

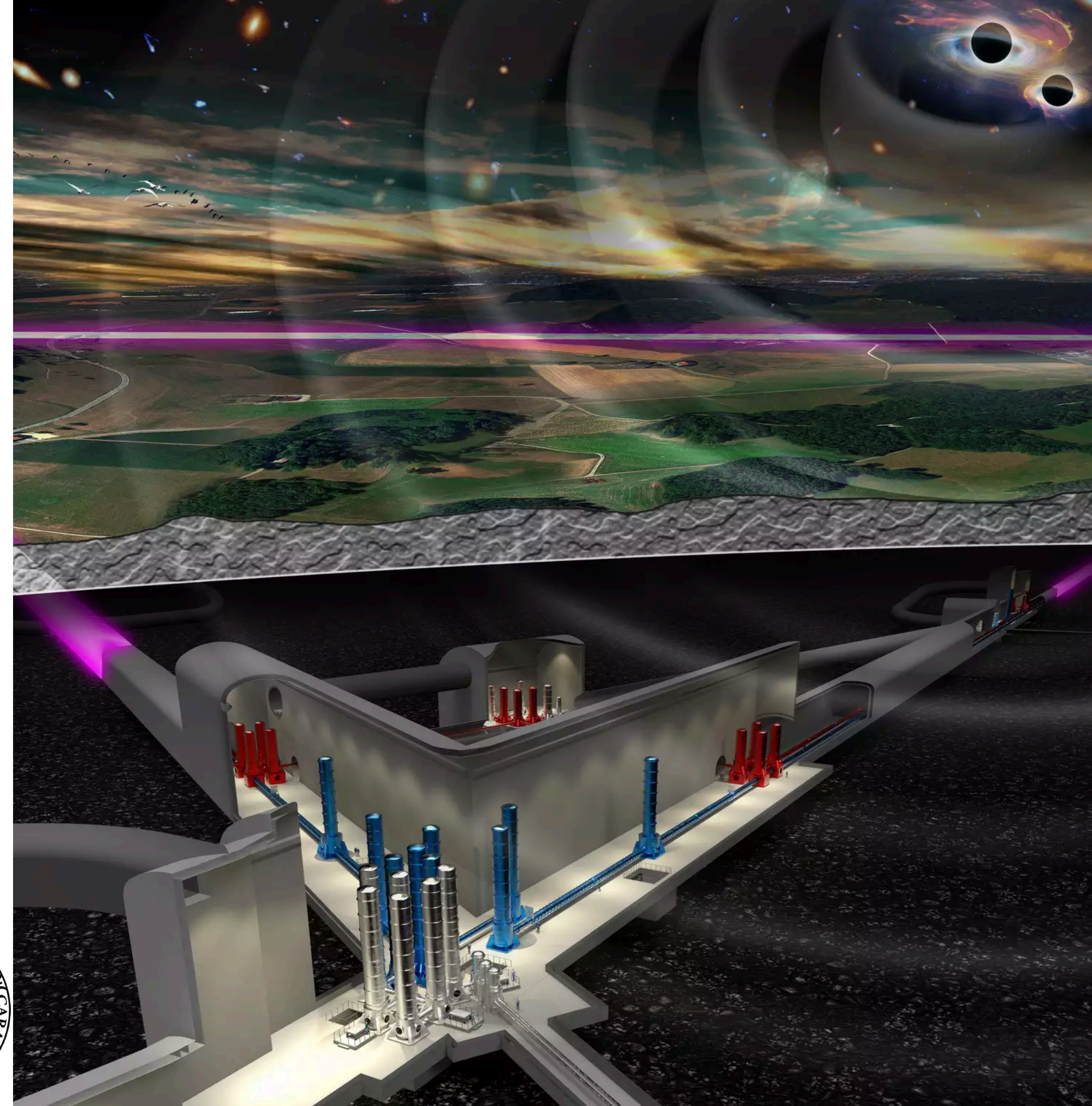
# WINES

Wind turbine Noise assessment in the Italian site candidate for Einstein Telescope

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*Istituto Nazionale di Geofisica e Vulcanologia (Italy)*

&

Sardinia Characterization Team from *UniSS, UniCA, INFN*

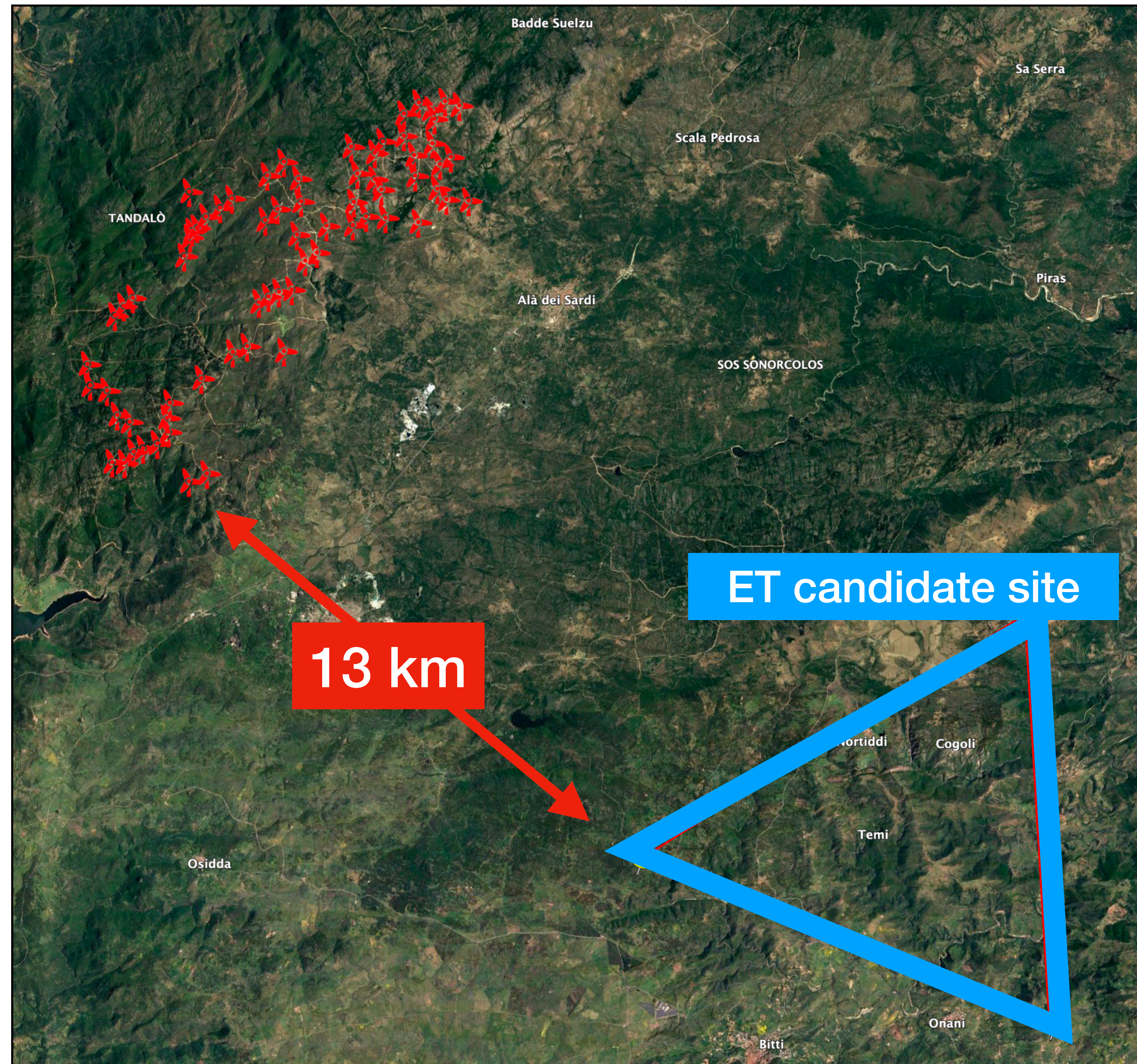


# The Buddusò Wind Park

- one of the largest Wind park in Italy
- 69 turbines (~2MW each)
- 130 MW installed

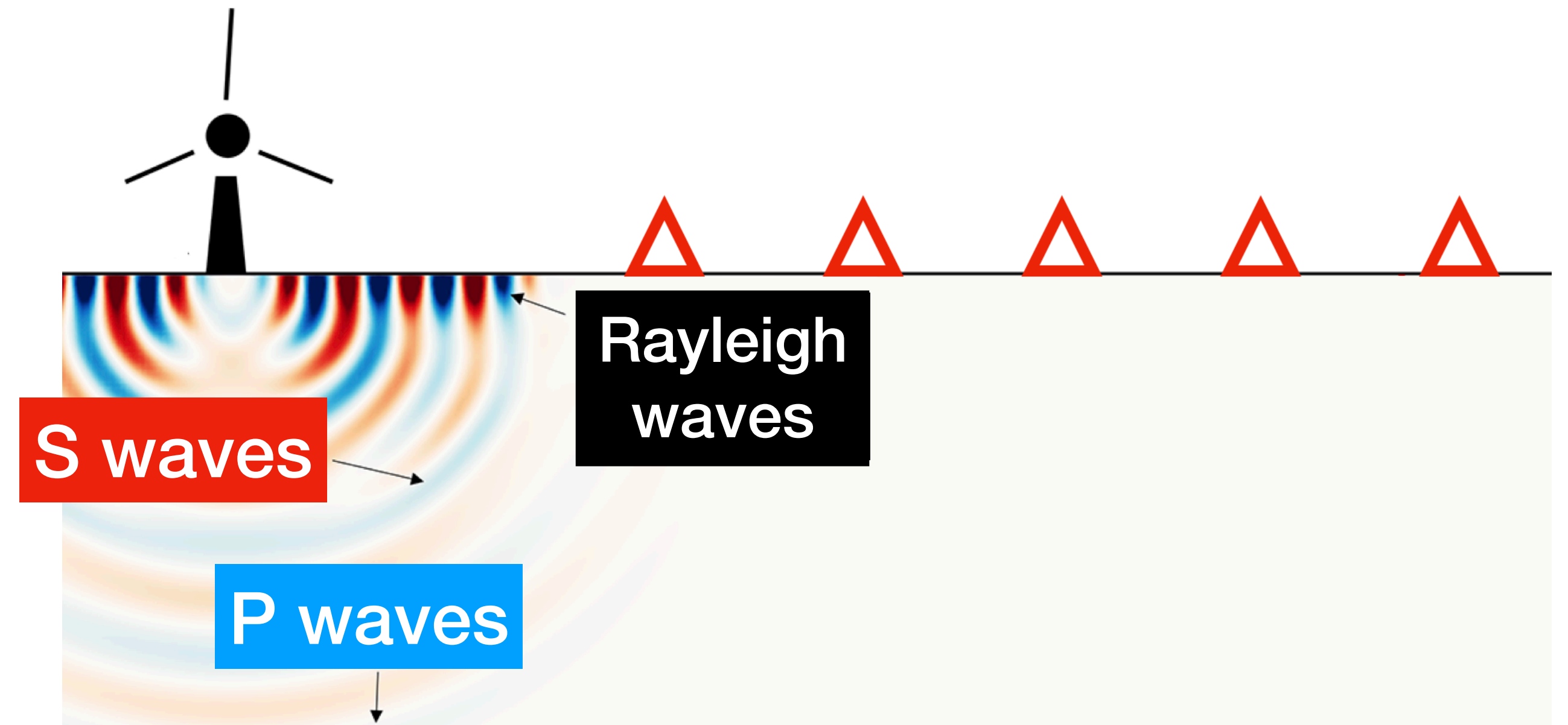
For wind energy exploitation, the area is **exceptionally promising** due to the:

- rural environment
- strong and persistent winds in any season



# Windpark and seismic noise...

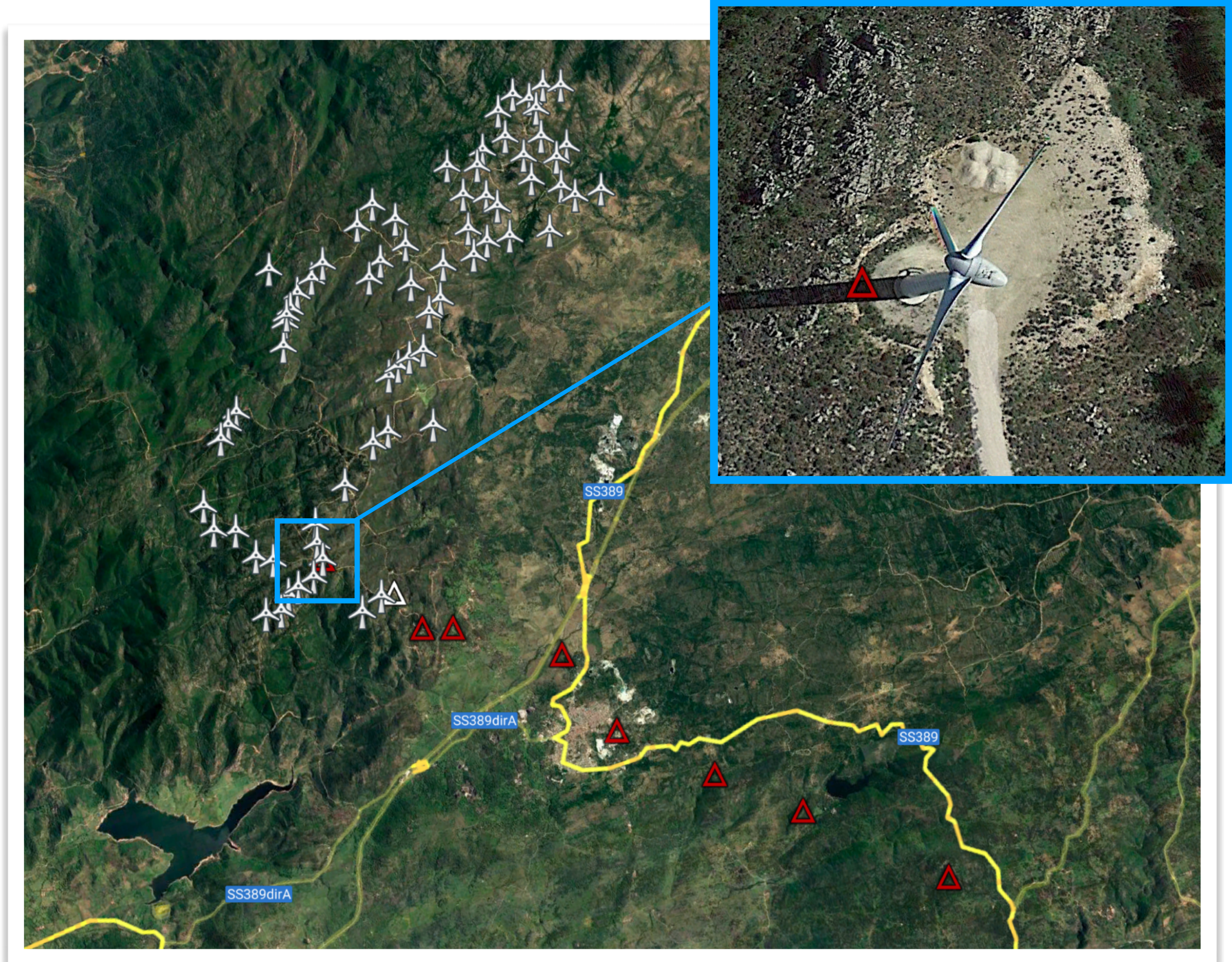
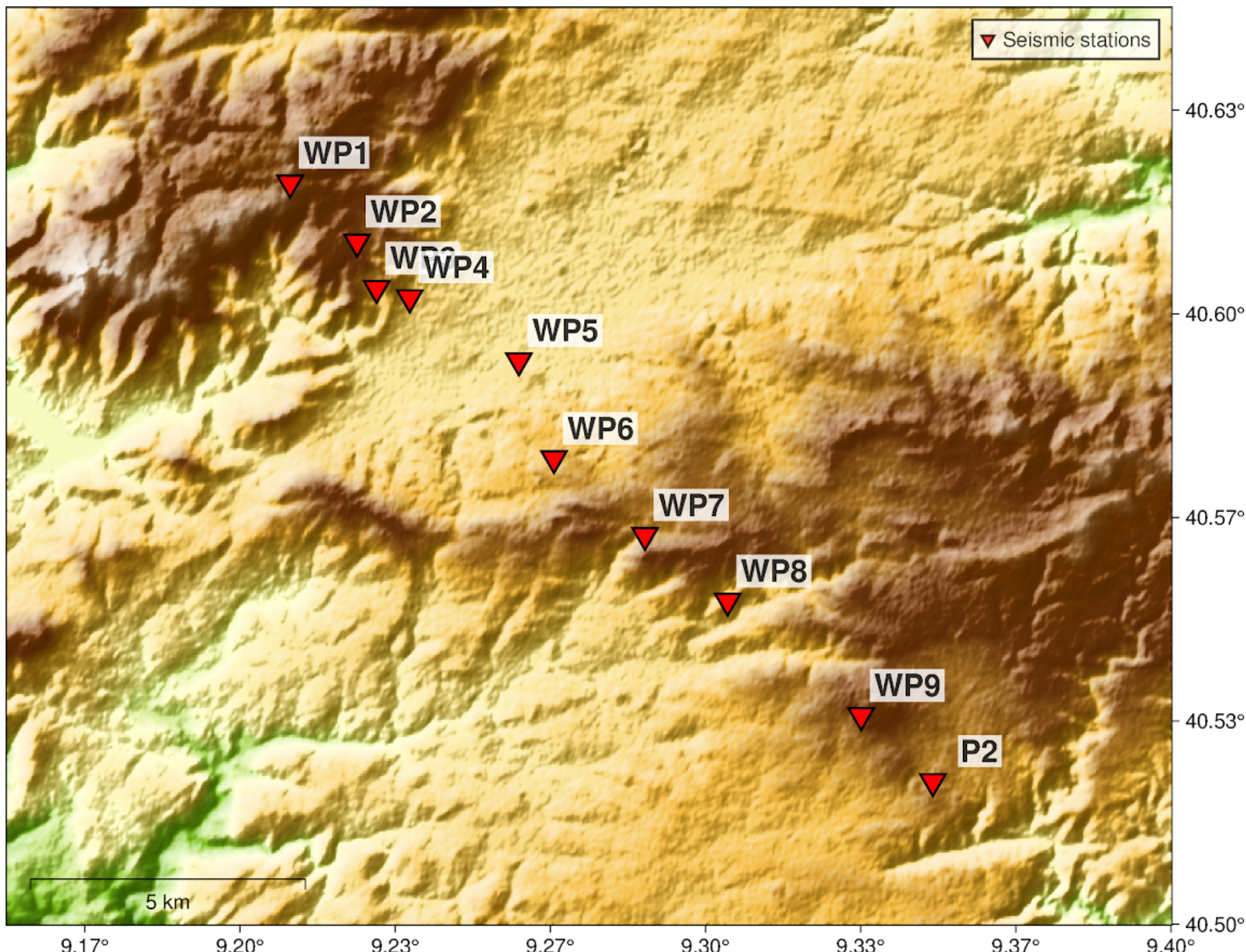
- **Blades motion** is transferred to tower, from tower **to the ground**
- Seismic noise propagates as **surface waves** (mainly Rayleigh waves)
- Generated noise is found in the **1-10 Hz frequency band...**



