

Recent developments in testable leptogenesis

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Theory Seminar Nikhef
November 23, 2023



Nikhef

1. Neutrino masses and type-I seesaw
2. Low-scale leptogenesis
3. Effects of flavour- and CP-symmetries
4. Take-home

Open questions in the Standard Model



Origin of flavours

[Sandbox Studio, Chicago]



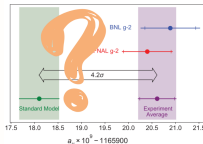
Hierarchy problem

[A. Stonebraker/APS]



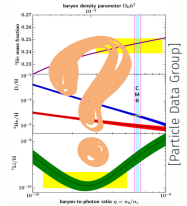
ν oscillations

Spin-1/2 fermions						Spin-1 bosons		Spin-0 Higgs boson
Quarks						Force carriers		
Left	u	Right	Left	c	Right	Left	t	Right
Left	d	Right	Left	s	Right	Left	b	Right
Left	ν_1	Right	Left	ν_2	Right	Left	ν_3	Right
Left	e	Right	Left	μ	Right	Left	τ	Right
						g		Z ⁰
						γ		
						W [±]		



Muon $g - 2$

[Muon $g-2$ collaboration]



Baryon asymmetry

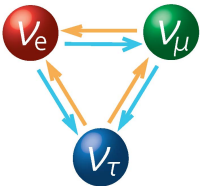
[Particle Data Group]



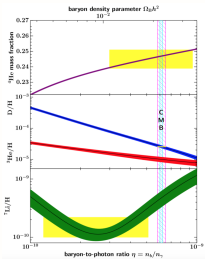
Dark Matter

[Chandra]

Right-handed neutrinos

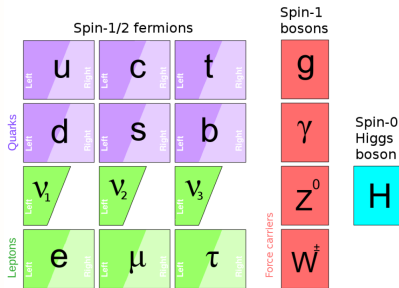


Neutrino oscillations/masses

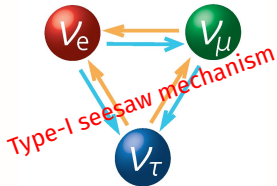


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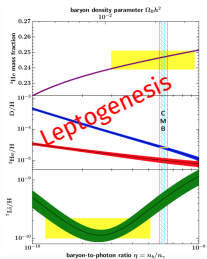
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Neutrino oscillations/masses



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Baryon asymmetry

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Quarks	Left	Right	Left	Right	Left	Right	Force carriers	g
	Left	Right	Left	Right	Left	Right		γ
	Left	Right	Left	Right	Left	Right		Z^0
Leptons	Left	Right	Left	Right	Left	Right	Force carriers	W^\pm
	Left	Right	Left	Right	Left	Right		H
	Left	Right	Left	Right	Left	Right		

Type-I seesaw mechanism

Type-I seesaw Lagrangian

$$\mathcal{L} \supset Y_{\alpha i} (\bar{\ell}_{\alpha} \tilde{\phi}) \nu_{Ri} + \frac{1}{2} \bar{\nu}_{Ri}^c (M_M)_{ij} \nu_{Rj} + \text{h.c.}$$

Yukawa

Majorana

Seesaw relation

$$m_{\nu} = -v^2 (Y \cdot M_M^{-1} \cdot Y^t)$$

$$\nu \simeq U_{\nu}^{\dagger} (\nu_L - \theta \nu_R^c) + \text{h.c.}$$

Light neutrinos

$$N \simeq U_N^{\dagger} (\nu_R + \theta^t \nu_L^c) + \text{h.c.}$$

Heavy neutrinos (HNL)

- $n \geq 2$ HNL generations needed to explain light neutrino masses



Type-I seesaw mechanism

Type-I seesaw Lagrangian (below EWSB)

$$\mathcal{L} \supset \underbrace{v Y_{\alpha i} \bar{\ell}_{\alpha} \nu_{Ri}}_{\text{Dirac}} + \frac{1}{2} \bar{\nu}_{Ri}^c \underbrace{(M_M)_{ij}}_{\text{Majorana}} \nu_{Rj} + \text{h.c.}$$

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Heavy neutrinos (HNL)

- $n \geq 2$ HNL generations needed to explain light neutrino masses
 - What is our prior on n ?
 - $n = 2$: Minimality (ν MSM)
 - $n = 3$: Flavour symmetries, gauge extensions,... (LRSM,...)

Type-I seesaw mechanism

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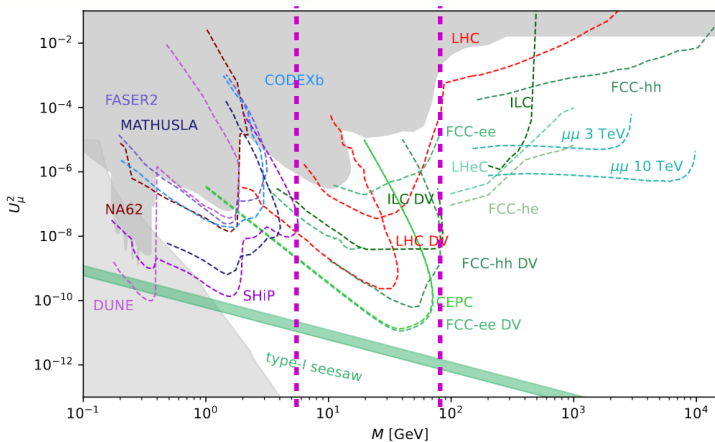
Heavy neutrinos (HNL)

- $n \geq 2$ HNL generations needed to explain light neutrino masses
- Experimental sensitivity expressed in terms of

$$U_{\alpha}^2 = \sum_i |\theta_{\alpha i}|^2 = \sum_i |v(Y \cdot M_M^{-1})_{\alpha i}|^2$$

Testing the type-I seesaw

Many different ways to probe HNLs:



[Bose et al; 2209.13128]

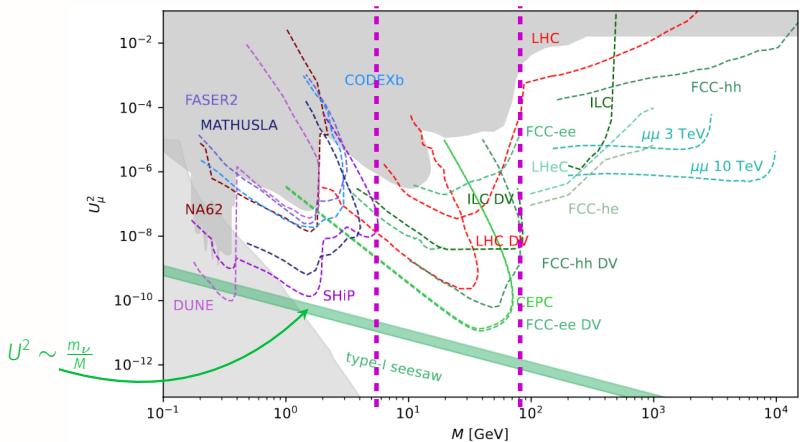
Meson decays

W/Z decays

Virtual W/Z exchange

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Naive seesaw bound

$$m_\nu = -v^2(Y \cdot M_M^{-1} \cdot Y^t) \Leftrightarrow U_i^2 \sim \frac{m_\nu}{M_i} \sim 10^{-10} \frac{\text{GeV}}{M_i}$$

How to reach large coupling ? B-L approximate symmetry

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B-L approximate symmetry

Majorana mass

$$\bar{M} \cdot \begin{pmatrix} 1 - \mu & 0 & 0 \\ 0 & 1 + \mu & 0 \\ 0 & 0 & \mu' \end{pmatrix}$$

Yukawa coupling

$$\begin{pmatrix} f_e(1 + \epsilon_e) & if_e(1 - \epsilon_e) & f_e \epsilon'_e \\ f_\mu(1 + \epsilon_\mu) & if_\mu(1 - \epsilon_\mu) & f_\mu \epsilon'_\mu \\ f_\tau(1 + \epsilon_\tau) & if_\tau(1 - \epsilon_\tau) & f_\tau \epsilon'_\tau \end{pmatrix}$$

Technically natural: Small m_ν from small symmetry breaking parameters $\mu, \epsilon, \epsilon' \ll 1$
Consistent with large U^2 .

