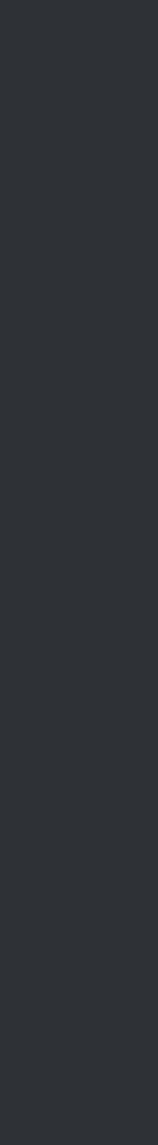


## a short introduction to our latest version of FINESSE

Daniel Brown, Andreas Freise for the Finesse team, 05.06.2023

## the new generation

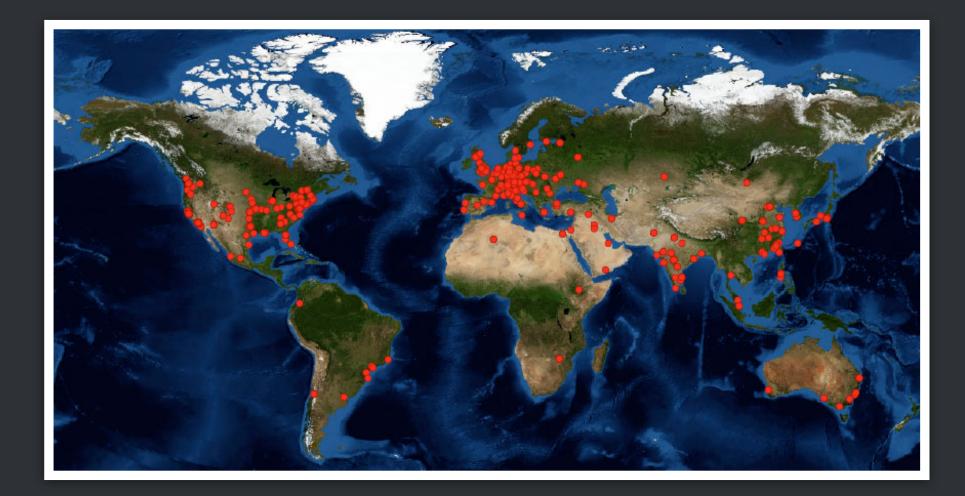






# Long history short

- Started 1997, PhD side project
- Used extensively worldwide
- Open sourced in 2012
- Continuously used and developed





Nikhef VU VRIJE UNIVERSITEIT AMSTERDAM 2





# What can FINESSE do?

FINESSE can simulate:

- Beam shapes
- **Optical losses**
- Quantum noise
- Squeezing
- Radiation pressure effects
- Diverse detectors
- Error signals
- Transfer functions

... so long as the model is frequency domain (i.e. static or quasi-static), paraxial, and suits modelling using a modal basis.



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## FINESSE 2.1

WADW, Elba, Italy, May 2016

GraWIToN 👪

ESSE is a frequency domain modelling tool for ptical experiments, specialised at gravitational ave laser-interferometry. It is free and oper source, released under the GPL license

PyKat is a free Python interface and set of tools for running FINESSE and for performing stand alone optical calculations. It is specialised for automating advanced optical simulations that involves multiple

## **FINESSE can Simulate:**

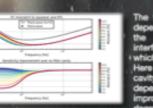
- Imperfect Beams
- Imperfect Optics
- Quantum Noise Squeezing
- **Radiation Pressure Effects**
- Parametric Instabilities
- Realistic Detectors

ezed vacuum in the fundamental e that scatters into anti-squeezed harmful anti-squeezing is reduce

However, to find out if we can achieve the necessary frequency dependent squeezed higher order modes with correct squeezing angles is a delicate ssue as the field passes through the omplex optical system of a ional wave detector. To the best knowledge, the above has nly been initiated, and worked on, by Jan Harms in our community

Finesse is being actively used in aLIGO ling. We have a ete file of the interferometer as in the design documents, as well as variants specific to the Livingston and ford sites. The core file is tuned to the operating point and locked using PyKat to automate much of the process

