



Nikhef



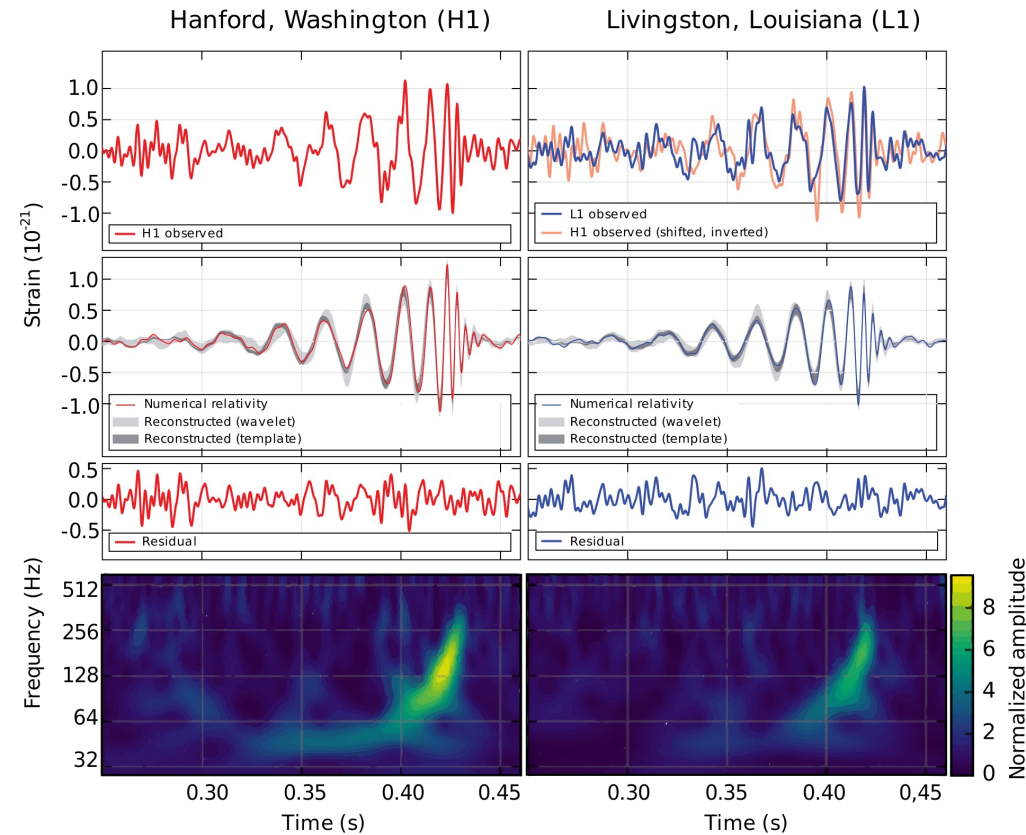
# Detection of anomalies amongst LIGO's glitch populations with autoencoders

Melissa Lopez

[m.lopez@uu.nl](mailto:m.lopez@uu.nl)

ArXiv: 2310.03453

# How are gravitational waves detected?

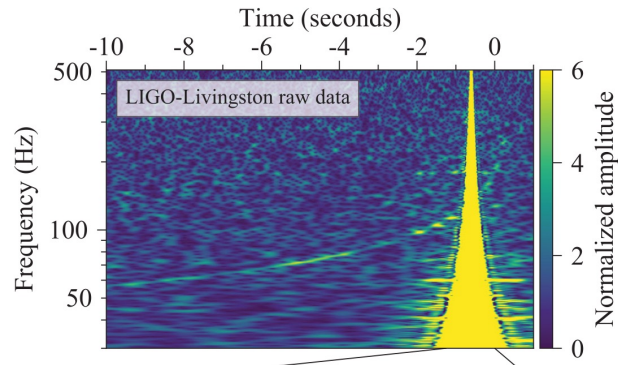


# Transient noise a.k.a. glitches in LIGO

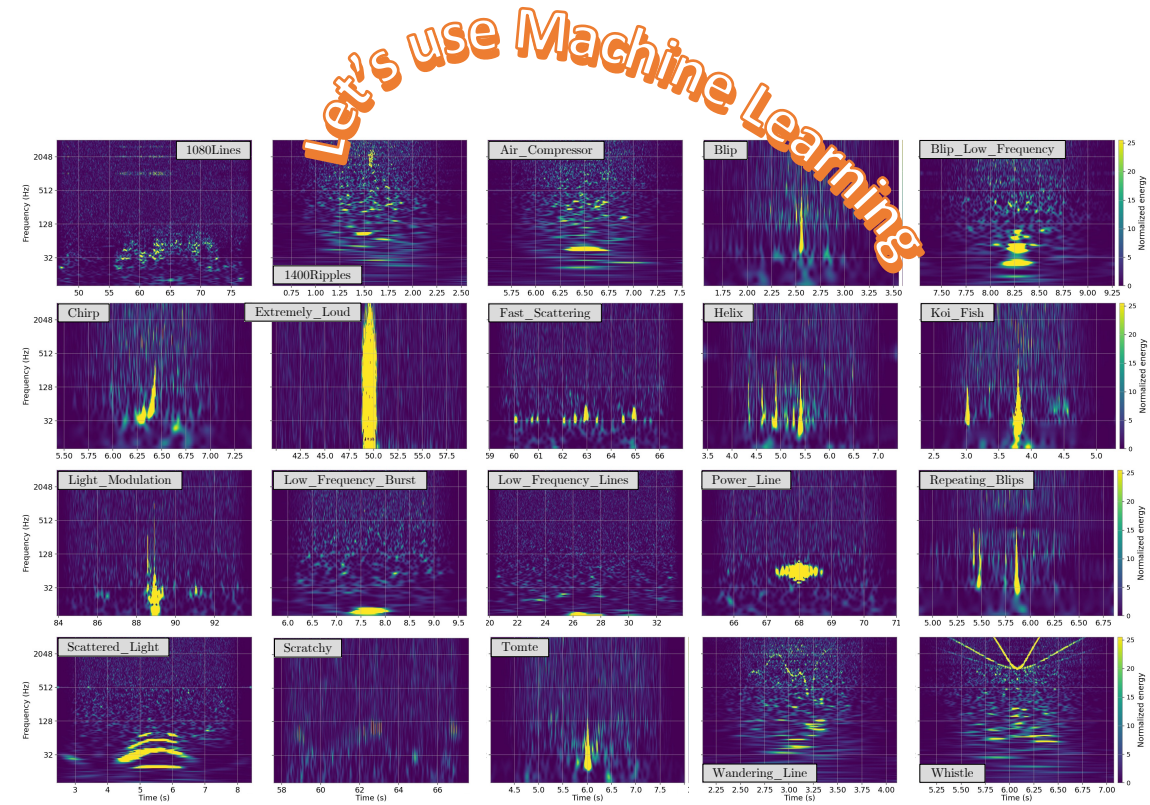
- Caused by instruments or environment (known or unknown)
- Diminish scientific data available
- Hinder GW detection (mask and/or mimic)
- Present in LIGO, Virgo and probably Einstein Telescope!

**Idea:** we need to mitigate them, so let's identify them first

But... too many glitches!  $\sim 1 \text{ min}^{-1}$  during O2

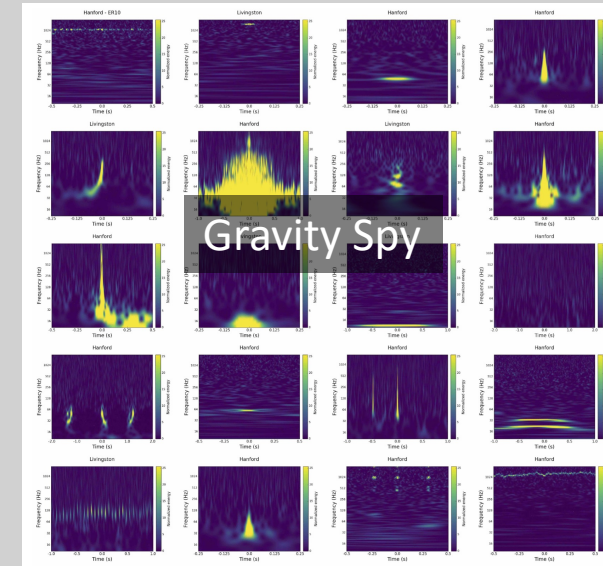


GW masked by glitch (GW170817)



# Machine learning for glitch identification

## Supervised learning: classification



## Challenges

- Representation in the main strain of the detector
- Classes are rigid and labels expensive
- The detector evolves over time

**Idea 1:** we can use information from the detector itself, ie. *auxiliary channels*? →  $\sim 10^6$  channels to process!

