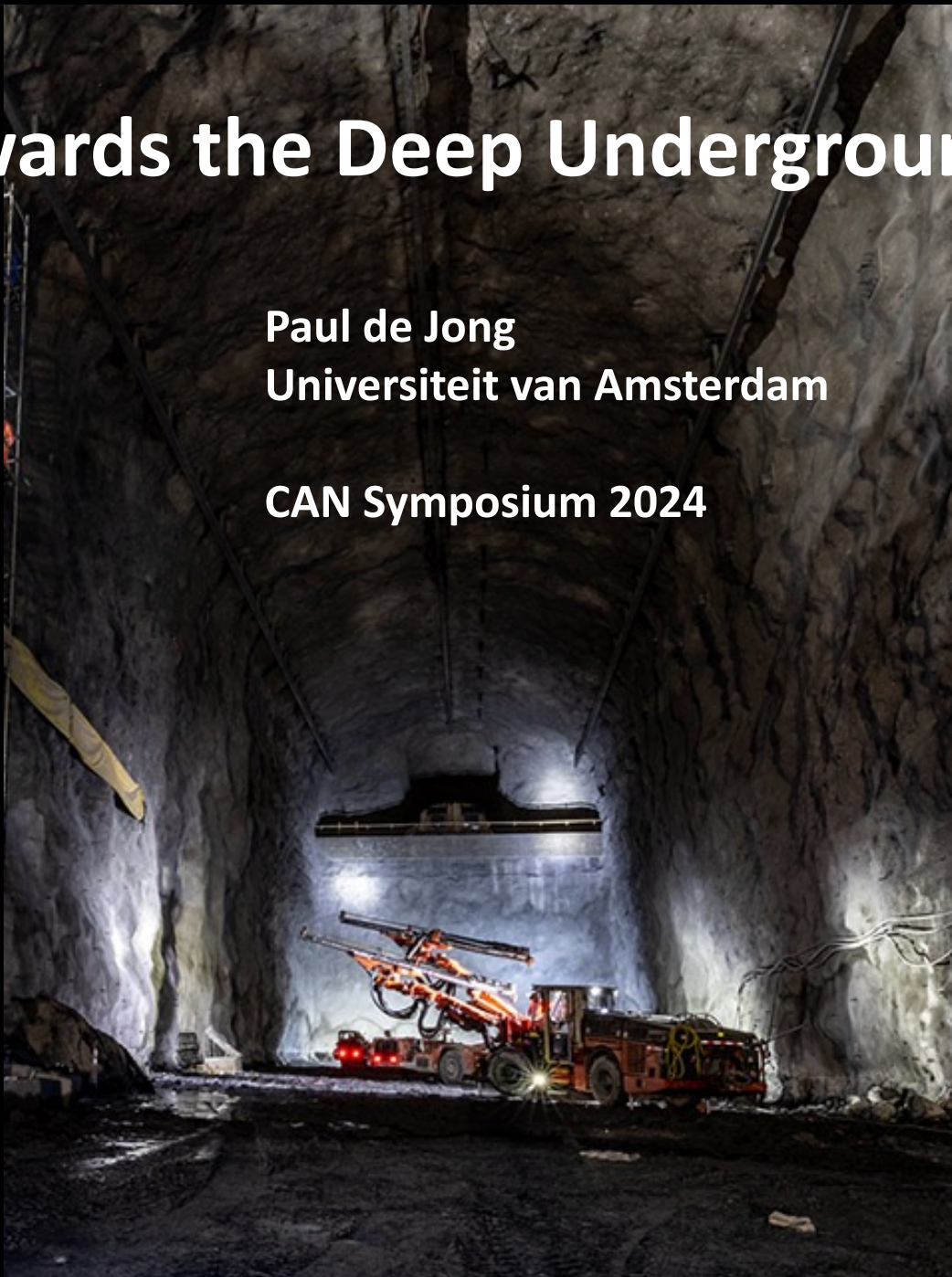


Towards the Deep Underground Neutrino Experiment (DUNE)

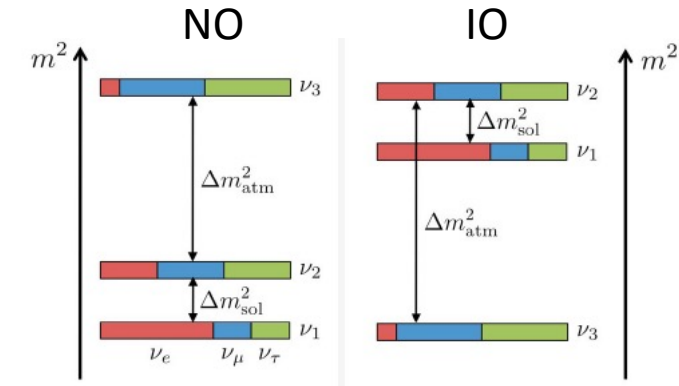
Paul de Jong
Universiteit van Amsterdam

CAN Symposium 2024



Oscillations with accelerator neutrinos

$$U_{\text{PMNS}} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & e^{-i\delta_{\text{CP}}} s_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{\text{CP}}} s_{13} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

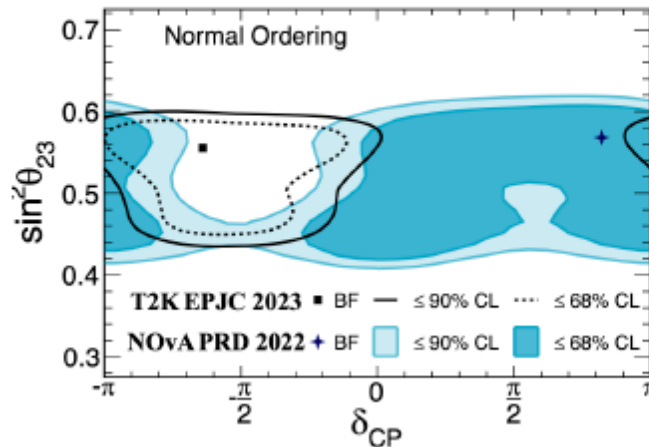


Long-baseline oscillations with accelerator neutrinos: sensitivity to θ_{23} , δ_{CP} , mass ordering, new physics. Fewer neutrinos than neutrino telescopes, but more control over beam, in particular ν vs $\bar{\nu}$

Testing CP-violation in the neutrino sector is top priority in neutrino physics.

Slight tension T2K & NOvA for NO

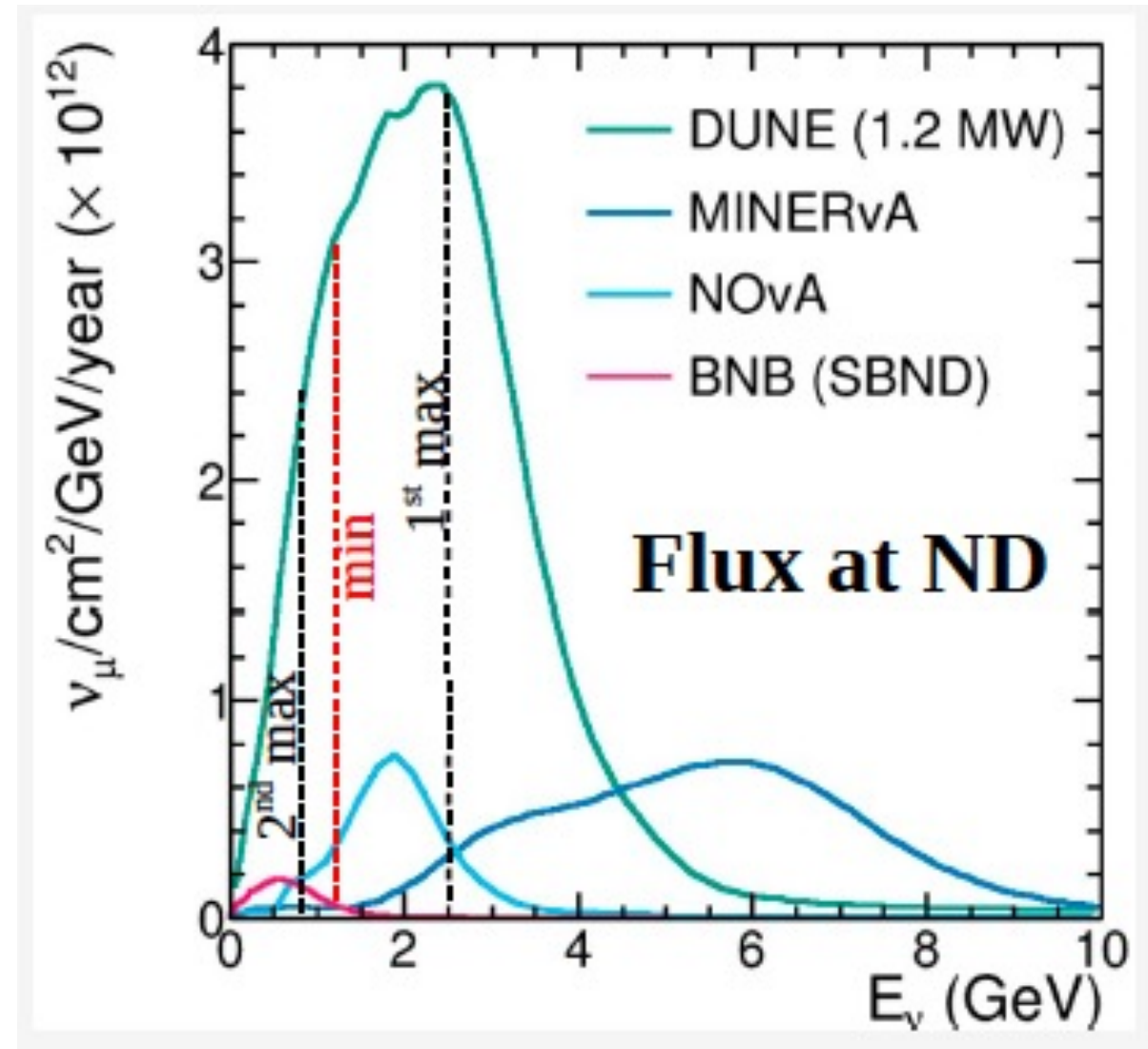
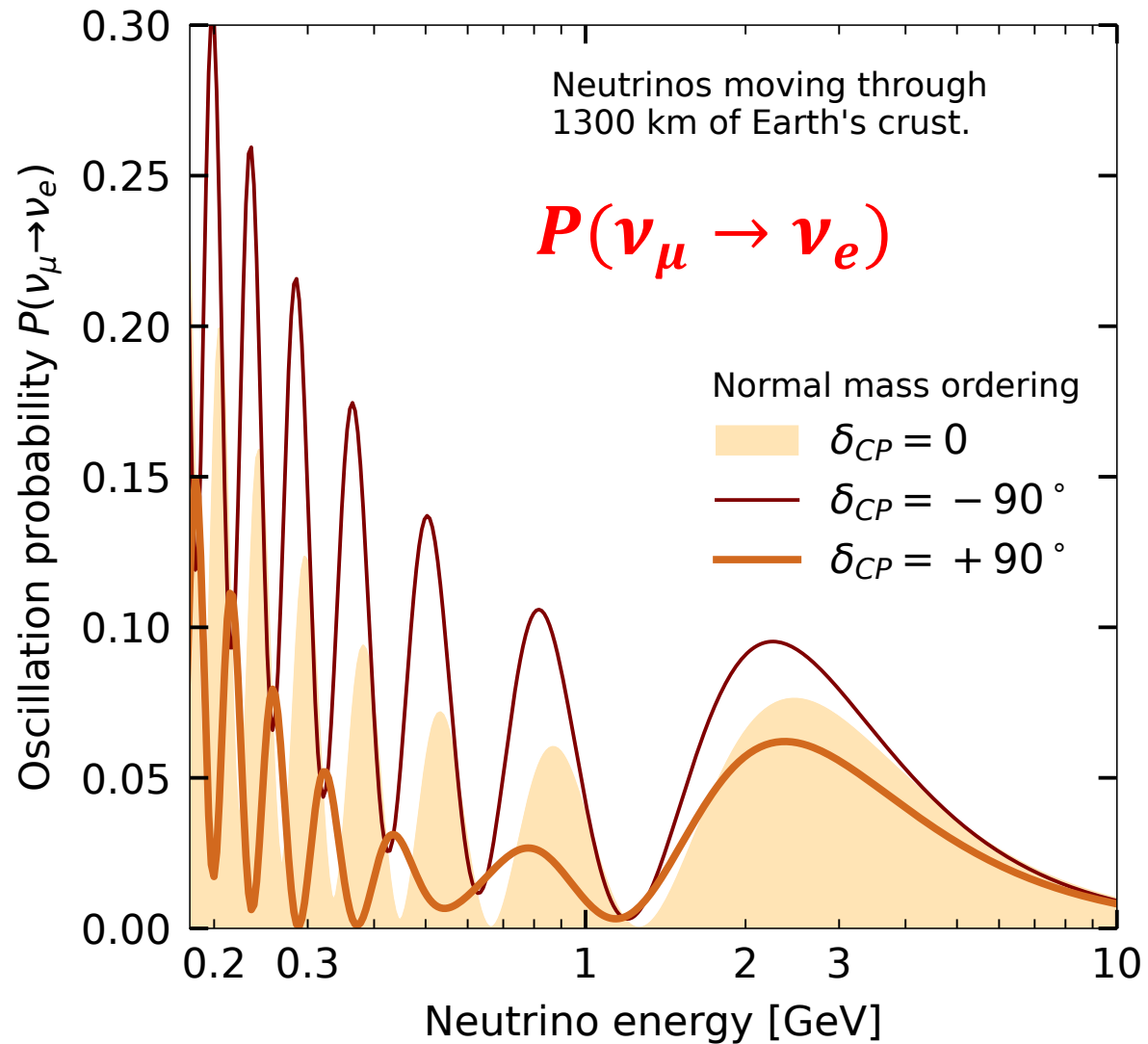
Current experiments: T2K, NOvA