- what are **the characteristics** of the generated noise signal?
- **how far can we track it?**
- how does the seismic noise signal **decay with distance?**

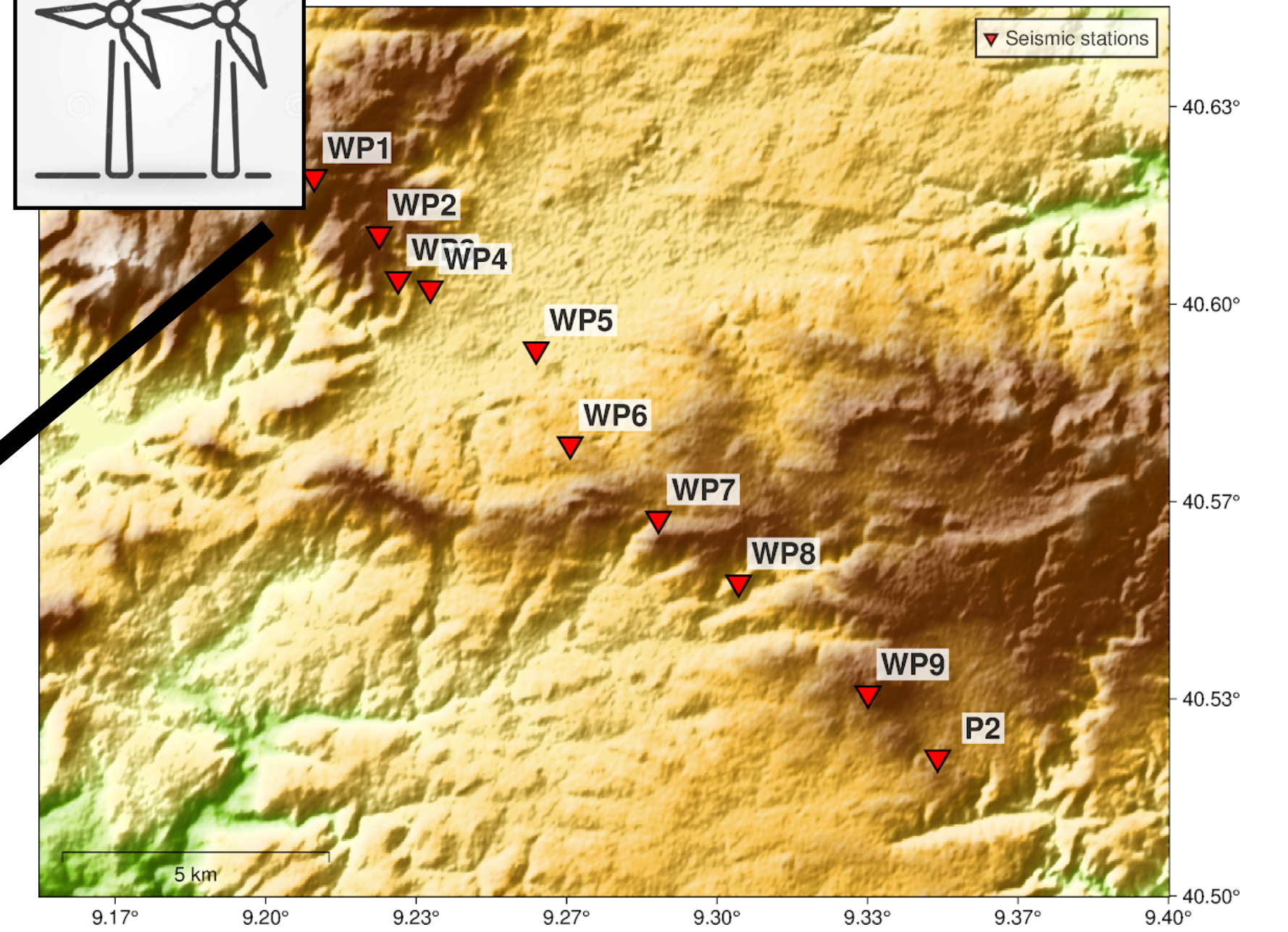
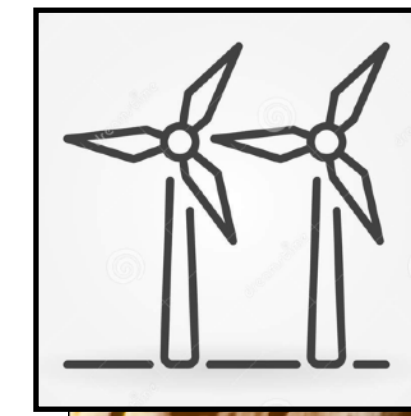
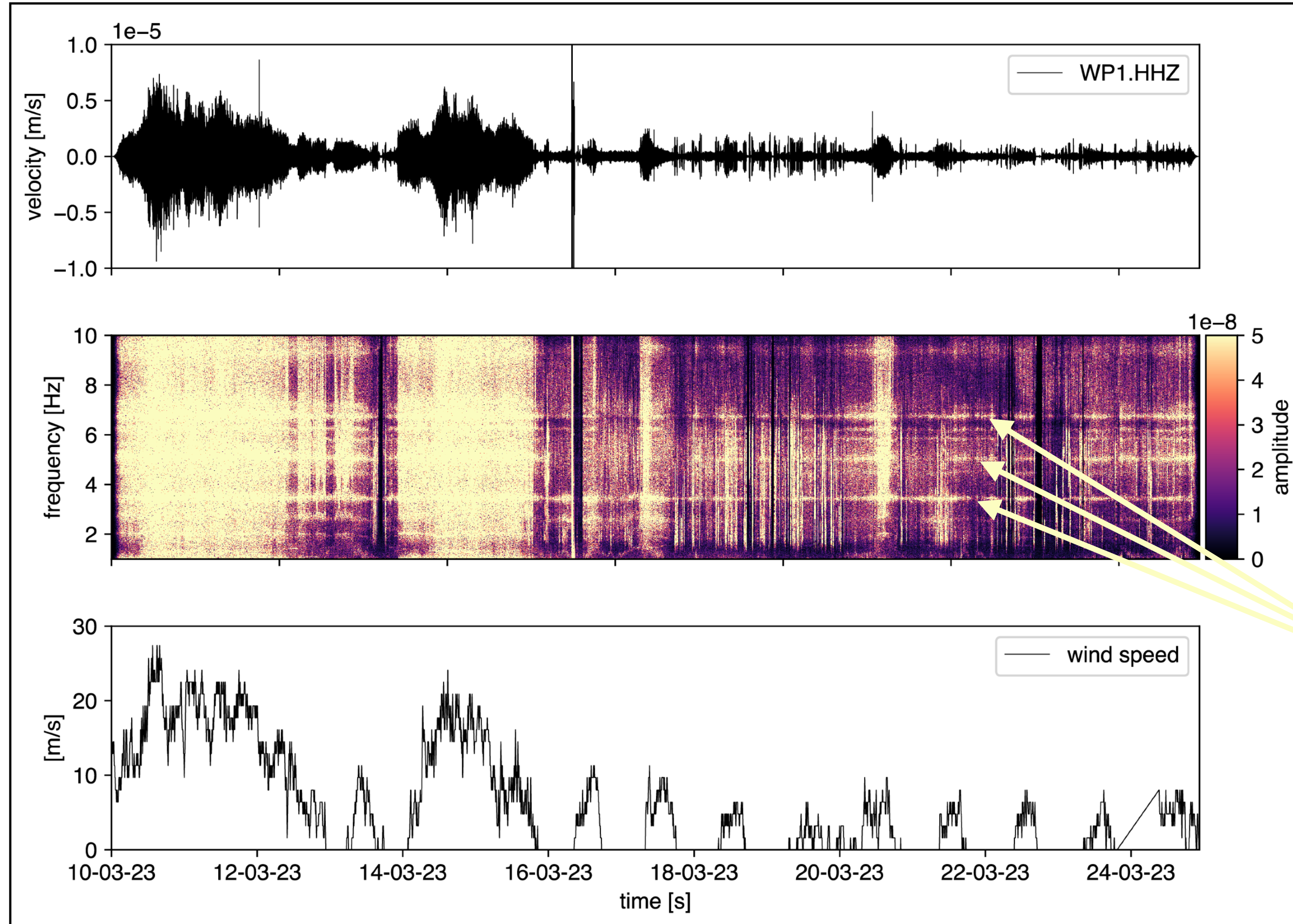
# The WINES experiment

- 9 broad-band seismic stations
- ~13 km linear array
- ~2 months of recording (8/04-30/05/2023)
- wind-speed data from a nearby meteorological station



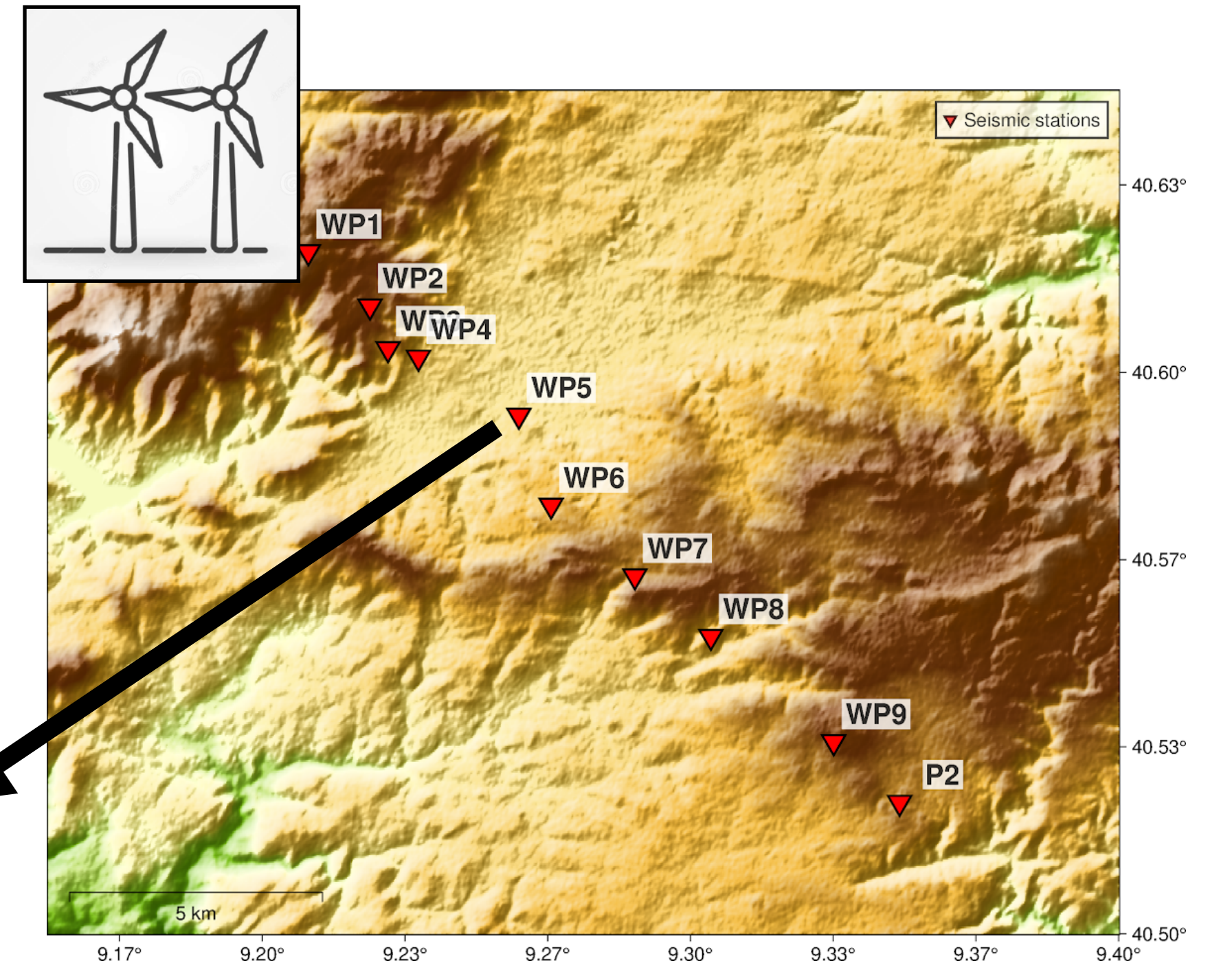
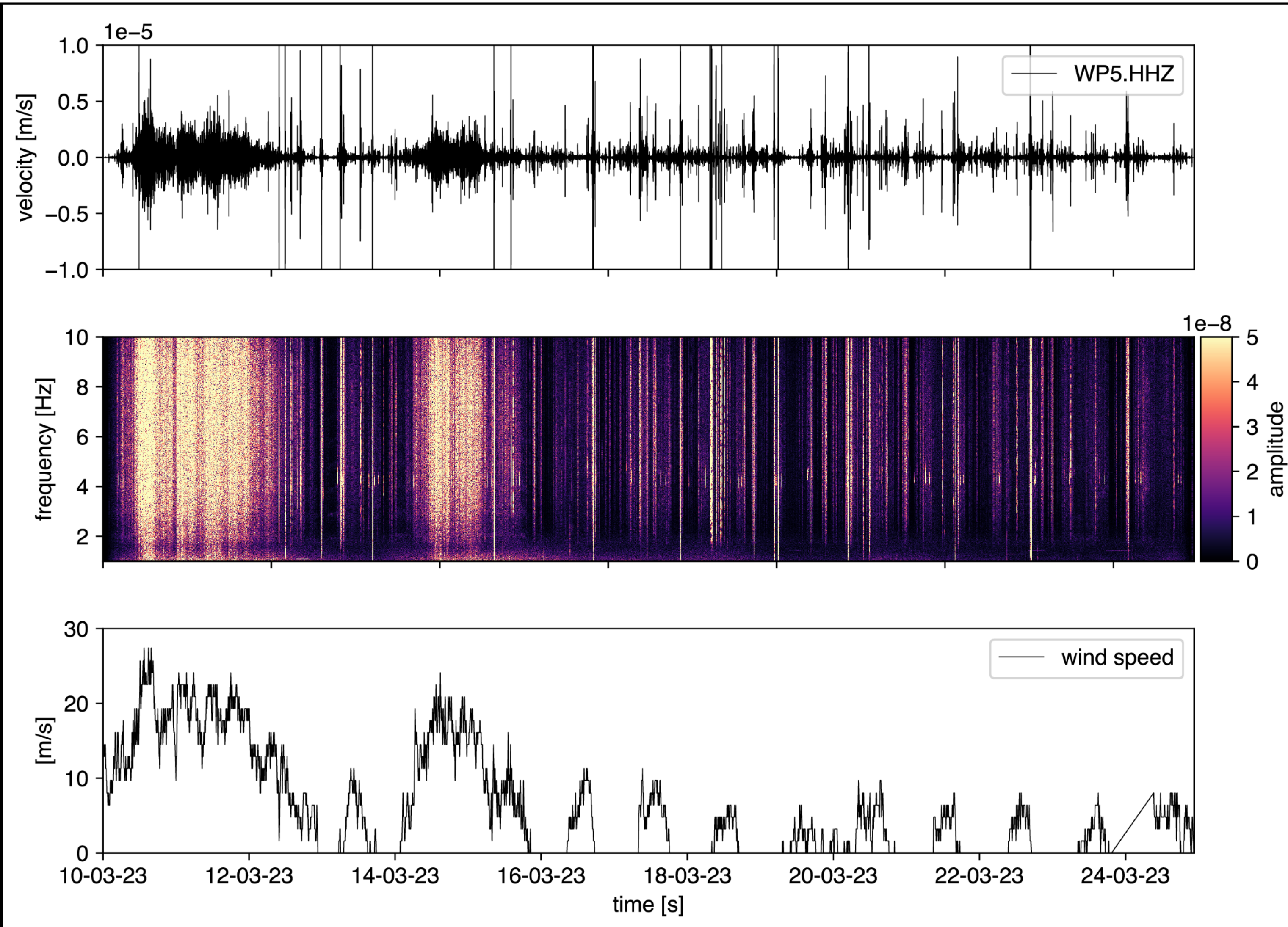
NB: in our analysis we also include the permanent stations **P2, P3**, located on the **two closest vertices** of the **ET candidate site**

