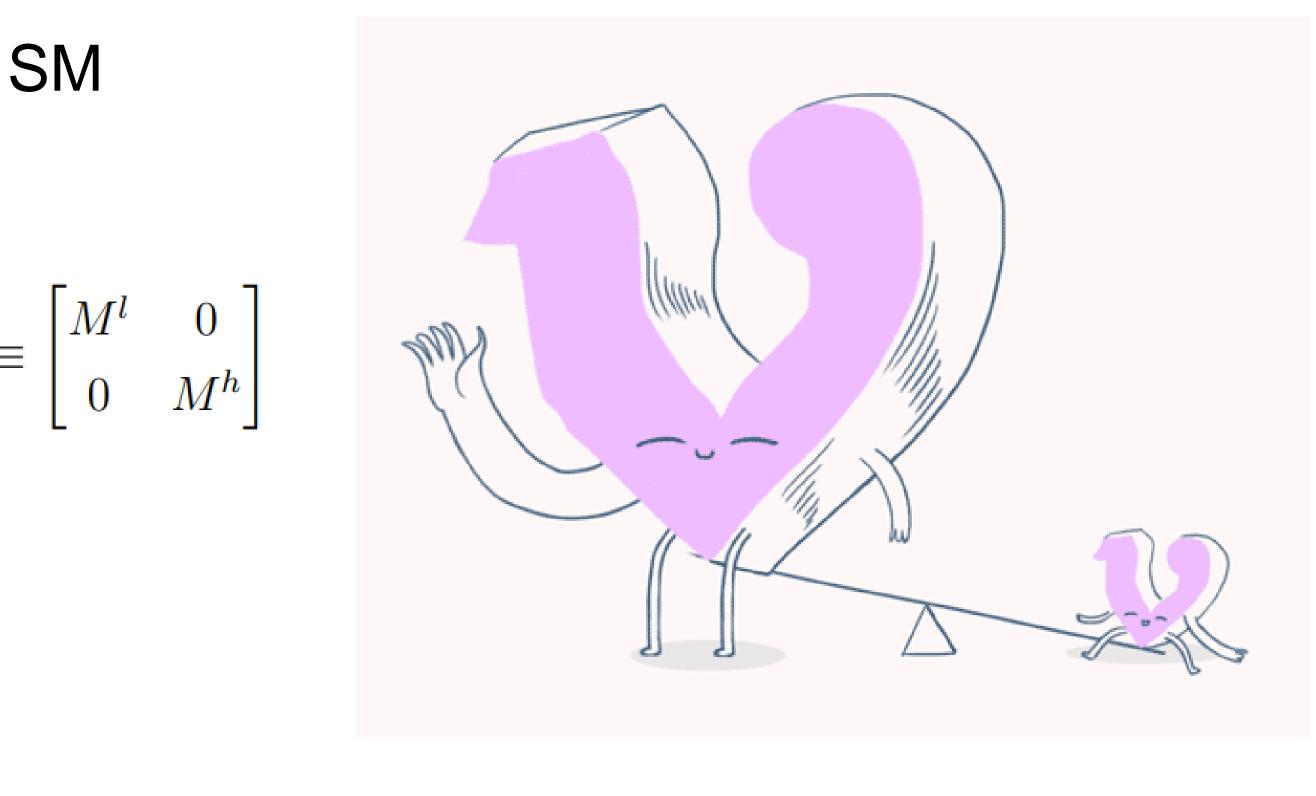
One way to explain the tiny masses We can add Majorana neutrinos to the SM

$$-\mathcal{L}_{M\nu} = \frac{1}{2} \bar{\nu}_l M^l \nu_l + \frac{1}{2} \bar{\nu}_h M^h \nu_h$$
$$M_{\nu} \rightarrow \begin{bmatrix} -V^T M_D^T M_N^{-1} M_D V & 0\\ 0 & U^T M_N U \end{bmatrix} \equiv$$

#### We have

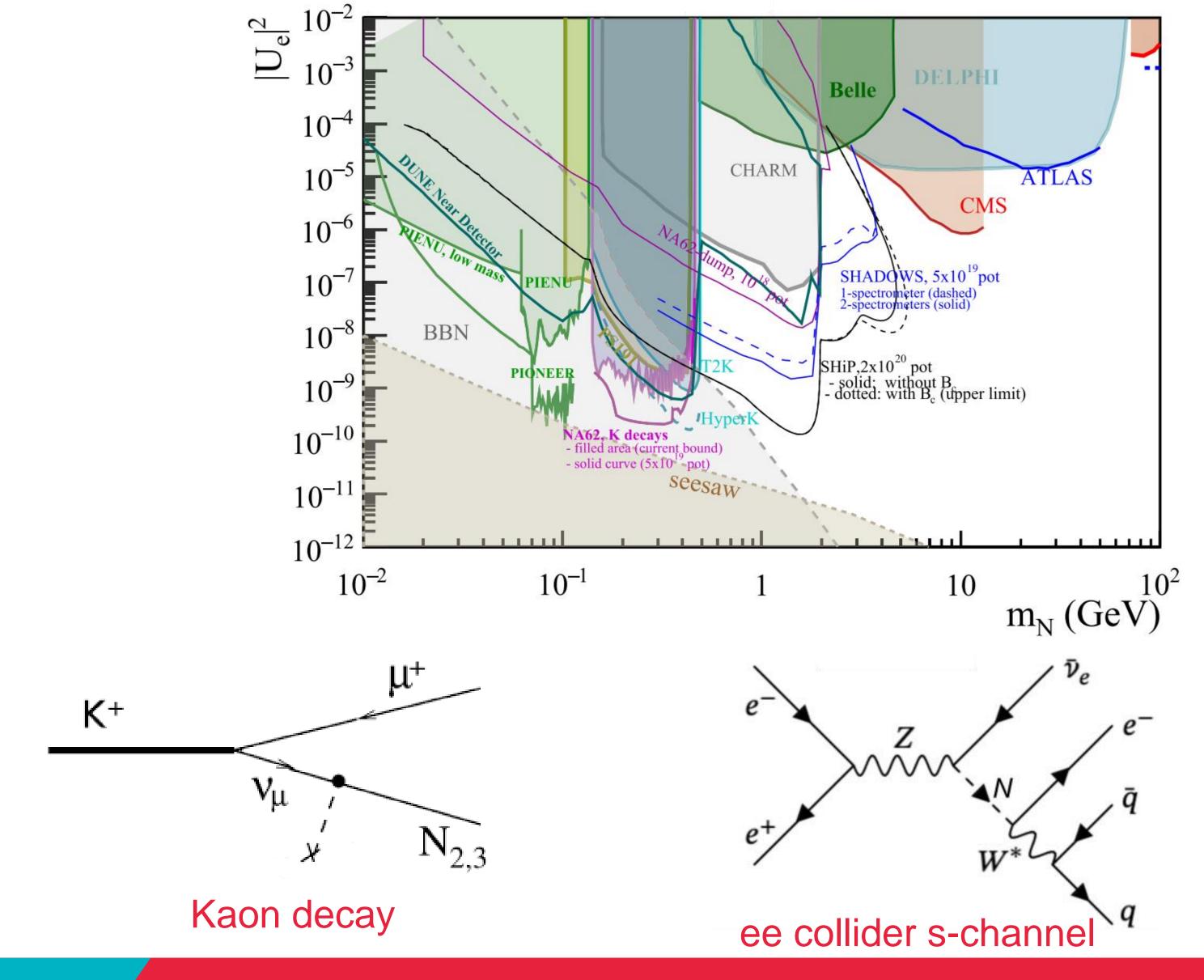
- Light neutrino mass ~  $M_N^{-1}$
- Heavy neutrino mass  $\sim M_N$