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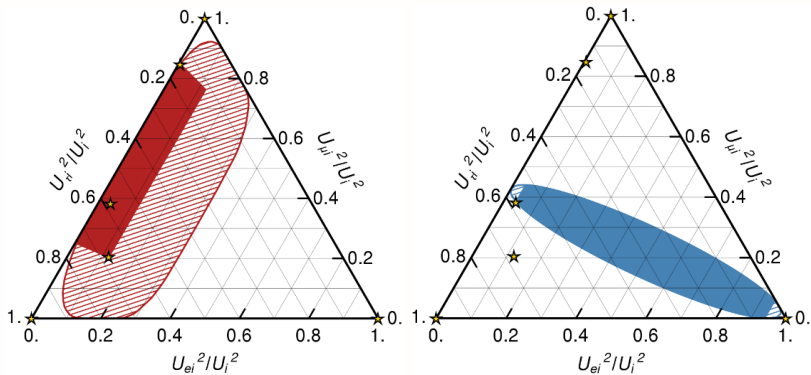
Pseudo-Dirac pair

Decoupled

$$\begin{pmatrix} f_e(1 + \epsilon_e) & if_e(1 - \epsilon_e) & f_e \epsilon'_e \\ f_\mu(1 + \epsilon_\mu) & if_\mu(1 - \epsilon_\mu) & f_\mu \epsilon'_\mu \\ f_\tau(1 + \epsilon_\tau) & if_\tau(1 - \epsilon_\tau) & f_\tau \epsilon'_\tau \end{pmatrix}$$

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Flavour triangle



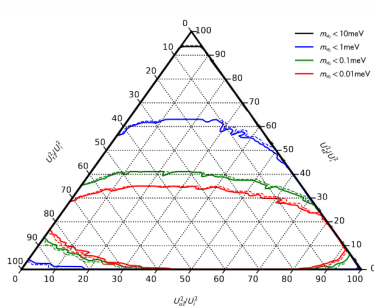
(a) Normal ordering.

(b) Inverted ordering.

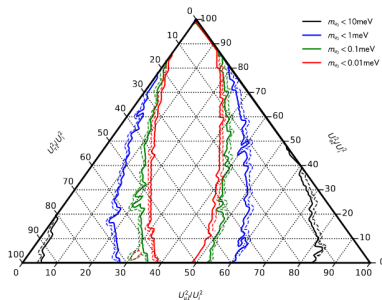
[Drewes/Hajer/Klarić/Lafranchi; 1801.04207]

- Branching ratios constrained for $n = 2$: Can test HNLs as origin of ν masses.
- For $m_0 \neq 0$ (only possible for $n = 3$), almost all flavour ratios are allowed.

Flavour triangle



Normal ordering



Inverted ordering

[Chraszcz et al; 1908.02302]

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Low-scale leptogenesis

Leptogenesis

Sakharov conditions:

- ★ C- and CP-violation
- ★ Deviation from thermal equilibrium
- ★ Baryon number violation

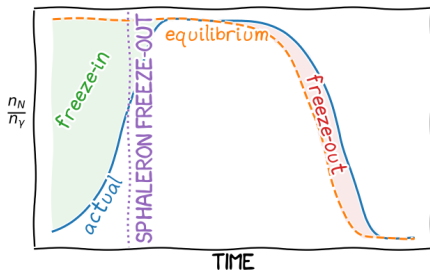
Sakharov conditions:

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Leptogenesis

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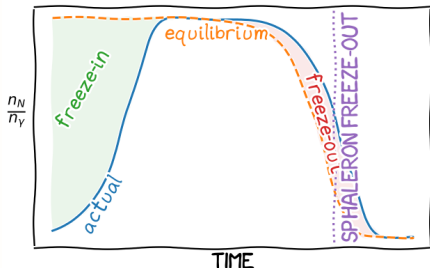


[Klarić/Shaposhnikov/Timiryasov, 2103.16545]

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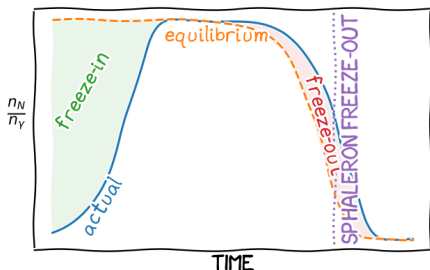


[Klarič/Shaposhnikov/Timiryasov, 2103.16545]

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[Klarič/Shaposhnikov/Timiryasov, 2103.16545]

- ★ Baryon number violation
 - Sphaleron process

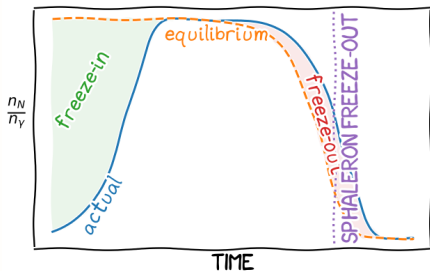
Efficient for $130 \text{ GeV} \lesssim T \lesssim 10^{12} \text{ GeV}$



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[Klarič/Shaposhnikov/Timiryasov, 2103.16545]



Vanilla thermal leptogenesis

Assumptions:

- * Asymmetry generated by heavy neutrino decays
- * Hierarchical mass spectrum $M_1 \ll M_i$
- * Unflavoured

Vanilla thermal leptogenesis

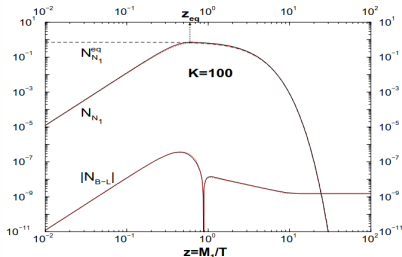
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Boltzmann equations

$$\frac{d}{dz} n_1 = -\frac{\Gamma_D}{HZ} (n_1 - n_1^{eq})$$

$$\frac{d}{dz} n_{B-L} = \epsilon_1 \frac{\Gamma_D}{HZ} (n_1 - n_1^{eq}) - \frac{\Gamma_W}{HZ} n_{B-L}$$



[Buchmüller/Di Bari/Plümacher, '04]

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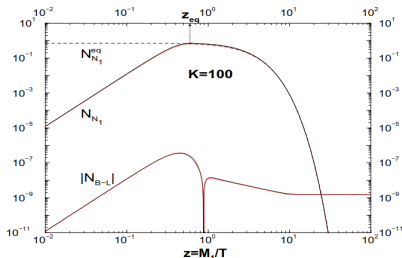
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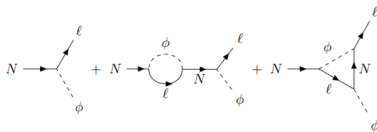
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- Decay asymmetry $\epsilon_1 \equiv \frac{\Gamma_{N_1 \rightarrow \ell + \phi} - \Gamma_{N_1 \rightarrow \bar{\ell} + \phi^*}}{\Gamma_{N_1 \rightarrow \ell + \phi} + \Gamma_{N_1 \rightarrow \bar{\ell} + \phi^*}}$



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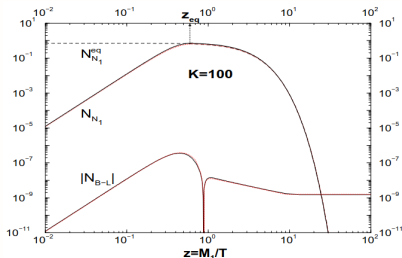
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- For large mass splittings $|\epsilon_1| \lesssim \frac{3}{8\pi} \frac{M_1}{v^2} \sqrt{\Delta m_{23}^2}$ leading to the Davidson-Ibarra bound

$$M_1 \gtrsim 4 \cdot 10^8 \text{ GeV}$$

↪ Direct detection ☹️

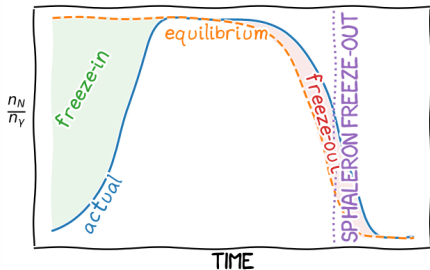


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Leptogenesis

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[Klarič/Shaposhnikov/Timiryasov, 2103.16545]



[Akhmedov/Rubakov/Smirnov '98, Pilaftsis/Underwood '03, Asaka/Shaposhnikov '05, ...]