# 15 days-long spectrogram

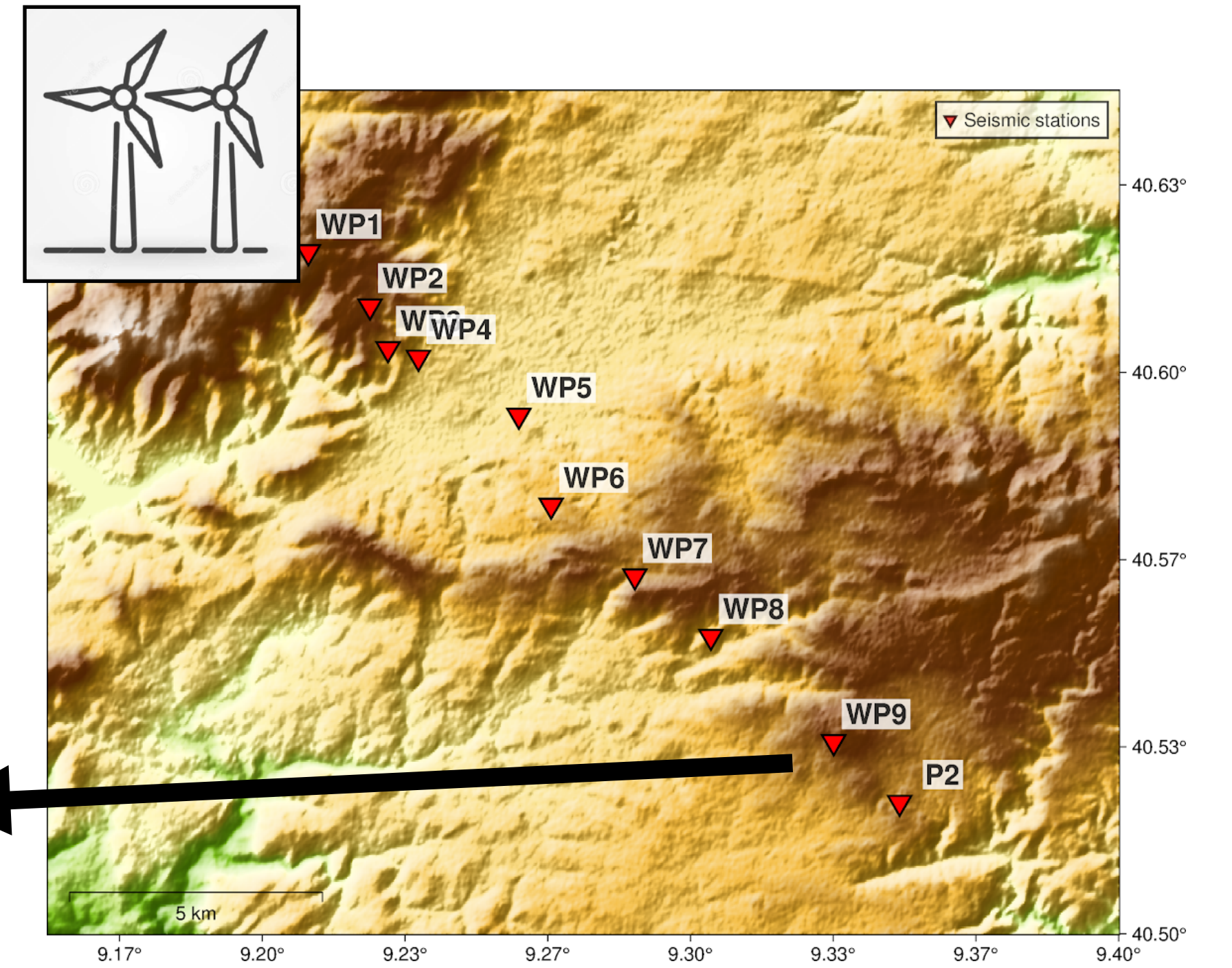
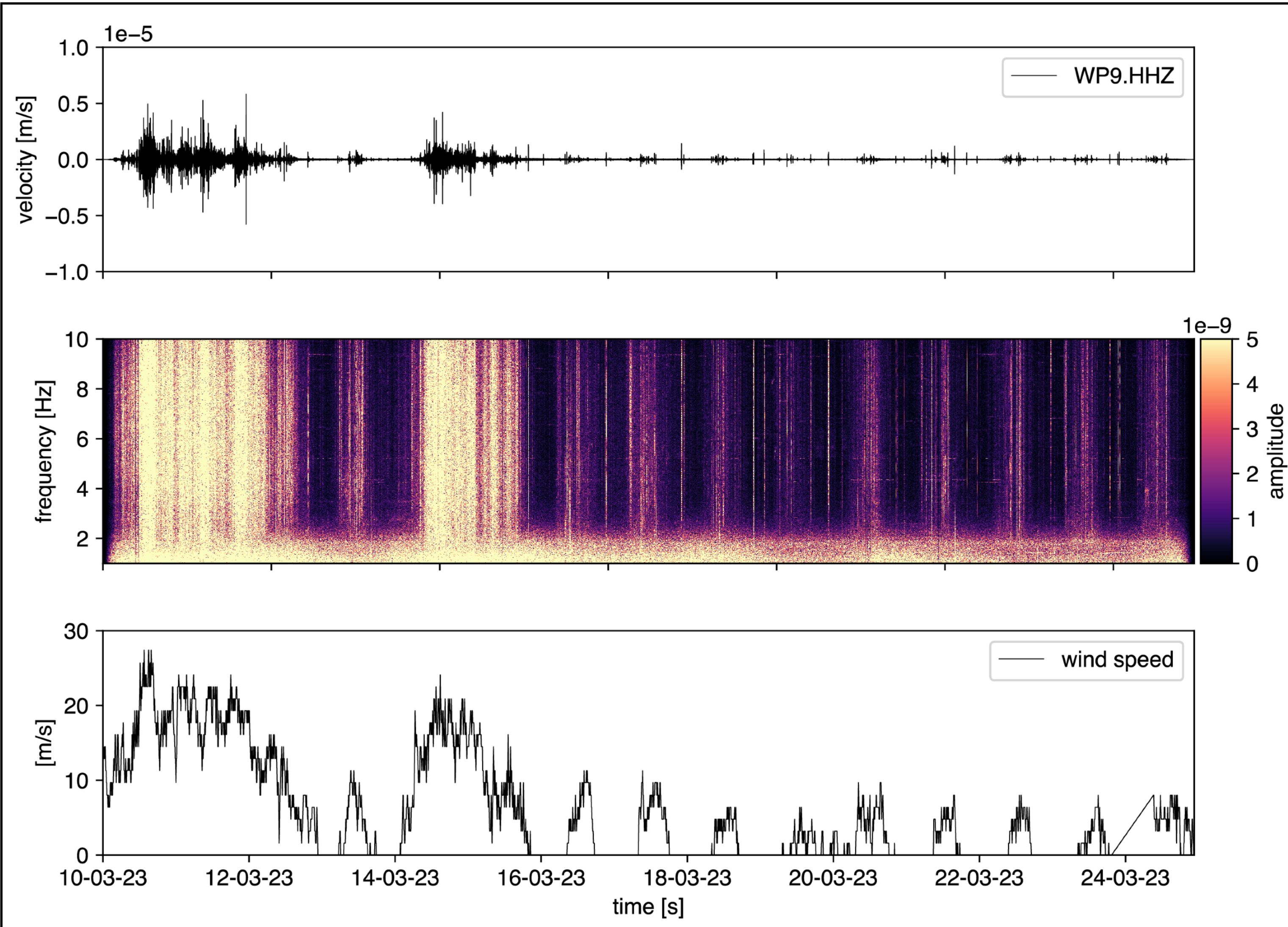