Simple, minimal assumption! Heavy particles, that's what we can search for in the LHC!





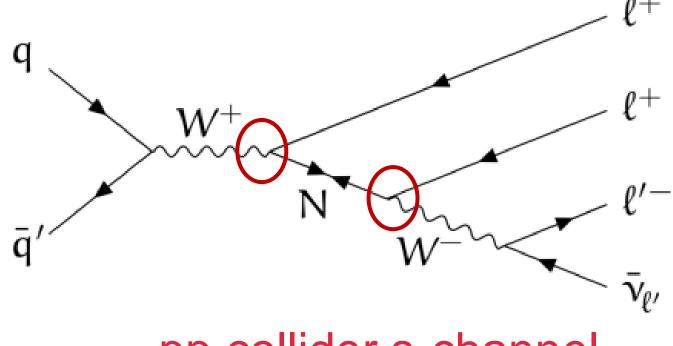
### **Current status and exclusion limits**



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#### Heavy Neutrino coupling often described by U<sub>e or</sub> V<sub>eN</sub>

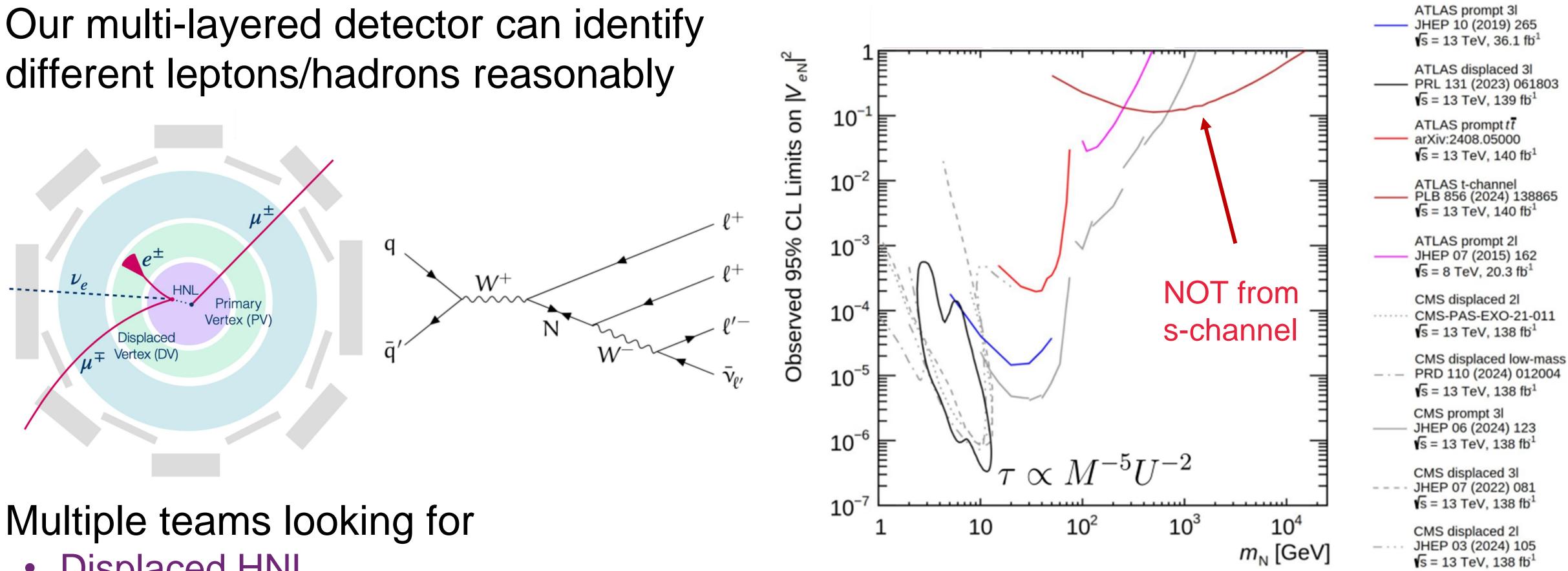


pp collider s-channel