[Fukugita/Yanagida '86]

- Traditionally, 2 main mechanisms:

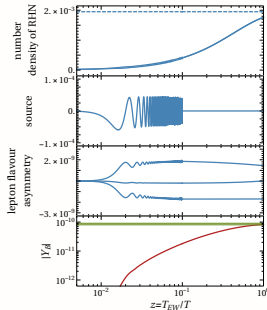
ARS Leptogenesis

Asymmetry produced during
freeze-in from CP-violating
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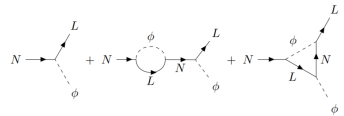


Resonant leptogenesis

Resonant enhancement of
CP-violation from small mass
splittings



[Drewes/Garbrecht/Gueter/Klarić; 1606.06690]



Decay asymmetry:

$$\epsilon_i \simeq \frac{\text{Im}(\gamma^\dagger \gamma)_{ij}^2 (M_{N_i}^2 - M_{N_j}^2) \cdot M_{N_i} \Gamma_N}{(\gamma^\dagger \gamma)_{ii} (\gamma^\dagger \gamma)_{jj} (M_{N_i}^2 - M_{N_j}^2)^2 + M_{N_i}^2 \Gamma_N^2}$$

Low-scale models

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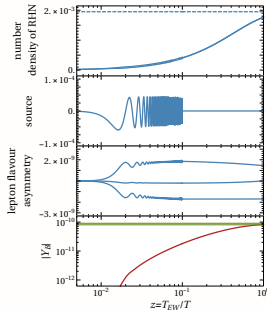
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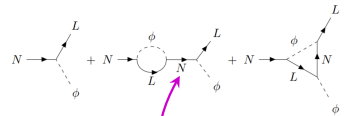


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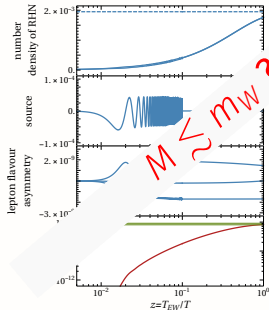
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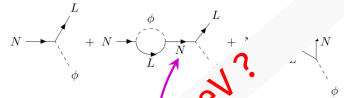


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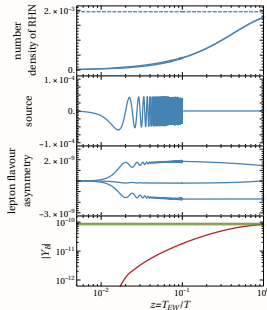
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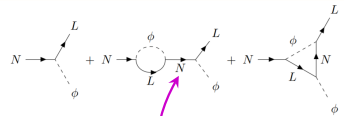
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[Drewes/Garbrecht/Gueter/Klarić; 1606.06690]

→ Two regimes of the same mechanism ! Represented by the same set of kinetic equations (cfr. [Garbrecht; 1812.02651] for a review)



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Quantum kinetic equations

$$i \frac{d}{dt} \rho = [H, \delta \rho] - \frac{i}{2} \{ \Gamma, \delta \rho \} - i \sum_{a \in \{e, \mu, \tau\}} \tilde{\Gamma}_a \frac{\mu_a}{T} f_F (1 - f_F),$$

$$i \frac{d}{dt} \bar{\rho} = -[H, \delta \bar{\rho}] - \frac{i}{2} \{ \Gamma, \delta \bar{\rho} \} + i \sum_{a \in \{e, \mu, \tau\}} \tilde{\Gamma}_a \frac{\mu_a}{T} f_F (1 - f_F),$$

$$\frac{d}{dt} n_{\Delta_a} = - \frac{2i \mu_a}{T} \int \frac{d^3 \vec{k}}{(2\pi)^3} \text{Tr}[\Gamma_a] f_F (1 - f_F) + i \int \frac{d^3 \vec{k}}{(2\pi)^3} \text{Tr}[\tilde{\Gamma}_a (\delta \bar{\rho} - \delta \rho)].$$

Density matrix

Effective Hamiltonian

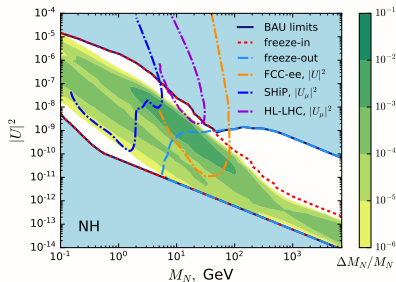
Lepton asymmetry

Interaction rates

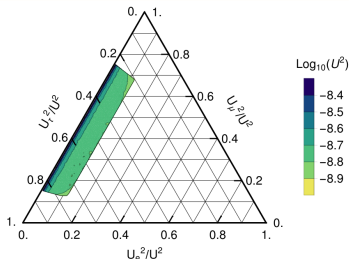
- **Interaction rates** can be
 - ★ Fermion number **conserving** $\sim (Y^\dagger Y) T$
 - ★ Fermion number **violating** $\sim (Y^\dagger Y^*) \frac{M^2}{T}$
- Refined calculation subject to intensive studies over the last years, e.g. Anisimov/Bedak/Bödeker '10, Garny/Kartavtsev/Hohenegger '11, Drewes/Garbrecht/Gueter/Klarič '16, Hernandez/Kekic/Lopez-Pavon/Racker/Salvado '16, Laine/Ghiglieri '16 '18, Klarič/Shaposhnikov/Timiryasov '21, ...

$n = 2$ (ν MSM) parameter space

- Parameter space for **freeze-in** and **freeze-out** are **connected**
- Sizeable fraction of the parameter space can be tested at colliders or fixed target experiments
- Relies on **flavour hierarchies** to reach large U^2
- IH parameter space larger than for NH for $M \lesssim \mathcal{O}(100)$ GeV due to stronger washout



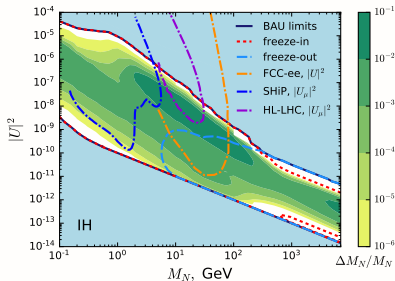
[Klarič/Shaposhnikov/Timiryasov; 2103.16545]



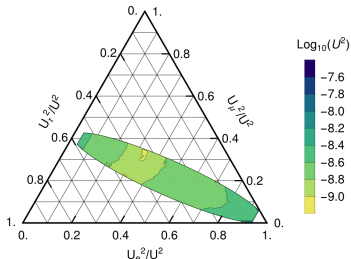
[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/Klarič; 1710.03744]

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- Relies on **flavour hierarchies** to reach large U^2
- IH parameter space **larger** than for NH for $M \lesssim \mathcal{O}(100)$ GeV due to **stronger washout**