harmonics

# 15 days-long spectrogram



# 15 days-long spectrogram

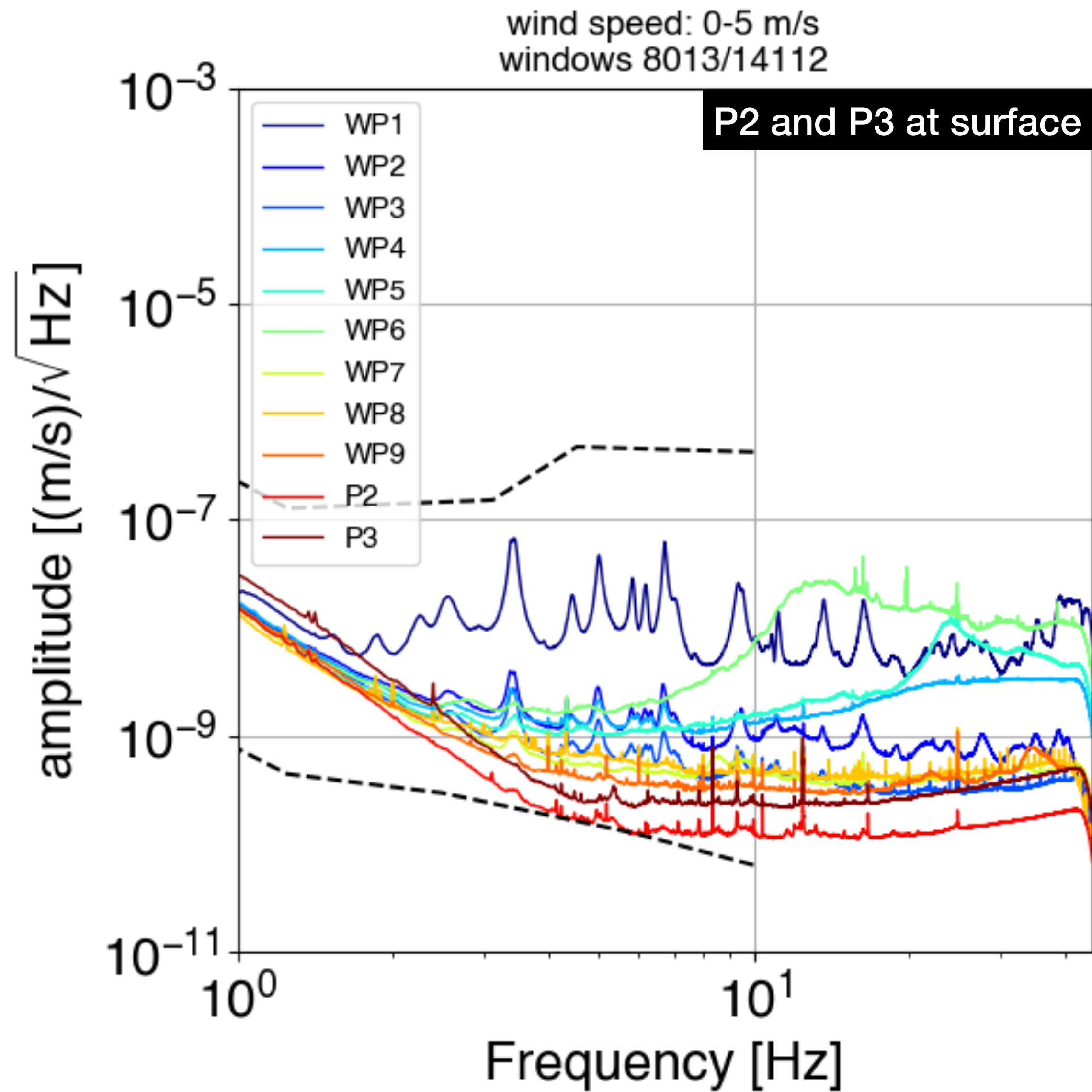


# Power spectral densities (PSDs)

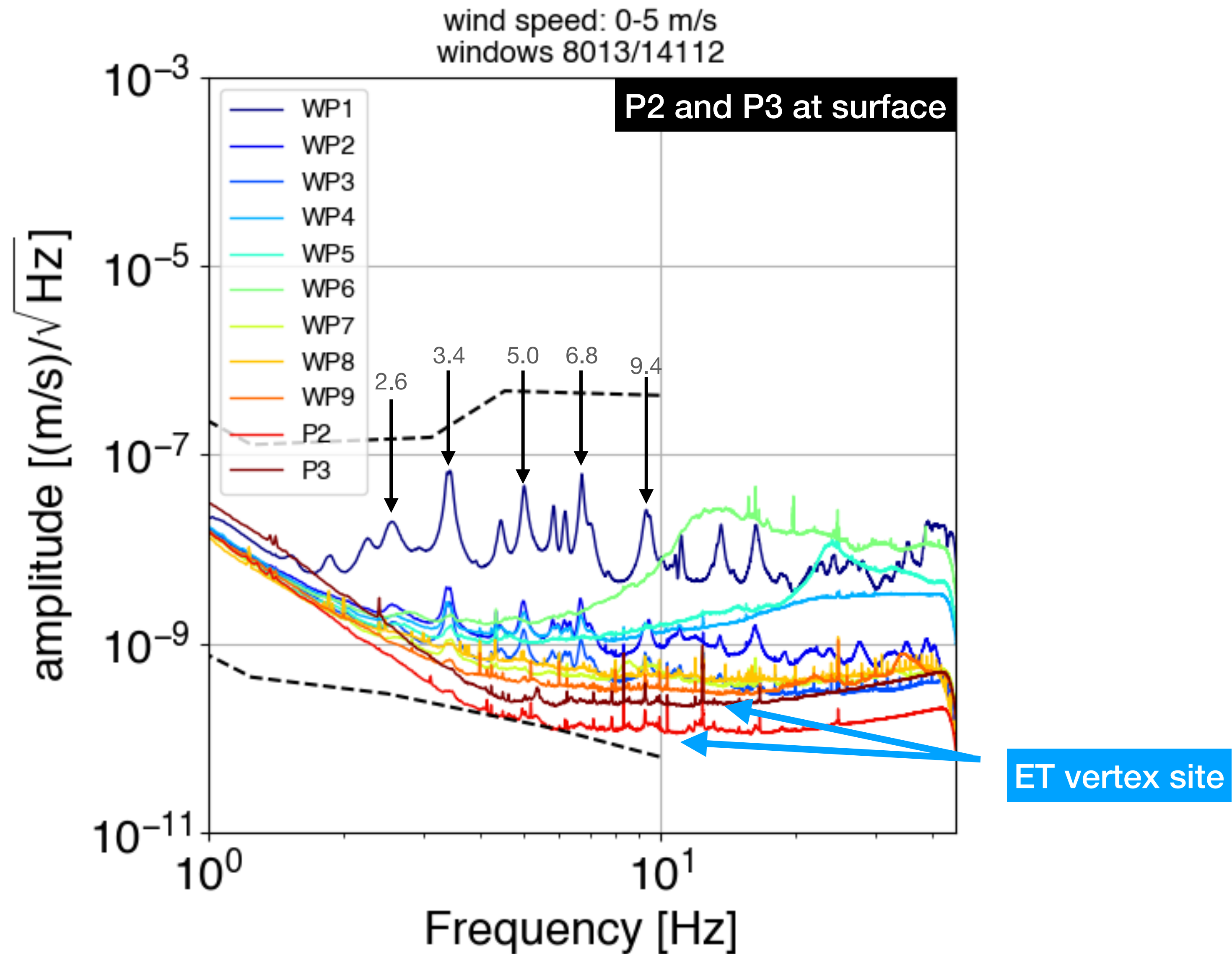
- **5 min-long window**
- **welch's method** for spectral estimation
- analysis of the **entire continuous recording**
- **median value** across time windows, **in the selected wind bands (i.e. 0-5, 10-20, 20-30 m/s)**



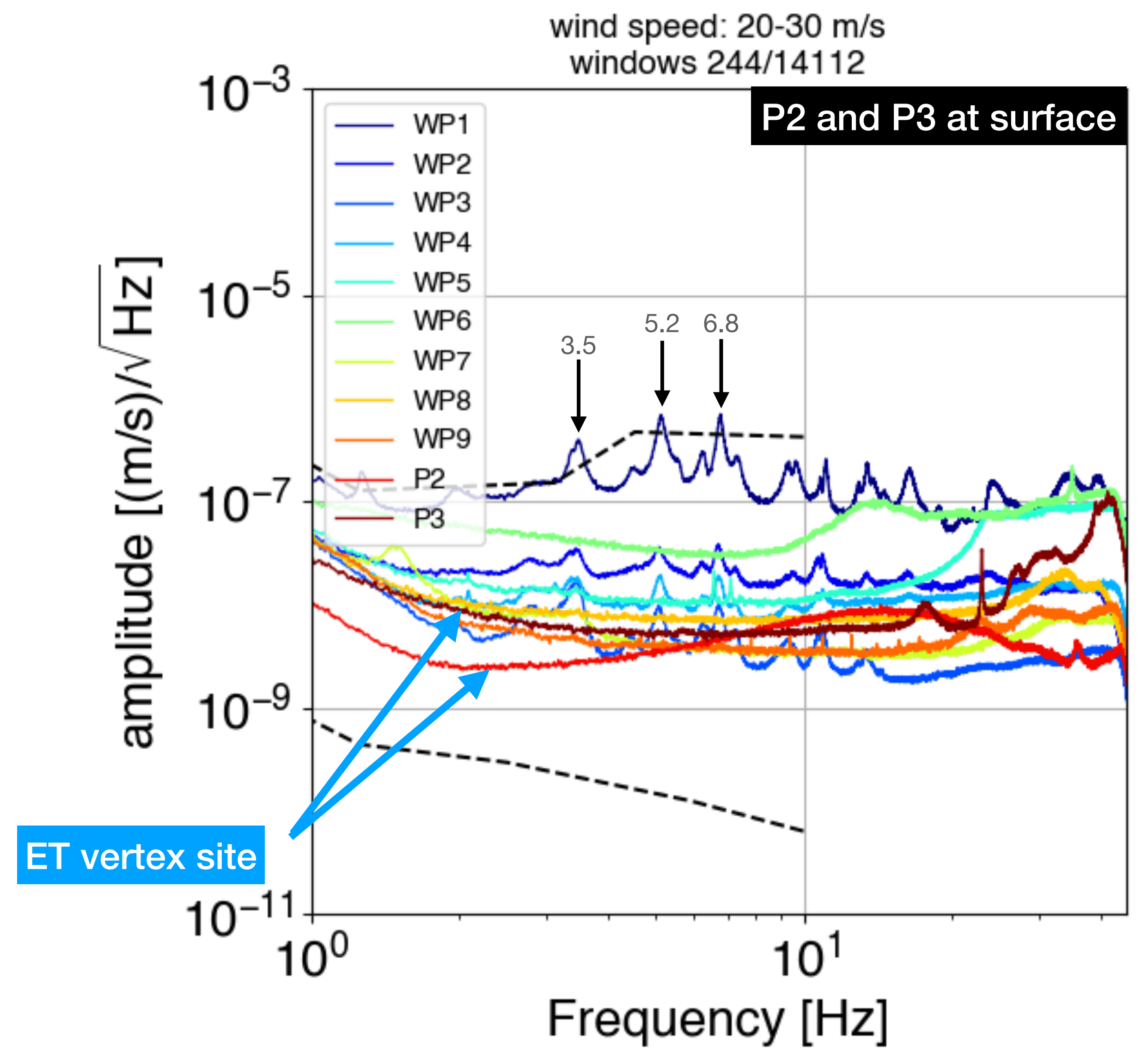
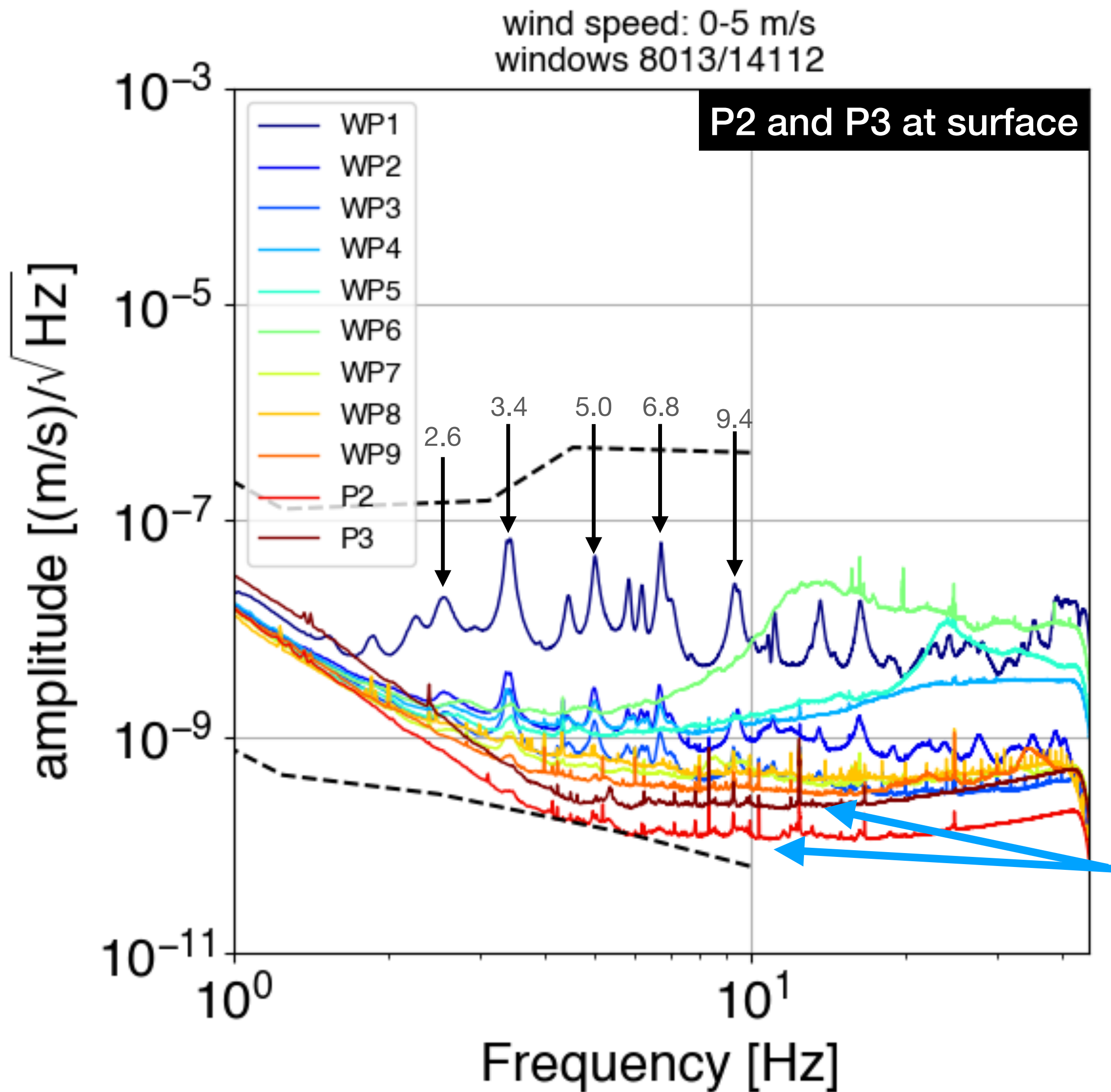
# Power spectral densities (PSDs)



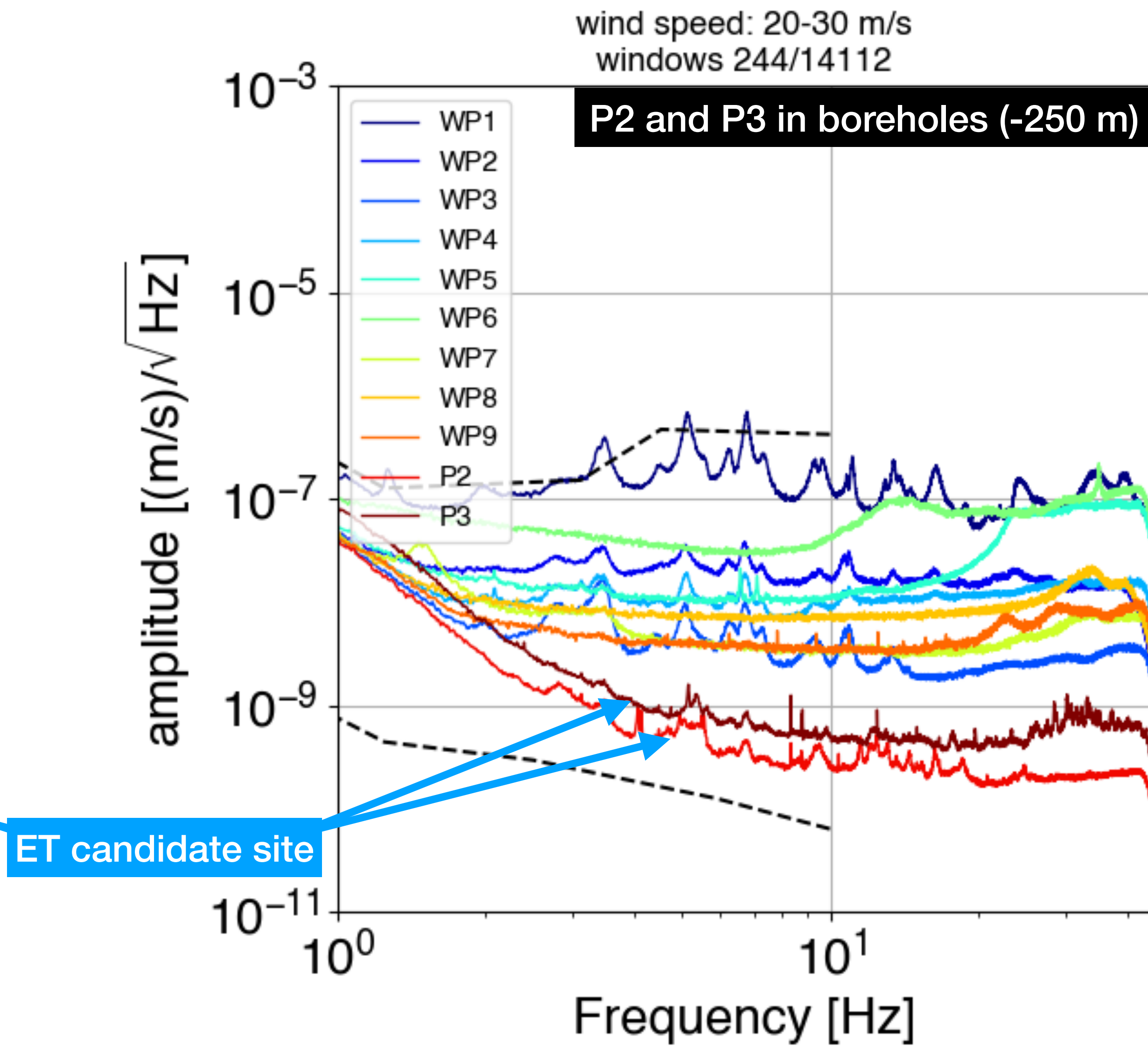
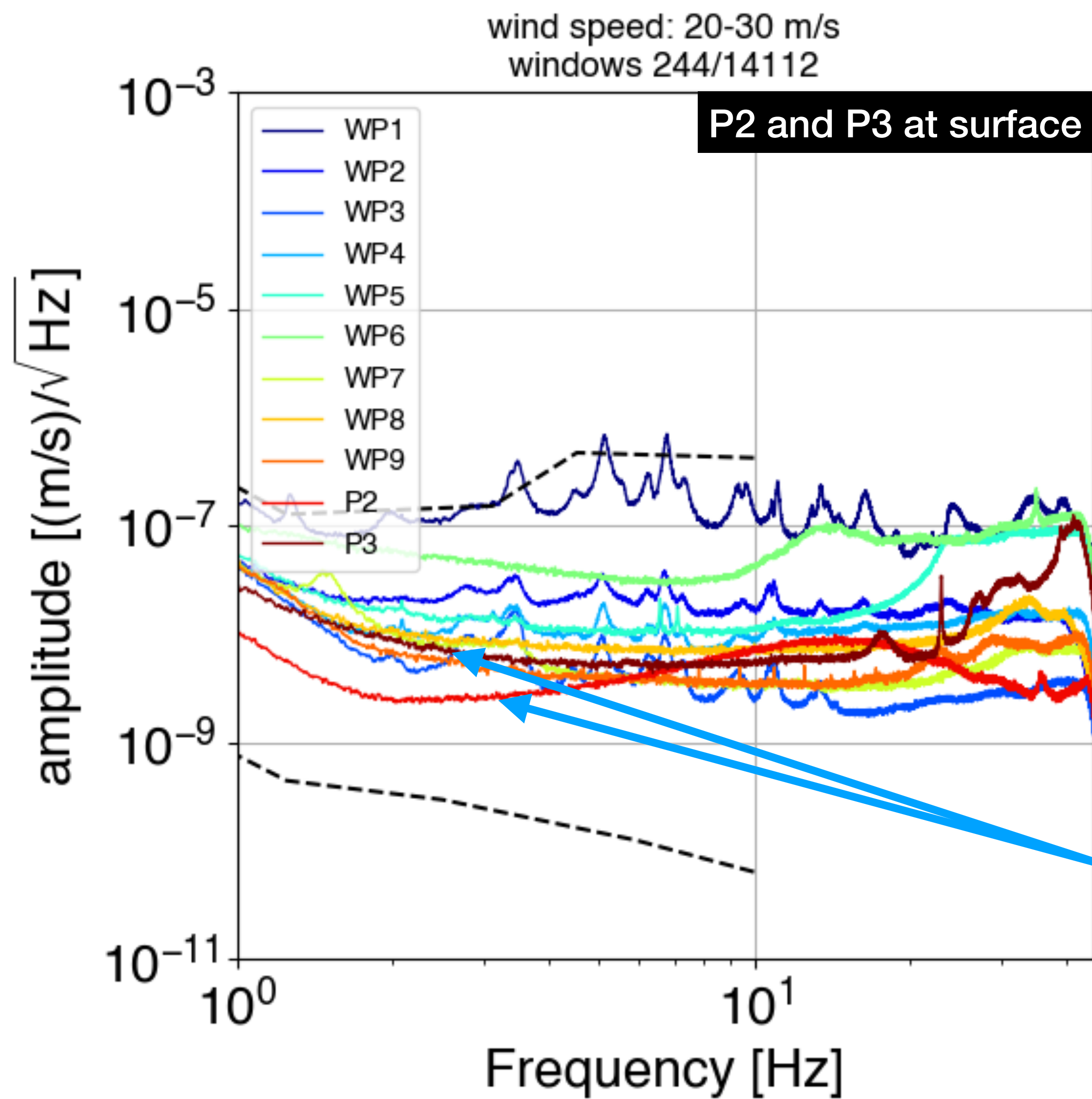
# Power spectral densities (PSDs)