Ve are working on Activ vefront Control (AWC) to

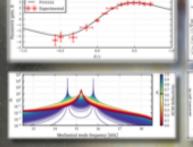






FINESSE, and this model has been against the experimental results ob

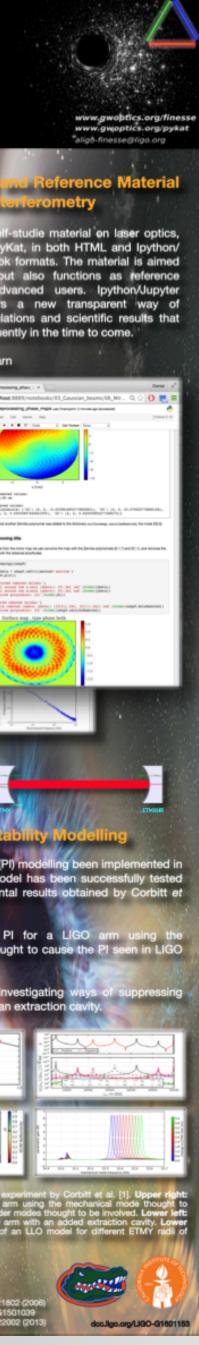
felled PI for a LIGO de thought to cause the PI :



## Image: 'Finesse 2.1' poster by D. Töyrä, LIGO-G1601153

Nik|hef

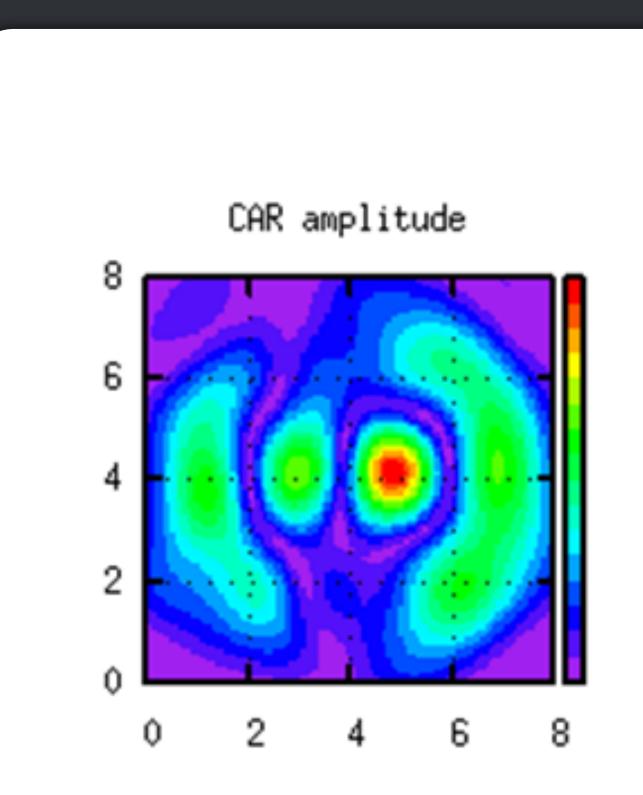




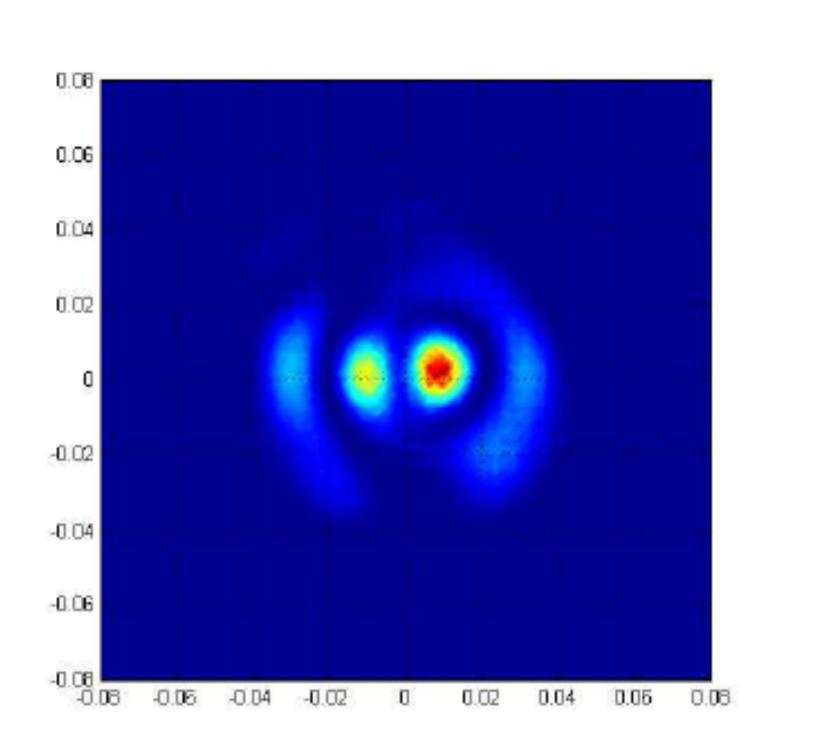


## Using FINESSE: match simulation to defective interferometers

The FINESSE team always spend significant effort to model imperfect interferometers. This was not yet crucial for the first generation of detectors, but has become essential for current and future instruments. Many FINESSE features, conventions and habits have been designed with this in mind.

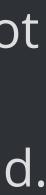


## Modal model (Finesse)



FFT propagation model (DarkF)