HNL in ATLAS

# What about ATLAS?

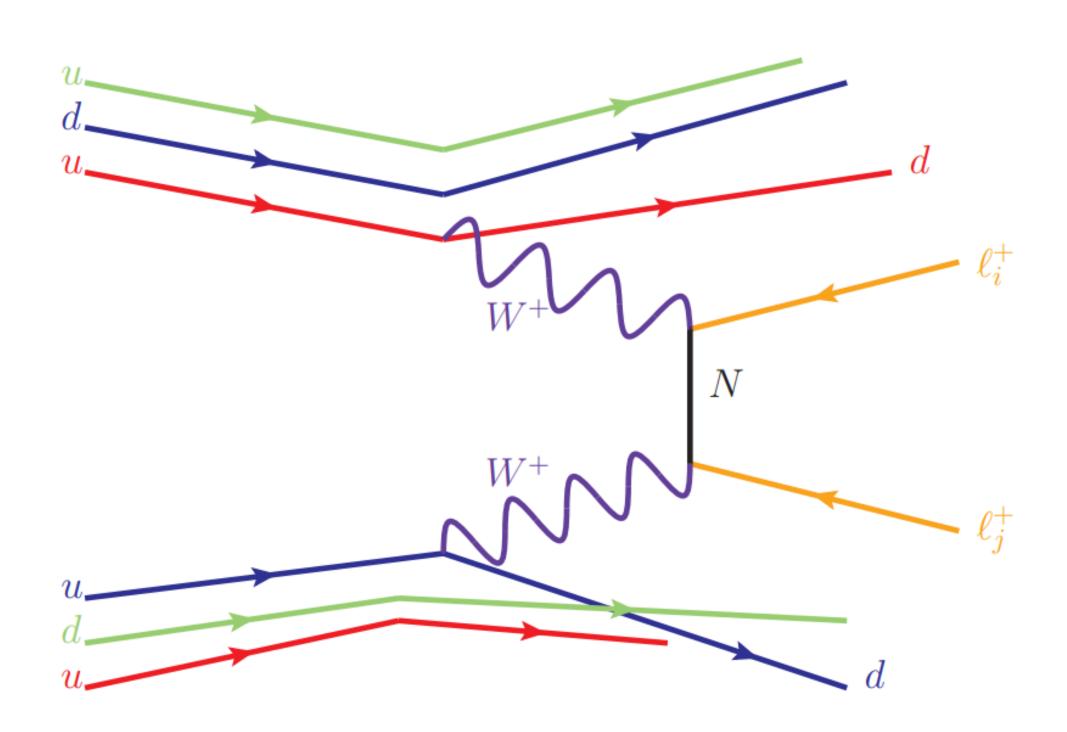


- Displaced HNL
- s-channel
- Good result for  $\mu$  and e

s-channel contributes up to  $m_N \sim 100 \text{ GeV}$ 

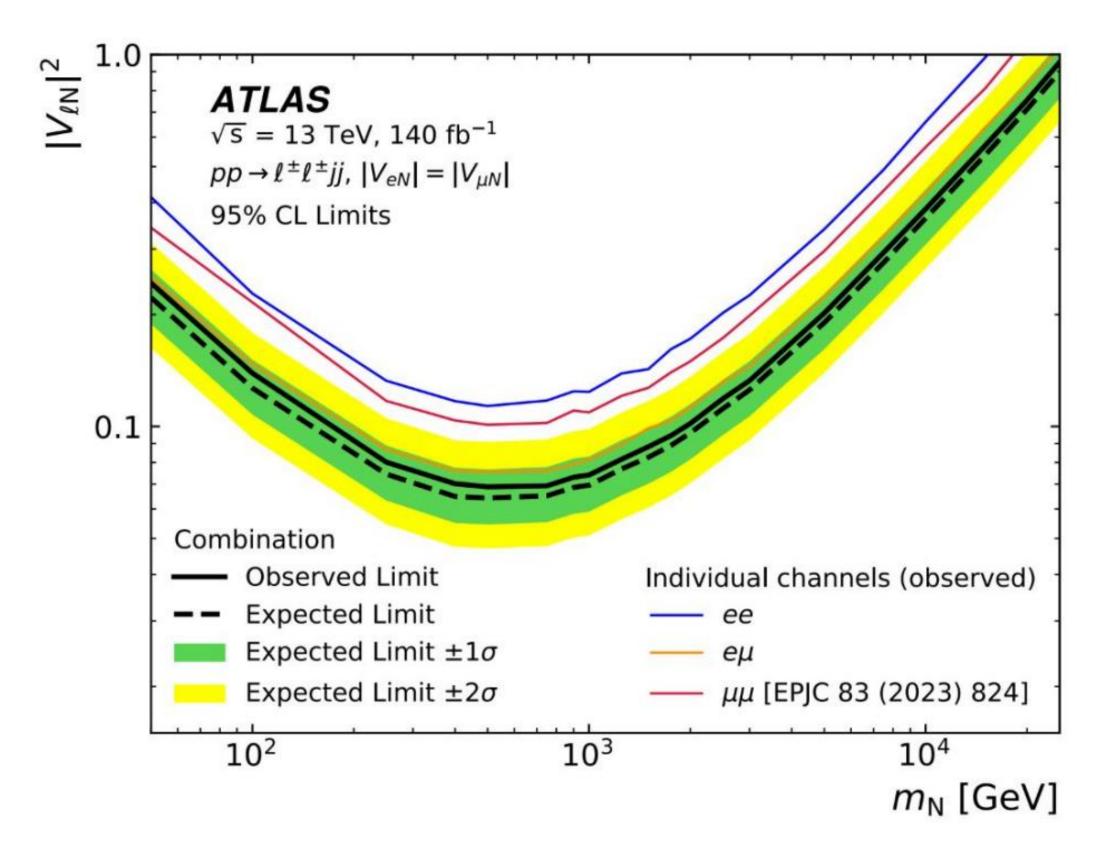


### t-channel HNL in ATLAS



#### Many interesting physics keywords

- VBS, HNL, LFV, Majorana neutrino
- May solve many problems at once!