Weak sensitivity to δ_{CP}
Tension between results T2K, NOvA

Next generation: HyperK, DUNE

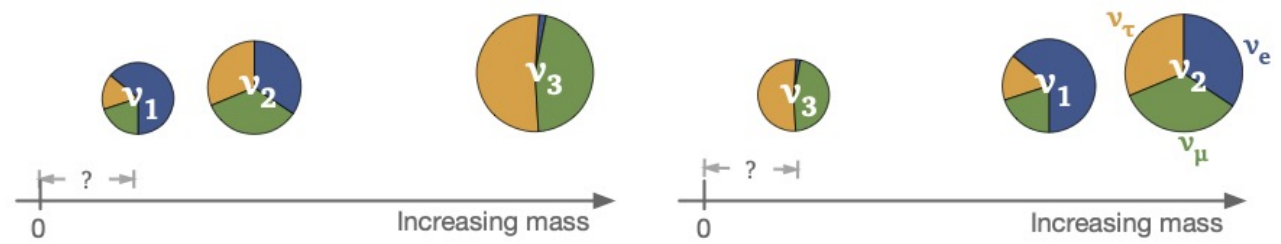
Oscillation probabilities and the DUNE wideband beam



Oscillation probabilities

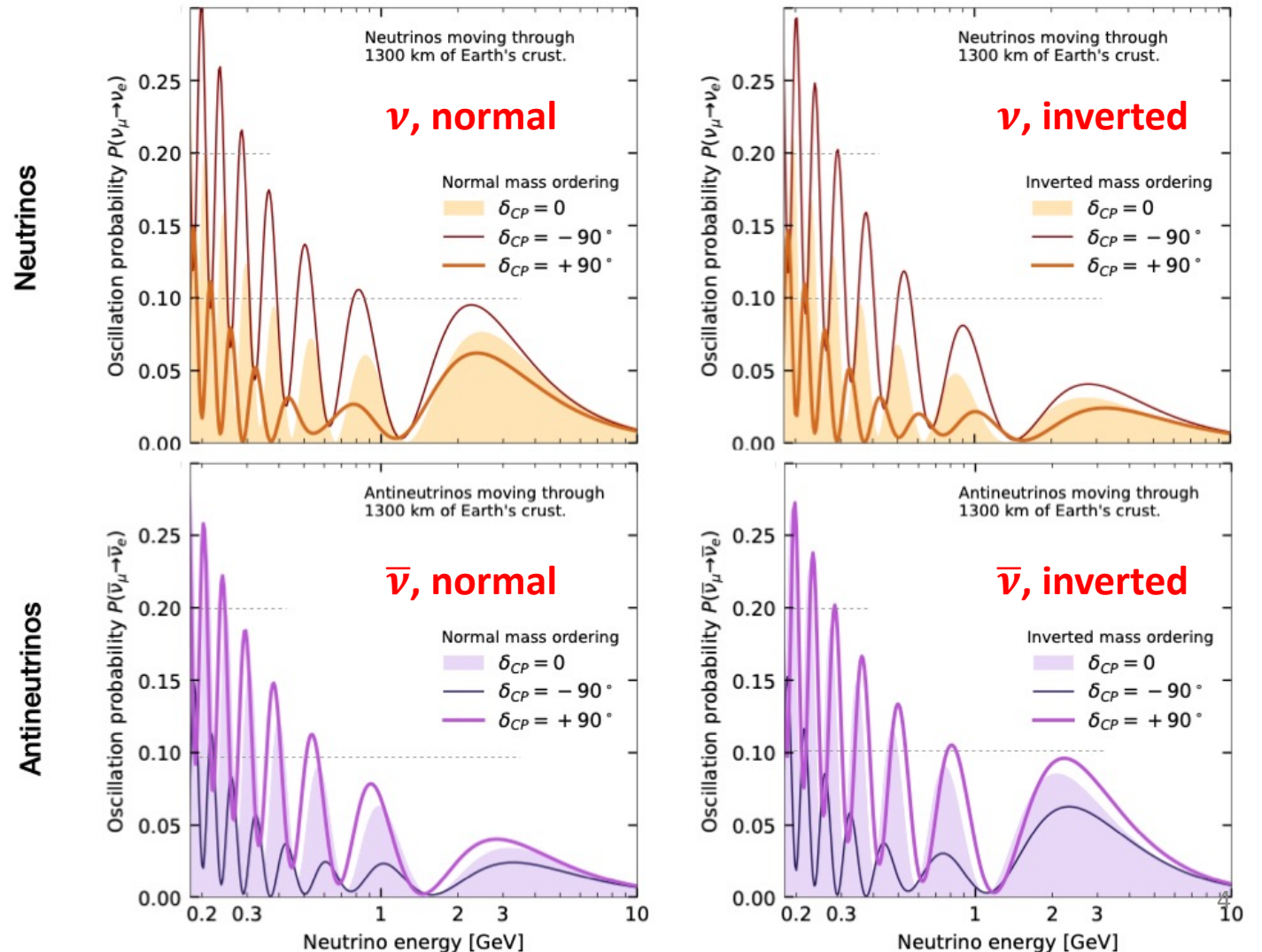
By using 1st and 2nd oscillation maximum, DUNE can disentangle neutrino mass ordering and CP-violation effects, in one expt. (Assuming SM, 3 generations)

If data cannot be fitted:
new physics?
(Sterile ν , BSM interactions)

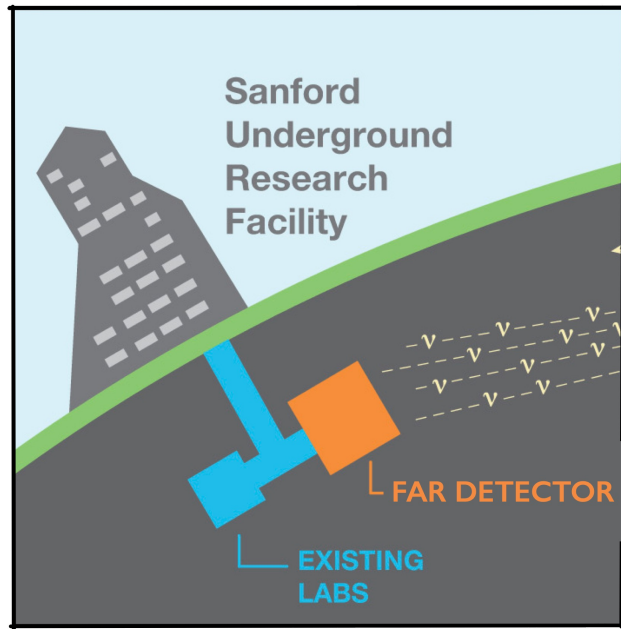


Normal Ordering ($m_3 > m_1$)