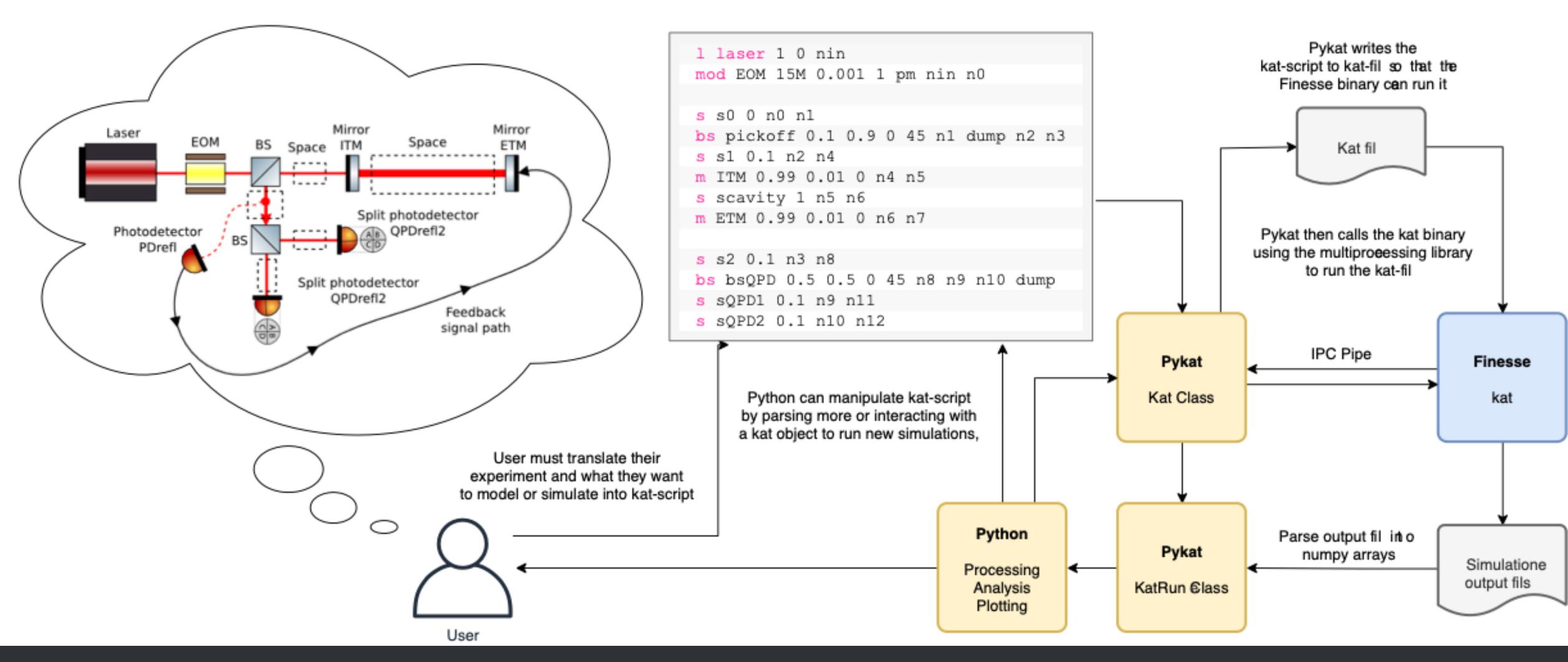
2009 a et Marque 





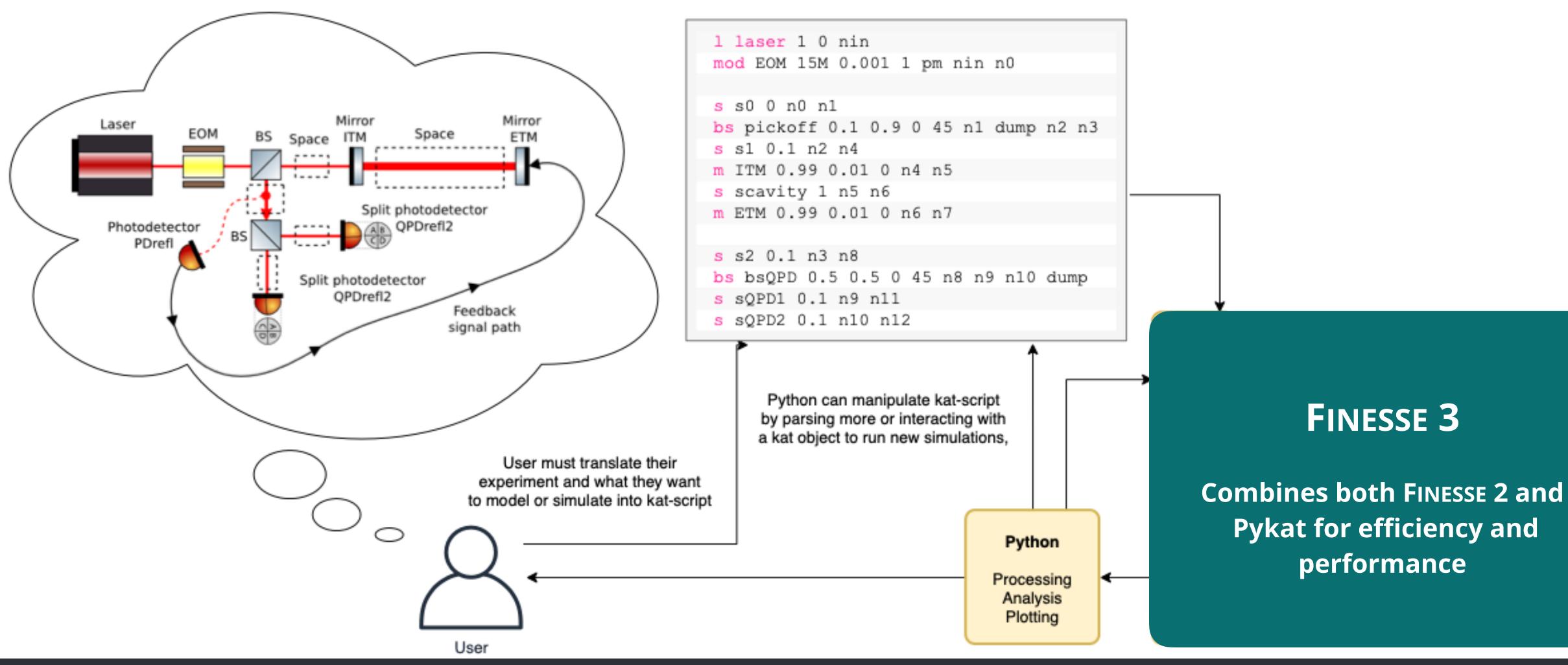


## FINESSE 2 + Pykat



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We are now encouraging the use of FINESSE 3 over previous versions!

The new FINESSE manual is not yet fully complete but is constantly evolving on the web page.

It contains many examples and documentation for various functions and commands.

If a part of the manual is not yet complete and you really need it, ask in the chat channel and we will aim to fill it in faster.