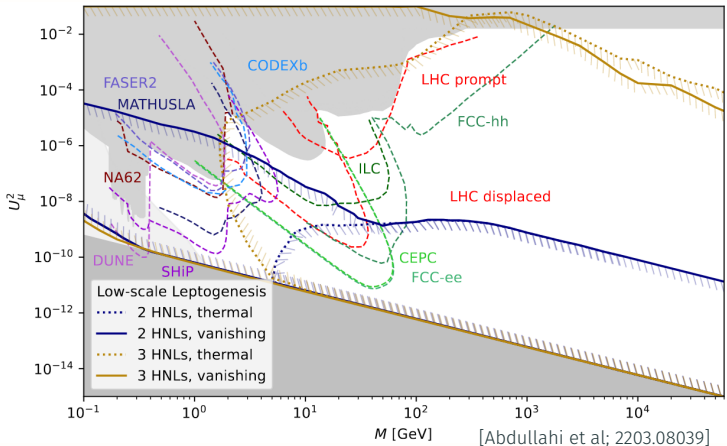


[Klarič/Shaposhnikov/Timiryasov; 2103.16545]



[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/Klarič; 1710.03744]

$n = 3$ parameter space, NH

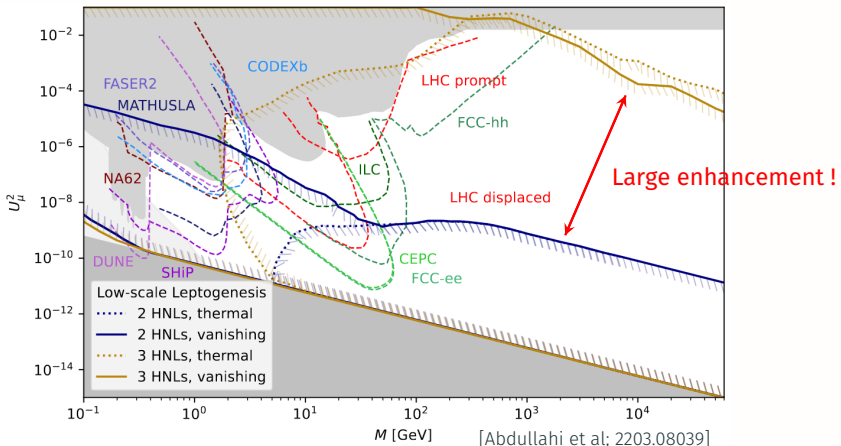


$n = 2$ lines from [Klarič/Shaposhnikov/Timiryasov, 2103.16545]

$n = 3$ lines from [Drewes/YG/Klarič; 2106.16226]

- Can potentially produce enough HNLs to **test leptogenesis** !

$n = 3$ parameter space, NH



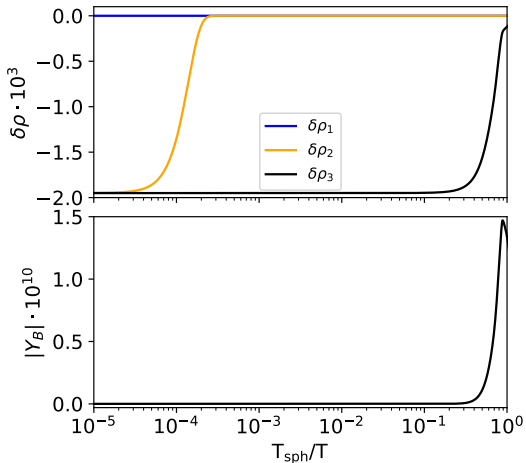
$n = 2$ lines from [Klarič/Shaposhnikov/Timiryasov, 2103.16545]

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- Can potentially produce enough HNLs to **test leptogenesis** !

Why such large mixings ?

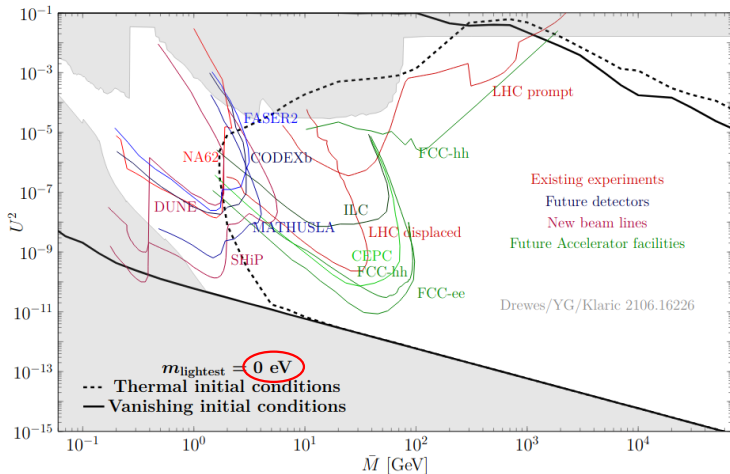
$$U^2 = 0.0248, \bar{M} = 100 \text{ GeV and } m_{\text{lightest}} = 0 \text{ eV}$$



[YG; 2305.06663]

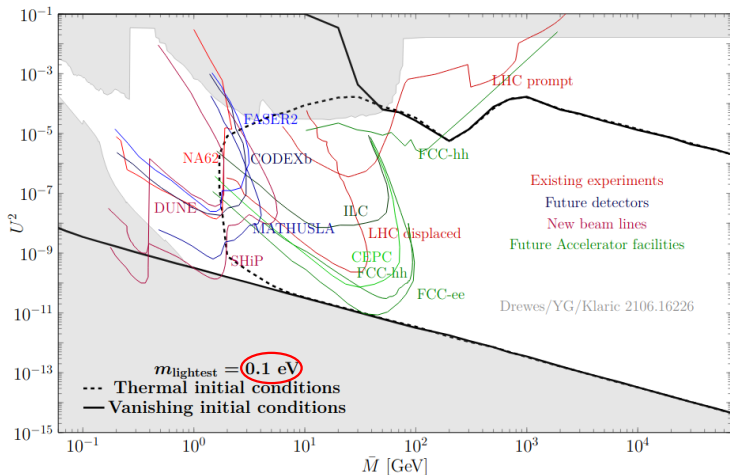
- Large mixing angles allow **late equilibration** of one HNL $U_i^2 \ll 1$
↳ Late BAU production, less time for washout

$n = 3$ parameter space, NH



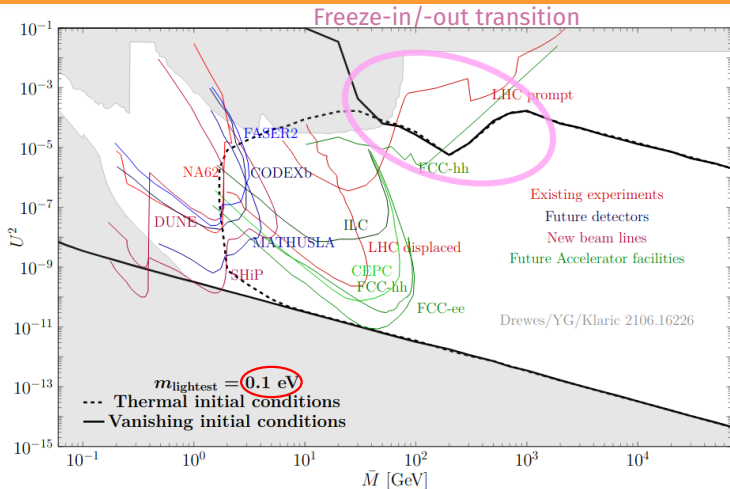
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$n = 3$ parameter space, NH



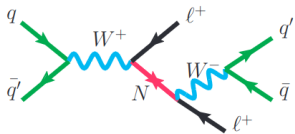
- Can potentially produce enough HNLs to test leptogenesis !

$n = 3$ parameter space, NH



- Can potentially produce enough HNLs to test leptogenesis !

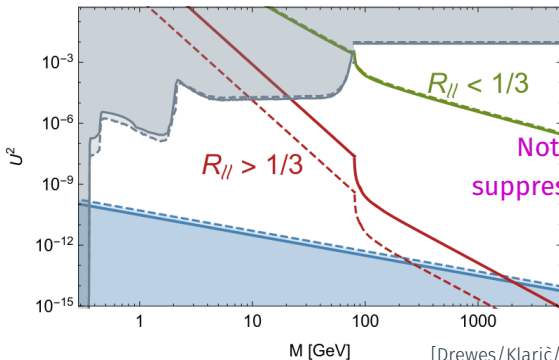
Lepton number violation at colliders



[CMS collaboration; 1806.10905]

- Large U^2 but lepton number conserved if $\mu, \epsilon \rightarrow 0$
- Ratio of lepton number **violating** to **conserving** decays parametrised by

$$R_{ll} = \frac{\Delta M_{\text{phys}}^2}{2\Gamma_N^2 + \Delta M_{\text{phys}}^2}$$



Not always strongly suppressed for sizeable U^2 !