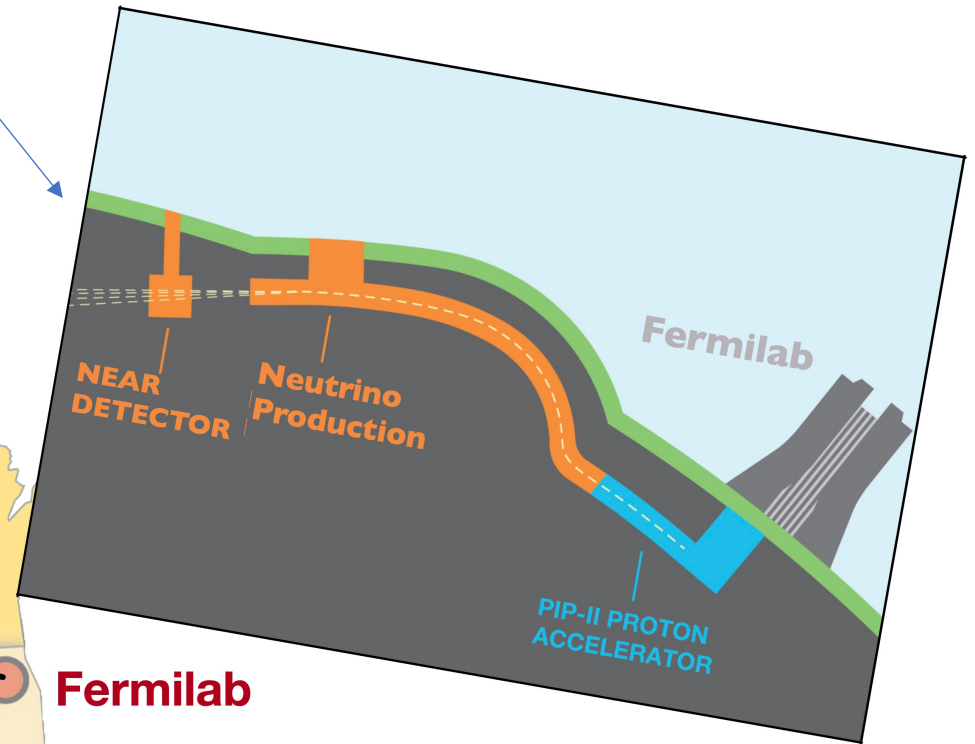
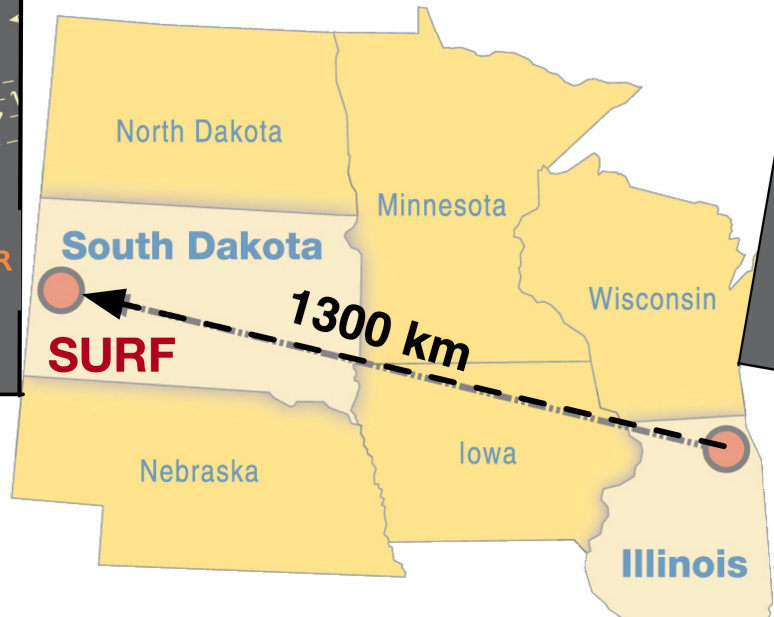
Inverted Ordering ($m_3 < m_1$)



Far Away (So Close)

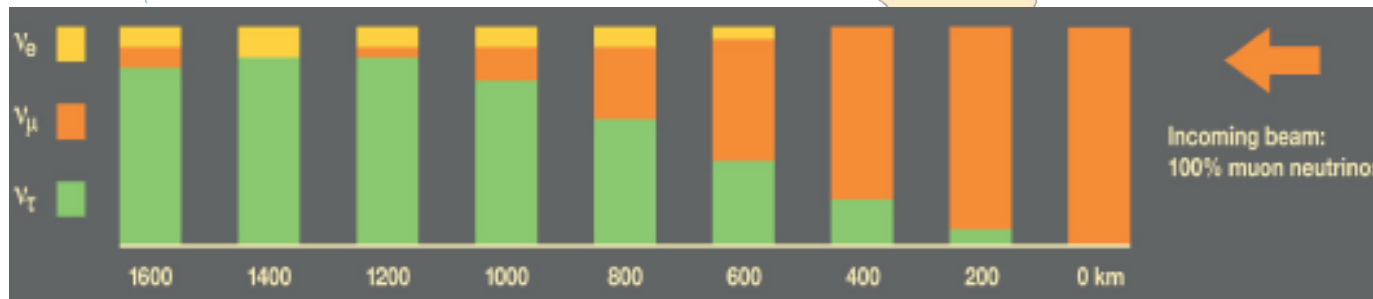


4 x 17 kton LAr detector
1500 m underground



New PIP-II accelerator

1.2 MW proton beam
to be upgraded to 2 MW

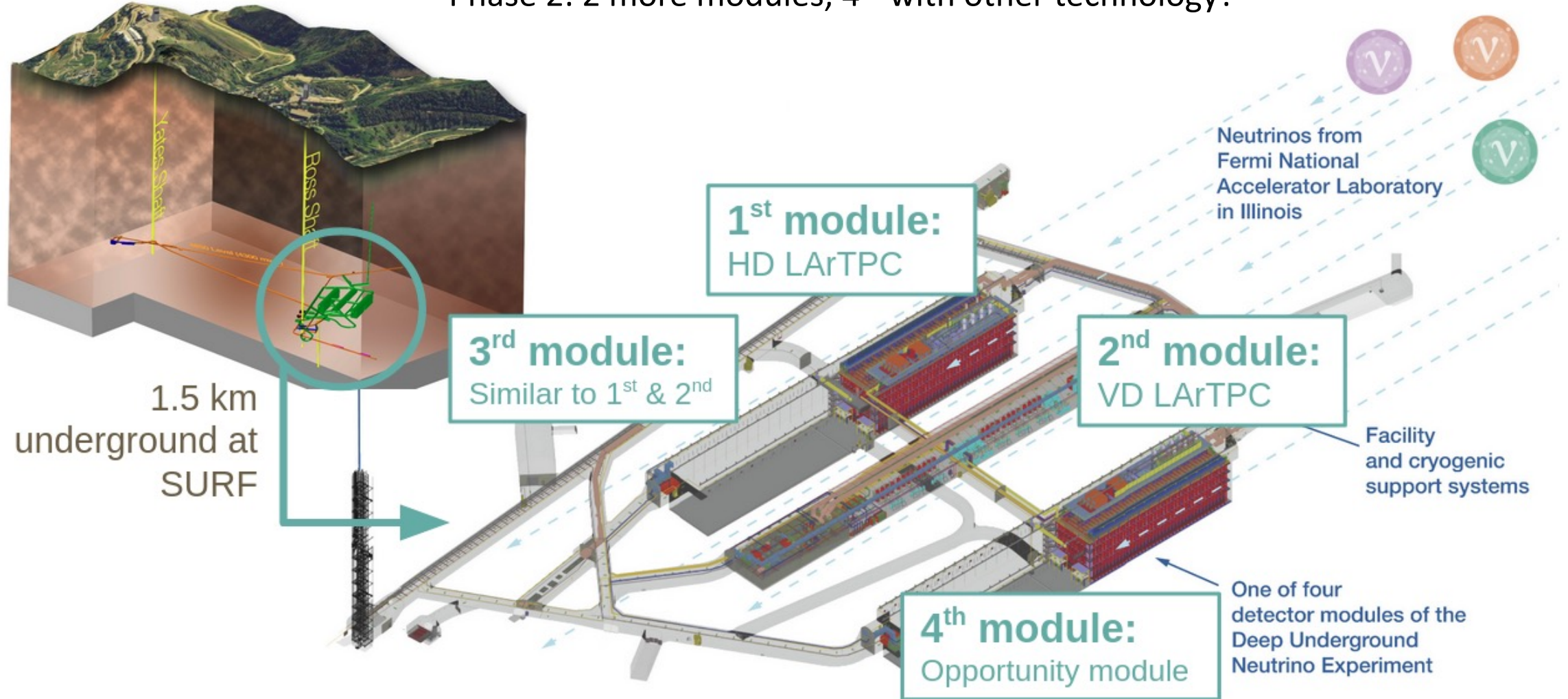


oscillations of a muon neutrino beam

Far site: Sanford Underground Research Facility, Lead, South Dakota

Phase 1: at least 2 modules

Phase 2: 2 more modules, 4th with other technology?