3.0a8-5-g3aee1c2d

Search docs

- Introduction to Finesse
- Getting started
- $\Box$  Examples
- $\oplus$  An optical cavity
- Pound-Drever-Hall locking
- $\oplus$  Near unstable cavities
- **Extracting geometrical parameters**
- Radio-frequency modulation and **Bessel functions**
- $\oplus$  Radiation pressure
- $\oplus$  Squeezing and homodyne detection
- $\oplus$  Optical springs
- Quantum-noise limited sensitivity of Advanced LIGO
- Aligning and translating a beam with steering mirrors
- $\oplus$  Cavity eigenmodes
- $\oplus$  Locking actions
- ⊞ Frequency dependant squeezing
- **Defining cavity eigenmodes with** the cavity command
- Using Finesse
- Physics of Finesse



A / Getting started / Installing Finesse / First Install

## **First Install**

This section provides information on getting started with your first installation of FINESSE. Although other installation methods are available, we recommend following this guide to set up your initial environment to avoid known issues.

## Python Installation

We recommend using Miniconda to install the Python ecosystem. We have created Conda packages to automate and easily install FINESSE in one command.

From the section titles, pick the one that suits your condition.

I am a Python user and have Anaconda/Conda installed already (Windows, OSX, or Linux)

I am a Python user and have my own environment already set up that I want to use

For Windows users without Conda installed

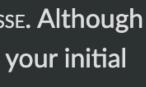
For MacOS users without Conda installed

For Linux users without Conda installed

I am a Python user and have Anaconda/Conda installed already (Windows, OSX, or Linux)

If you already use Conda on your system, then installation is very easy! You can proceed to Installing Finesse and Jupyter with Conda.









# Python vs KatScript

- FINESSE 2 only worked with KatScript (which limited) what could be done).
- PyKat added a Python wrapper which essentially wrote KatScript for you.
- FINESSE 3 is firstly a Python program, it has a Python interface for everything. KatScript is a wrapper around the full Python programming interface.
- This means that you can use both to make models and run simulations. KatScript is often easier or more compact, Python is more powerful.
- Some features are not (yet) supported in KatScript and can only be used through Python (e.g surface maps).

```
model = finesse.script.parse(""""
laser l1 P=1
space s1 l1.p1 m1.p2
mirror m1 R=0.5 T=0.5
power_detector_dc P m1.p2.o
```

# or you can use a Python interface... from finesse.components import Laser, Space, Mirror from finesse.detectors import PowerDetector

```
model_python = finesse.Model()
model_python.add(Laser('l1', P=1))
model_python.add(Mirror('m1', R=0.5, T=0.5))
model_python.add(
   PowerDetector(
       'P'
       model_python.m1.p2.o
model_python.add(Space(
    's1',
    model_python.l1.p1,
    model_python.m1.p1
```

# 'ipt



## FINESSE 2

m2 PRAR 0 \$L\_PRAR 0 nPR1 nPRsub1 s sPRsub 0.1003 \$nsilica nPRsub1 nPRsub2 m1 PR \$T\_PR \$L\_PR 0 nPRsub2 nPR2 # Measured cold IFO PR RoC [VIR-0029A-15] #attr PR Rc -1477 # Design value to have good matching (sho attr PR Rc -1430 # Measured PR AR RoC [VIR-0029A-15] attr PRAR Rc -3.62

# Space between PR and POP. Length from TDR. s lPR\_POP 0.06 1 nPR2 nPOP1

# Pick off plate. The angle of incidence and the physical distanc # propagates inside POP are computed from thickness of 3.5 cm [TD # tilt [TDR], and refractive index of \$nsilica. POP AR is wedged, # the AR-reflectivity is set as a loss. bs2 POP\_AR 0 \$L\_POP2 0 6.0 nPOP1 nPOPunused1 nPOPsub1 nPOPsub3 s sPOPsub 0.03549 \$nsilica nPOPsub1 nPOPsub2 bs2 POP \$R\_POP1 0 0 4.135015 nPOPsub2 nPOPsub4 nPOP2 nB4 s sB4\_att 0 nB4 nB4att bs B4\_attenuator 0.7344 0.2656 0 0 nB4att dump nB4b dump

# Changes to KatScript

## FINESSE 3

**m** PRAR R=0.0 L=160u Rc=-3.62

s sPRsub PRAR.p2 PR.p1 L=0.1003 nr=nsilica

m PR T=0.04835 L=30u Rc=-1430.0

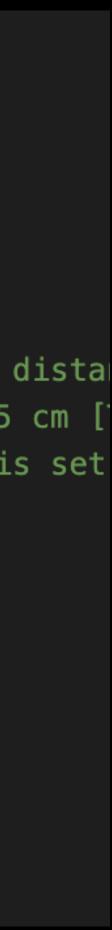
# Space between PR and POP. Length from TDR.