Although no signal for  $e/\mu$ , But best limit is set at ~0.05 for 1TeV

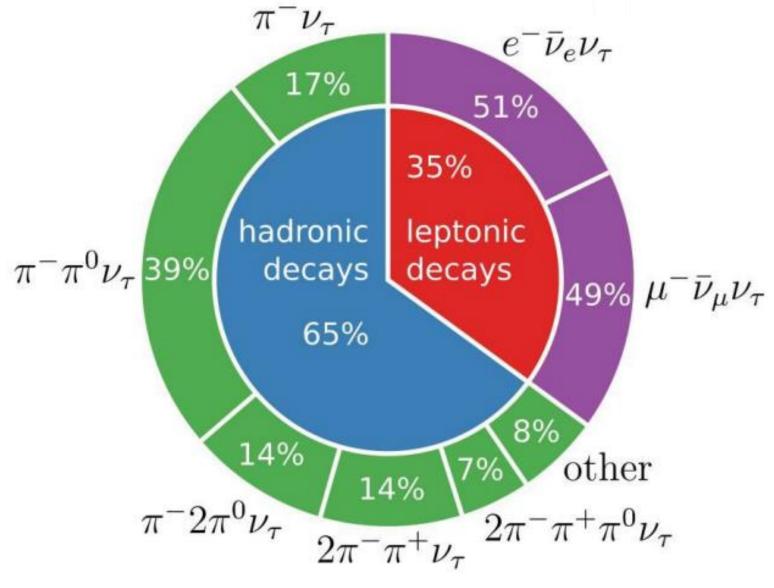


### t-channel with $\tau$

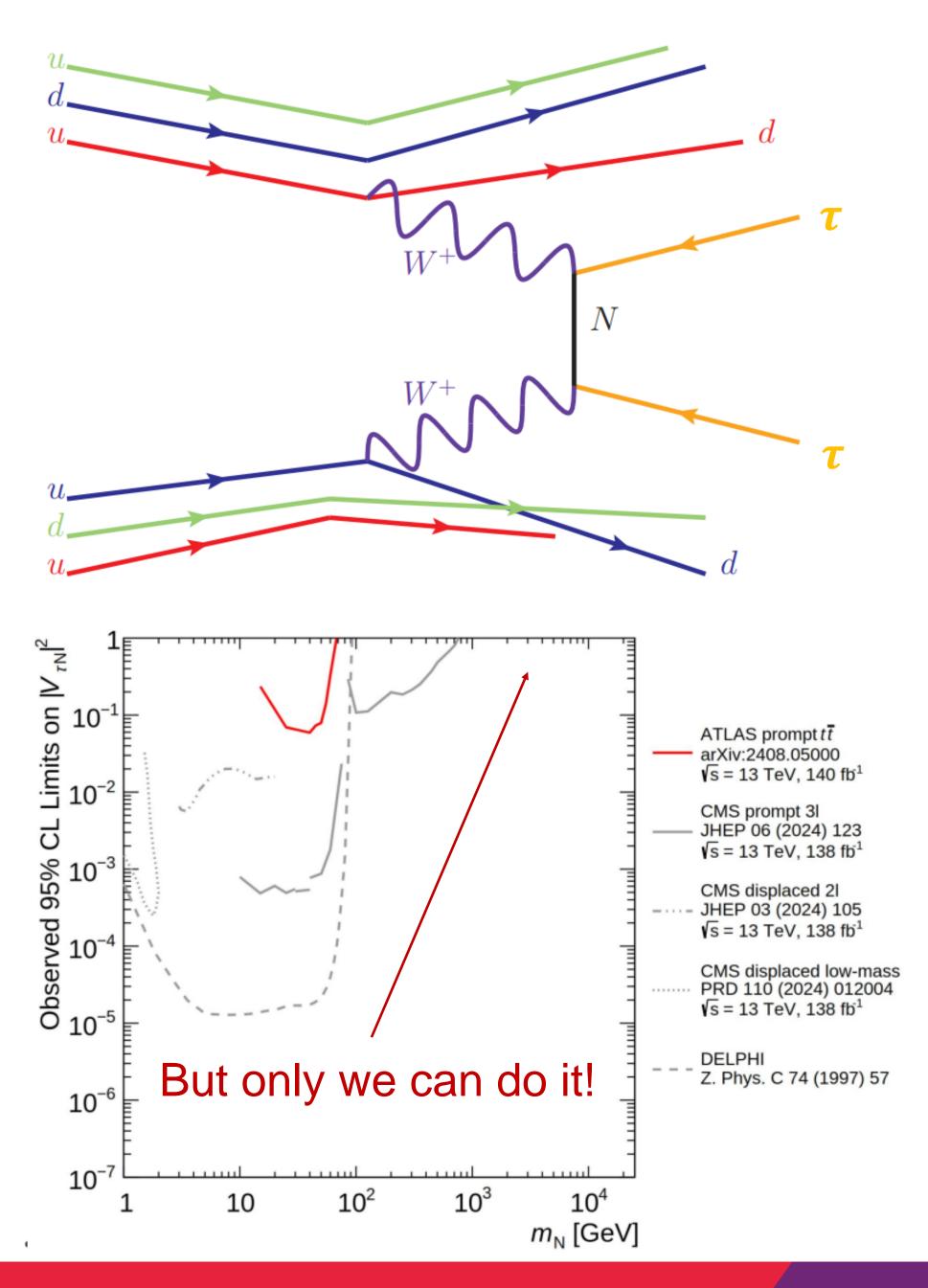
### Uncovered $\tau$ channel

- Much harder channel
- Worse reconstruction
- More fakes
- Complicated decays

(2  $\tau$  means many combinations...)



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## t-channel basic strategy

Some terminologies

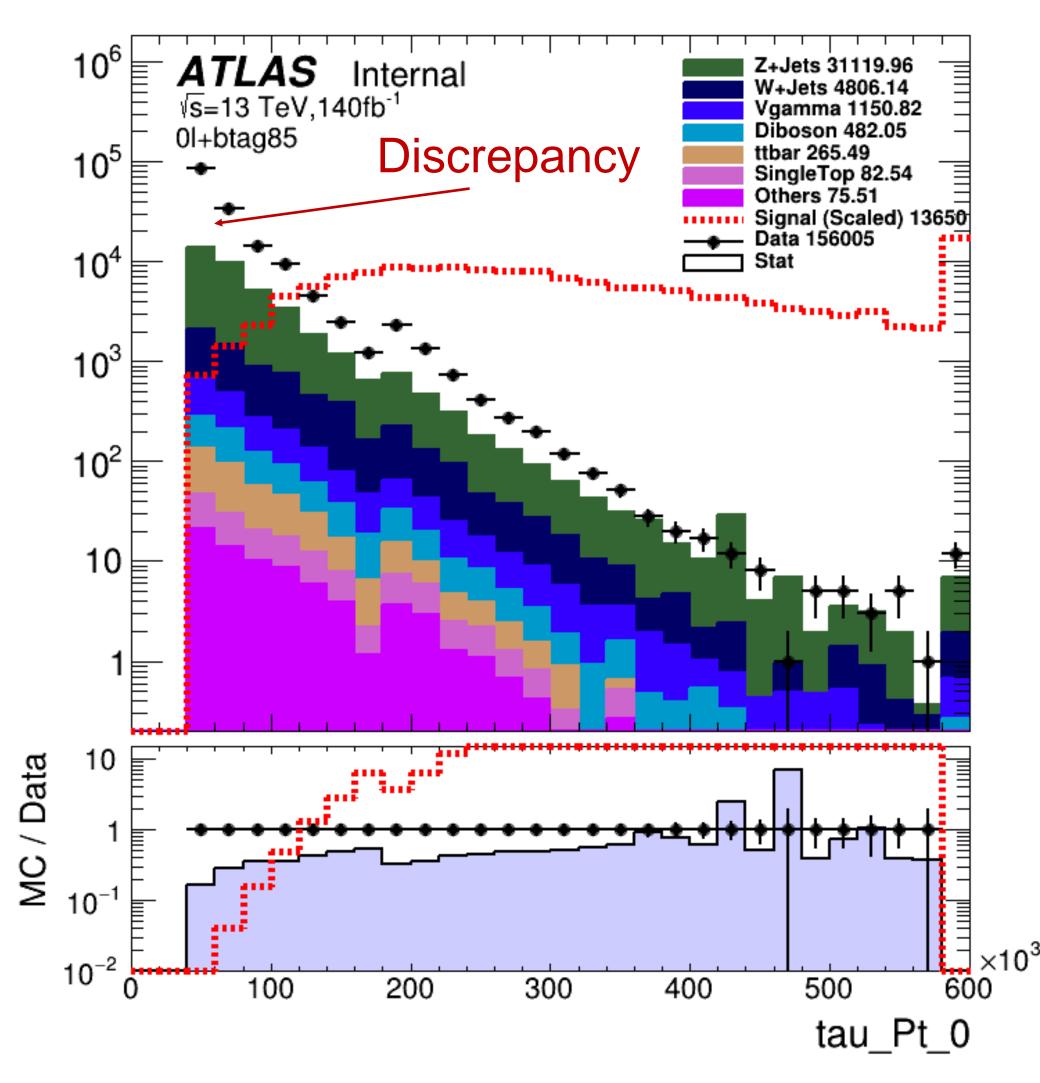
- SS = same sign pair
- OS = opposite sign pair
- MC = simulation

Regions

- $SS\tau\tau$  regions blinded
- $OS\tau\tau$  regions for validation • (as a proxy for  $SS\tau\tau$  background)

First look

- Simply selecting events with 2 jets and 2  $\tau$
- Comparison to data in  $OS\tau\tau$  region
- Very high  $p_T$  signature
- A huge discrepancy...