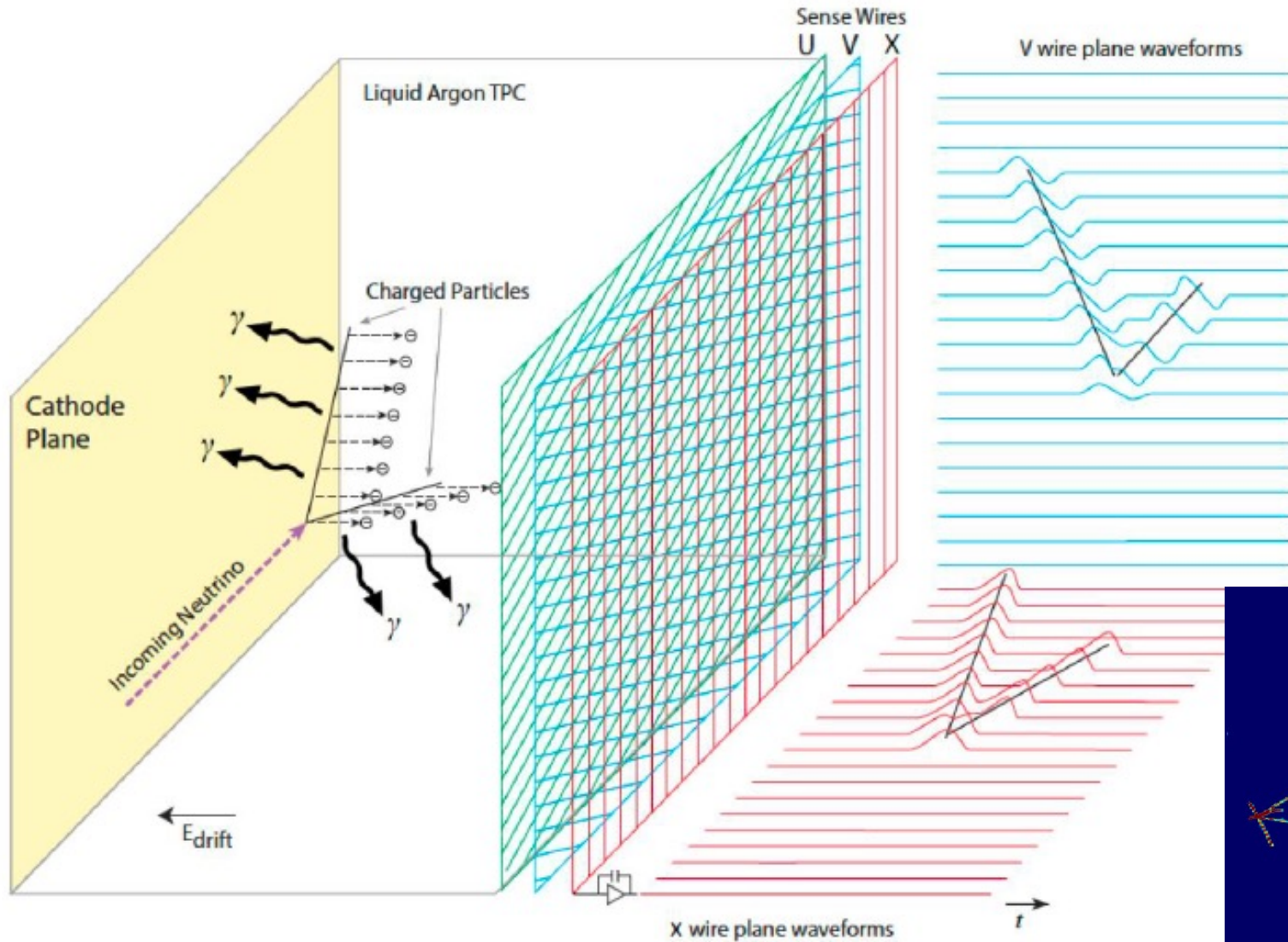


DUNE Far Site: excavation and infrastructure

First cryostat elements arrive from CERN



Liquid Argon Time Projection Chamber

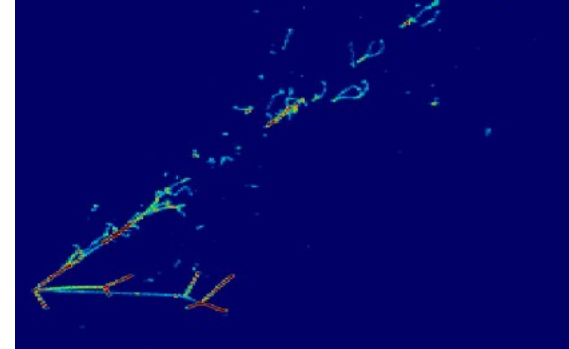


3D-tracking

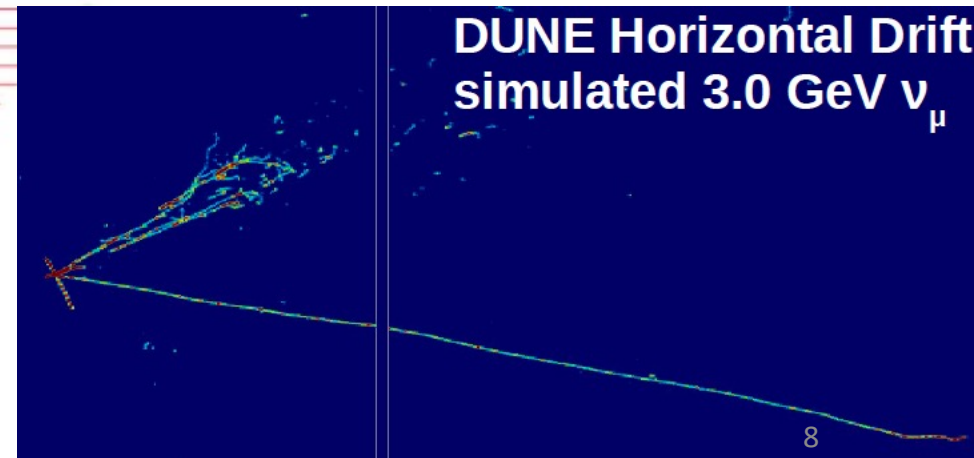
Reconstruction of all charged particles

dE/dx measurement, particle id

DUNE Horizontal Drift
simulated 2.5 GeV ν_e



DUNE Horizontal Drift
simulated 3.0 GeV ν_μ



Challenges

Or: what could possibly go wrong?

- detector technology and scale of the experiment
- neutrino cross section systematics
- beam uncertainties

Near Detector at Fermilab site

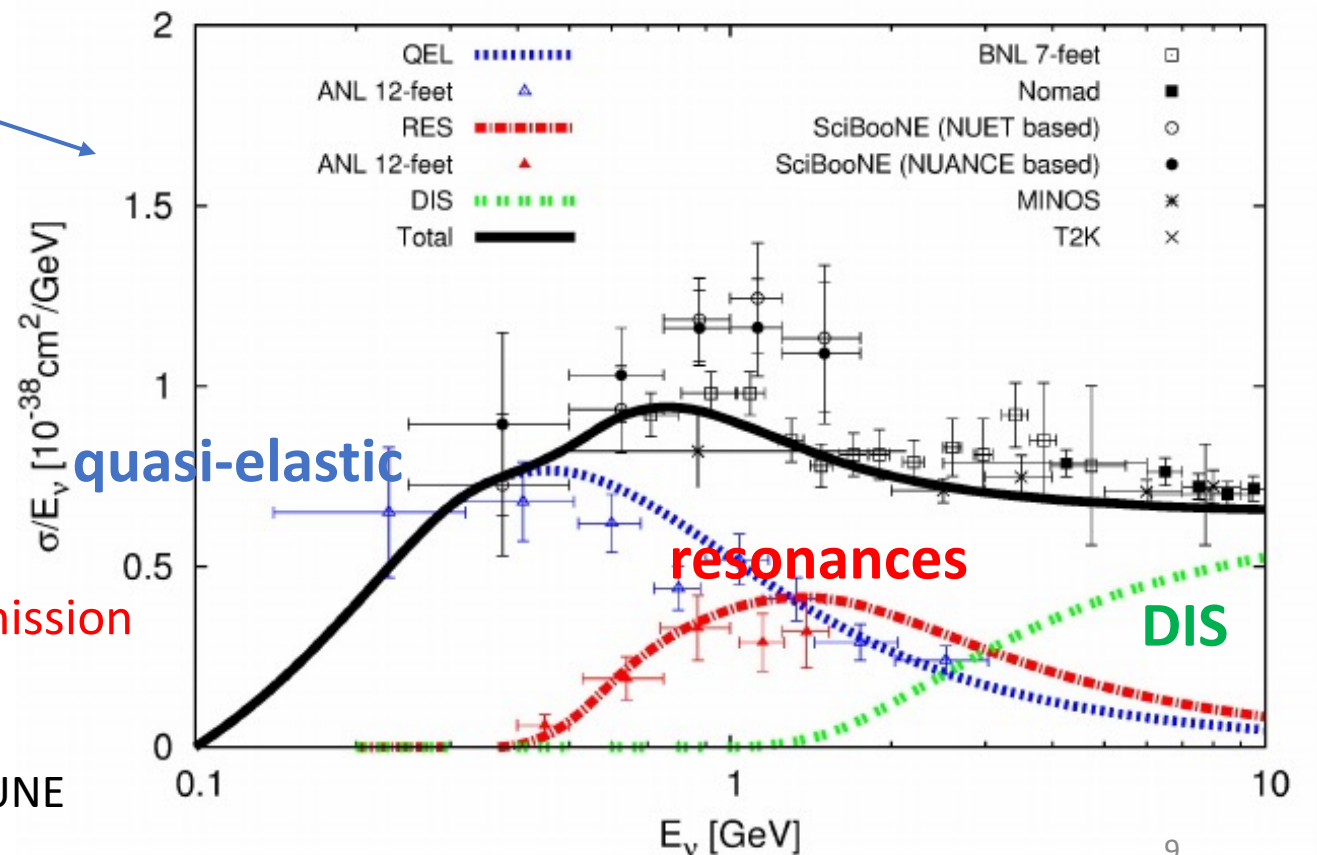
quasi-elastic: Ar nucleus recoils

resonances: nucleus excited, decays via p , π^\pm , π^0 emission

DIS: deep inelastic scattering, nucleus breaks up

input from ArgoNeuT, MicroBoone, Minerva, T2K, ProtoDUNE

Prototype: **ProtoDUNE**



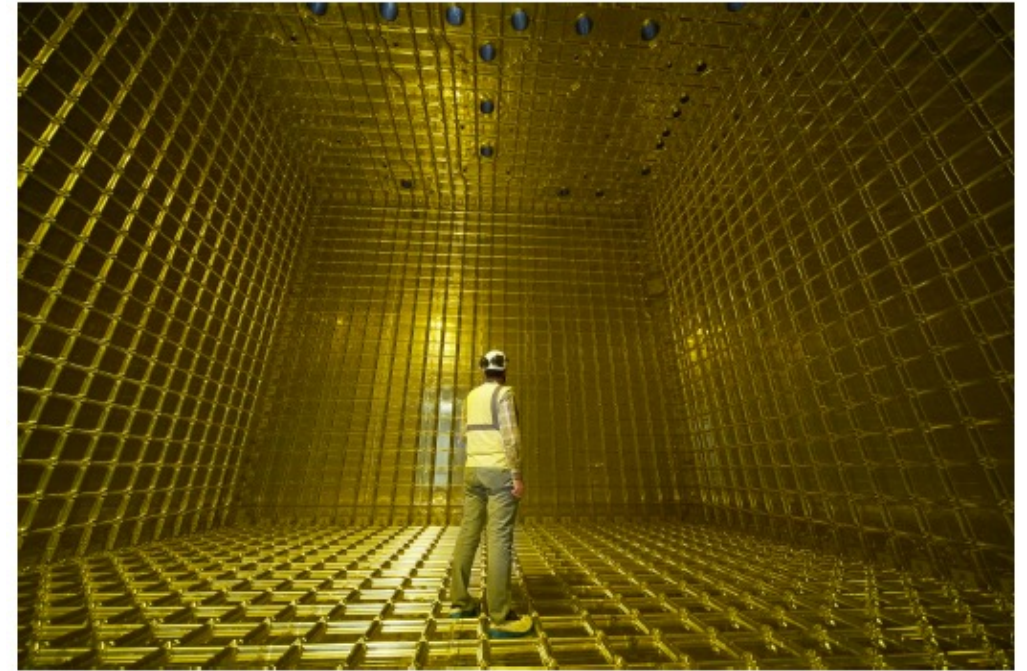
ProtoDUNE

CERN North Area

- Actually 2 ProtoDUNEs, 6x6x6 meter, each 800 ton LAr
- NP4 took beam data in 2018, NP2 took cosmics data
- Different technologies