**Idea 2:** Let's the data speak for itself  
→ unsupervised learning



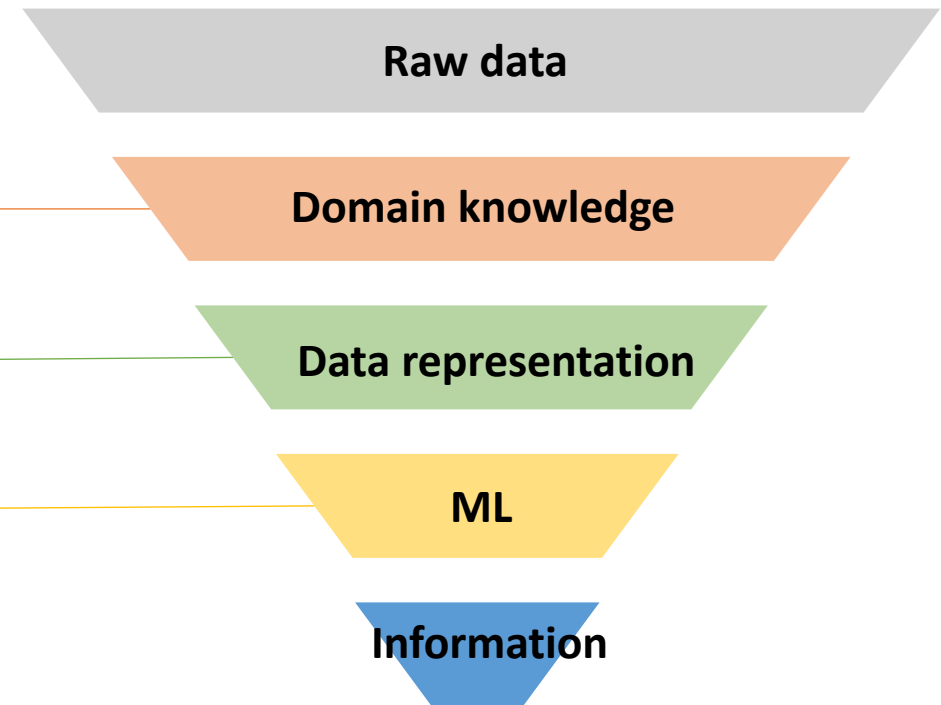
# Encode, encode, encode

How can we reduce a  $10^6$  auxiliary channels (ac)?

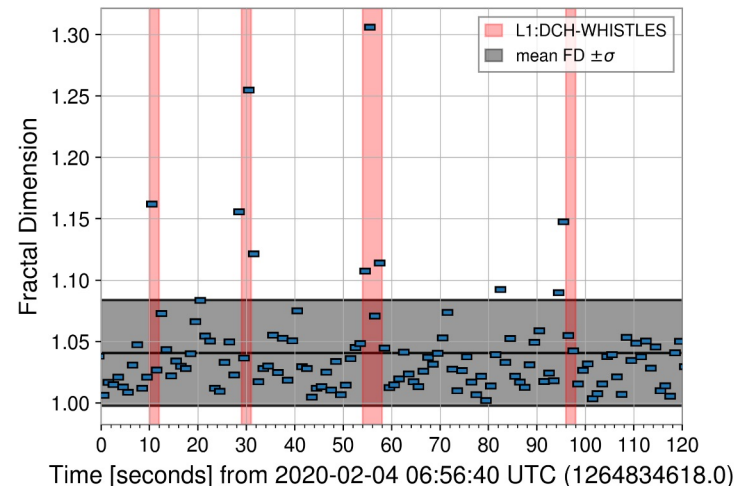
Select safe channels, i.e. not affected by GW (350 ac)

Encode with fractal dimension, i.e. measure complexity of the data

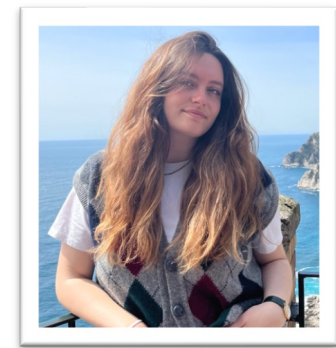
Use convolutional autoencoders



- M. Cavaglia 2022 → 1h of data encoded in 1h
- Our work → 1h of data encoded in 11s



Robin van der Laag (UU)  
Expert in high performance computing



Paloma Laguarda (UM)  
Expert in ML  
PhD at LHCb

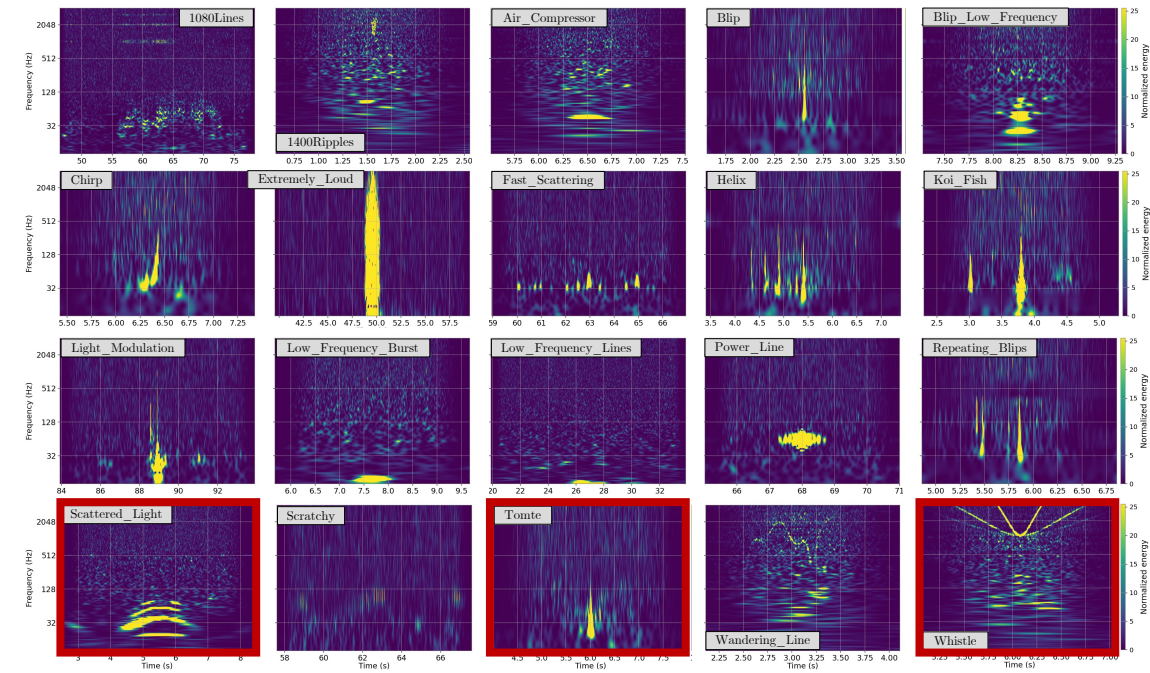
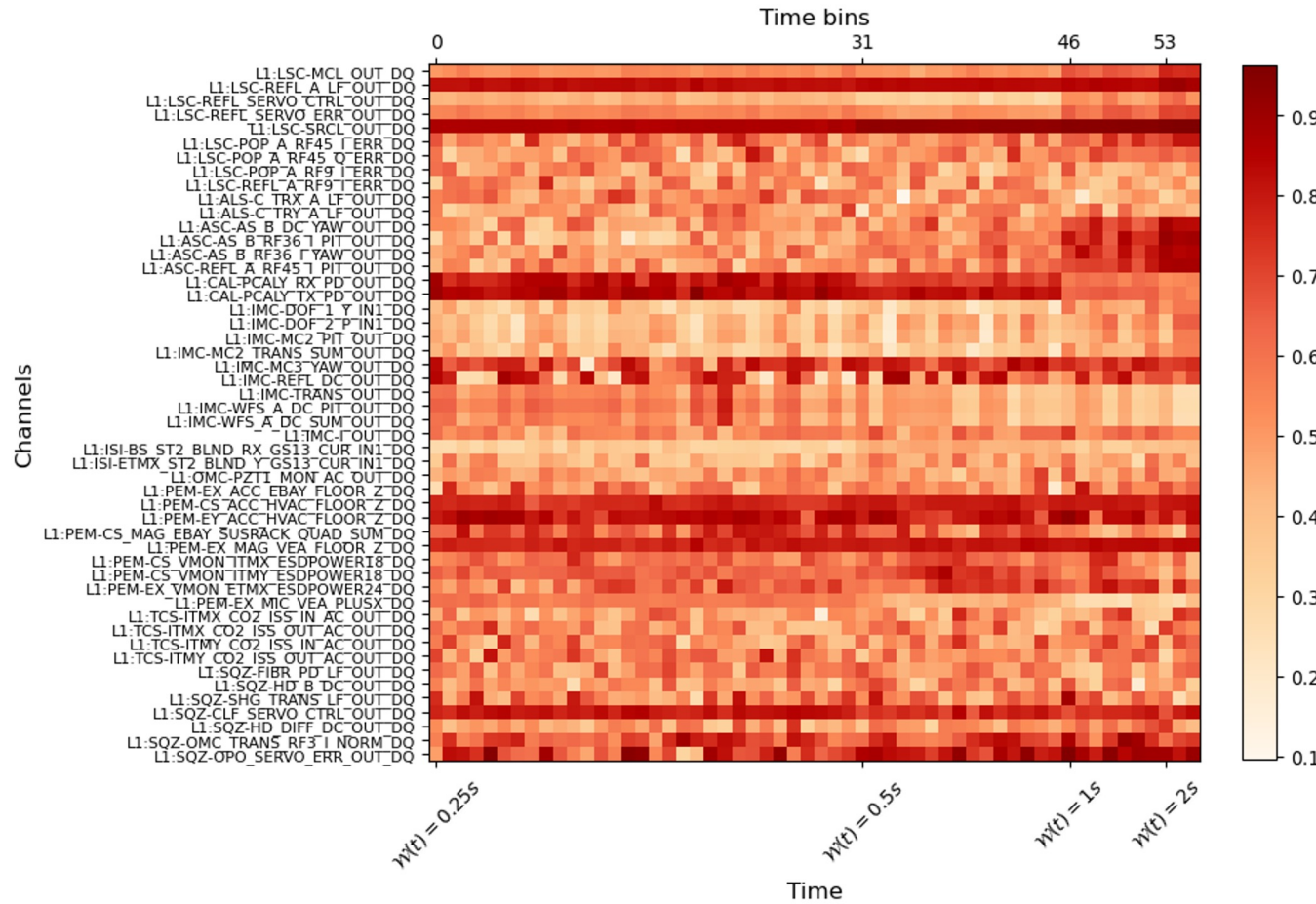
# The input data