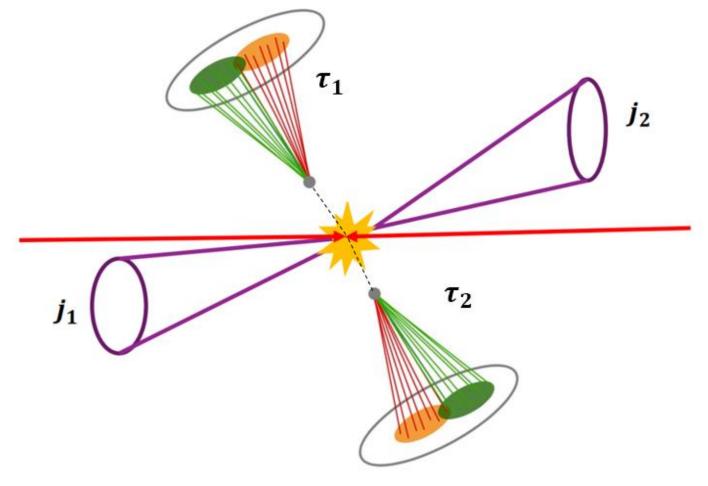


 $OS\tau\tau$  (data, MC) vs  $SS\tau\tau$  (signal)

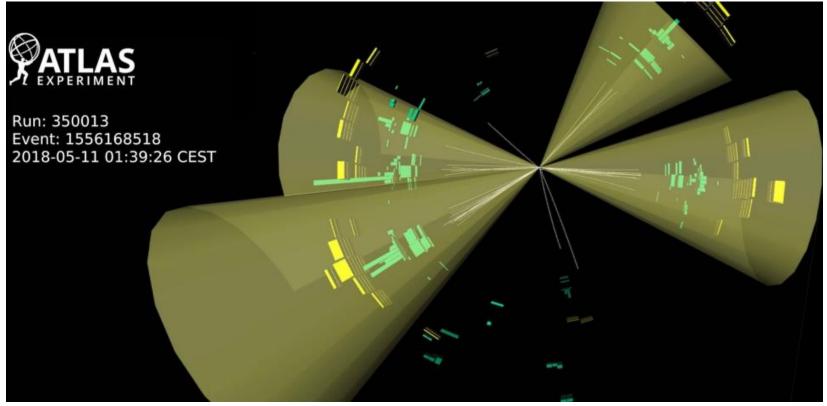


# Explaining the discrepancy

#### The reason is FAKES from QCD (mislabeling $j \rightarrow \tau$ )

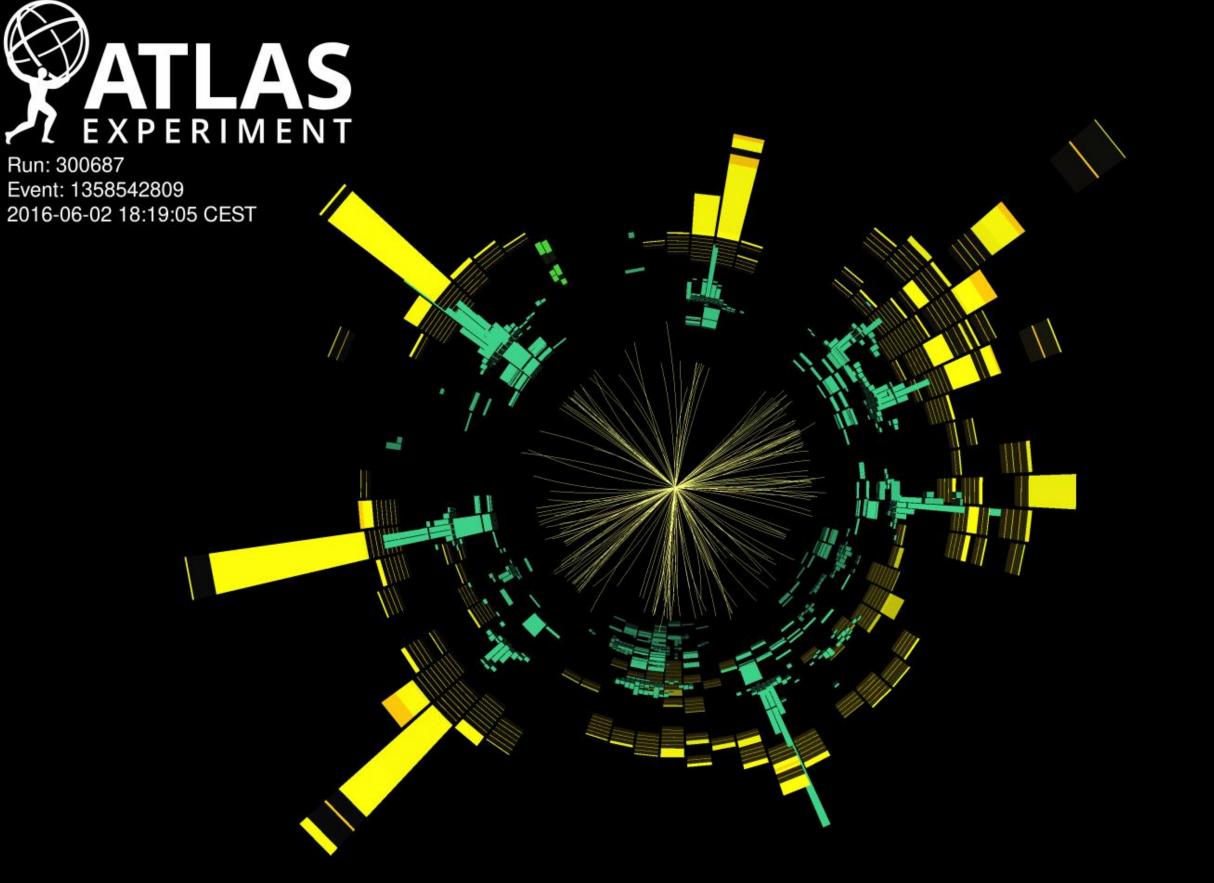






#### We want 2 $\tau$ and 2 jets

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But we often record this, how do we tell which is which? This is also hard to simulate correctly...