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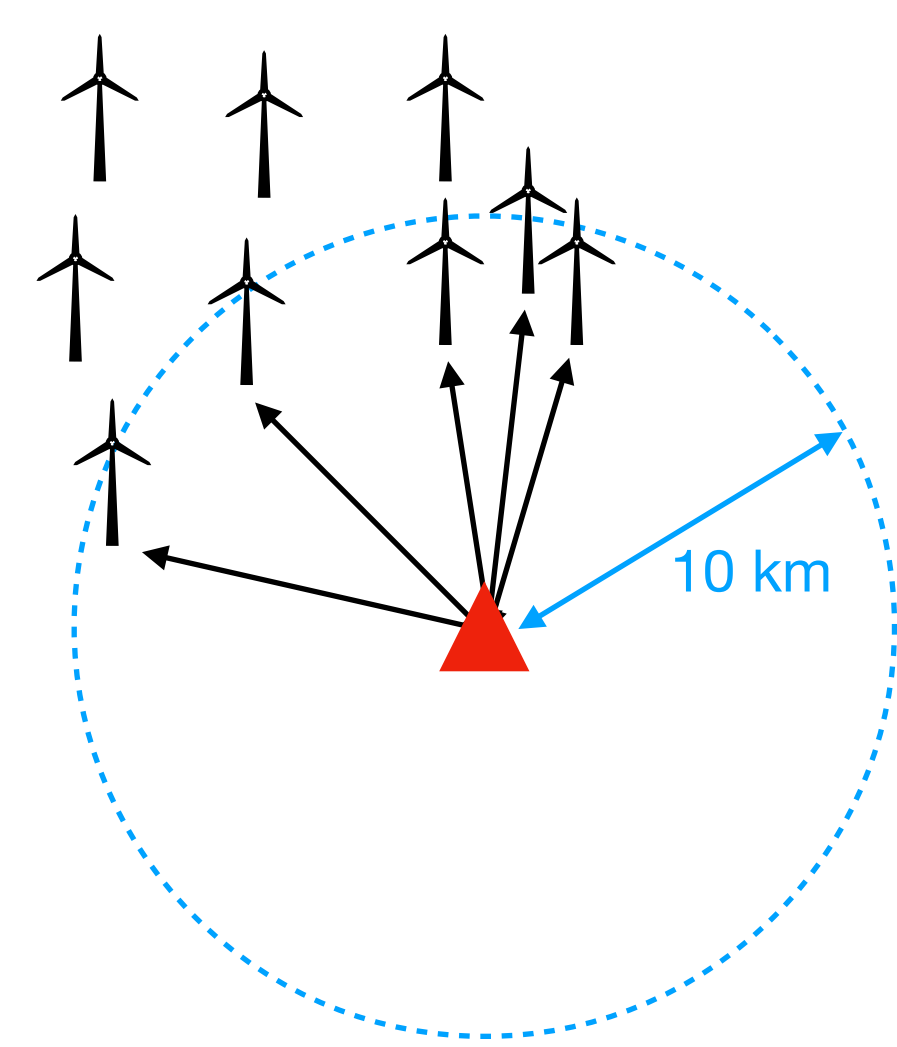
# Power spectral densities (PSDs)



ET candidate site

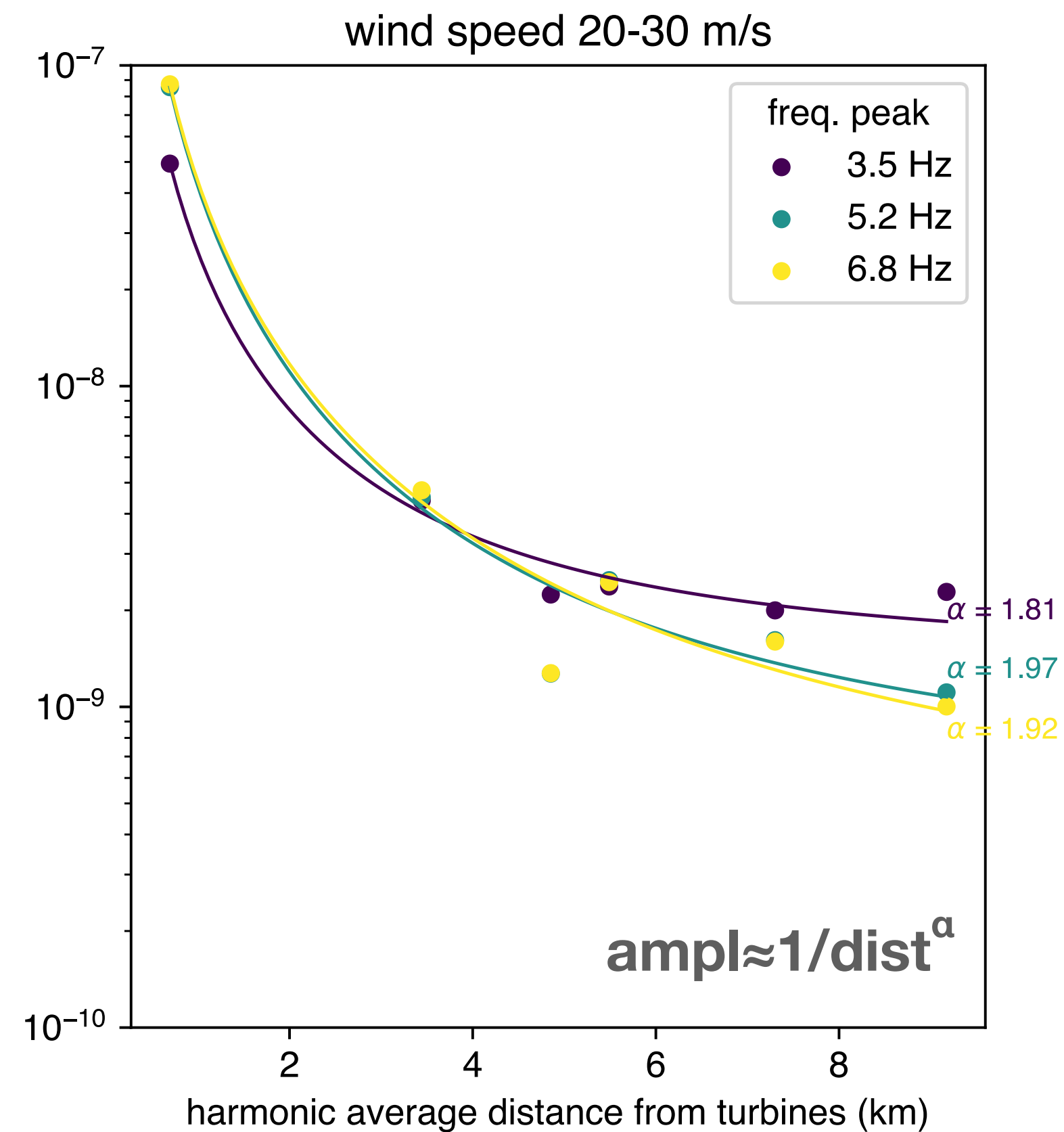
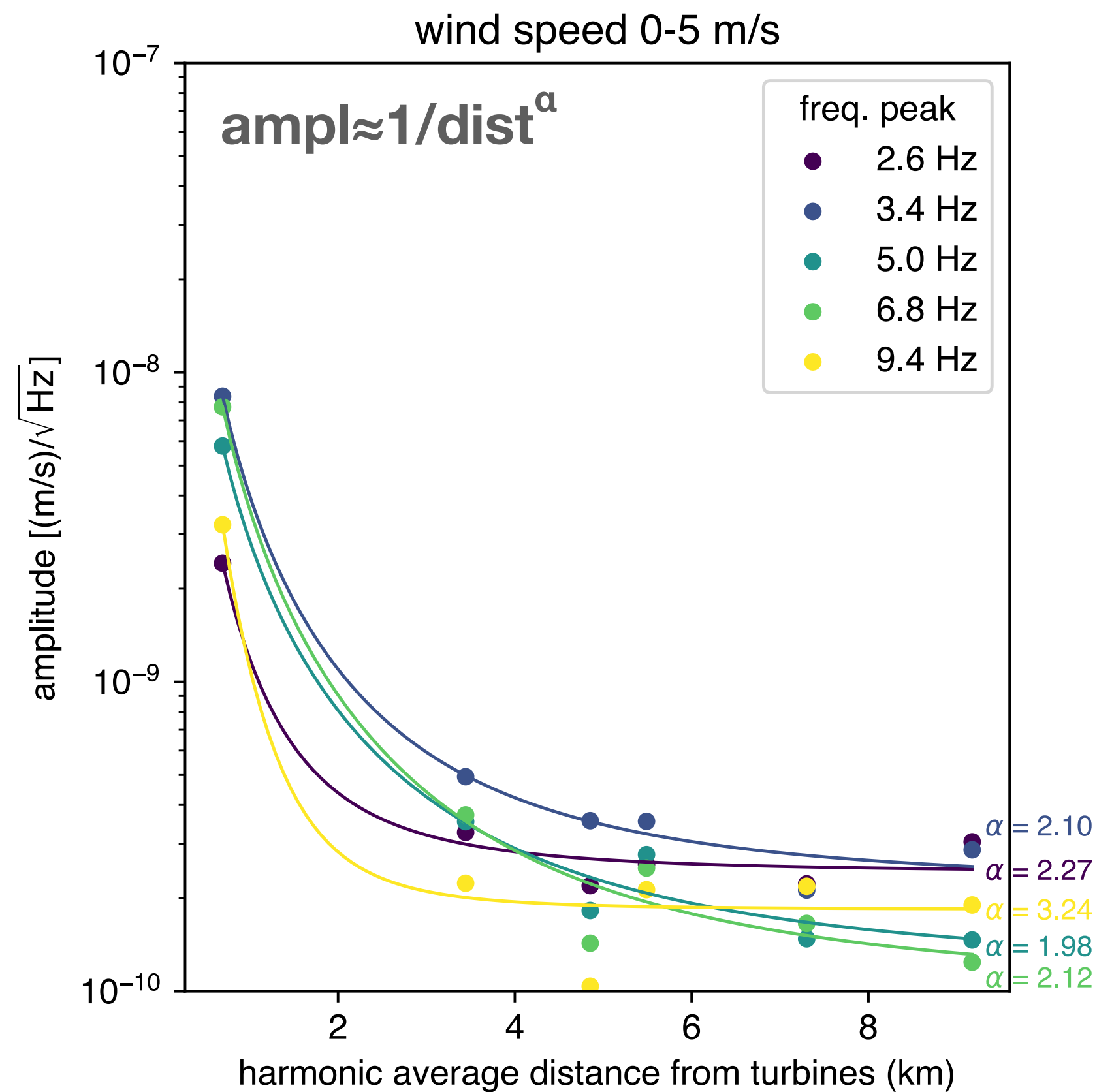
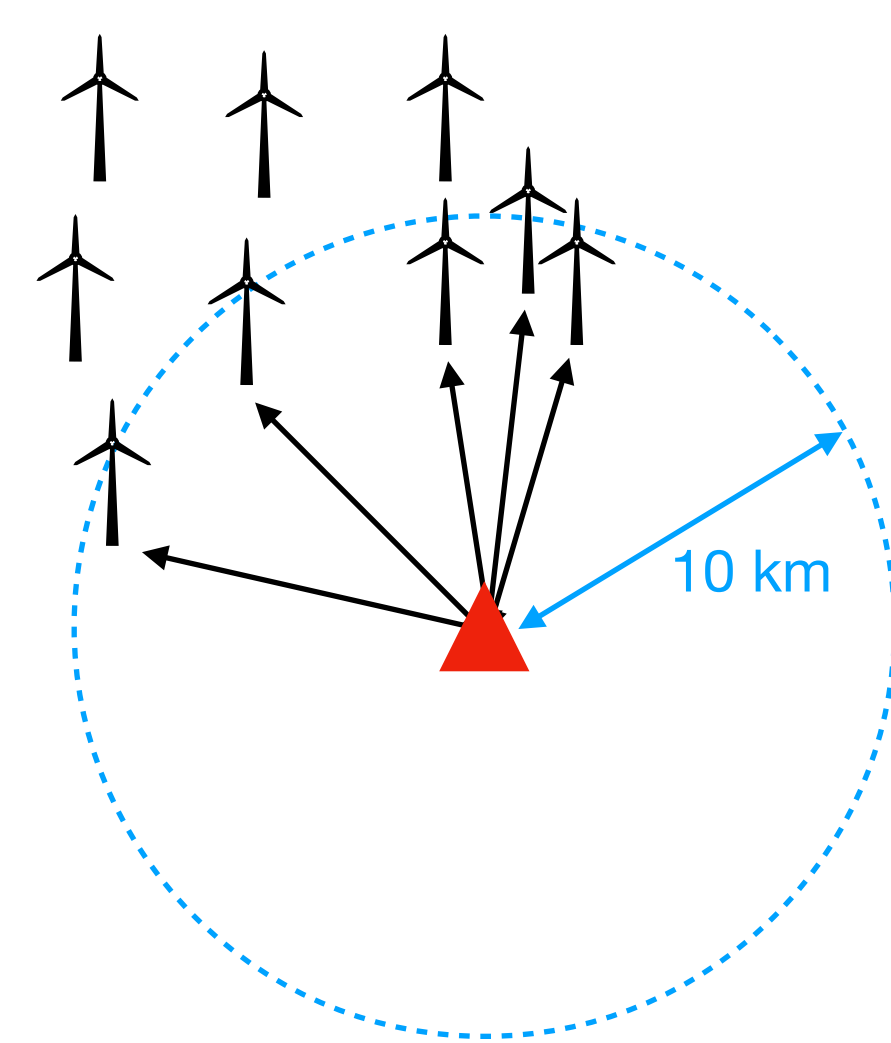
# Amplitude decays vs. distance