We preprocess 350 auxiliary channels  $\rightarrow$  encode 50 auxiliary channels

SCAN ME

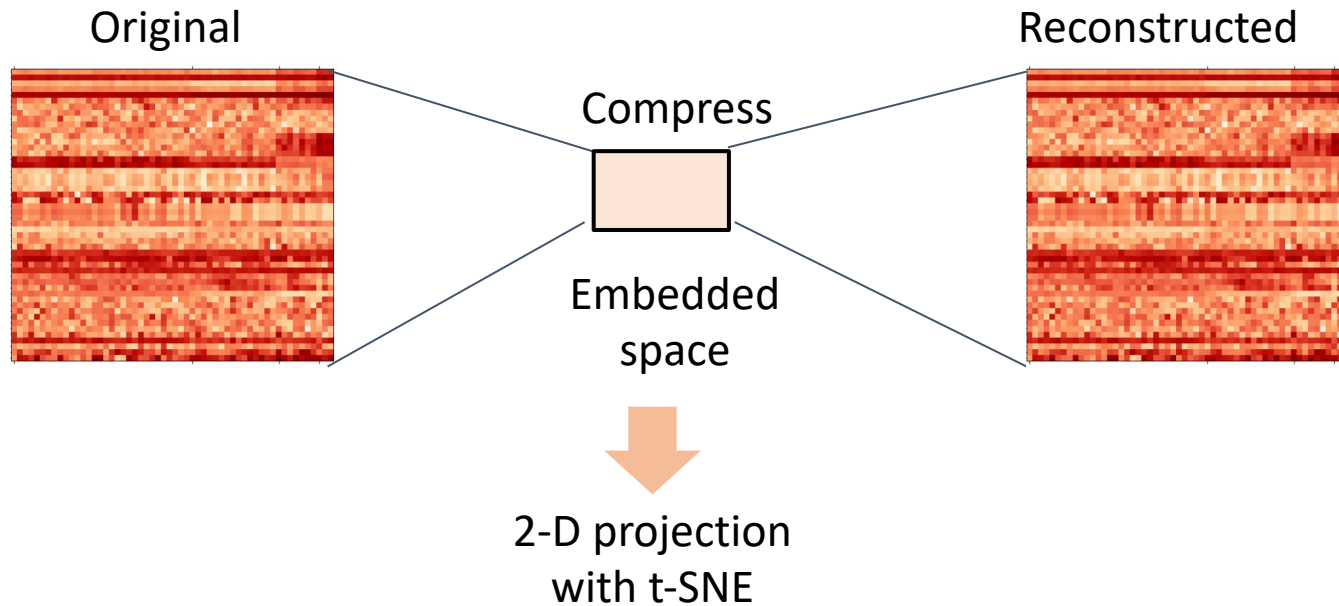


ArXiv: 2310.03453



Classified with supervised learning

# Results of compressed data

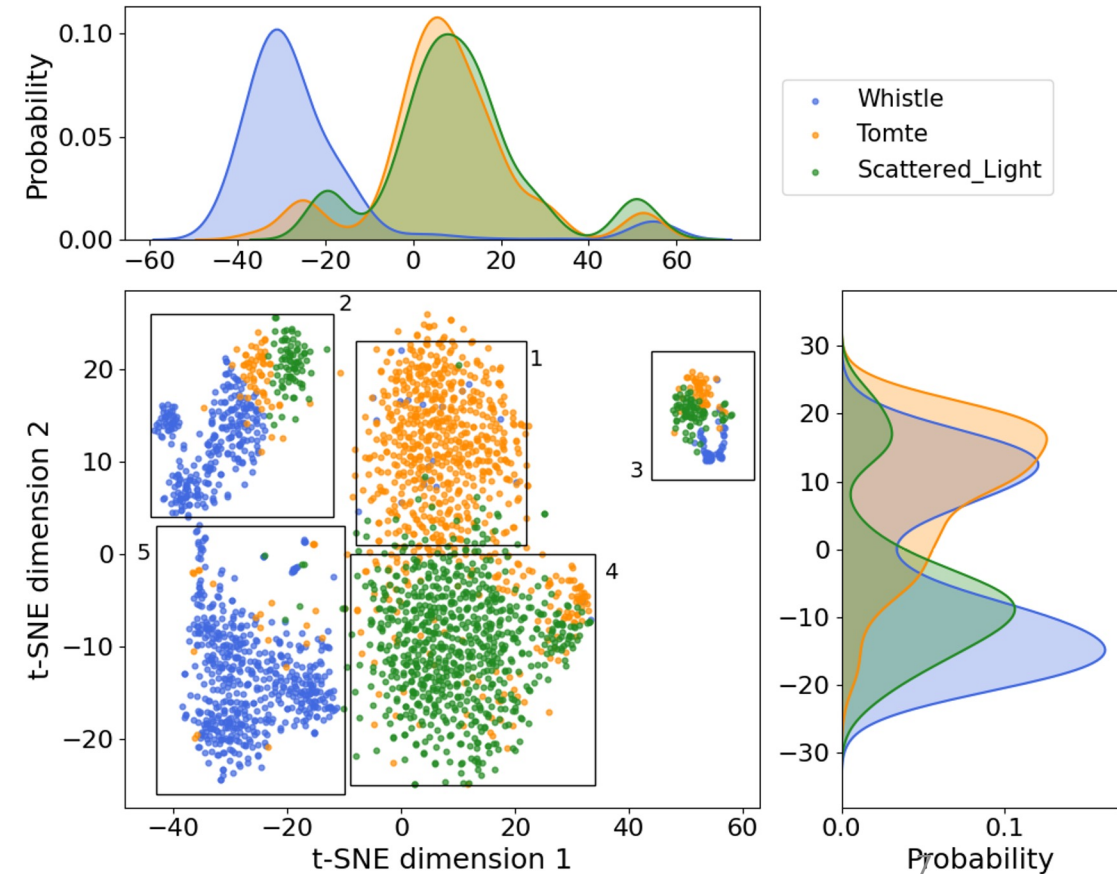


## Benchmarking against supervised learning:

Clusters consistent with Gravity Spy, but

- Gravity Spy → spectrograms of  $h(t)$
- Our work → fractal dimension with auxiliary channels