### Quality selection

 $OS\tau\tau$  and  $SS\tau\tau$  regions could be complicated Let's look at 1 reconstructed  $\tau$  first.

What about using data to estimate the fakes? Use ID Score of  $\tau$ !

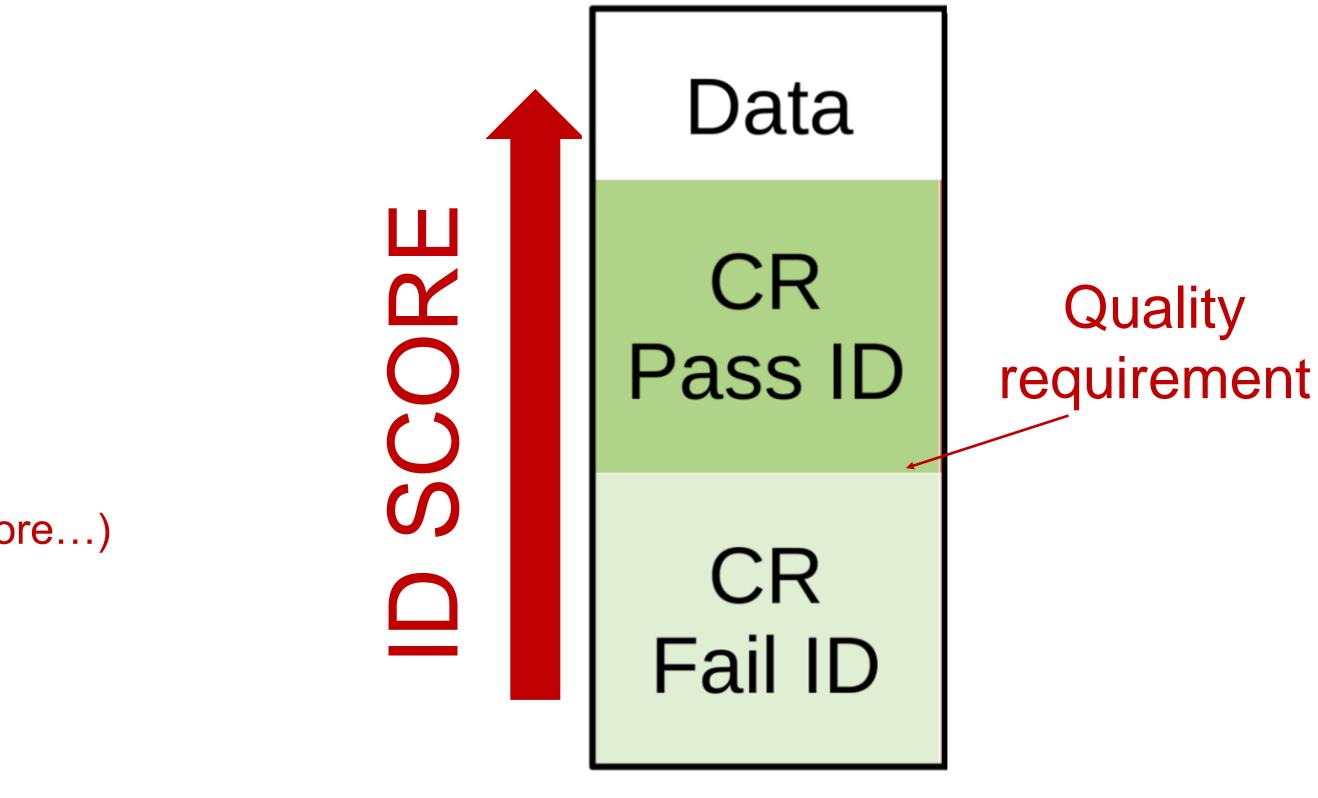
What is ID Score?

- RNN based Identification score (take into account tracks, calorimeter deposition, and more...)
- Quality measurement of reconstructed  $\tau$