s lPR\_POP PR.p2 POP\_AR.p1 L=0.06

# Pick off plate. The angle of incidence and the physical dista # propagates inside POP are computed from thickness of 3.5 cm [ # tilt [TDR]. POP AR is wedged, thus one AR-reflectivity is set # POP reflectivities [VIR-0027A-15]

```
bs POP_AR R=0.0 L=125u alpha=6.0
s sPOPsub POP_AR.p3 POP.p1 L=0.03549 nr=nsilica
bs POP R=184u L=0.0 alpha=4.135015
# B4' port is POP.p4, attenuated B4 is B4_attenuator.p3
```

s sB4\_att POP.p4 B4\_attenuator.p1 **bs** B4\_attenuator R=0.7344 T=0.2656



- Main syntax style stays the same, i.e. m=mirror, one line per component
- No 'attribute' command any more
- Values are always assigned using the parameter name
- No need to specify nodes explicitly
- Instead spaces connect components directly (using 'ports')

## Changes to KatScript

## FINESSE 3

**m** PRAR R=0.0 L=160u Rc=-3.62

s sPRsub PRAR.p2 PR.p1 L=0.1003 nr=nsilica

m PR T=0.04835 L=30u Rc=-1430.0

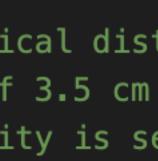
# Space between PR and POP. Length from TDR. s lPR\_POP PR.p2 POP\_AR.p1 L=0.06

# Pick off plate. The angle of incidence and the physical dist # propagates inside POP are computed from thickness of 3.5 cm # tilt [TDR]. POP AR is wedged, thus one AR-reflectivity is se # POP reflectivities [VIR-0027A-15]

bs POP\_AR R=0.0 L=125u alpha=6.0 s sPOPsub POP\_AR.p3 POP.p1 L=0.03549 nr=nsilica bs POP R=184u L=0.0 alpha=4.135015

# B4' port is POP.p4, attenuated B4 is B4\_attenuator.p3

s sB4\_att POP.p4 B4\_attenuator.p1 **bs** B4\_attenuator R=0.7344 T=0.2656



- You can do math with numbers, variables and references in every command
- New 'degree of freedom' (dof) command
- New `readout' commands to work with `dof' for sensing matrices and noise projection

# Changes to KatScript

## FINESSE 3

# Useful frequencies var fsrN (0.5 \* c0 / LN.L) var fsrW (0.5 \* c0 / LW.L) var fsrPRC (0.5 \* c0 / lPRC) var fsrSRC (0.5 \* c0 / lSRC) var f1\_arm (125.5 \* fsrN - 300.0) # Definition of f1, TDR section 2.3 var f1\_SRC (3.5 \* fsrN) var f1\_PRC (3.5 \* fsrN)

## # DOFs

## # position

dof DARM NE.dofs.z -1 WE.dofs.z +1 dof CARM NE.dofs.z +1 WE.dofs.z +1 dof MICH NI.dofs.z -1 NE.dofs.z -1 WI.dofs.z +1 WE.dofs.z +1 dof PRCL PR.dofs.z +1 dof SRCL SR.dofs.z -1