Lepton number violation at colliders

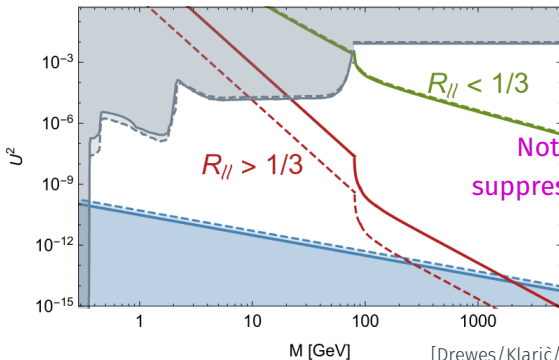
Approximate B-L symmetry

$$M_M = \begin{pmatrix} \bar{M}(1 - \mu) & 0 \\ 0 & \bar{M}(1 + \mu) \end{pmatrix},$$

$$Y = \begin{pmatrix} f_e(1 + \epsilon_e) & if_e(1 - \epsilon_e) \\ f_\mu(1 + \epsilon_\mu) & if_\mu(1 - \epsilon_\mu) \\ f_\tau(1 + \epsilon_\tau) & if_\tau(1 - \epsilon_\tau) \end{pmatrix}$$

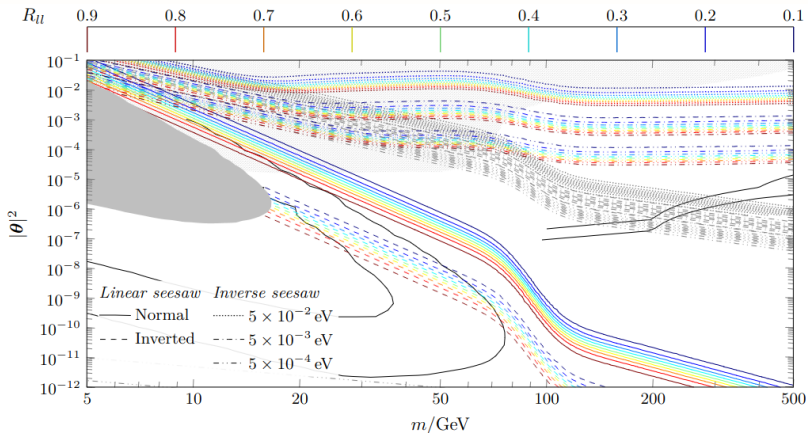
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Lepton number violation at colliders

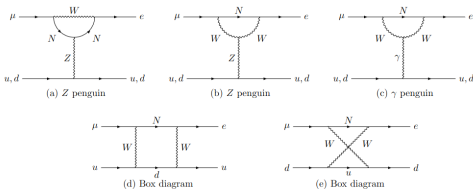


[Antusch/Hajer/Roskopp, 2307.06208]

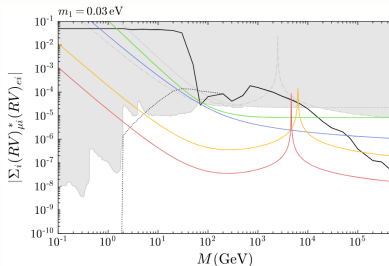
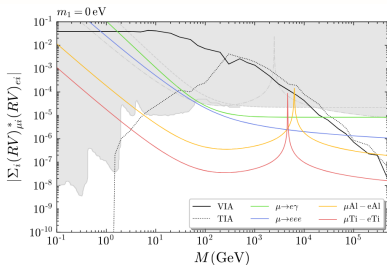
In practice, decoherence effects can make testability prospects even more optimistic !

Testing leptogenesis through CLFV experiments

- HNLs also lead to charge lepton flavour violation.



[Urquia-Calderon/Timiryasov/Ruchayskiy; 2206.04540]

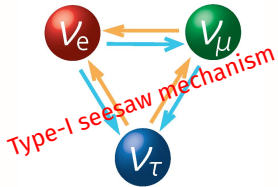


[Granelli/Klaric/Petcov; 2206.04342]

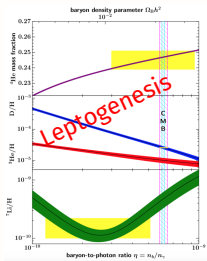
- Right-handed neutrinos provide minimal solution for ν masses + baryon asymmetry
- Parameter space largely enhanced for $n = 3$ due to decoupled 3rd HNL
- Large mixing angle opens up the possibility of testing leptogenesis by combining information from colliders, $0\nu\beta\beta$, ν oscillations, ...
- Collider testability of $n = 3$ scenario to be further explored

Effects of flavour and CP-symmetries

Right-handed neutrinos



Neutrino masses

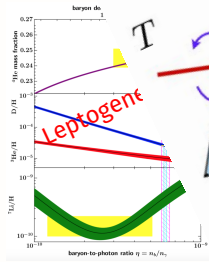
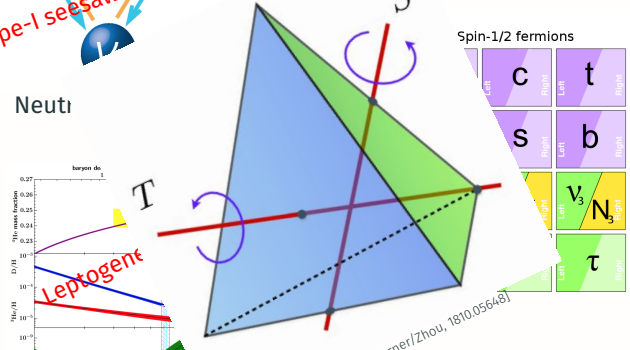
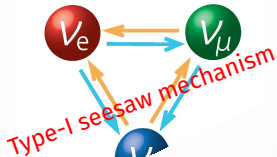


[Particle Data Group]

Baryon asymmetry

Spin-1/2 fermions						Spin-1 bosons						
Quarks	Left	u	Right	Left	c	Right	Left	t	Right	Force carriers	Spin-0 Higgs boson	
	Left	d	Right	Left	s	Right	Left	b	Right			g
	Left	ν_1	N_1	Right	Left	ν_2	N_2	Right	Left			ν_3
Leptons	Left	e	Right	Left	μ	Right	Left	τ	Right	Z^0	H	
	Left			Right			Left		Right	W^\pm		

Right-handed neutrinos



Baryon asymmetry

Spin-1/2 fermions

Left	c	Right	t
Left	s	Right	b
Left	ν_3	Right	N_3
Left	τ	Right	

Spin-1 bosons

g
γ
Z^0
W^\pm

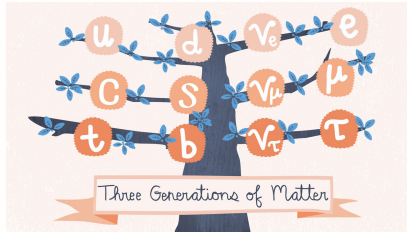
Spin-0 Higgs boson

H

Force carriers

Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?



[Sandbox Studio, Chicago]

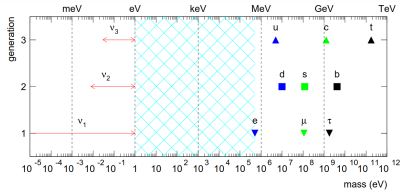
Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?
- Hierarchy in the CKM matrix structure ?

$$|U_{\text{CKM}}| \approx \begin{pmatrix} 0.97 & 0.22 & 0.004 \\ 0.22 & 0.99 & 0.04 \\ 0.008 & 0.04 & 1.01. \end{pmatrix}$$

Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?
- Hierarchy in the CKM matrix structure ?
- Hierarchy in the fermion masses ?