Cutting at higher quality

• tighter selection, fewer fakes, fewer data Cutting at lower quality

looser selection, more fakes, more data



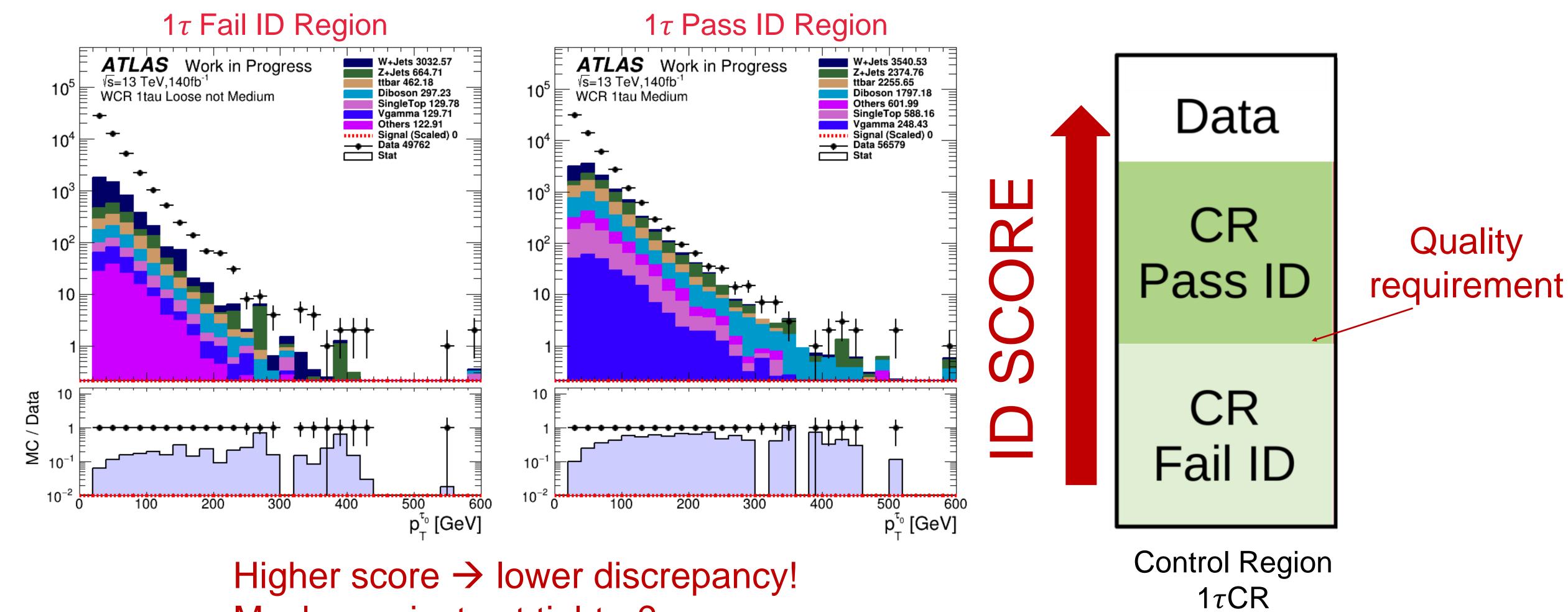
#### **Control Region** $1\tau CR$







## Quality selection

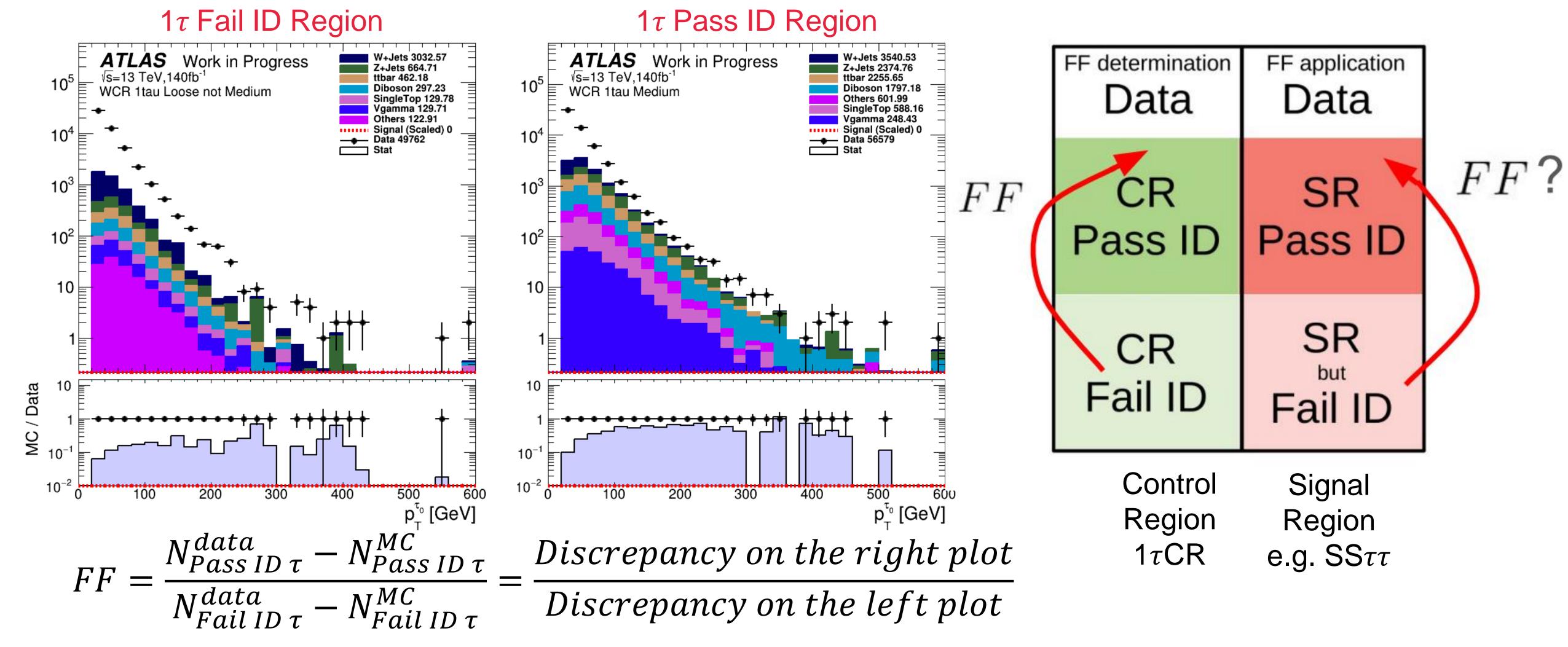


Maybe we just cut tighter? But we still have discrepancies...