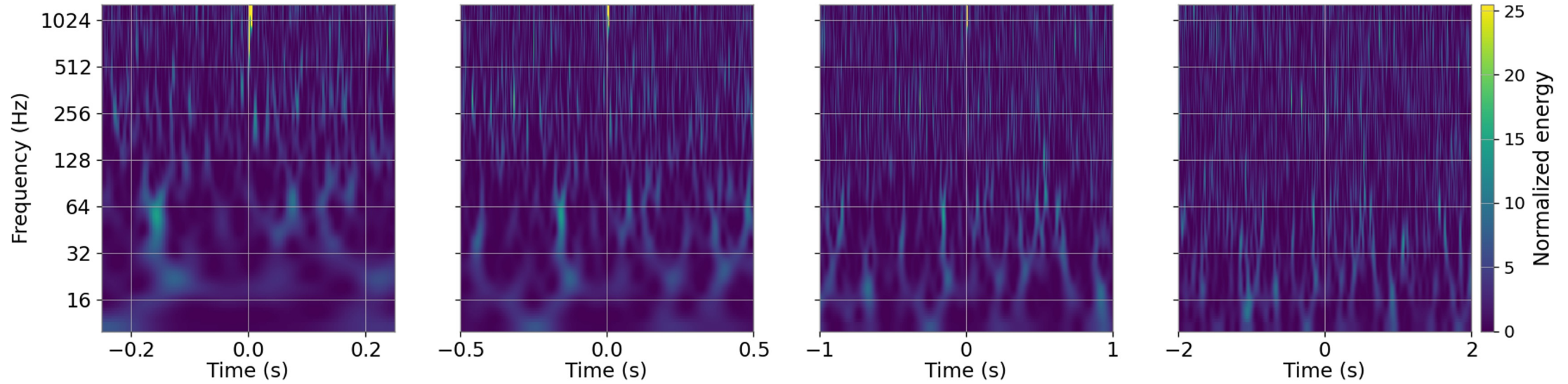
OK, now let's represent anomalies in spectrograms of  $h(t)$