[A. de Gouvea; hep-ph/0411274]

Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?
- Hierarchy in the CKM matrix structure ?
- Hierarchy in the fermion masses ?
- Why such neutrino mixing pattern ? In particular, why the PMNS matrix

$$|U_{\text{PMNS}}| \approx \begin{pmatrix} 0.82 & 0.55 & 0.15 \\ 0.29 & 0.59 & 0.75 \\ 0.49 & 0.59 & 0.64 \end{pmatrix}$$

is so close to a tri-bimaximal mixing

$$U_{\text{TB}} = \begin{pmatrix} \sqrt{2/3} & \sqrt{1/3} & 0 \\ -\sqrt{1/6} & \sqrt{1/3} & \sqrt{1/2} \\ -\sqrt{1/6} & \sqrt{1/3} & -\sqrt{1/2} \end{pmatrix}.$$

Motivations for flavour symmetries

- Why 3 generations in the Standard Model ?
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- Hierarchy in the fermion masses ?
- **Why such neutrino mixing pattern ?** In particular, why the PMNS matrix

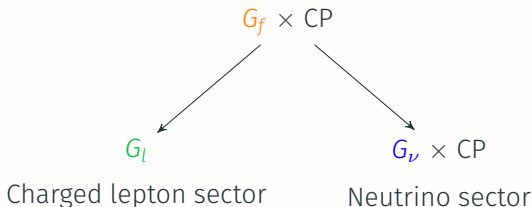
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Discrete flavour symmetries

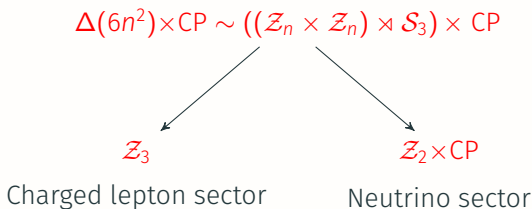
- Discrete symmetry G_f at high scale, broken at low scale into residual symmetries $G_l, G_\nu \subset G_f$.



- What group to choose ?
 - * G_f discrete subgroup of $U(3)$ (not always necessary)
 - * G_f non-abelian to avoid texture zero
 - * G_l abelian and minimal to avoid imposing too strong constraints on the charged lepton masses
 - * G_ν as minimal as possible

Discrete flavour symmetries

- Discrete symmetry G_f at high scale, broken at low scale into residual symmetries $G_l, G_\nu \subset G_f$.



In our setup

Prediction

$$U_{\text{PMNS}} = \Omega(\mathbf{3}) R_{ij}(\theta_L) K_\nu$$

$$Y = \Omega(\mathbf{3}) R_{ij}(\theta_L) \text{diag}(y_1, y_2, y_3) P_{kl}^{ij} R_{kl}(-\theta_R) \Omega(\mathbf{3}')^\dagger$$

Parametrisation of flavour symmetries

- 4 qualitatively different scenarios:

Case 1), Case 2), Case 3 a) and Case 3 b.1).

- 13 \rightarrow 6 or 7 free parameters: For Case 1),

$$\phi_S, \theta_R, M_1 \approx M_2 \approx M_3, m_0.$$

\rightarrow Better **analytical understanding** of the parameter space.

- Total coupling proportional to

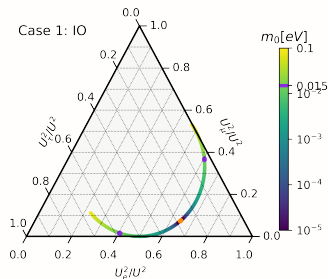
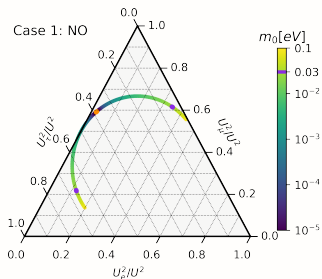
$$U^2 \propto \frac{1}{|\cos(2\theta_R)|}, \frac{1}{|\sin(2\theta_R)|}.$$

$\Leftrightarrow \theta_R \rightarrow k\frac{\pi}{4}, k \in \mathbb{Z}$ (but enhanced residual symmetry) leads to experimentally testable scenarios !

- Can relate low- and high-scale parameters. For Case 1):

$$\sin(\delta) = 0, |\sin(\alpha)| = |\sin(6\phi_S)|, \sin(\beta) = 0.$$

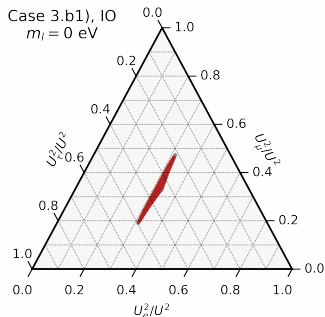
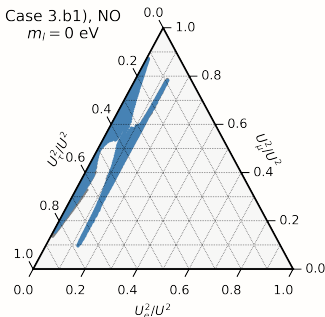
Ternary plots for Case 1)



[Drewes/Hagedorn/YG/Klarić; 24xx.xxxxx]

- Enhanced predictivity compared to the agnostic scenario
- Branching ratio fixed (or 2 possibilities) for fixed m_0
↪ Can pinpoint m_0 at colliders just by measuring the HNs branching ratio.
- Other cases are slightly less predictive.

Ternary plots for Case 3 b.1)



[Drewes/Hagedorn/YG/Klaric; 24xx.xxxxx]

- Enhanced predictivity compared to the agnostic scenario
- Branching ratio fixed (or 2 possibilities) for fixed m_0
↪ Can pinpoint m_0 at colliders just by measuring the HNs branching ratio.
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Quantum kinetic equations

$$i \frac{d}{dt} \rho = [H, \delta\rho] - \frac{i}{2} \{\Gamma, \delta\rho\} - i \sum_{a \in \{e, \mu, \tau\}} \tilde{\Gamma}_a \frac{\mu_a}{T} f_F(1 - f_F),$$

$$i \frac{d}{dt} \bar{\rho} = -[H, \delta\bar{\rho}] - \frac{i}{2} \{\Gamma, \delta\bar{\rho}\} + i \sum_{a \in \{e, \mu, \tau\}} \tilde{\Gamma}_a \frac{\mu_a}{T} f_F(1 - f_F),$$

$$\frac{d}{dt} n_{\Delta_a} = -\frac{2i\mu_a}{T} \int \frac{d^3\vec{k}}{(2\pi)^3} \text{Tr}[\Gamma_a] f_F(1 - f_F) + i \int \frac{d^3\vec{k}}{(2\pi)^3} \text{Tr}[\tilde{\Gamma}_a (\delta\bar{\rho} - \delta\rho)].$$

Density matrix (points to ρ)
Effective Hamiltonian (points to H)
Interaction rates (points to $\tilde{\Gamma}_a$)
Lepton asymmetry (points to μ_a/T)
Source term (points to $\tilde{\Gamma}_a (\delta\bar{\rho} - \delta\rho)$)

- **Interaction rates** can be
 - ★ Fermion number **conserving** $\sim (Y^\dagger Y)T$
 - ★ Fermion number **violating** $\sim (Y^\dagger Y^*) \frac{M^2}{T}$
- Refined calculation subject to intensive studies over the last years, e.g. Anisimov/Bedak/Bödeker '10, Garny/Kartavtsev/Hohenegger '11, Drewes/Garbrecht/Gueter/Klarič '16, Hernandez/Kekic/Lopez-Pavon/Racker/Salvado '16, Laine/Ghiglieri '16 '18, Klarič/Shaposhnikov/Timiryasov '21, ...