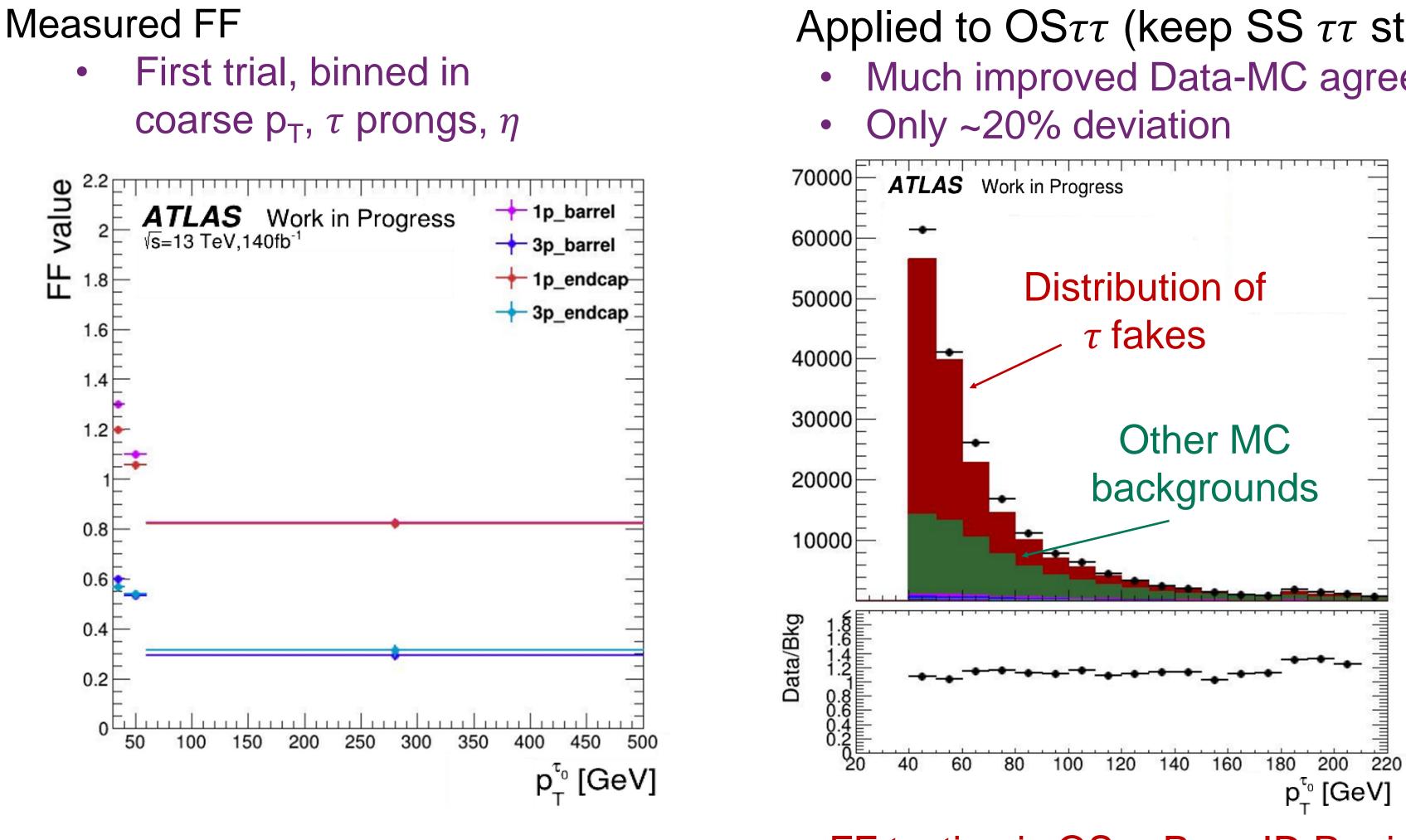
### Fake Factors definition



While we cannot predict the amount of fakes from pure simulation, this is a data-driven ratio!



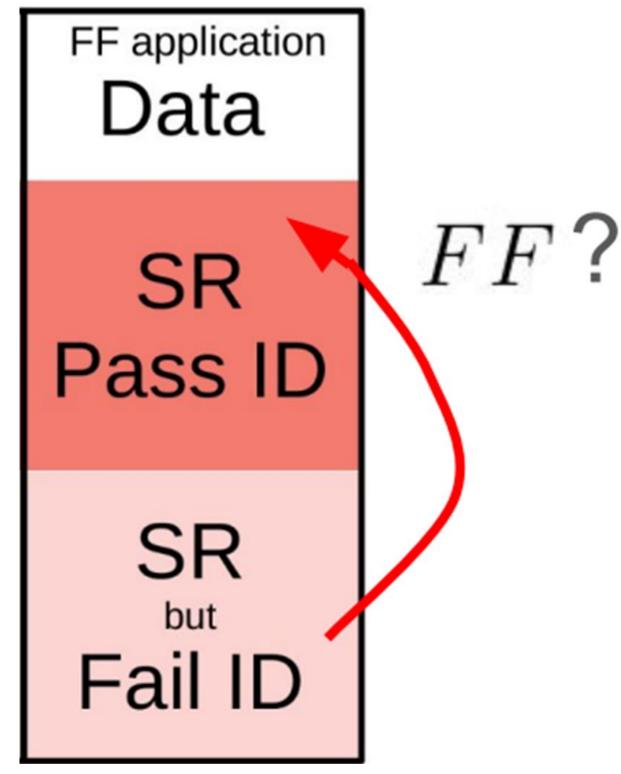
### Fake Factor application



FF measured in bins of  $p_{T}$ 

### Applied to $OS\tau\tau$ (keep SS $\tau\tau$ still blinded) Much improved Data-MC agreements



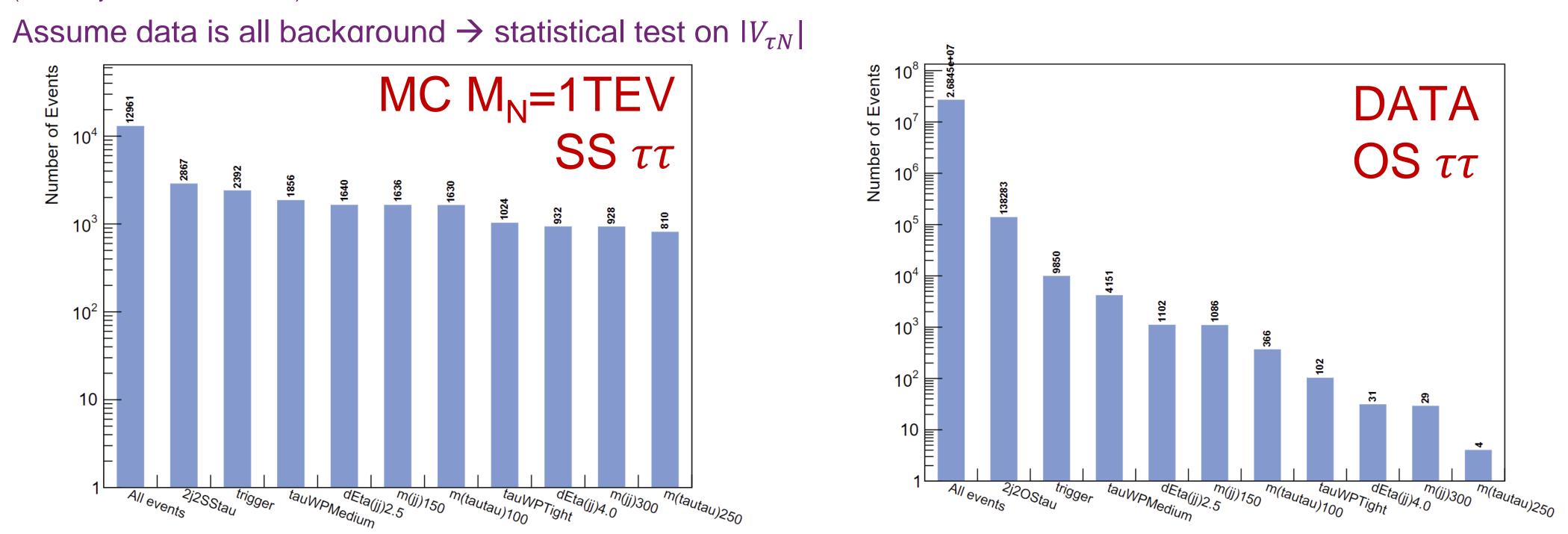




# Sensitivity estimate

Fake Factor is not entirely ready... Can we estimate our sensitivity? Testing our selections

- SS region is fully blinded  $\rightarrow$  Use OS distributions as proxy of our background
- Place progressive cuts on kinematics
- S/B ratio goes from 10<sup>-3</sup> to 10<sup>-2</sup> (raw MC yields vs data events)





# Summary

- Heavy neutral lepton above TeV scale is something unique in LHC
- Our team is working directly with theorists for both predictions and measurements
- Currently all the channels  $\{s t\} \times \{e \mu \tau\}$  are limited by statistics, significant improvement guaranteed for run 3 and HL-LHC