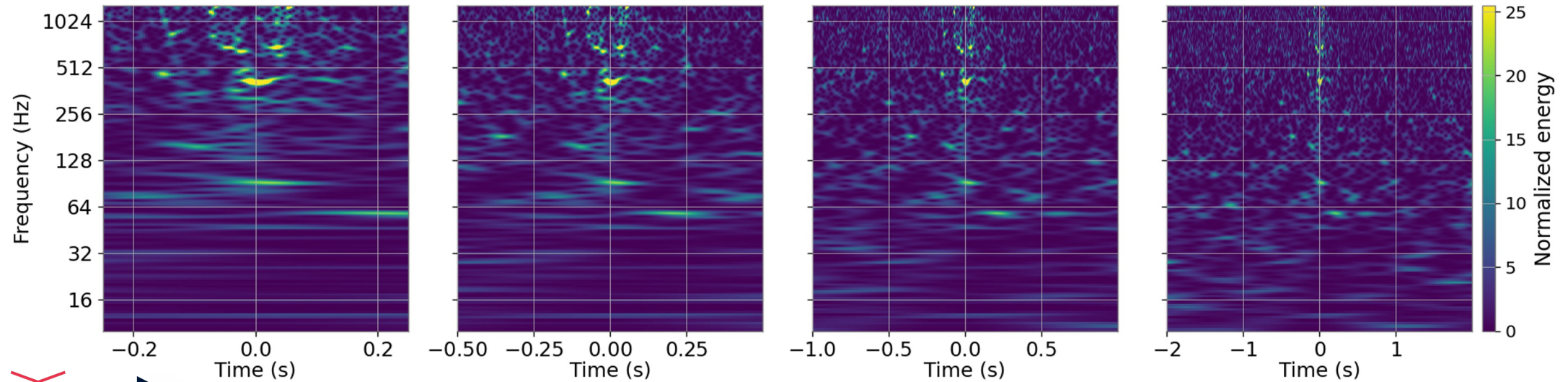


# Section 1

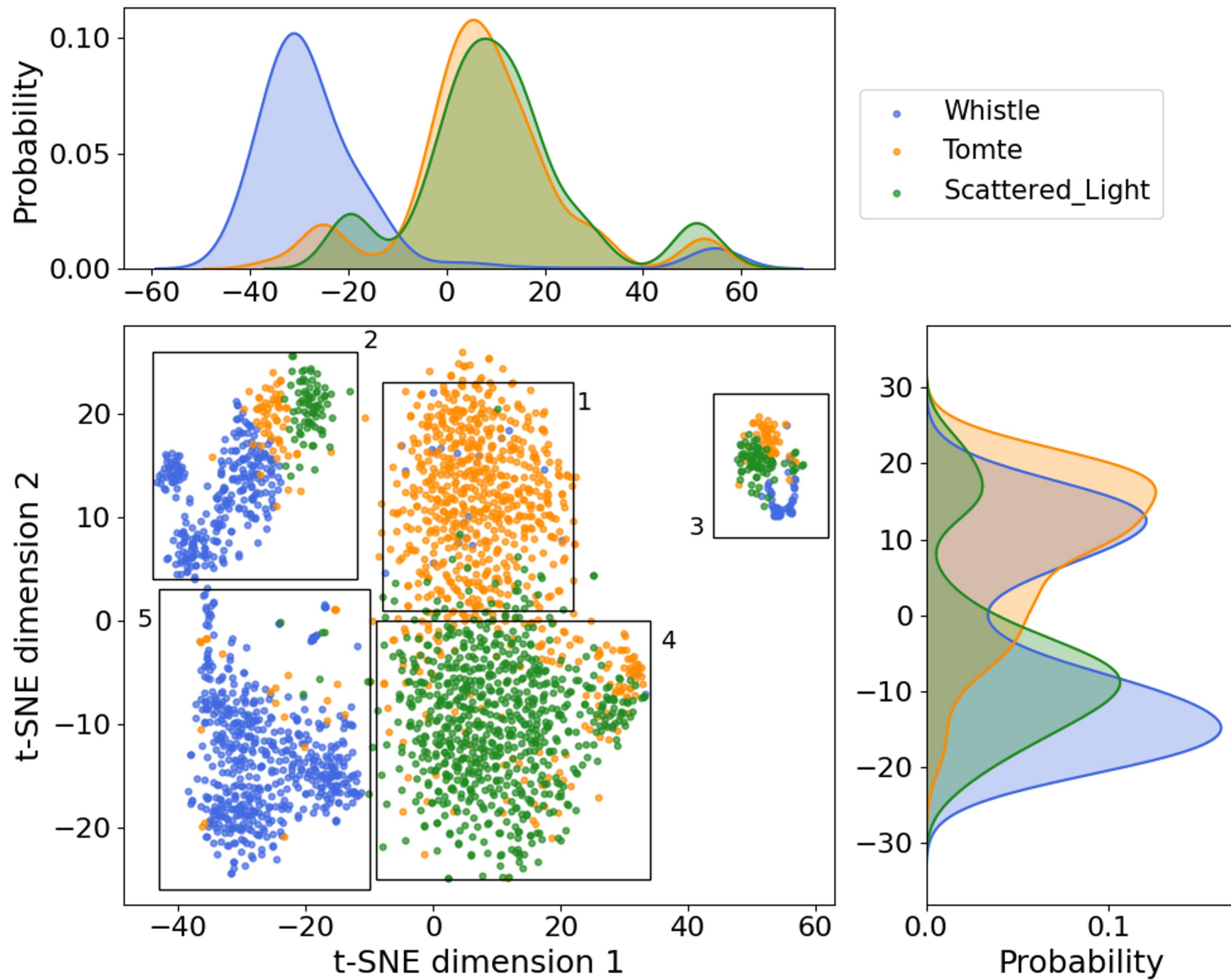
“Standard” Whistle



Labelled as Whistle but anomalous Whistle morphology

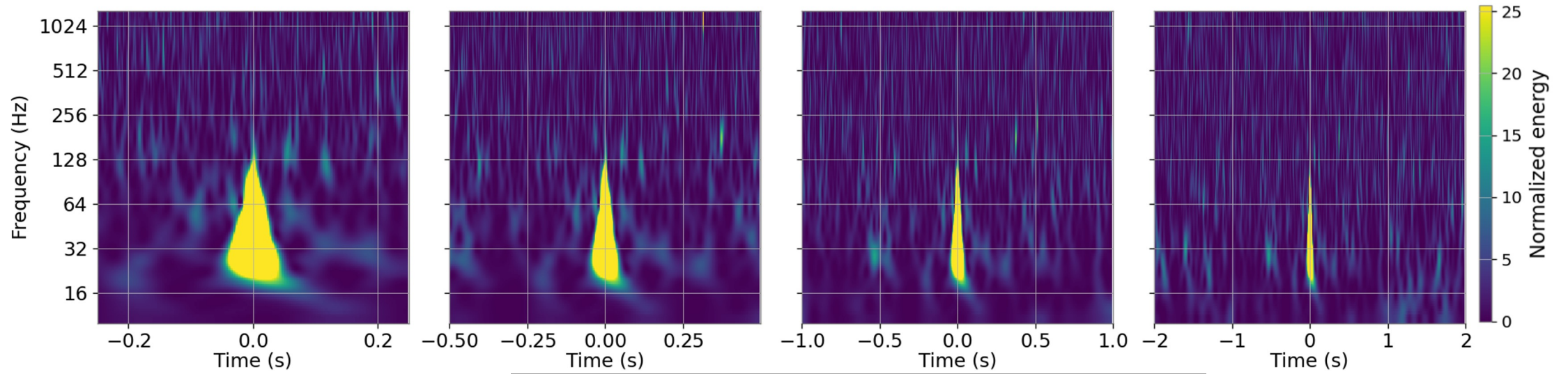




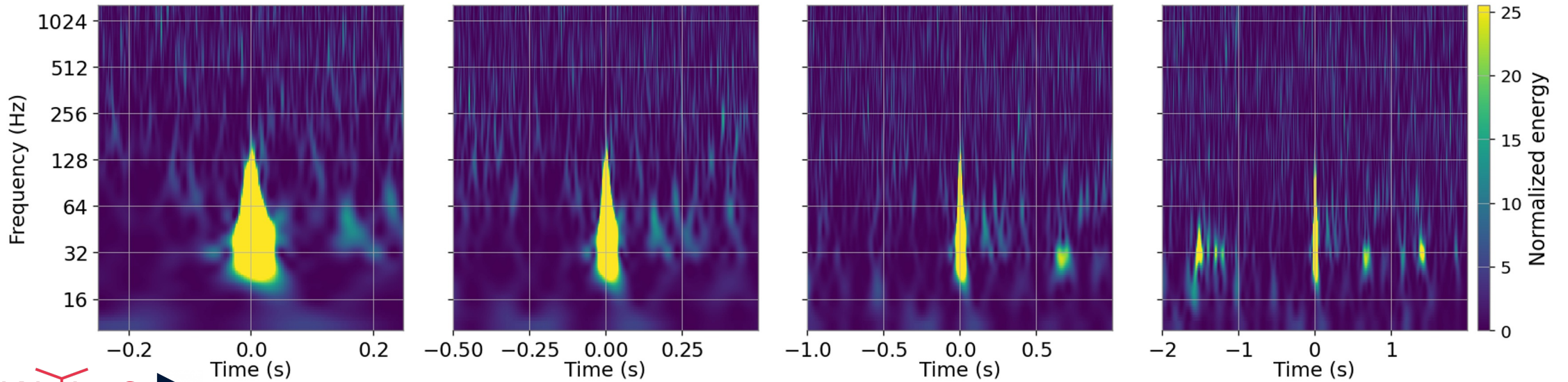


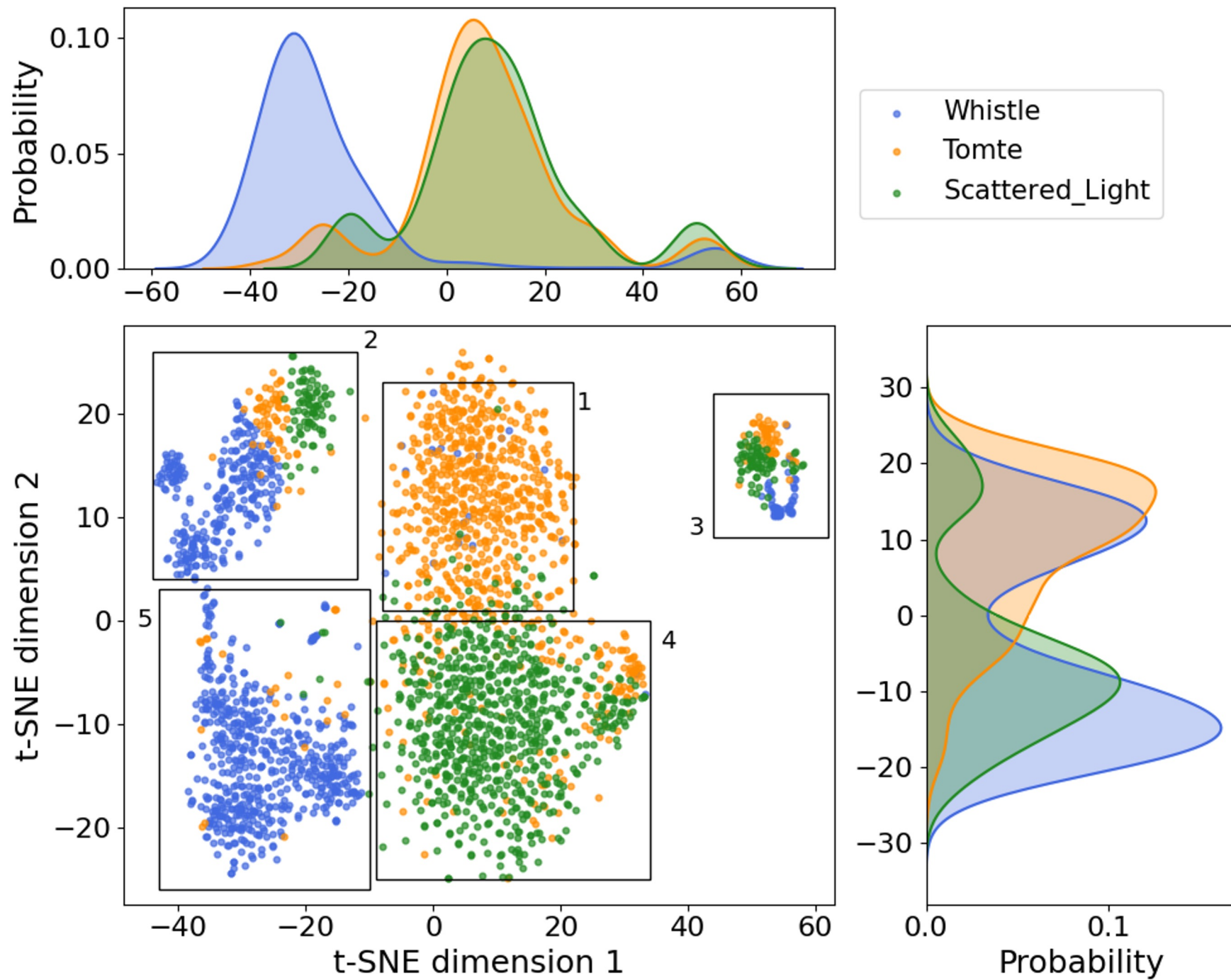