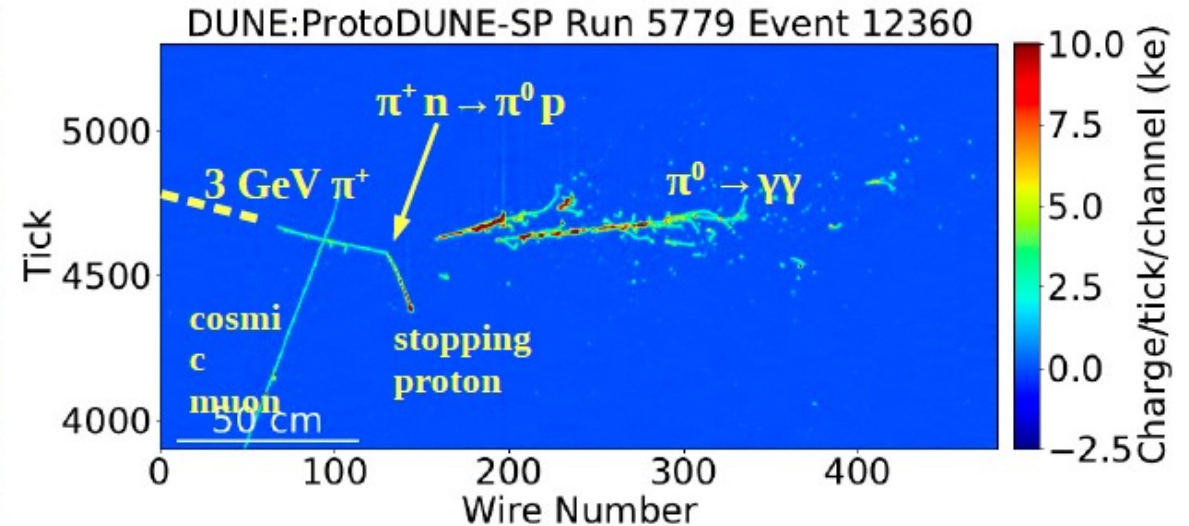
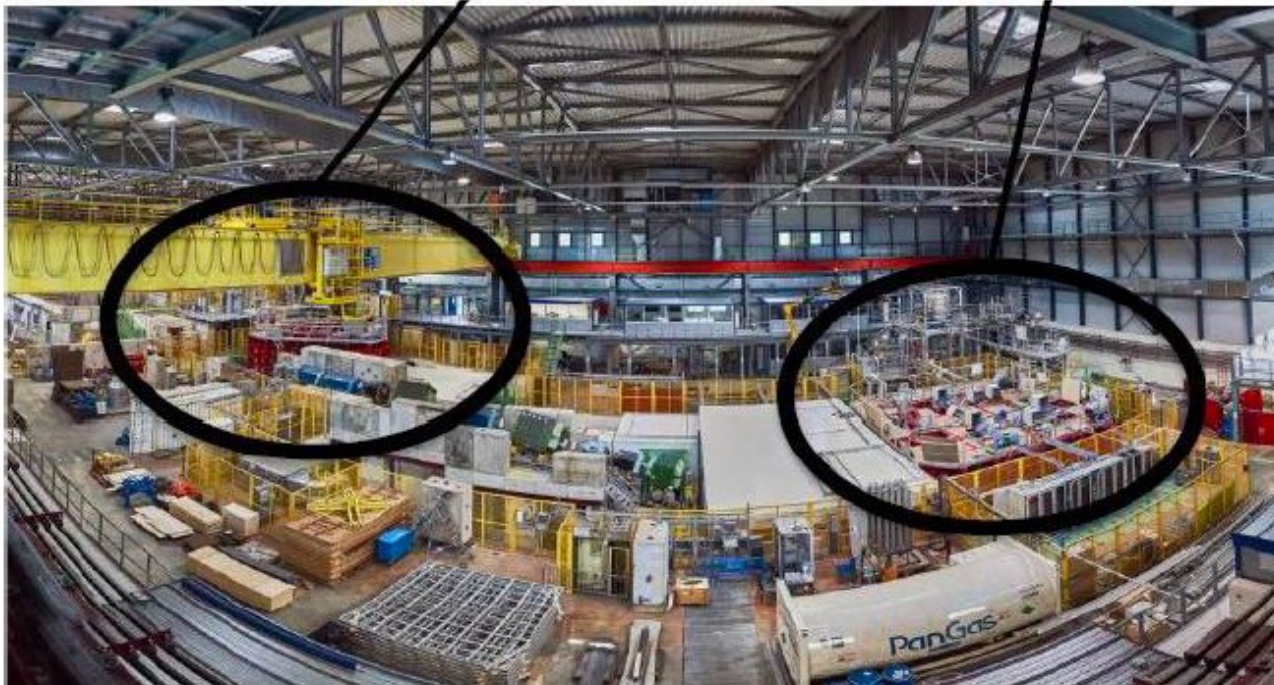
Nikhef contributed to DAQ for NP4

Inside of the ProtoDUNE Neutrino Platform 2 cryostat



Neutrino Platform 2

Neutrino Platform 4



p, e, μ, π beam

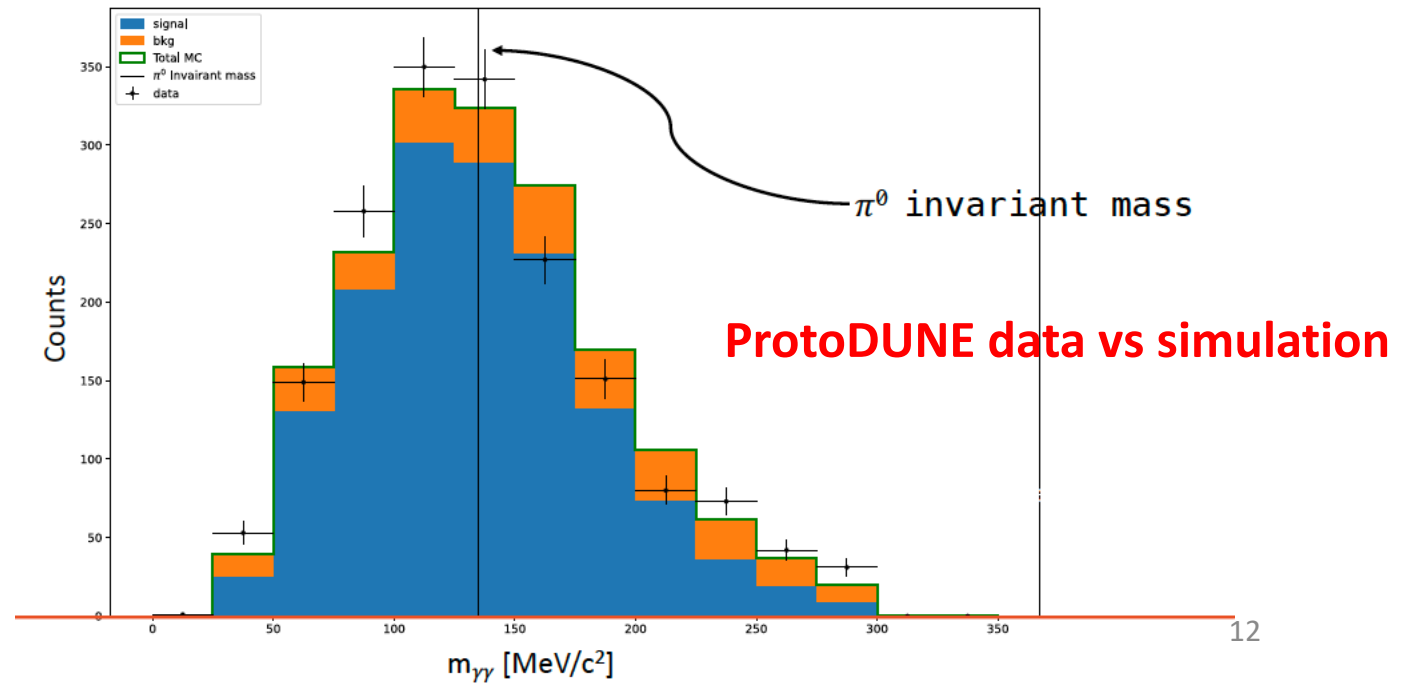
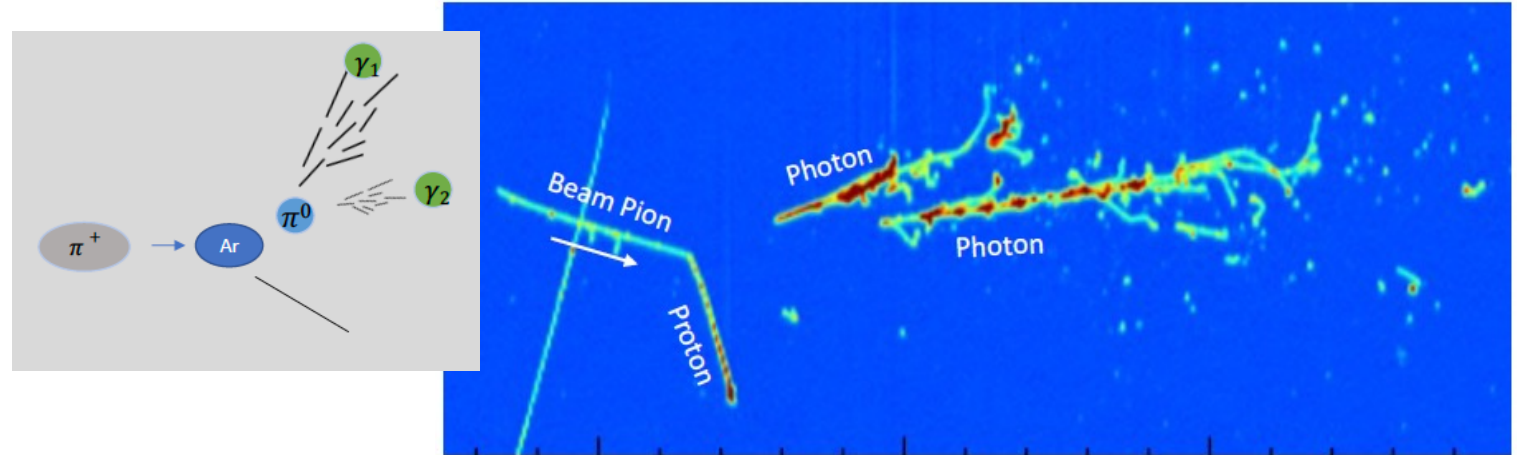
ProtoDUNE detector performance

<i>Detector parameter</i>	<i>ProtoDUNE-SP performance</i>	<i>DUNE specification</i>	
Average drift electric field	500 V/cm	250 V/cm (min) 500 V/cm (nominal)	✓
LAr e-lifetime	> 20 ms	> 3 ms	✓
TPC+CE			
Noise	(C) 550 e, (I) 650 e ENC (raw)	< 1000 e ENC	✓
Signal-to-noise ⟨SNR⟩	(C) 48.7, (I) 21.2 (w/CNR)		
CE dead channels	0.2%	< 1%	✓
PDS light yield	1.9 photons/MeV (@ 3.3 m distance)	> 0.5 photons/MeV (@ cathode distance — 3.6 m)	✓
PDS time resolution	14 ns	< 100 ns	✓