- for each station, we take the **number (N) of turbines within 10 km**
- calculate **harmonic mean distance** of all **N stations**
- **divide** spectral amplitude by  **$N^{1/2}$**



# Amplitude decays vs. distance

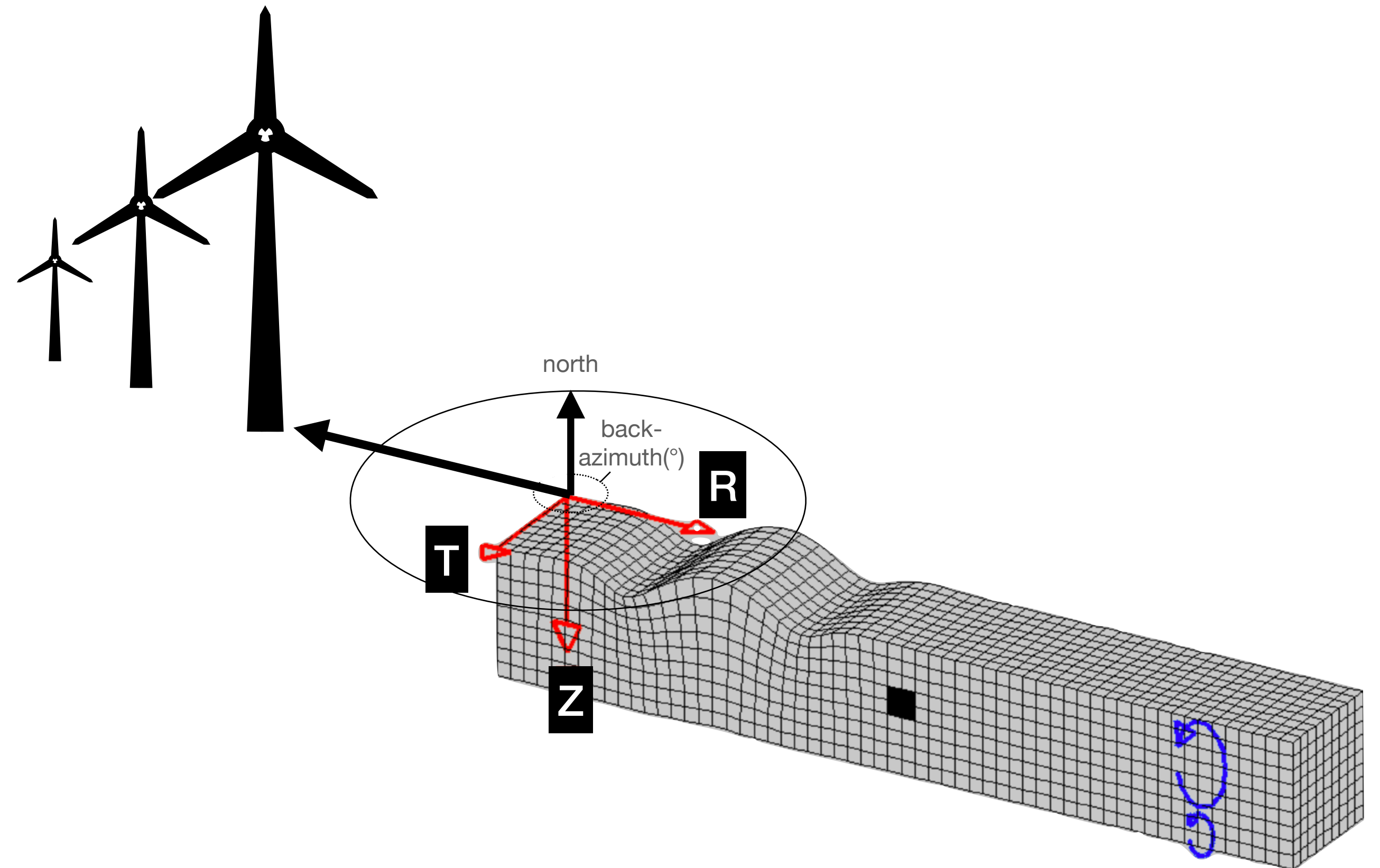
- for each station, we take the **number (N) of turbines within 10 km**
- calculate **harmonic mean distance** of all N stations
- **divide** spectral amplitude by  **$N^{1/2}$**



- clear **exponential decay**
- **stronger decay** for **higher frequencies**
- **~1 order of magnitude** decay in **10 km**
- **strong wind condition** increase noise **by ~1 order of magnitude**

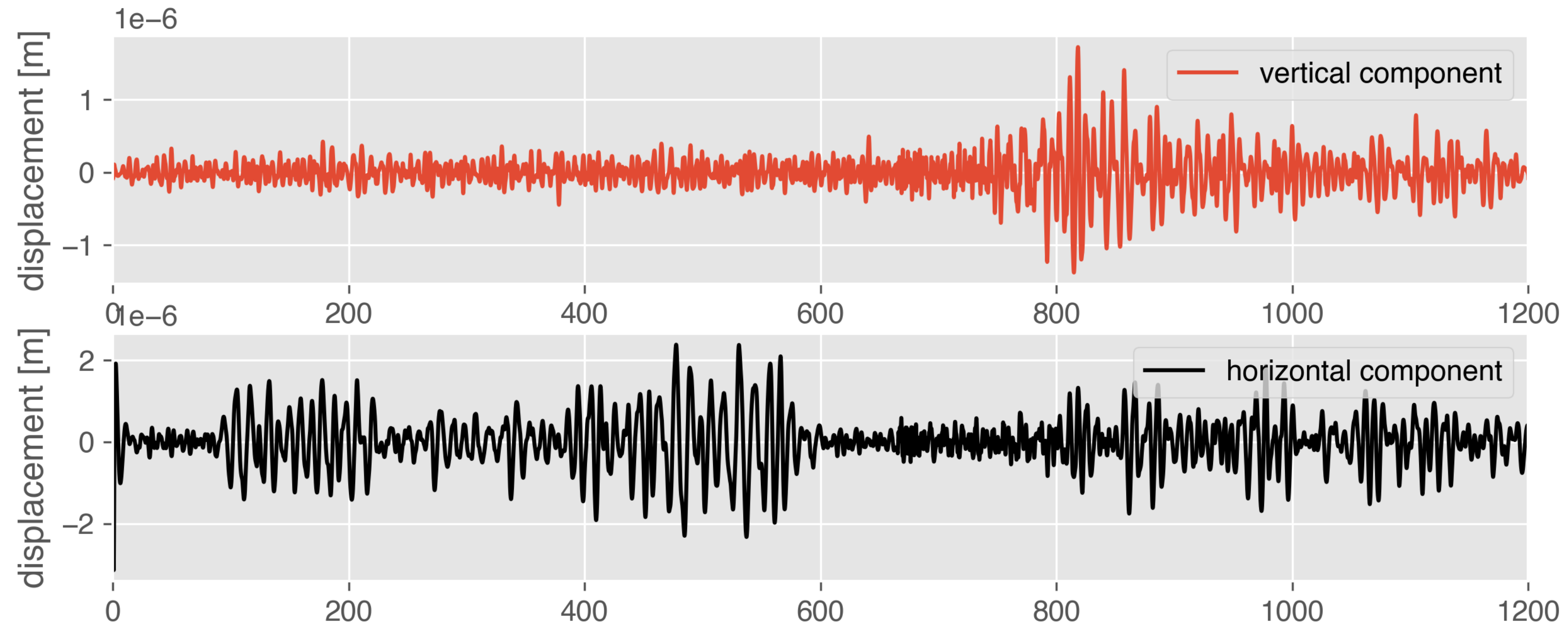
# Analysis of noise directionality

- For **Rayleigh waves**, **vertical** and **radial** components are (ideally) **90° phase shifted**
- The **back-azimuth** of a seismic source can be obtained by:
  - rotating the **N-** and **E-** components to **R-** and **T-** components, for various back-azimuth angle (0-360°).
  - Find the angle that **maximize the cross-correlation between the Z-component and the Hilbert-transformed R-components** (to compensate for the 90° phase shift).
  - **Cross-correlation is maximum** when the **rotation angle is equal to the direction of the incoming Rayleigh wave**



# Analysis of noise directionality

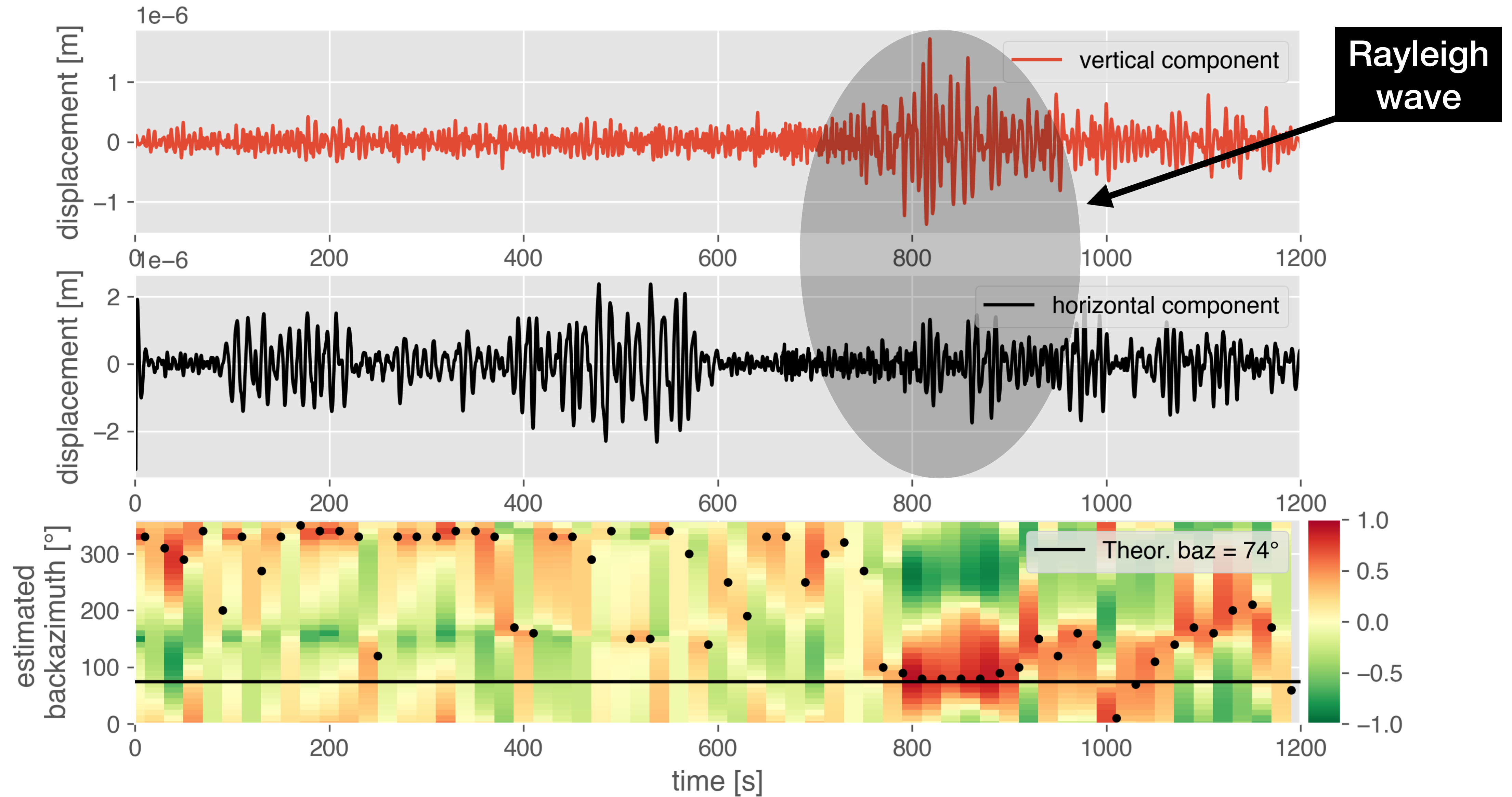
Application with a known source:  
**M4.6 earthquake** occurred on **28-03-2023**, with a **back-azimuth of 74°**





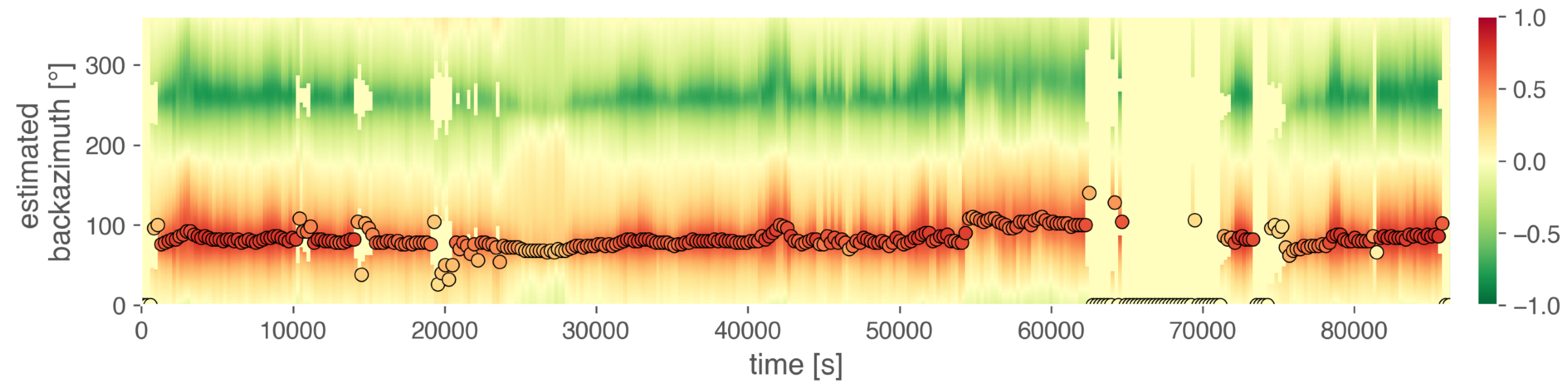
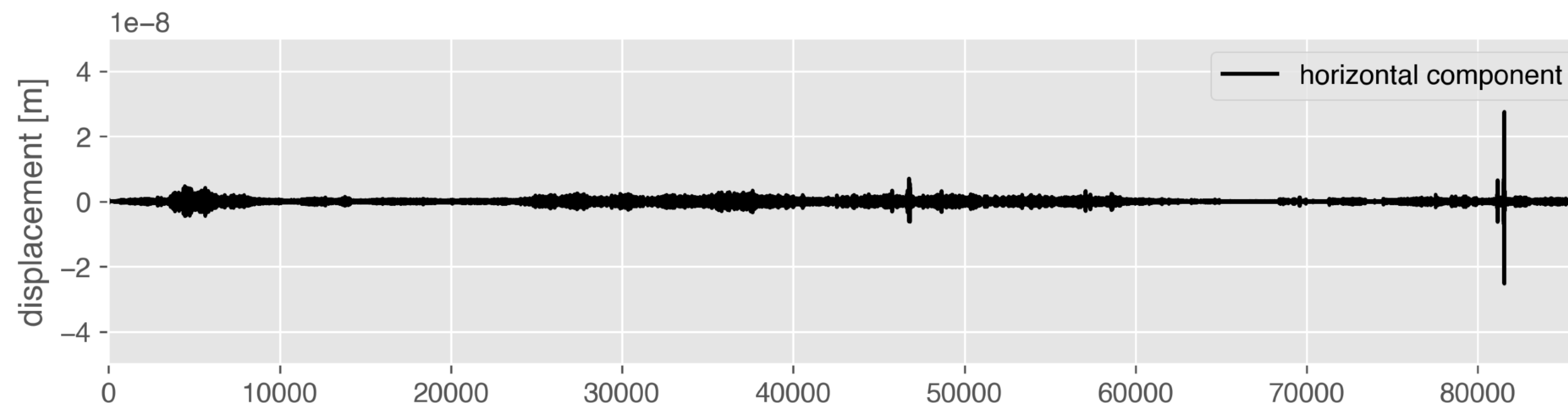
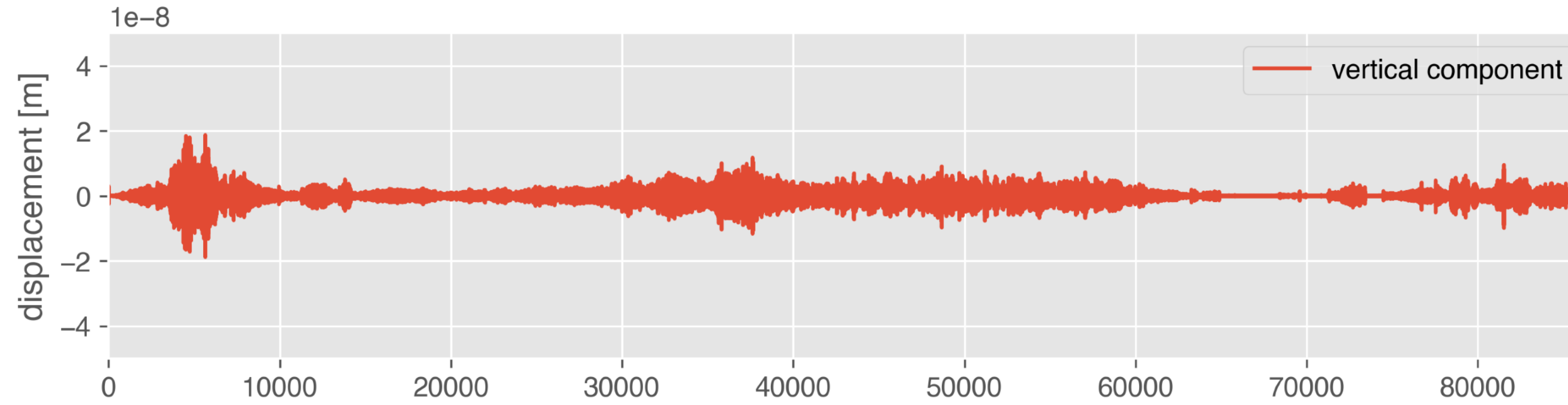
# Analysis of noise directionality: a test