- Perturbatively,

$$Y_B \propto \text{Tr} \left(\tilde{\Gamma}_\alpha (\delta\rho - \delta\bar{\rho}) \right) \propto \text{Tr} \left(\tilde{\Gamma}_\alpha [H_N, \Gamma] \right)$$

$$H_N = \frac{M_M^2}{2E} + h_+(T)Y^\dagger Y + h_-(T)Y^t Y^*, \quad \Gamma, \tilde{\Gamma} = \pm\gamma_+(T)Y^\dagger Y + \gamma_-(T)Y^t Y^*$$

- BAU production governed by

$$C_{\text{LFV},\alpha} = i \text{Tr} \left([M_M^2, Y^\dagger Y] Y^\dagger P_\alpha Y \right),$$

$$C_{\text{LNV},\alpha} = i \text{Tr} \left([M_M^2, Y^\dagger Y] Y^T P_\alpha Y^* \right),$$

$$C_{\text{DEG},\alpha} = i \text{Tr} \left([Y^T Y^*, Y^\dagger Y] Y^T P_\alpha Y^* \right),$$

CP-violation combinations

- Perturbatively,

$$Y_B \propto \text{Tr} \left(\tilde{\Gamma}_\alpha (\delta\rho - \delta\bar{\rho}) \right) \propto \text{Tr} \left(\tilde{\Gamma}_\alpha [H_N, \Gamma] \right)$$

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- BAU production governed by

Flavour violating only $\mathcal{C}_{\text{LFV},\alpha} = i \text{Tr} \left([M_M^2, Y^\dagger Y] Y^\dagger P_\alpha Y \right),$

$$\sum_\alpha \mathcal{C}_{\text{LFV},\alpha} = 0$$

$\mathcal{C}_{\text{LNV},\alpha} = i \text{Tr} \left([M_M^2, Y^\dagger Y] Y^T P_\alpha Y^* \right),$

$\mathcal{C}_{\text{DEG},\alpha} = i \text{Tr} \left([Y^T Y^*, Y^\dagger Y] Y^T P_\alpha Y^* \right),$

Flavour violating only, can be $\neq 0$ for $\Delta M = 0!$

Violates lepton number

$$\sum_\alpha \mathcal{C}_{\text{LNV},\alpha} \neq 0$$

- Perturbatively,

$$Y_B \propto \text{Tr} \left(\tilde{\Gamma}_\alpha (\delta\rho - \delta\bar{\rho}) \right) \propto \text{Tr} \left(\tilde{\Gamma}_\alpha [H_N, \Gamma] \right)$$

$$H_N = \frac{M_M^2}{2E} + h_+(T)Y^\dagger Y + h_-(T)Y^t Y^*, \quad \Gamma, \tilde{\Gamma} = \pm\gamma_+(T)Y^\dagger Y + \gamma_-(T)Y^t Y^*$$

- For Case 1,

$$C_{\text{LFV},\alpha} \sim \frac{8}{3} M^2 \kappa y_2 y_3 (y_2^2 - y_3^2) \sin \theta_{L,\alpha} \sin \theta_R \cos 3 \phi_S,$$

$$C_{\text{LNV},\alpha} \sim \frac{8}{3} M^2 \kappa y_2 y_3 (y_3^2 \cos(2\theta_R) - y_2^2) \sin \theta_{L,\alpha} \sin \theta_R \cos 3 \phi_S,$$

$$C_{\text{DEG},\alpha} = 0,$$

where

$$\theta_{L,\alpha} = \theta_L + \rho_\alpha \frac{4\pi}{3} \text{ with } \rho_e = 0, \rho_\mu = +1, \rho_\tau = -1.$$

- Perturbatively,

$$Y_B \propto \text{Tr} \left(\tilde{\Gamma}_\alpha (\delta\rho - \delta\bar{\rho}) \right) \propto \text{Tr} \left(\tilde{\Gamma}_\alpha [H_N, \Gamma] \right)$$

$$H_N = \frac{M_M^2}{2E} + h_+(T)Y^\dagger Y + h_-(T)Y^t Y^*, \quad \Gamma, \tilde{\Gamma} = \pm\gamma_+(T)Y^\dagger Y + \gamma_-(T)Y^t Y^*$$

- For Case 1,

$$C_{\text{LFV},\alpha} \sim \frac{8}{3} M^2 \kappa y_2 y_3 (y_2^2 - y_3^2) \sin \theta_{L,\alpha} \sin \theta_R \cos 3\phi_S,$$

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where

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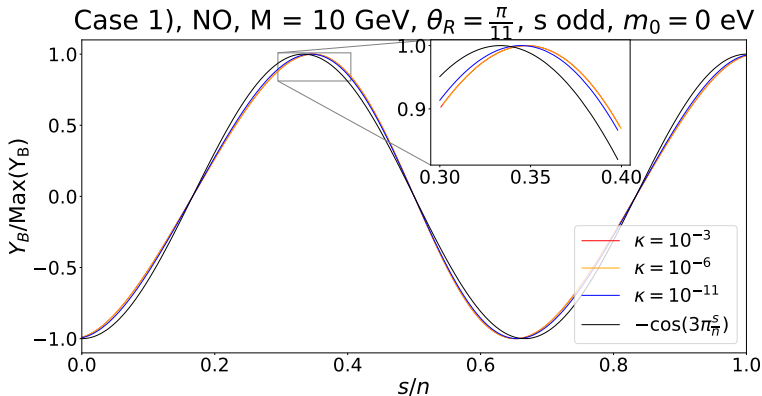


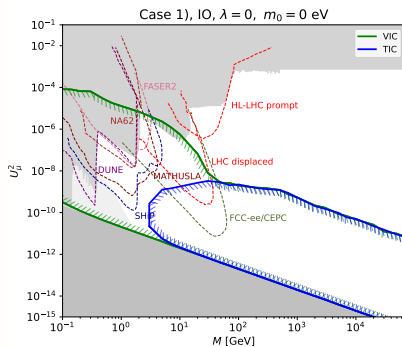
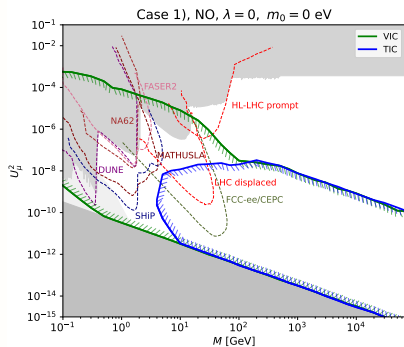
Figure 1: Vanishing initial conditions, $\lambda = 0$

[Drewes/Hagedorn/YG/Klaric; 2203.08538]

- Correlation between Y_B and low-energy observables. Here,

$$\sin(\alpha) = \sin(6\pi \frac{S}{n}).$$

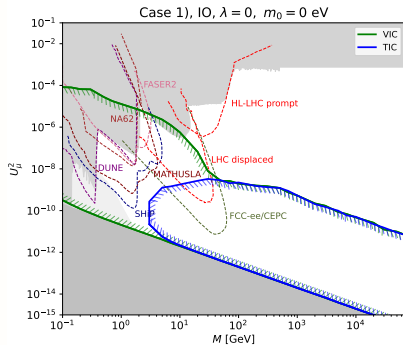
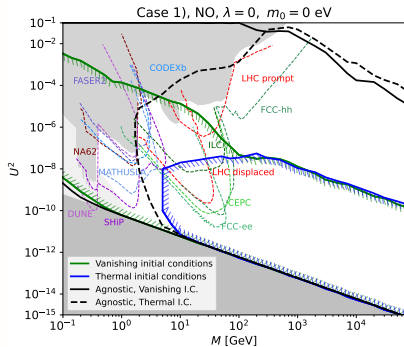
Leptogenesis with flavour symmetries



[Drewes/YG/Hagedorn/Klarić; 24xx.xxxxx]