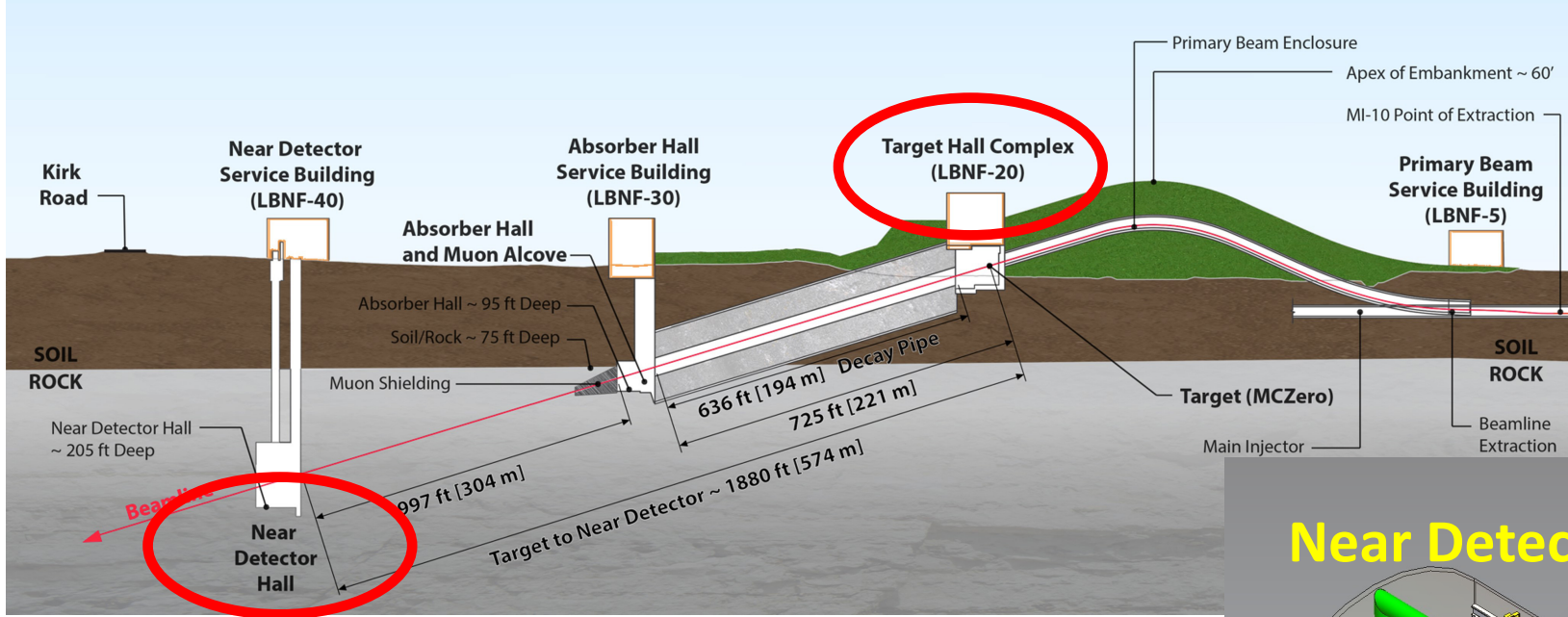
A second run of (slightly) modified detectors takes place in 2024, starting now!

ProtoDUNE analysis

π^0 reconstruction: energy scale, e- π separation, cross-section



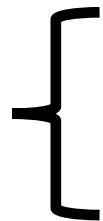
Characterizing the beam: Near Detector



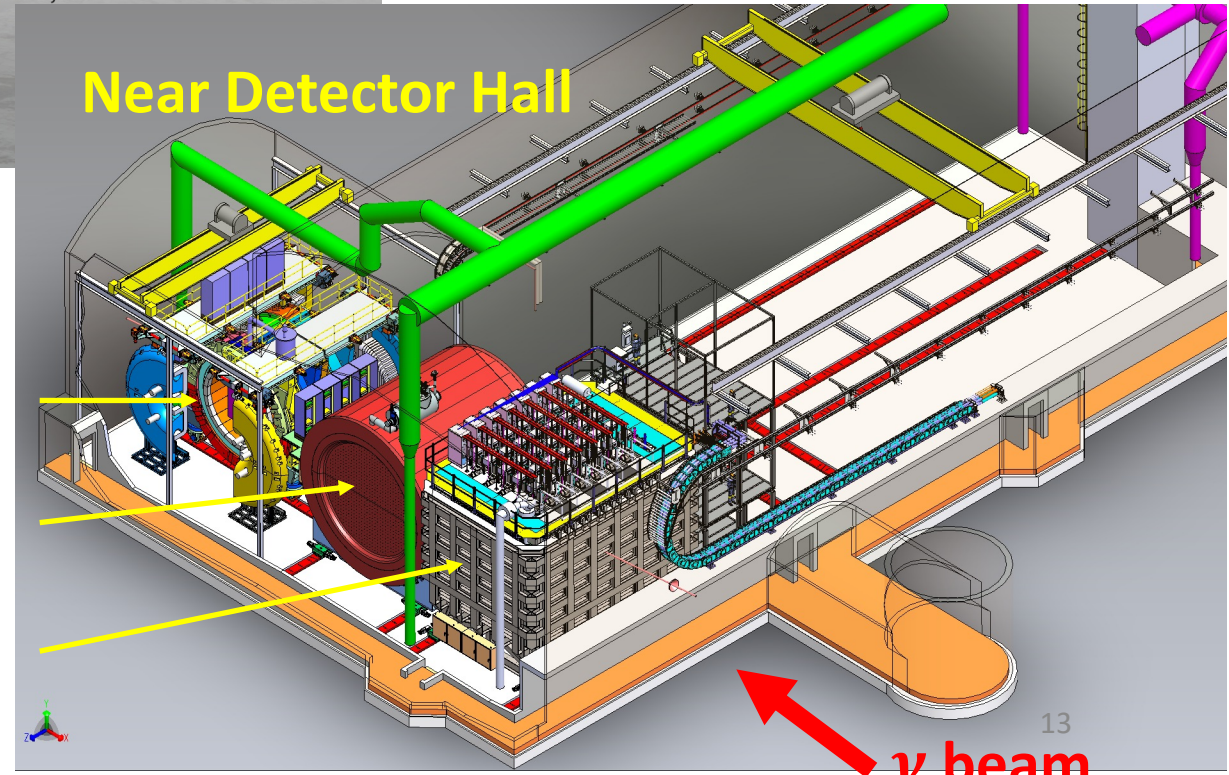
Near Detector, 574 meter from target, 60 m underground

SAND detector (magnet and calorimeter from KLOE)

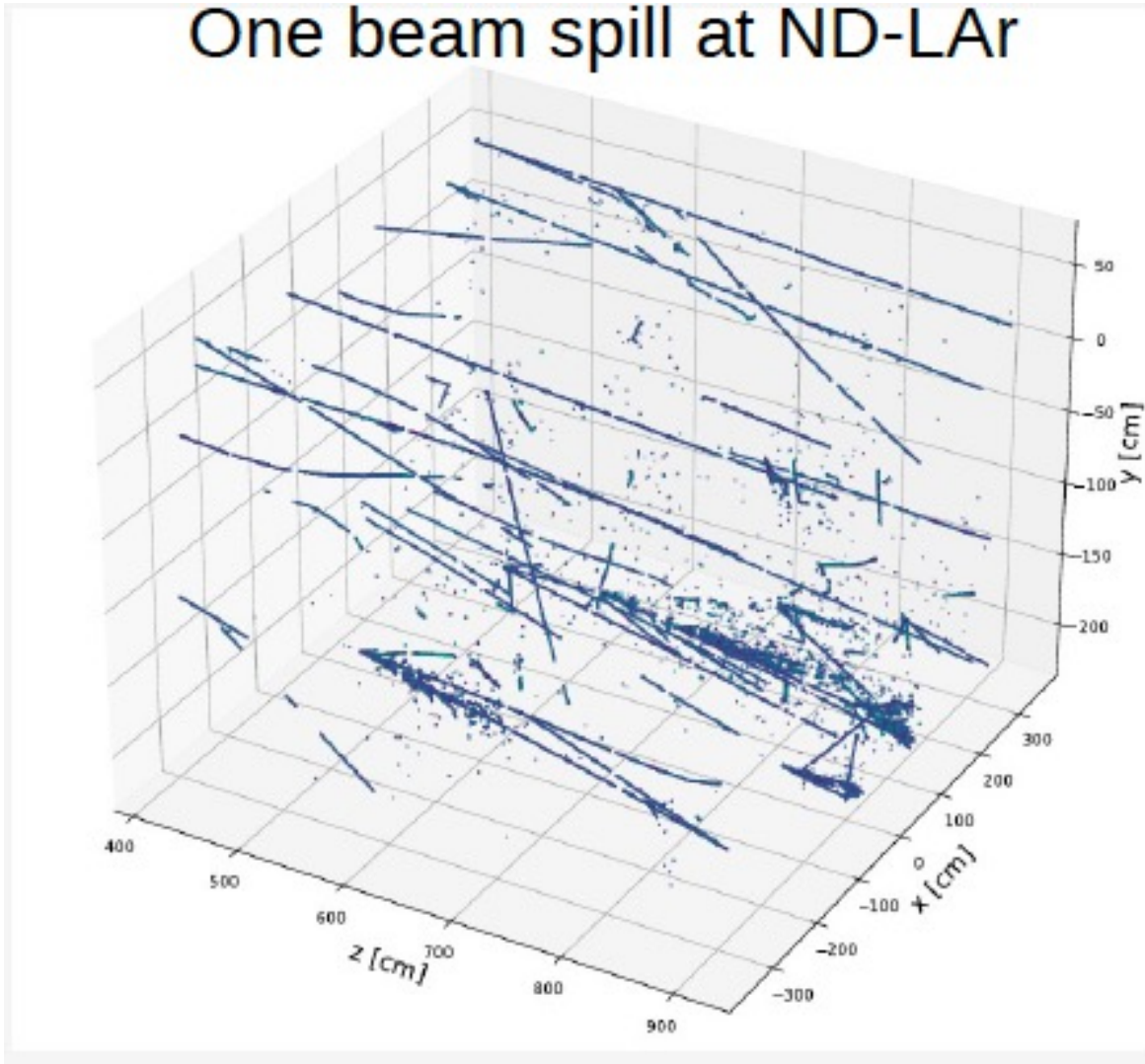
Movable
on-axis/off-axis
→ beam profile,
energy dependence