# Section 2

“Standard” Tomte



Label Tomte, but Tomte overlap with Scratchy

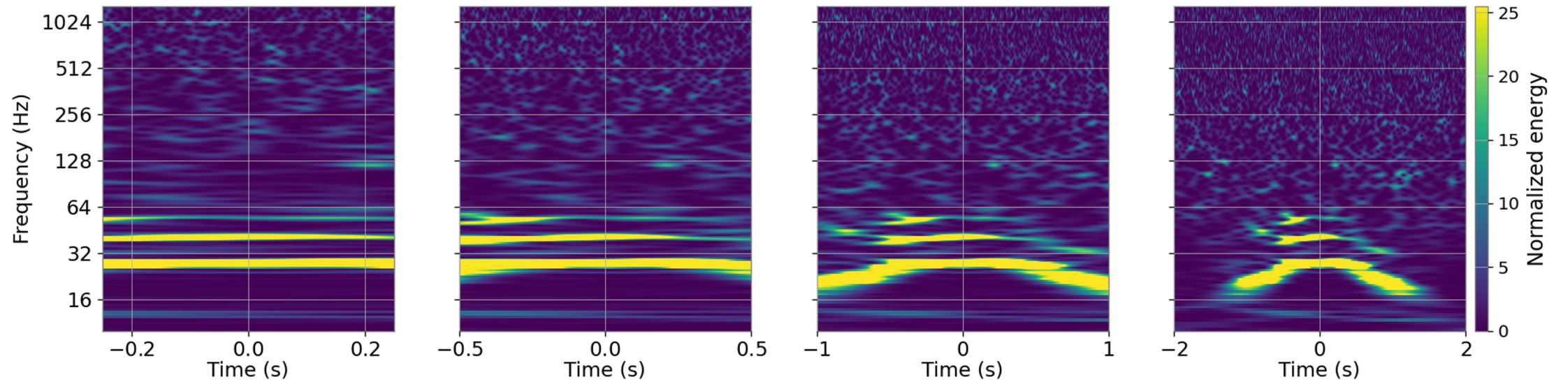




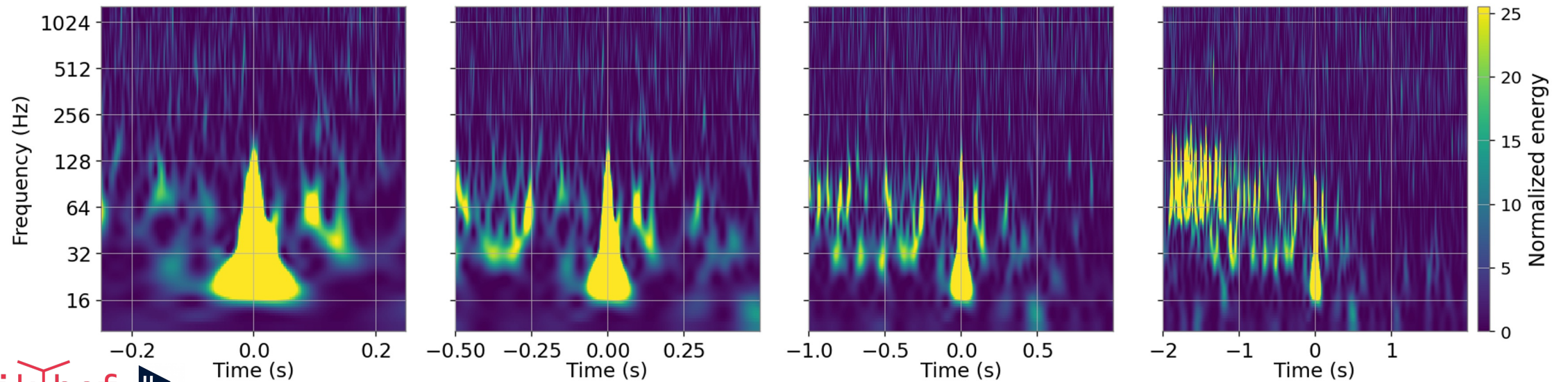


# Section 3

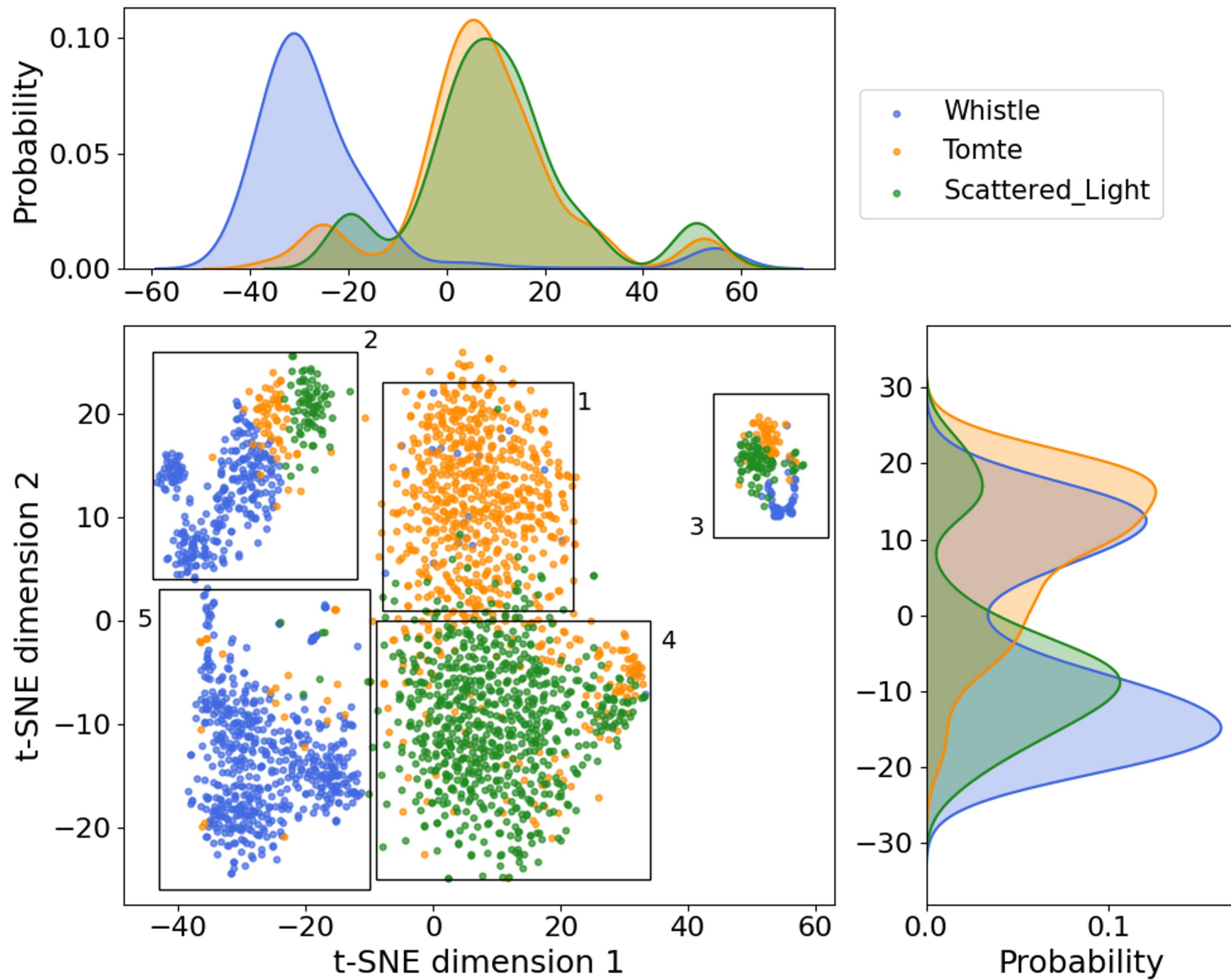
“Standard” Scattered Light