Application with a known source:  
**M4.6 earthquake** occurred on **28-03-2023**, with a **back-azimuth of 74°**



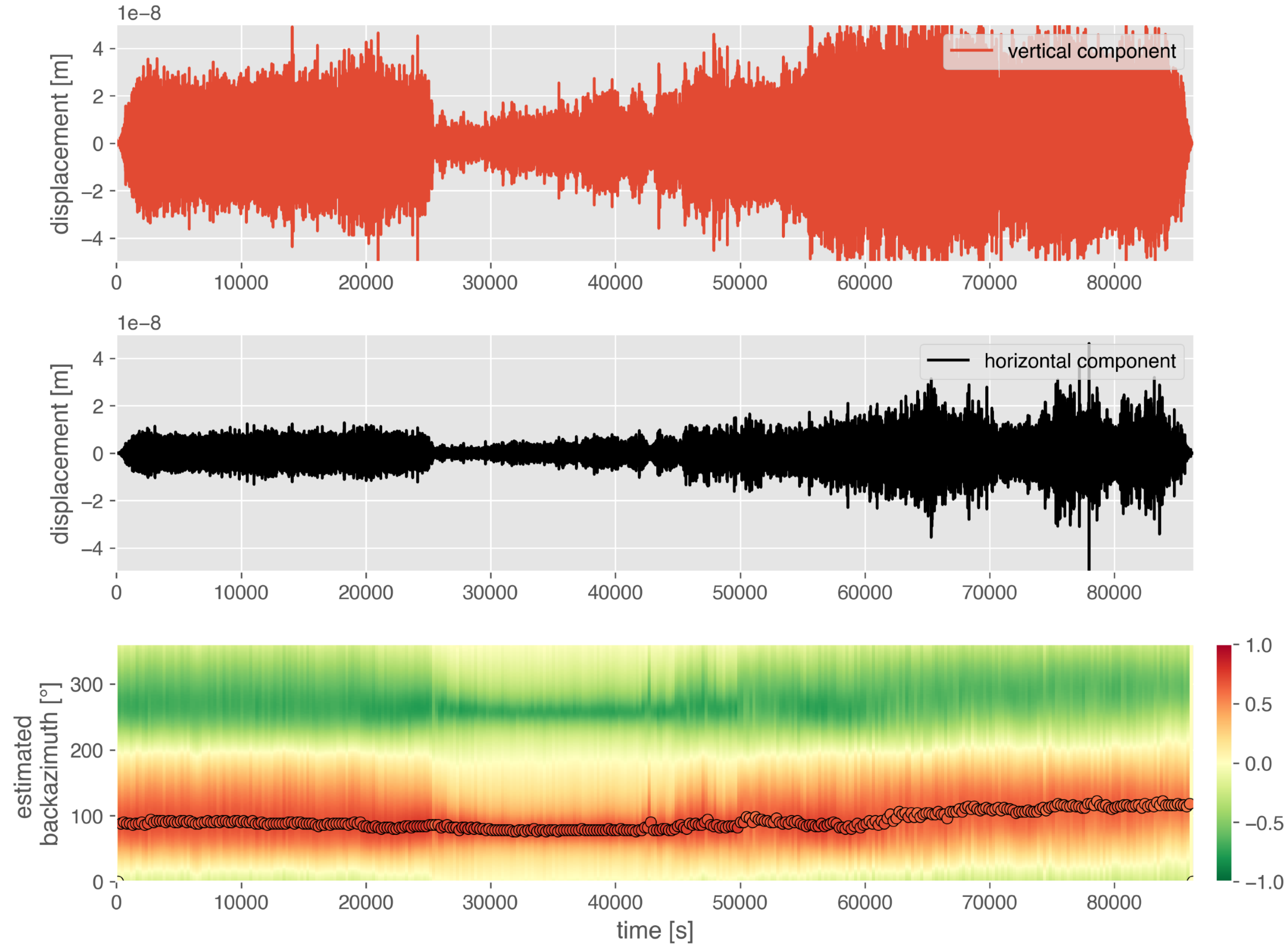
# Analysis of noise directionality

station: **WP1**  
duration: **24 hrs**  
wind: **almost absent**



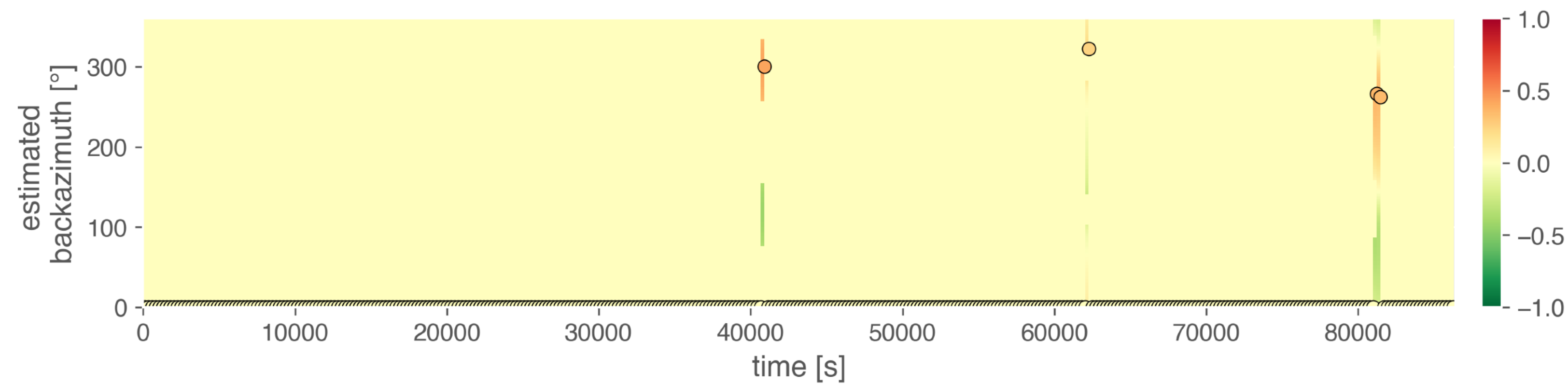
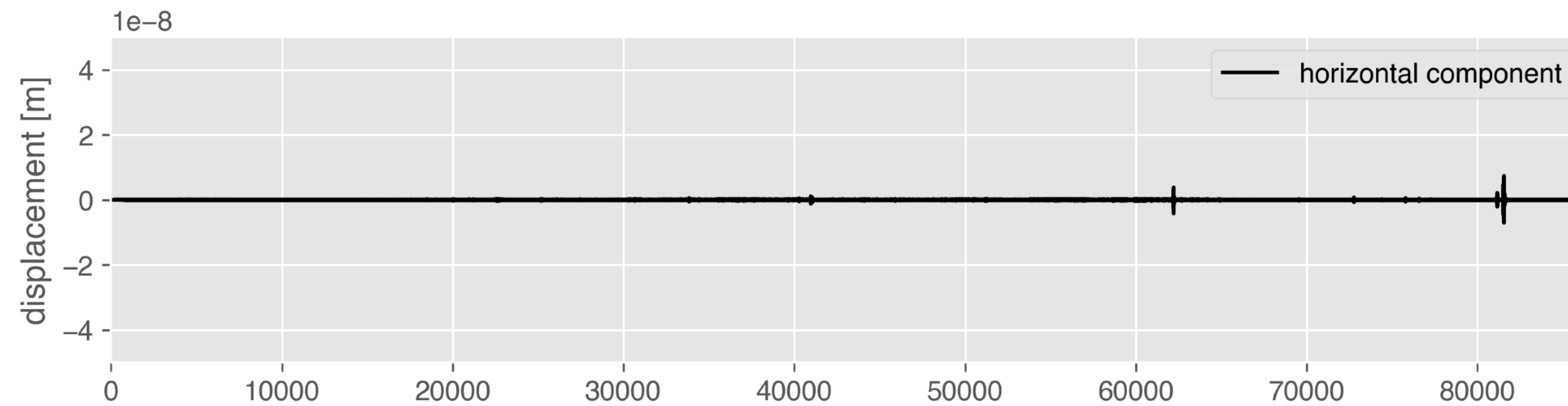
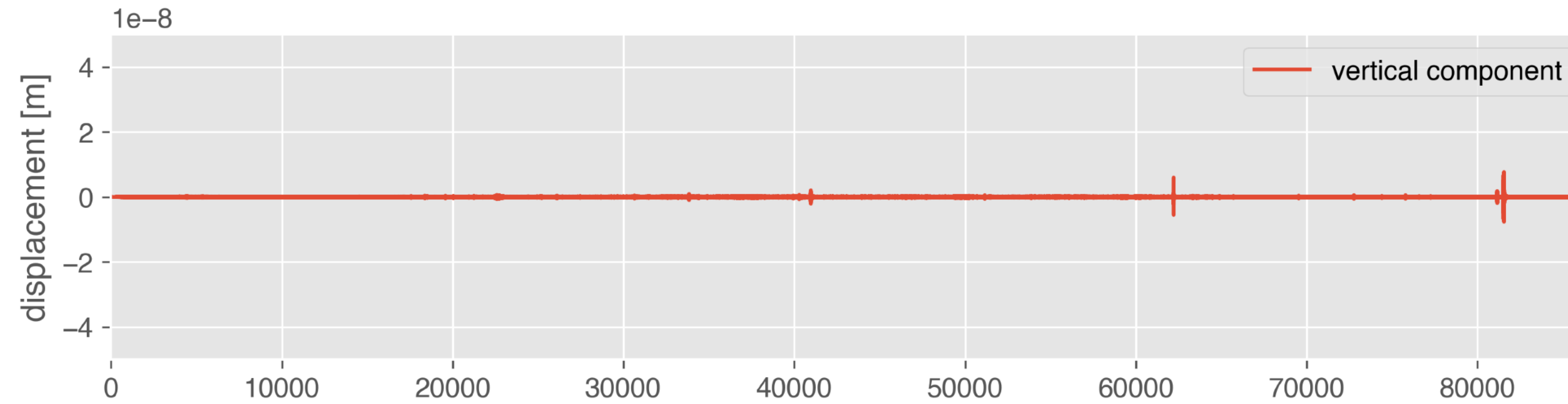
# Analysis of noise directionality