Muon spectrometer
ND-LAr: modular liquid argon TPC



The challenge of the Near Detector



Beam spill lasts $10 \mu\text{s}$, and repeats at 1 Hz

Maximum drift time of electrons: $300 \mu\text{s}$

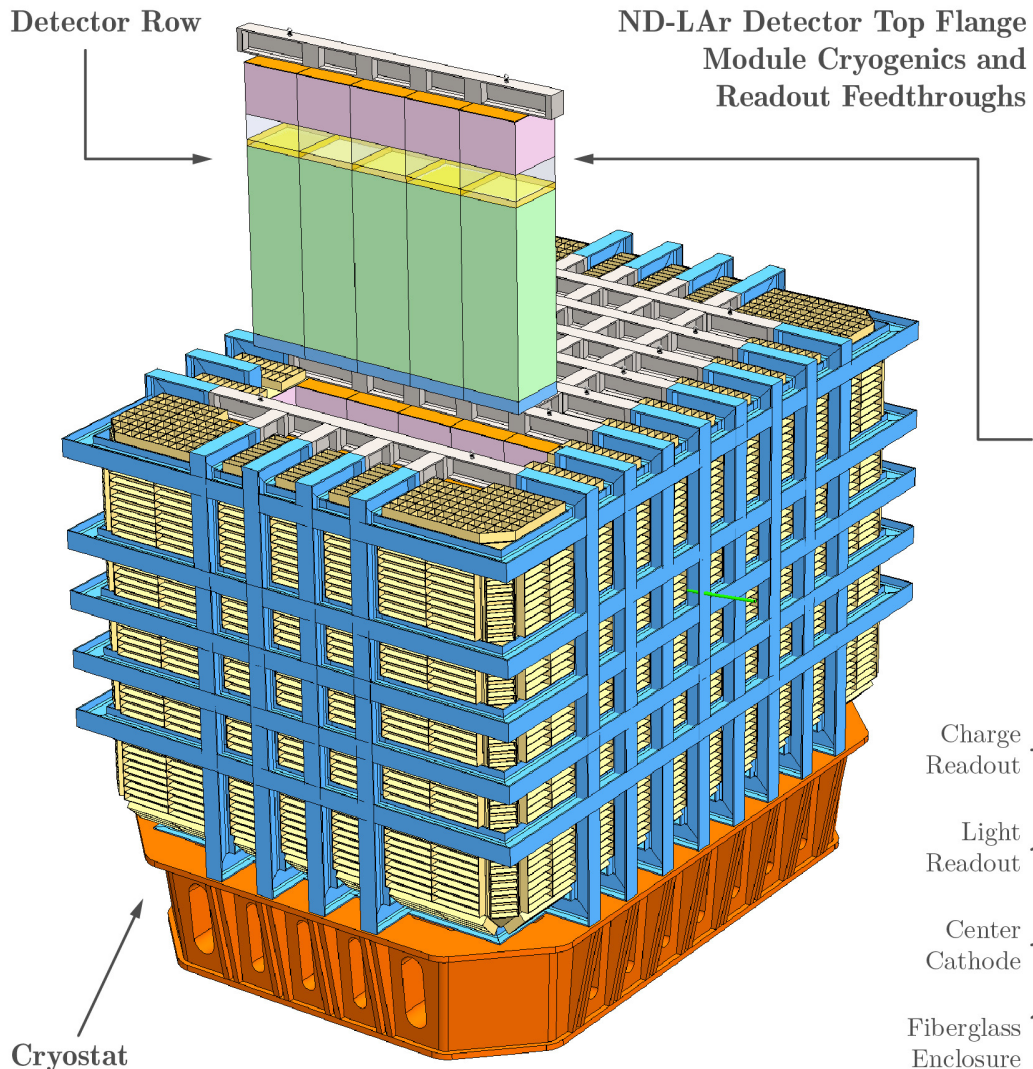
Every spill: up to 50 interactions
(neutrinos in LAr and in rock leading to muons)

Disentangle interactions:

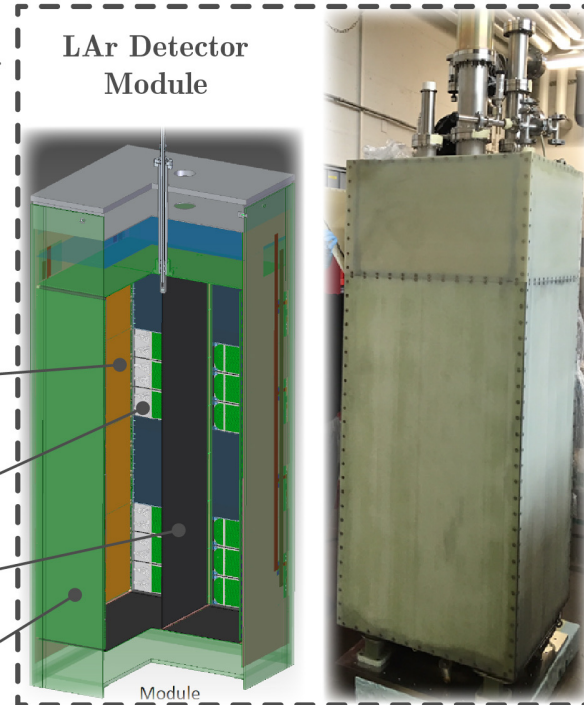
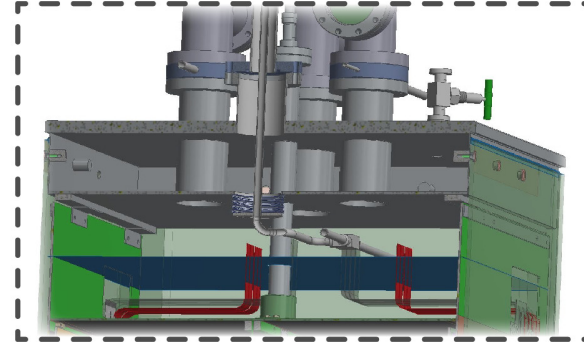
- make small light-tight compartments
- scintillation light determines event time
- match scintillation light and charge

(But also: unprecedented event sample for physics!)

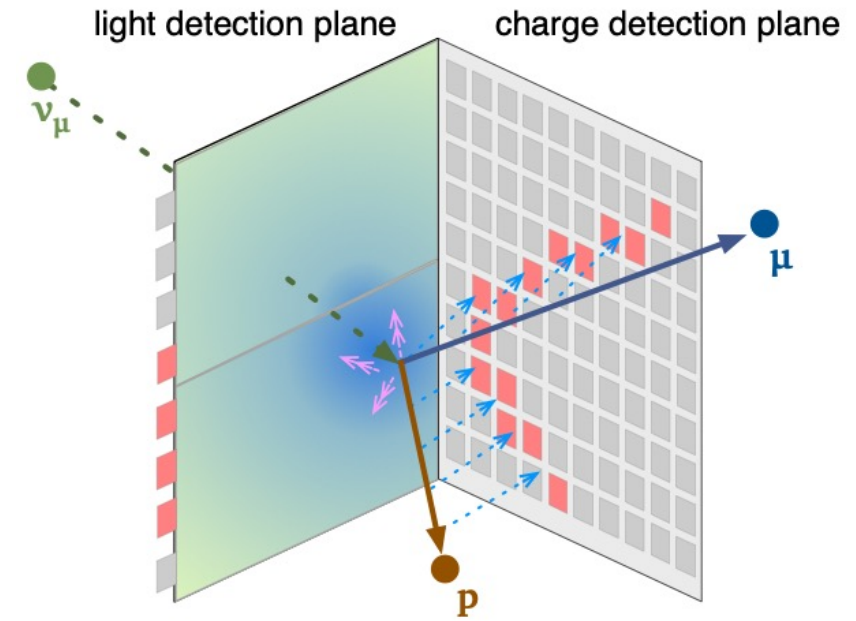
ND-LAr TPC



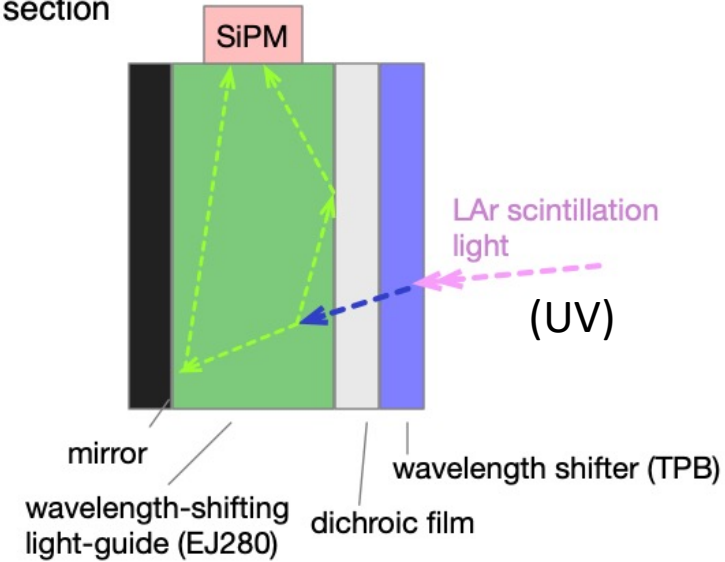
7 rows of 5 modules each



1 x 1 x 3 meter



light detection plane cross section



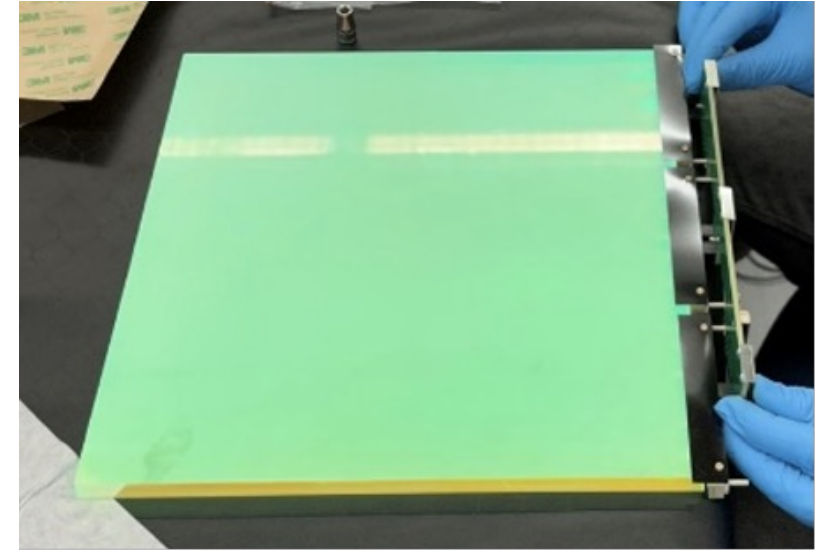
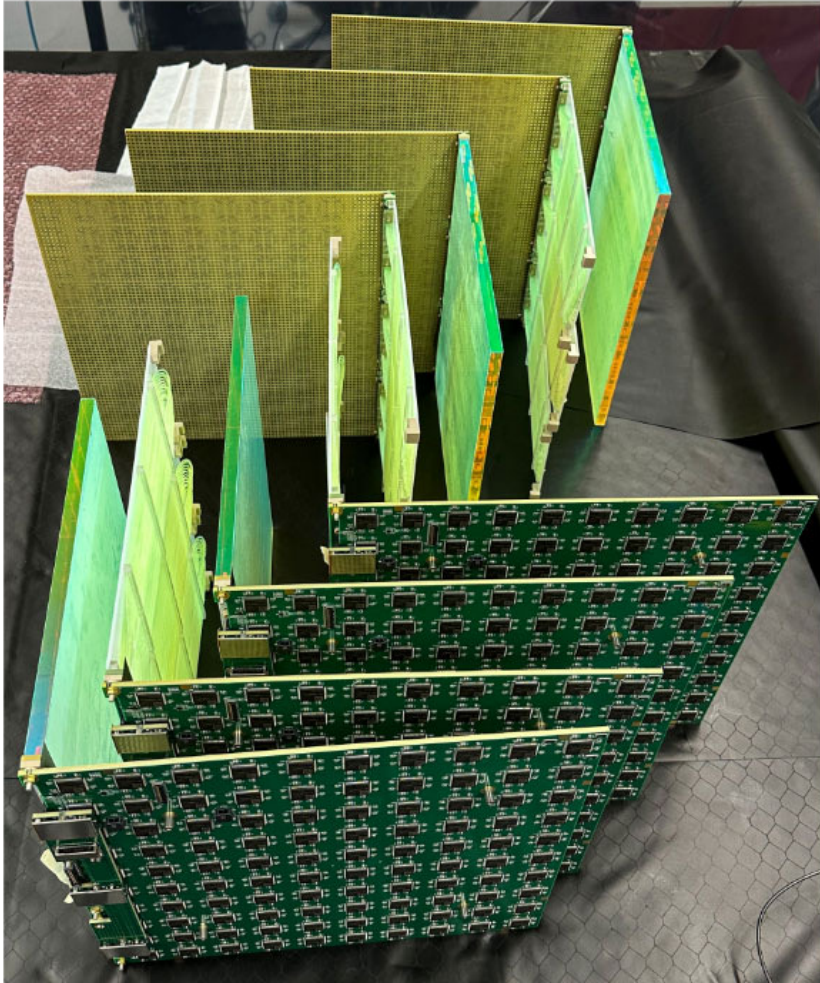
The 2x2 Prototype