## # Detectors

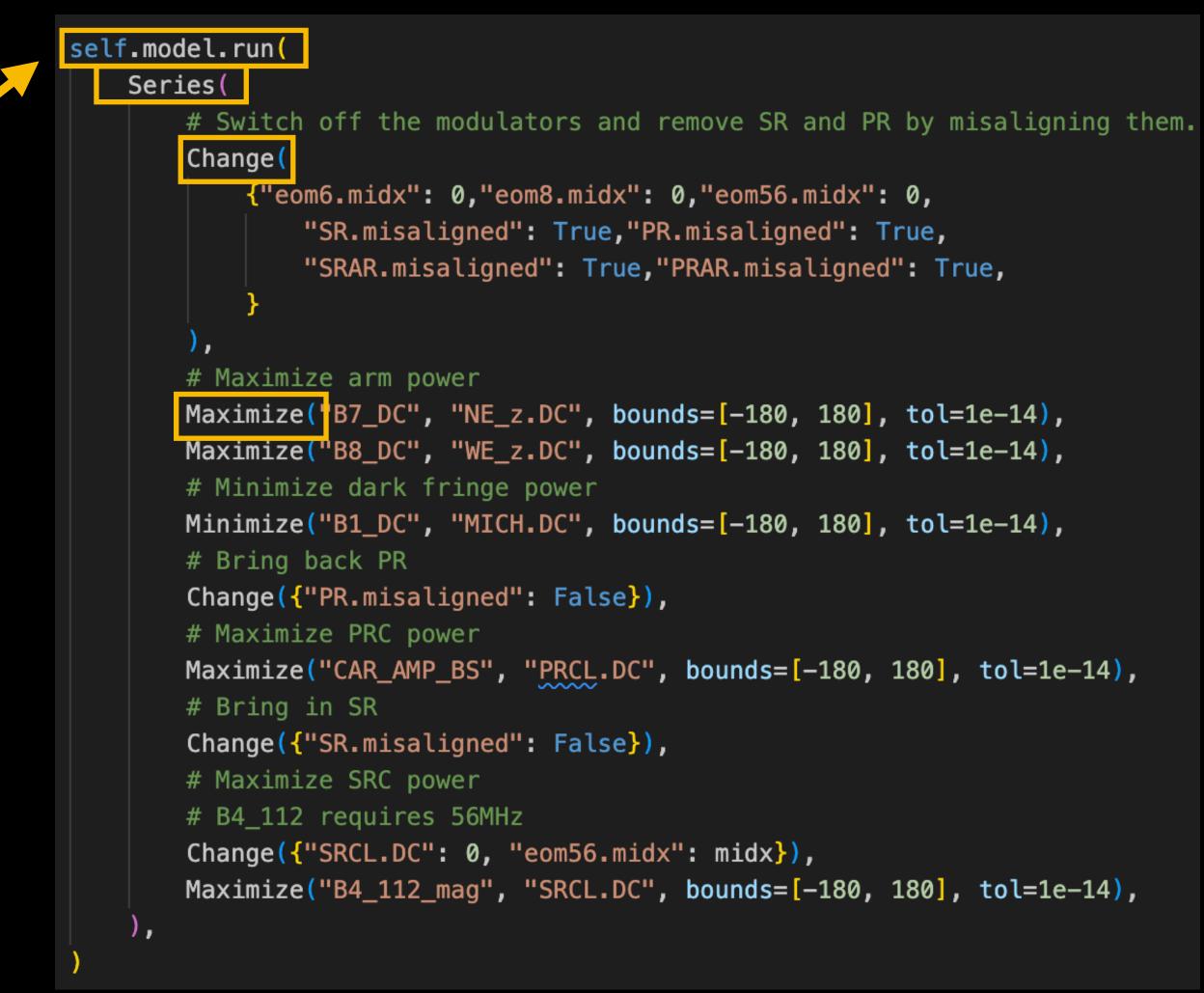
readout\_dc B1 OMC1\_2.p3.o output\_detectors=true readout\_dc B2 B2\_attenuator.p3.o output\_detectors=true readout\_dc B4 B4\_attenuator.p3.o output\_detectors=true readout\_dc B7 NEAR.p2.o output\_detectors=true readout\_dc B8 WEAR.p2.o output\_detectors=true



# Actions

- `Actions' are new Python functions to run FINESSE tasks.
- The actions pre-define all task before they are run. This allows FINESSE 3 to optimise the model (i.e. remove all tuning options that are not required by any action).
- Each action can do either a single simple task or execute a complex task.
- Most users would just use existing actions, but they are easy to write/expand with a bit of Python knowledge.

## FINESSE 3



Nodes are quite different in FINESSE 3! Each component have multiple *ports*, each port has multiple *nodes*.

How many and what nodes there are at a port depends o physical type: *Optical or Signal*.

Optical nodes have an input and an output, optical fields travel in both directions. Optical ports are typically named where N is the node number, pN.i and pN.o are the input output optical fields from the component.

Signal nodes represent electrical or mechanical states in system (these are where you can inject and read signals) like a GW signal or a small mirror oscillation).