station: **WP1**  
duration: **24 hrs**  
wind: **strong (>20 m/s)**



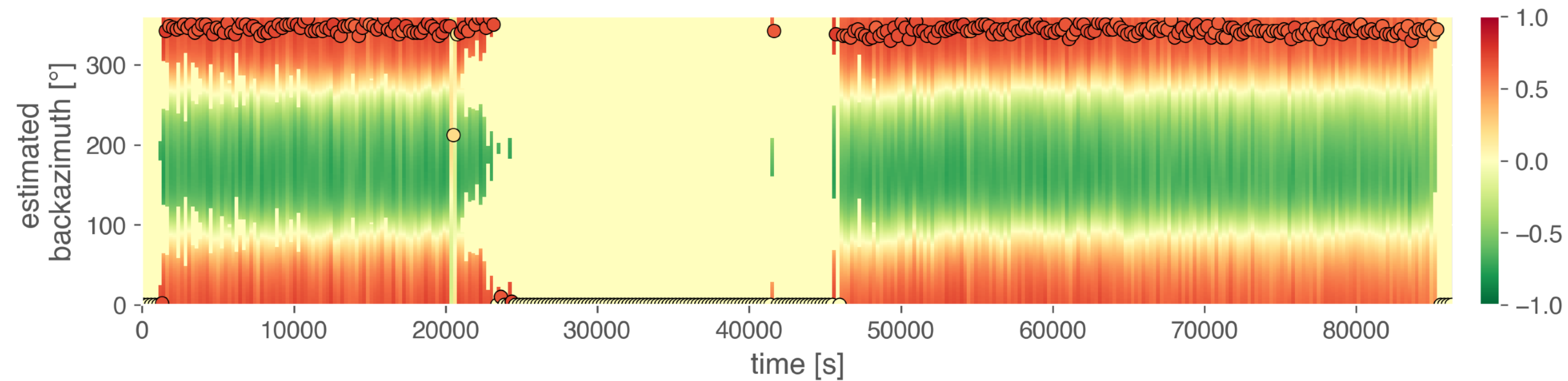
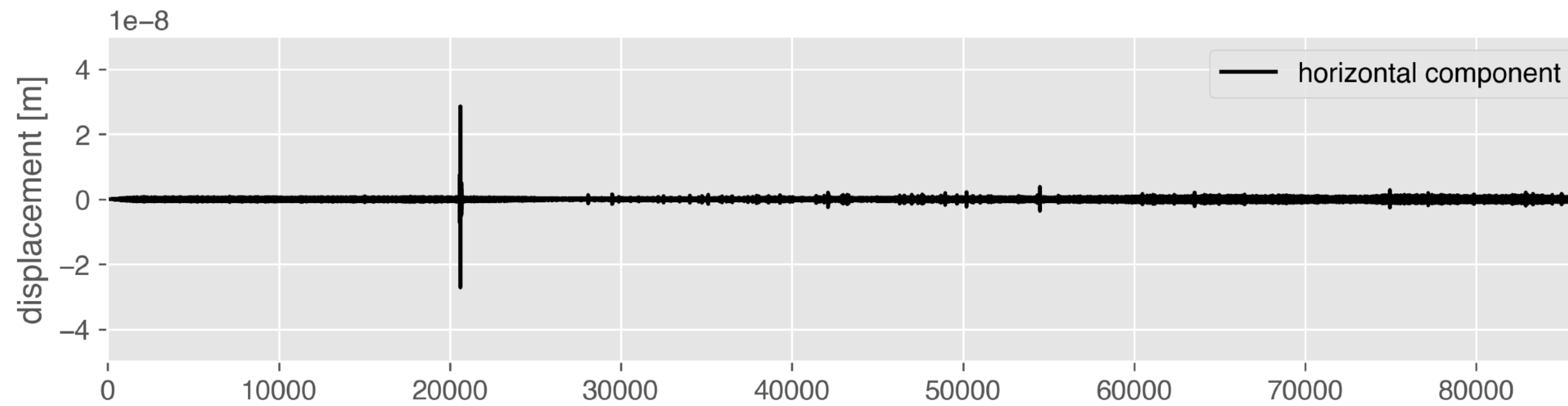
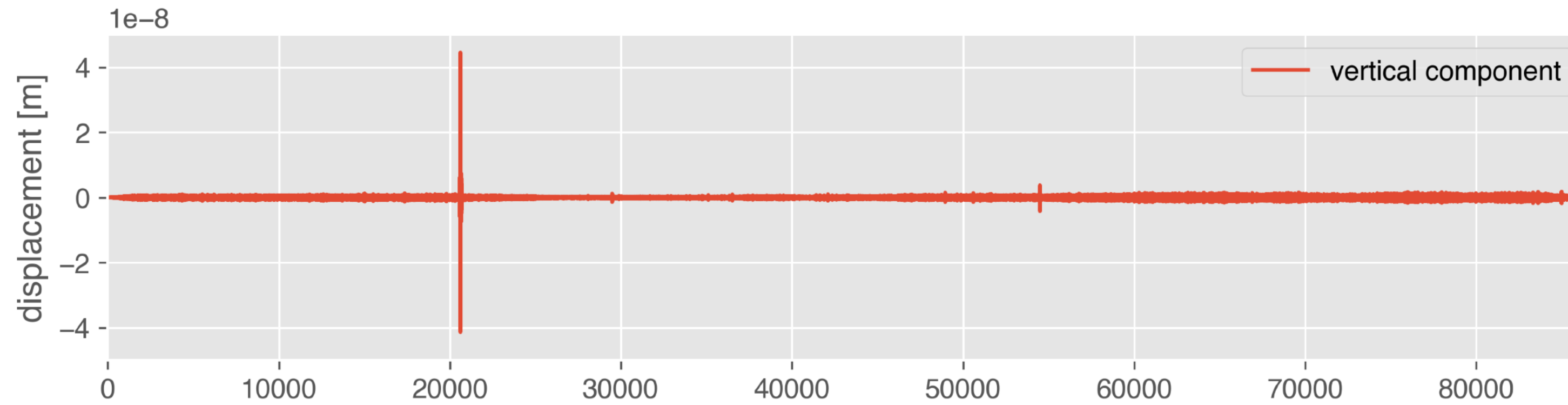
# Analysis of noise directionality

station: **WP3**  
duration: **24 hrs**  
wind: **almost absent**

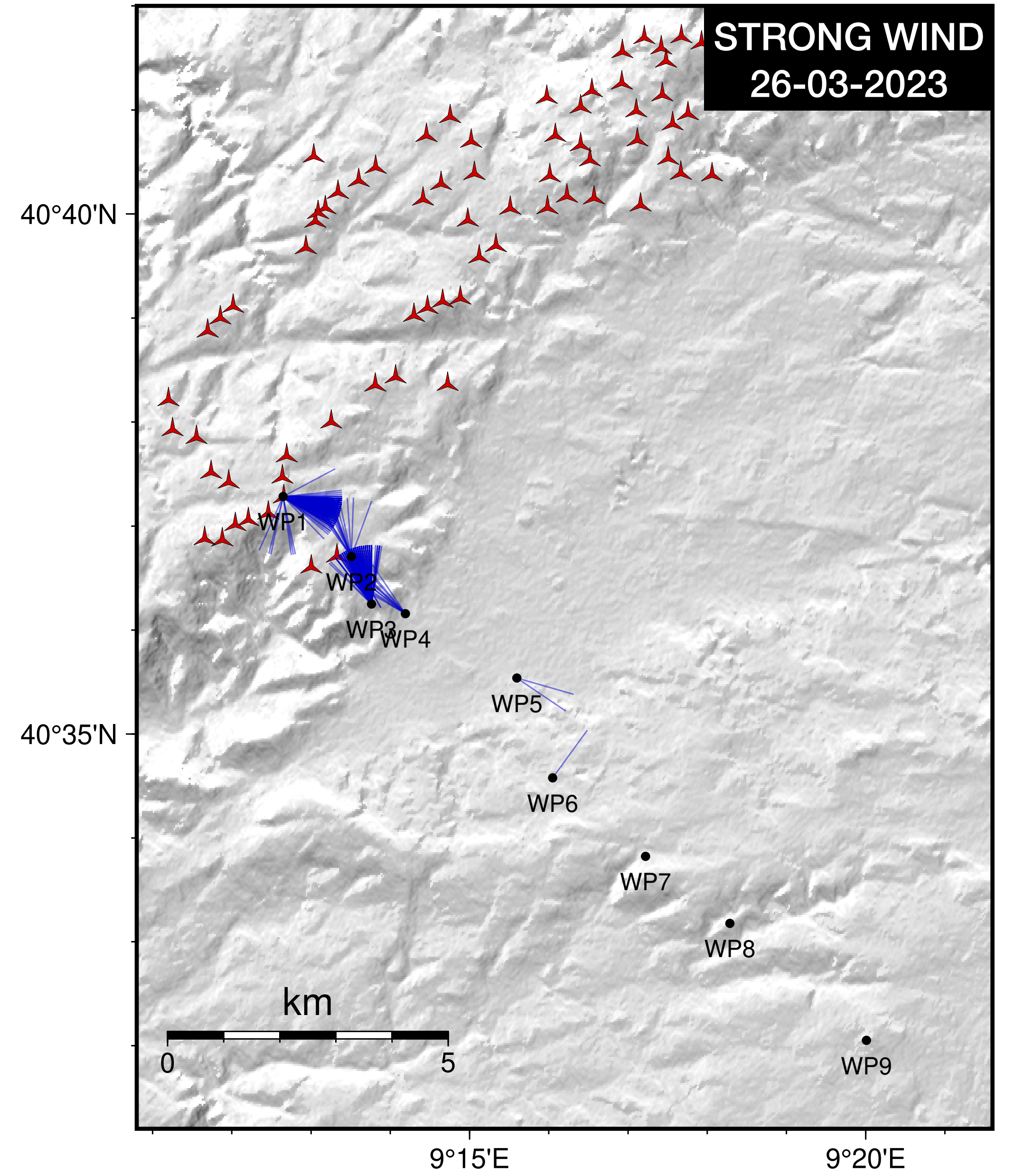
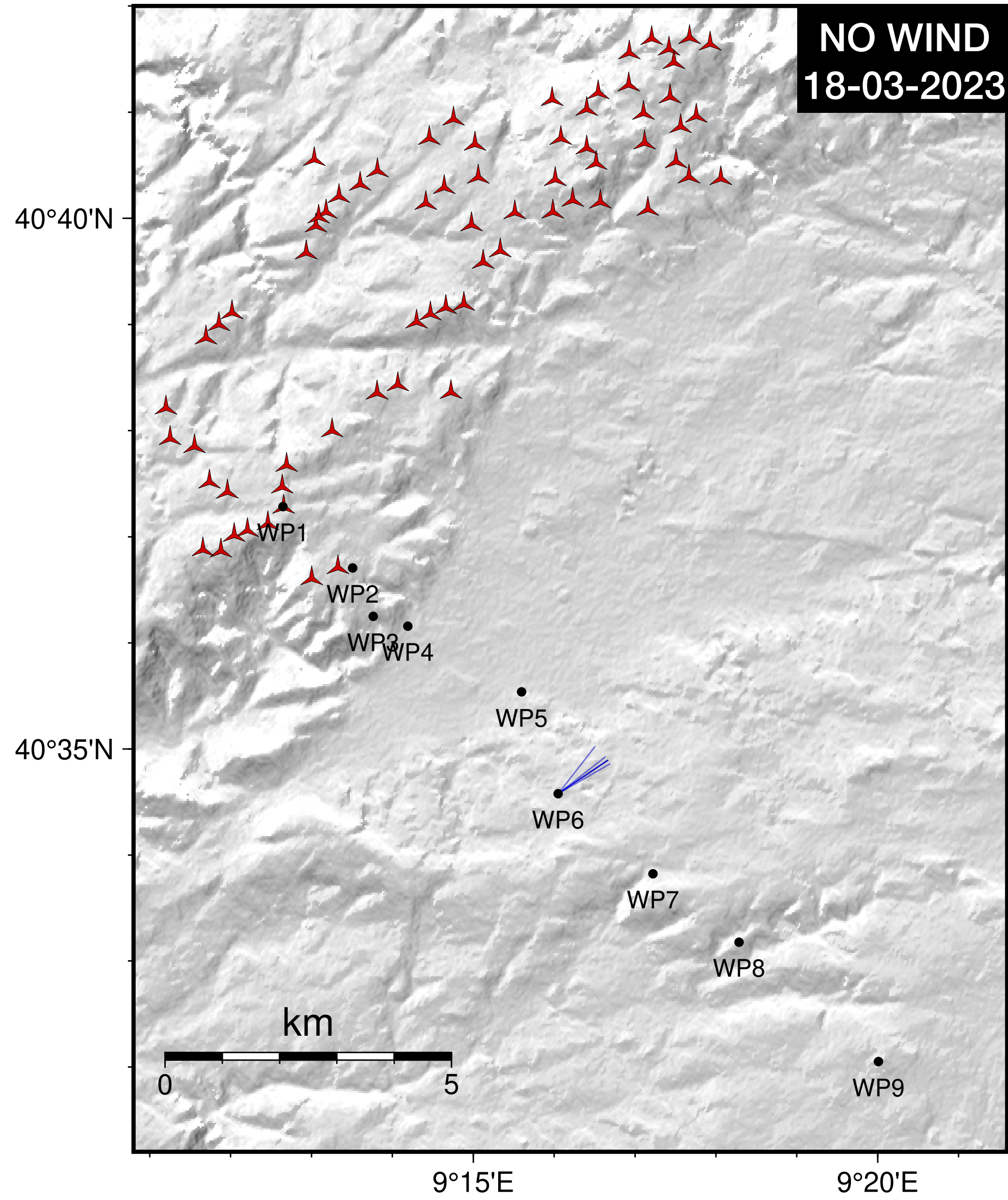


# Analysis of noise directionality

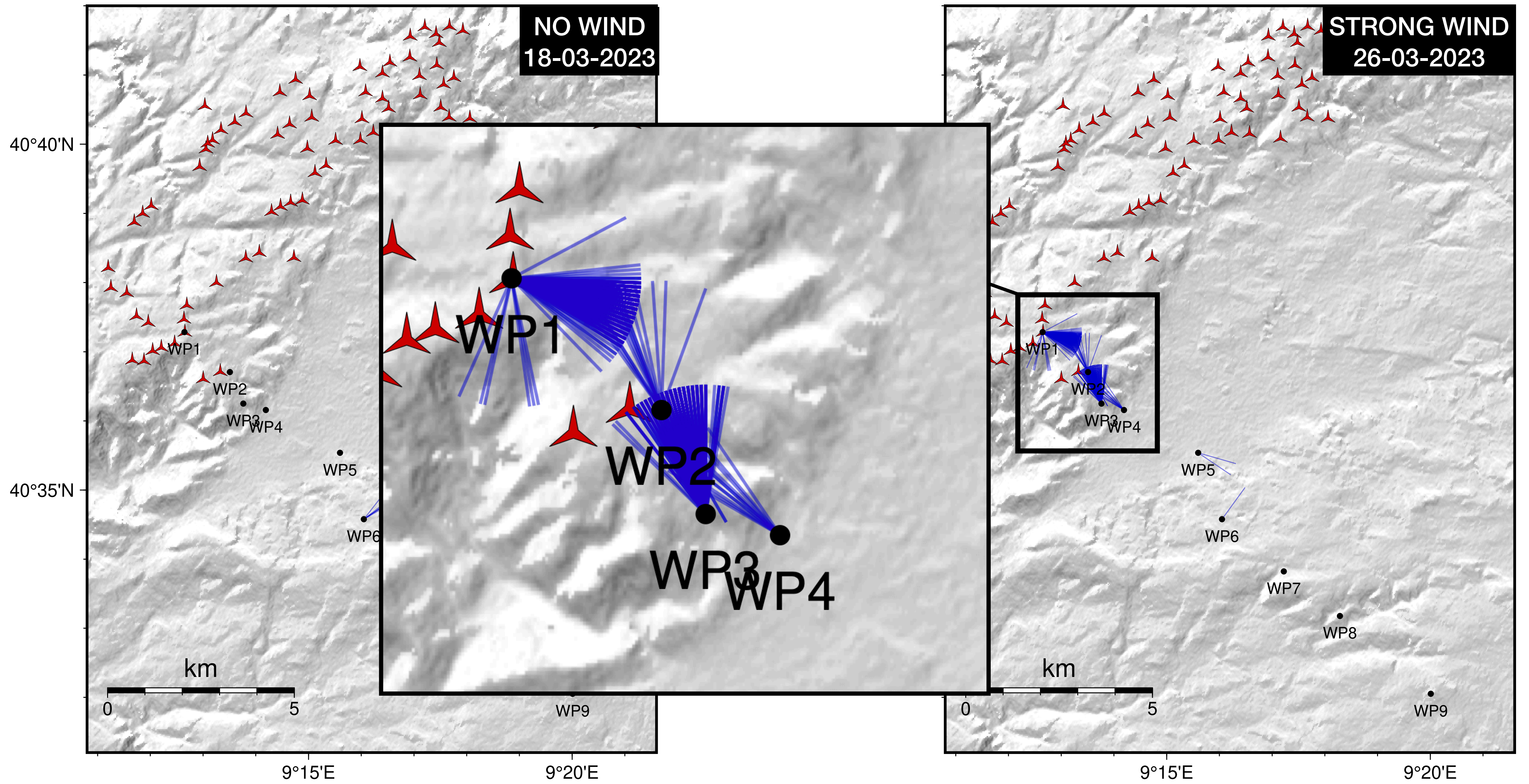
station: **WP3**  
duration: **24 hrs**  
wind: **strong (>20 m/s)**



# Analysis of noise directionality



# Analysis of noise directionality



# Conclusions and future directions

- As expected, **Buddusò wind park** generates **substantial seismic noise** with distinct frequency peaks (2.6, 3.4, 5.0, 6.8, 9.4 Hz)
- In case of **strong wind (>20 m/s)**, noise amplitude **increases** by roughly **one order of magnitude** w.r.t **weak or no wind (0-5 m/s)**
- Seismic amplitude is **exponentially damped** with distance, **decreasing by an order of magnitude in about 10 km.**
- **Boreholes stations (P2, P3 at ~250 m depth):** with strong wind **spectral peaks are visible. Influence from wind park or effect of wind at local scale?**
- Analysis of **noise polarization** works up to **5 km from windpark**, showing a direction that is **fully compatible** with the position of the wind park w.r.t. the deployed array.
- The **analysis will be improved** by using the **full dataset** (RPM, power throughput, wind speed at each turbine) just released by the wind turbine operator.