Currently in the NuMi neutrino beam at Fermilab



Nikhef plans

Light detection system for ND-LAr modules (with U. Bern)



Light collectors	}	cold
SiPM		
Preamps		
Cabling	}	warm
Amplifiers		
TDCs		

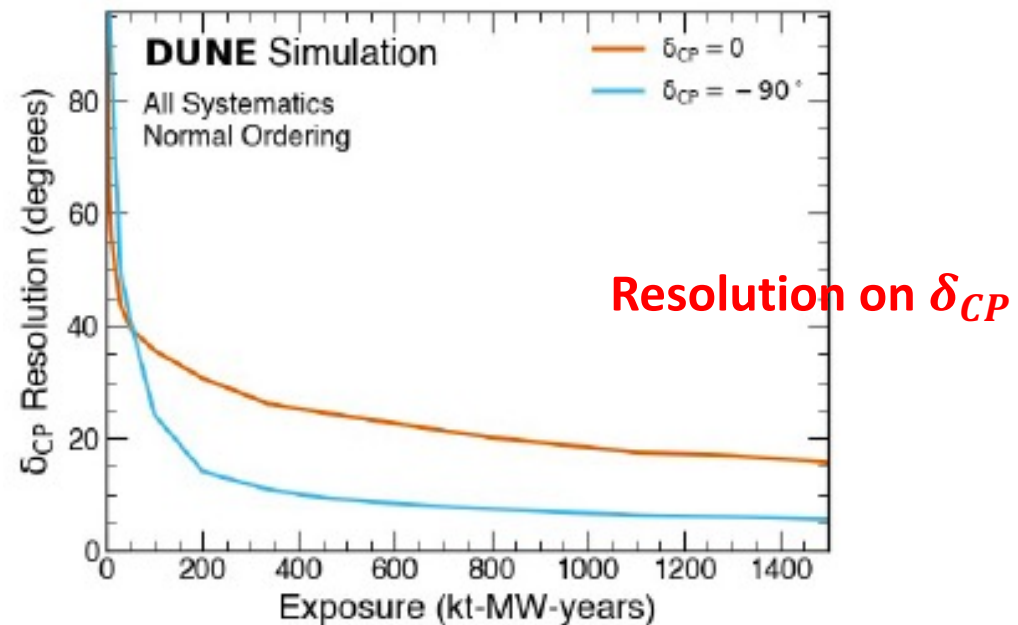
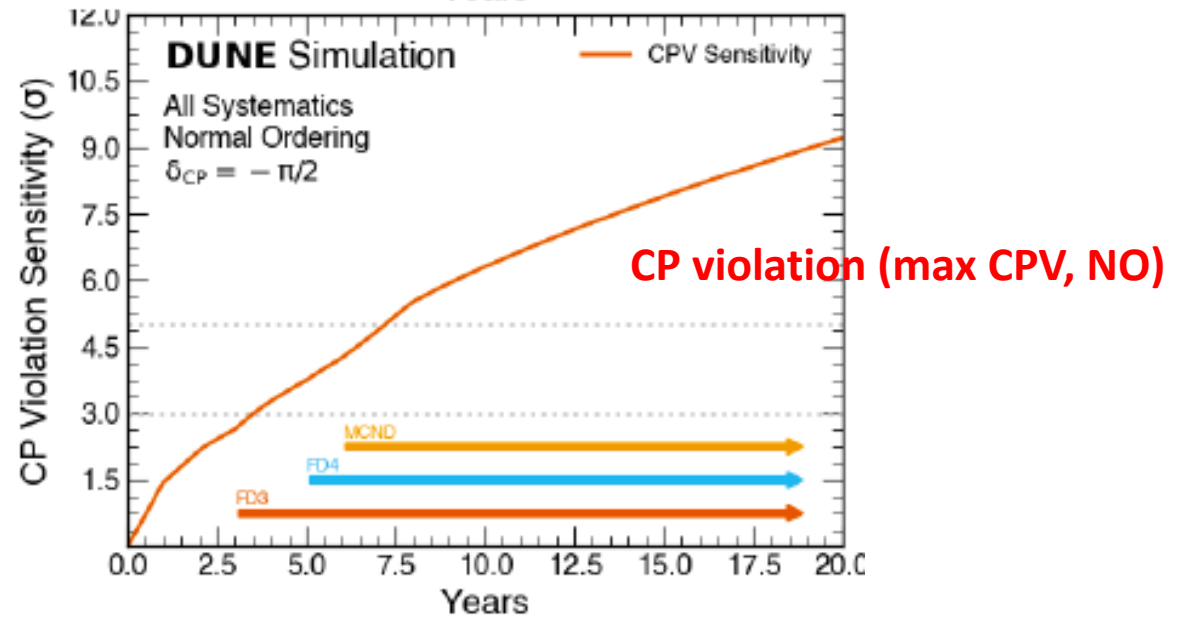
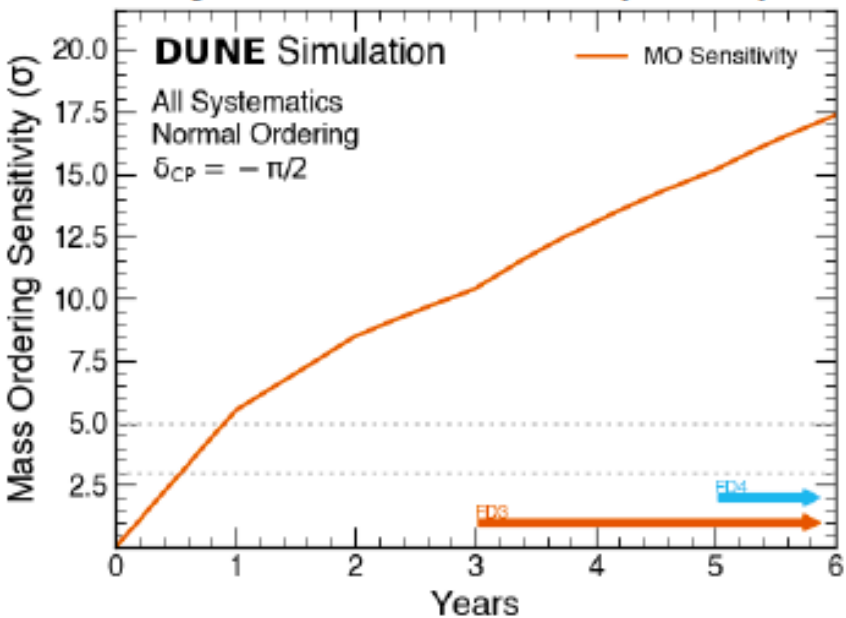
+ software, analysis

35 modules, 8400 electronics channels

Physics expectations

Assuming realistic detector staging and beam performance.

Mass ordering (in case of max CPV, NO)



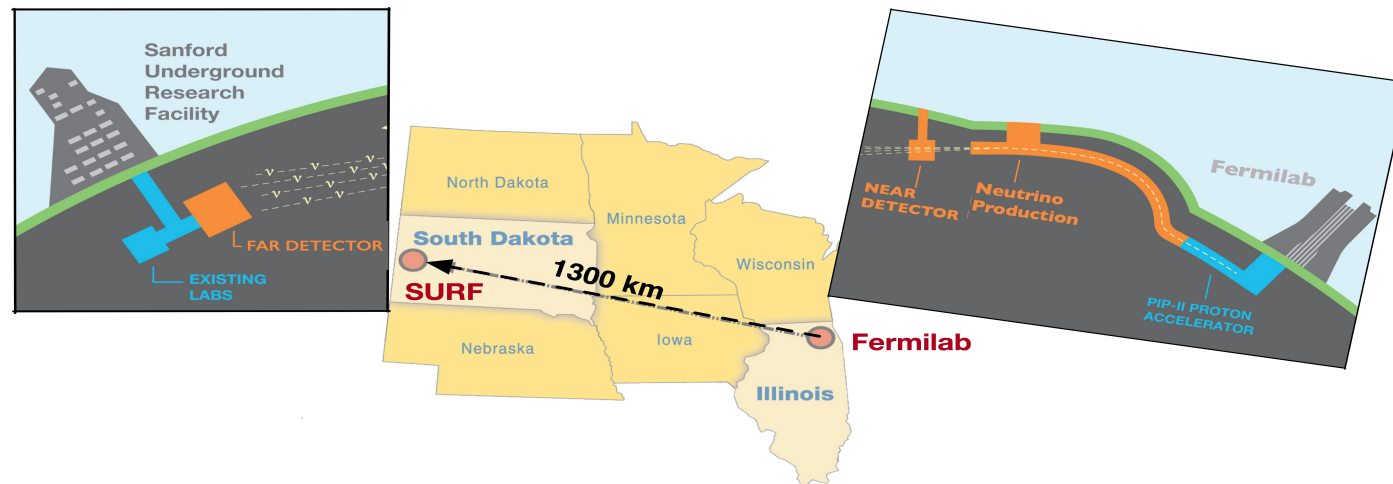
Timeline

Far Detector:

- 2024-2025: Infrastructure
- 2025-2026: Cryostats
- 2026-2027: First 2 modules
- 2028: fill with Ar and purify
- 2029: first data with atm. ν
- 2031: beam

Near Detector LAr:

- 2024-2025: 2x2 tests
- 2024: first full size demonstrator module
- 2026: first row of 5 modules
- 2028-2029: completion 35 modules
- 2029-2030: installation, commissioning
- 2031: beam



The Nikhef DUNE group



Paul de Jong



Tina Pollmann



Patrick Decowski



Frank Filthaut



Auke Pieter Colijn



James Mead



Vikas Gupta Marjolein van Nuland-Troost



Wessel Krah



Jasper Paul



Corryenne Groen



Dagmar Salomons

+Nikhef and RU electronic engineers