# Nodes and ports

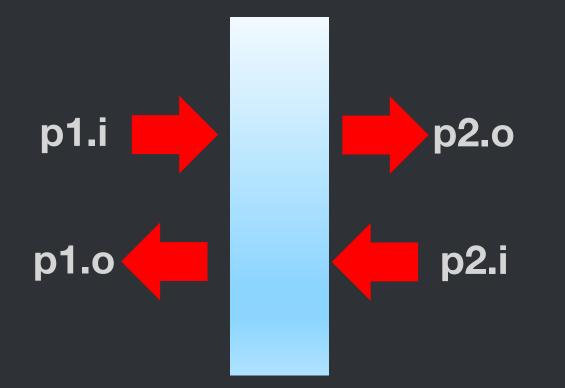
can	<pre>model = finesse.script.parse("""" laser l1 P=1 """") model.ll.nodes <!-- 0.0s</pre--></pre>
on the	OrderedDict([('l1.p1.i', <0pticalNode l1.p1.i @ 0x1678eb8b0 ('l1.p1.o', <0pticalNode l1.p1.o @ 0x1678eb490
	('ll.amp.i', <signalnode 0x1078eb490<br="" @="" ll.amp.i="">('ll.phs.i', <signalnode 0x1678eb94<="" @="" ll.phs.i="" td=""></signalnode></signalnode>
can d`pN` t and	<pre>('l1.frq.i', <signalnode 0x1678eb73<br="" @="" l1.frq.i="">('l1.dx.i', <signalnode 0x1678eb7c0="" @="" l1.dx.i=""> ('l1.dy.i', <signalnode 0x1678eba00="" @="" l1.dy.i=""> ('l1.yaw.i', <signalnode 0x1678eb9d<br="" @="" l1.yaw.i="">('l1.pitch.i', <signalnode 0x1678eb<br="" @="" l1.pitch.i="">('l1.mech.z', <signalnode 0x1678eb<br="" @="" l1.mech.z="">('l1.mech.x', <signalnode 0x1678eb<="" @="" l1.mech.x="" pre=""></signalnode></signalnode></signalnode></signalnode></signalnode></signalnode></signalnode></pre>
n the s from,	('l1.mech.y', <signalnode 0x1678eb<br="" @="" l1.mech.y="">('l1.mech.yaw', <signalnode 0x16<br="" @="" l1.mech.yaw="">('l1.mech.pitch', <signalnode @<="" l1.mech.pitch="" td=""></signalnode></signalnode></signalnode>



0>)**,** 0>)**,** 50>)**,** 40>)**,** 30>), **⊲>),** d0>), 8eb8e0>)**,** ba60>), ba90>), bbe0>), 678ebaf0> 0x1678eb



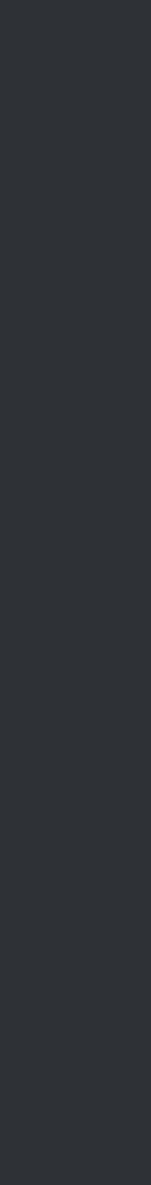
# Make more examples



mech.z mech.yaw mech.pitch

mech.F\_z mech.F\_yaw mech.F\_pitch

Nikhef VU



- FINESSE 3 main page: https://finesse.ifosim.org/
- FINESSE 3 code repository: https://gitlab.com/ifosim/finesse/finesse3
- FINESSE 3 anaconda package: https://anaconda.org/conda-forge/finesse
- Chat channel for FINESSE 3: https://matrix.to/#/#finesse:matrix.org

## Links to resources

- IFOsim logbooks https://logbooks.ifosim.org/
- Interferometer techniques for gravitational wave detection https://link.springer.com/article/10.1007/ <u>s41114-016-0002-8</u>
- GWIC 3G 'Simulations and Control', see https://dcc.ligo.org/LIGO-G1800565
- IFOsim mailing list: https://grouper.ligo.org/mailinglists/ifosim