Labelled as a Scattered Light, but overlap between Tomte and Scratchy

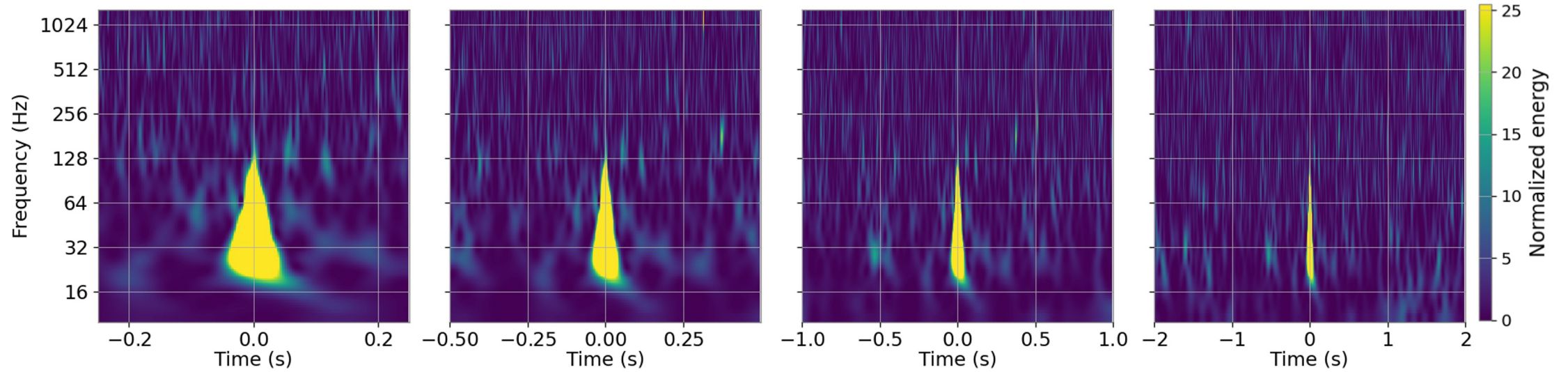




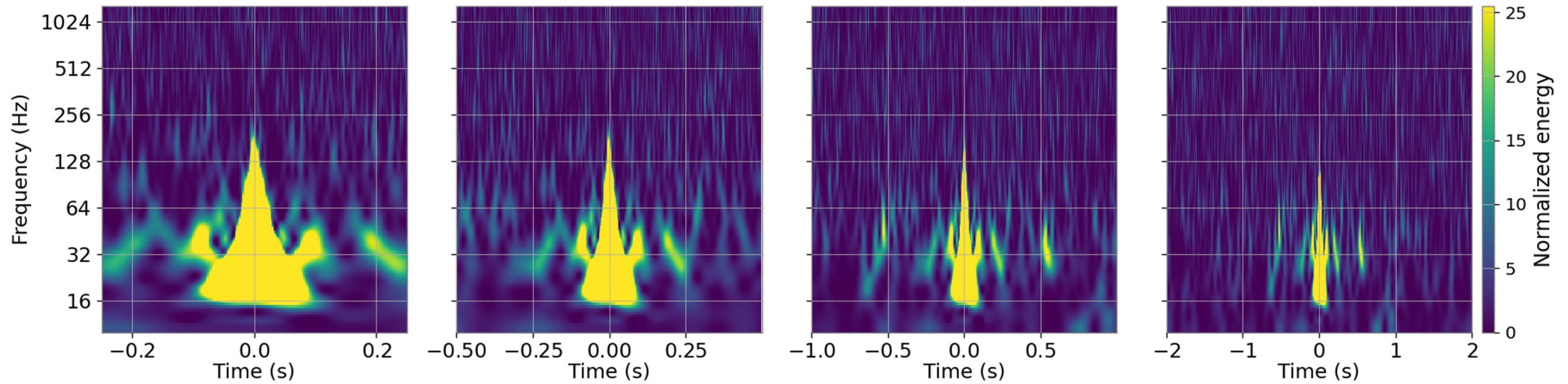


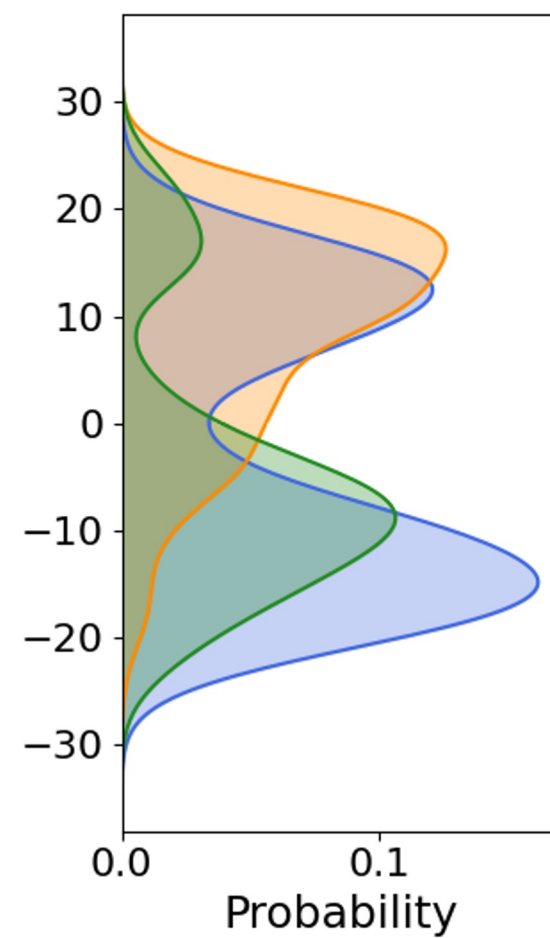
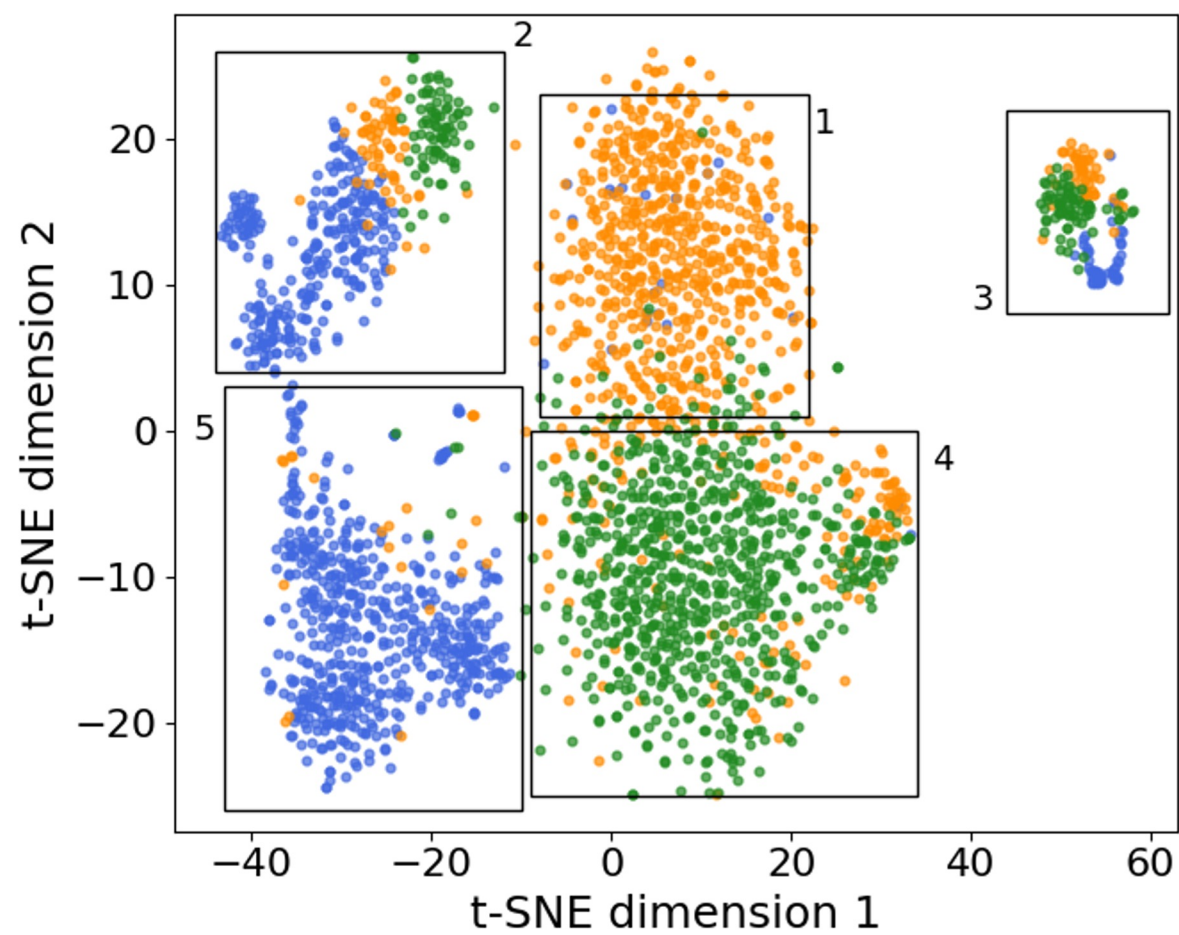
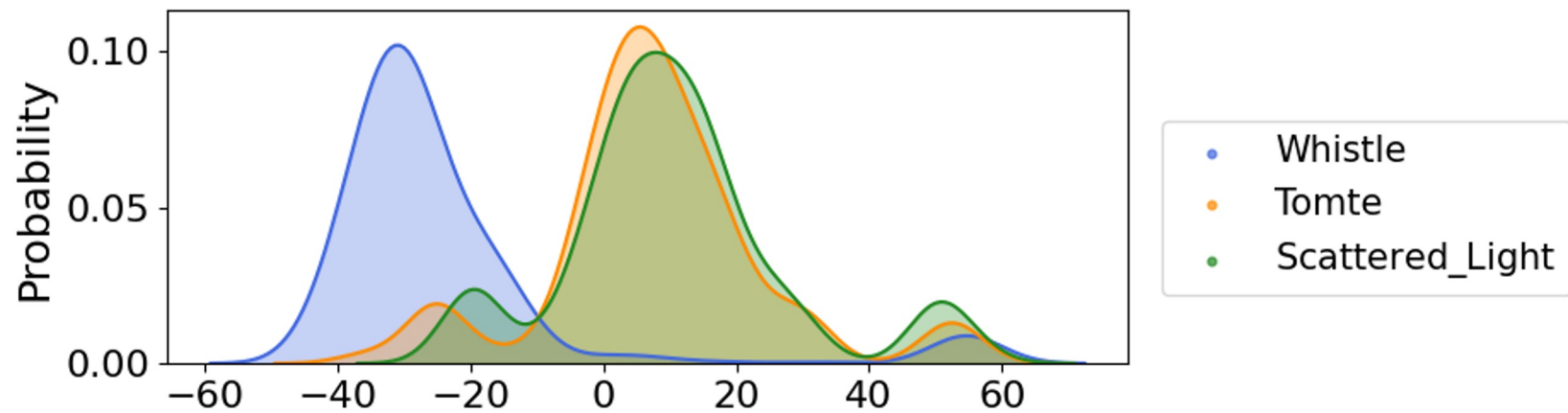
# Section 4