- Reduced parameter space but remains testable

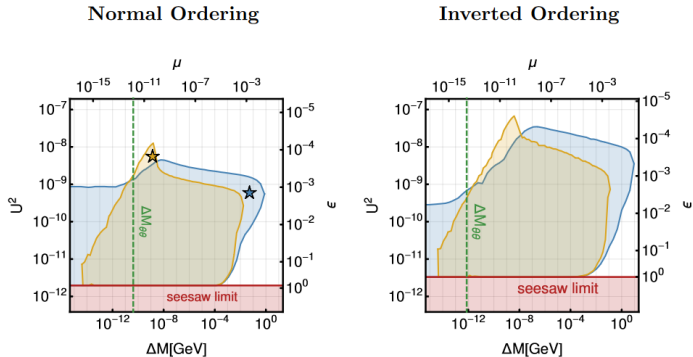
Leptogenesis with flavour symmetries



[Drewes/YG/Hagedorn/Klarić; 24xx.xxxxx]

- Reduced parameter space but remains testable

Leptogenesis in the mass degenerate case



[Antusch/Cazzato/Drewes/Fischer/Garbrecht/Gueter/Klarić; 1710.03744]

See also [Sandner/Hernandez/Lopez-Pavon/Rius; 2305.14427]

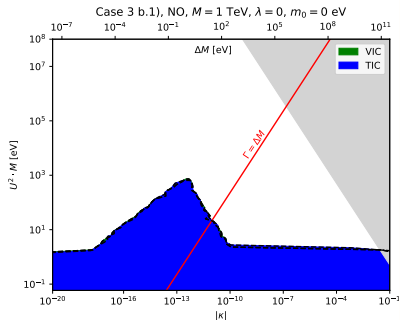
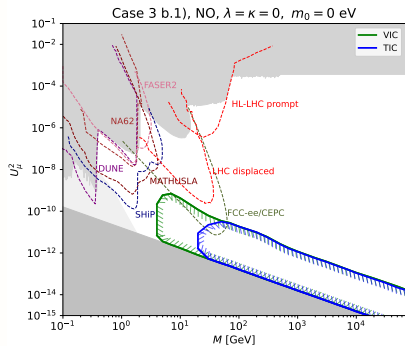
- Leptogenesis possible for $\Delta M = 0$ thanks to Higgs and thermal mass splittings

$$\Delta M_{\text{phys}} \sim h_+(T)Y^\dagger Y + h_-(T)Y^t Y^*$$

- Lepton asymmetry proportional to CP-violating combination

$$\text{Tr} \left(\tilde{\Gamma}_\alpha [H_N, \Gamma] \right) \sim \text{Tr} \left(\left[\hat{Y}^t \hat{Y}^*, \hat{Y}^\dagger \hat{Y} \right] \hat{Y}^t P_\alpha \hat{Y}^* \right) \neq 0!$$

Flavour symmetries and degenerate leptogenesis



[Drewes/YG/Hagedorn/Klarić; 24xx.xxxxx]

- For Case 3 b.1), $C_{\text{DEG},\alpha} \neq 0$! Leptogenesis viable for $\Delta M_M = 0$.

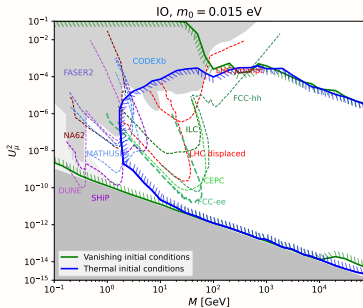
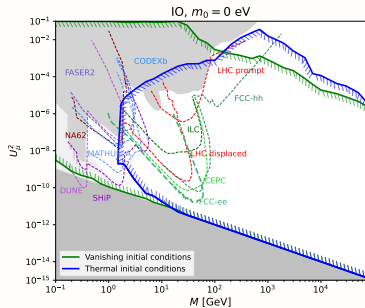
What should I take home ?

- Right-handed neutrinos provide minimal solution for ν masses + baryon asymmetry
- Leptogenesis parameter space largely enhanced for $n = 3$
- Large mixing angle opens up the possibility of testing leptogenesis by combining information from colliders, $0\nu\beta\beta$, ν oscillations, ...
- Combined with flavour symmetric explanation of PMNS: very predictive !
- Degenerate leptogenesis possible due to Higgs and thermal effects
- Collider testability of $n = 3$ scenario to be further explored

Thanks for your attention!

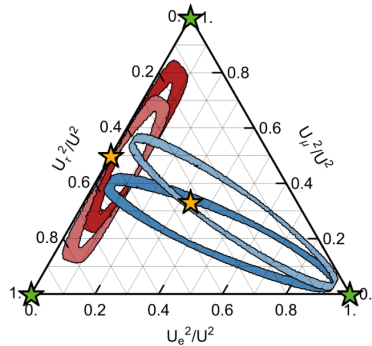
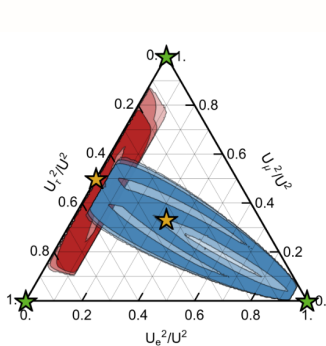
Appendix

$n = 3$ parameter space, IO



- Similar enhancement of the parameter space for IO.

Impact of low energy measurements on $\frac{U_{\alpha}^2}{U^2}$



Current ν oscillation data

DUNE projections

- New (more realistic) benchmarks proposed beyond the 1-flavour approximation
- DUNE measurement of δ could constrain the mixing to each SM flavour, hence leptogenesis

Seesaw parameter space

Consistency with ν -oscillation data induced by Casas-Ibarra parametrisation

$$F = \frac{i}{\sqrt{v}} U_\nu \sqrt{m_\nu^{diag}} R \sqrt{M_M}$$

$$\text{Seesaw relation: } m_\nu = -v^2 F \cdot M_M^{-1} \cdot F^t.$$

Seesaw parameter space

Consistency with ν -oscillation data induced by Casas-Ibarra parametrisation

$$F = \frac{i}{\sqrt{V}} U_\nu \sqrt{m_\nu^{diag}} R \sqrt{M_M}$$

R is a complex rotation matrix

n=3

3 CP-violating phases

3 PMNS angles (fixed)

3 light neutrino masses (2 fixed)

3 complex Euler angles

3 Majorana masses

13 free parameters

Seesaw parameter space

Consistency with ν -oscillation data induced by Casas-Ibarra parametrisation

$$F = \frac{i}{\sqrt{V}} U_\nu \sqrt{m_\nu^{diag}} R \sqrt{M_M}$$

n=2

- 2 CP-violating phases
- 3 PMNS angles (fixed)
- 2 light neutrino masses (fixed)
- 1 complex Euler angle
- 2 Majorana masses

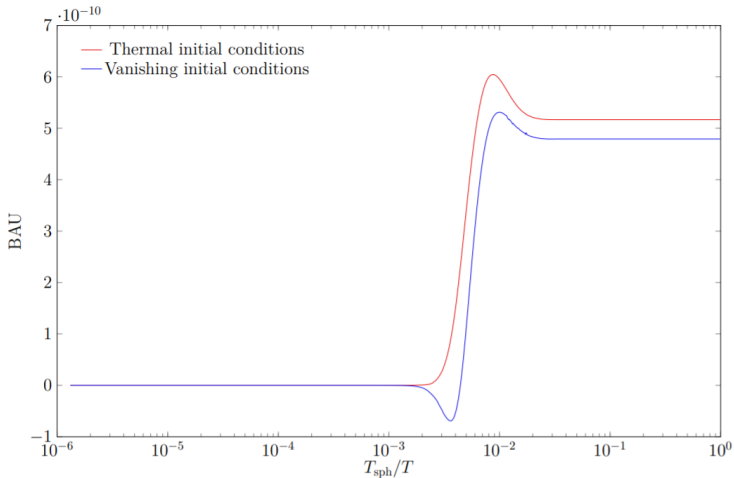
6 free parameters

n=3

- 3 CP-violating phases
- 3 PMNS angles (fixed)
- 3 light neutrino masses (2 fixed)
- 3 complex Euler angles
- 3 Majorana masses

13 free parameters

Thermal vs vanishing initial conditions



At large \bar{M} , parameter space for thermal I.C. is larger because **asymmetry** produced during **freeze-in** and **freeze-out** have **opposite** signs.