“Standard” Tomte



Labelled as Tomte but it is a Koi Fish

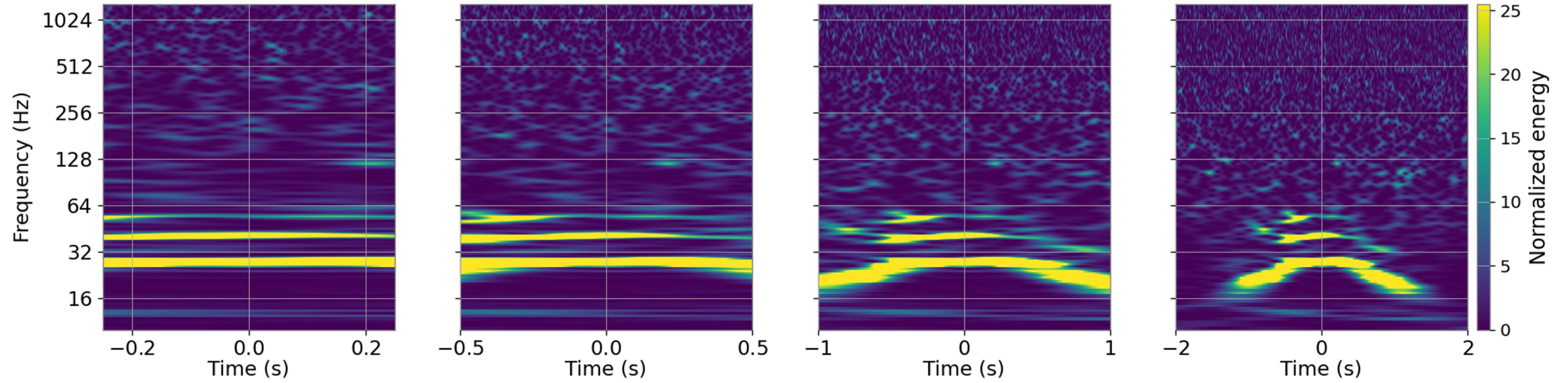




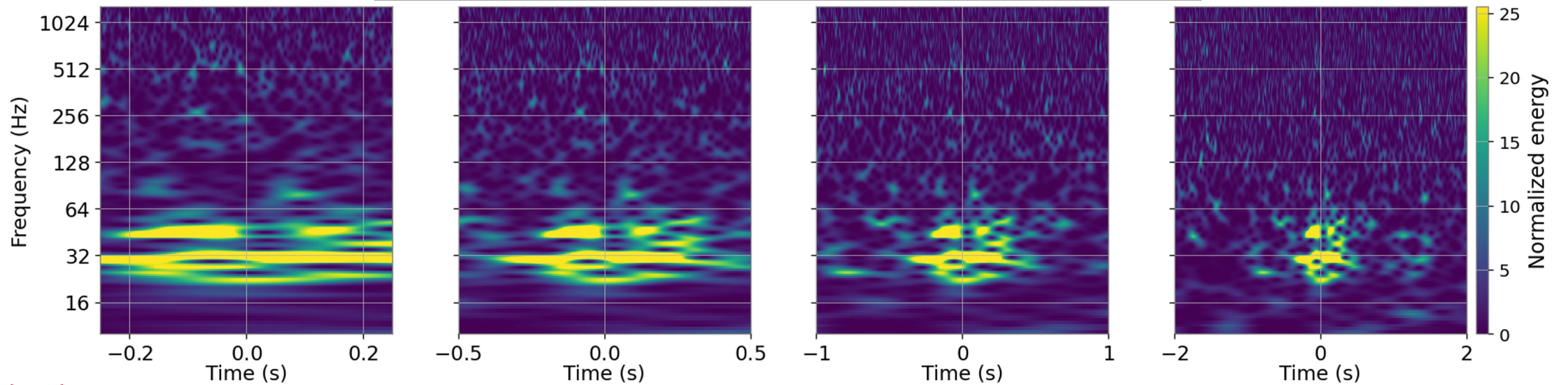


# Section 5

“Standard” Scattered Light



Labelled as Scattered Light but unknown morphology





# Results of compressed data

In total 177 anomalies were found, which constitute 6,6% of the data.

- **Anomalous whistles (49):**
  - 45% unknown morphologies, 28% misclassifications, 27% overlaps.
- **Anomalous Tomtes (57):**
  - 32% unknown morphologies, 21% misclassifications, 47% overlaps.
- **Anomalous Scattered Lights (71):**
  - 28% unknown morphologies, 72% misclassifications, 1 overlap.



# Conclusions and future work

- ✓ Fractal dimension representation is complementary to  $h(t)$
- ✓ Unsupervised learning can reveal misclassifications of supervised learning, glitch overlaps and novel morphologies
- Extend to glitch populations of GW detectors
- Relate glitches to auxiliary channels via explainable ML



# Thank you for listening! Questions?

m.lopez@uu.nl

