Cryogenic Coating Thermal Noise direct measurement

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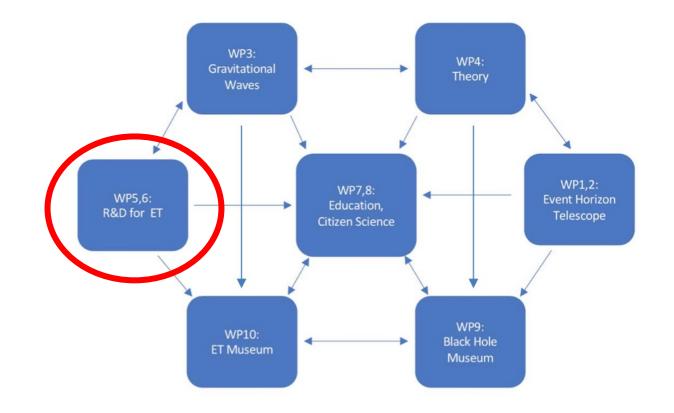




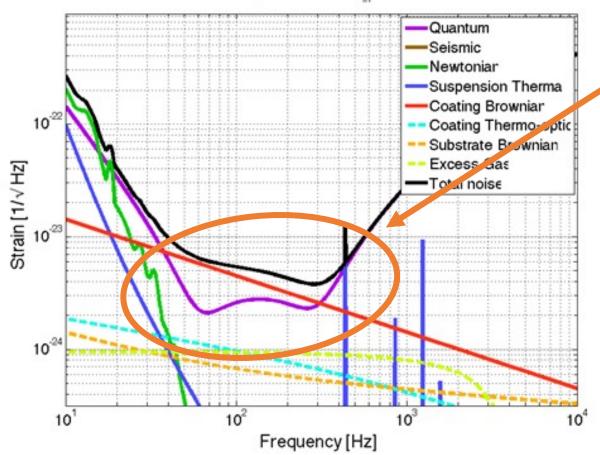


Dutch Black Hole Consortium

A group of 30+ scientists, carrying out an interdisciplinary black hole research program throughout the Netherlands: <u>https://www.dbhc.nl/</u>



Sensitivity of Gravitational Wave Detectors



Advanced Virgo Noise Curve: P ... = 125.0 W

Mid-range Frequencies:

Sensitivity limited by mirror coating thermal noise

$$x_c^2(\omega) = rac{4k_BT}{\omega} rac{1-\sigma^2}{\sqrt{\pi}E_0 w} \phi_{eff,c}(\omega)$$

- directly proportional to material properties (E_0 , σ , $\phi_{eff,c}$)
- inversely proportional to the beam spot size on the mirror (w)
- directly proportional to temperature (T)

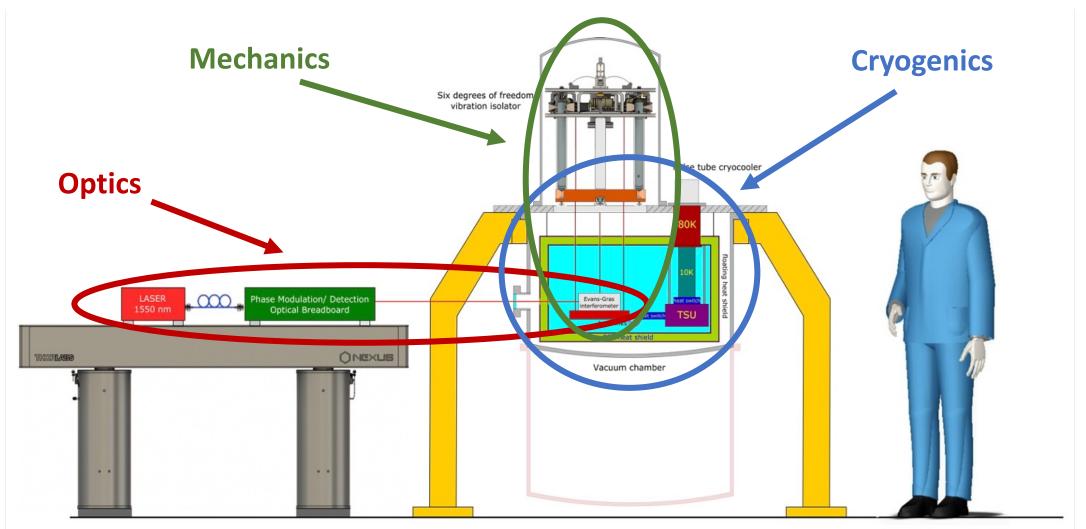
Different technique to reduce the thermal noise impact -> decrease the working temperature -> new material needed for coating

Motivations & Goals

- TN of new coatings directly measured using an interferometric method in a cryogenic environment to select the material to be used in the thirdgeneration detectors;
- A low noise fast turnaround cryogenic test bench essential to characterize new sensors and actuators compatible with low temperature environment.
- R&D on new low noise technology is of wide interest -> Liquid free low noise cryogenics more and more required in many fields of fundamental physics.

Cryogenic test bench with ~ 10^{-13} m / sqrt (Hz) residual motion above 30 Hz

Motivations & Goals



A flexible test facility available to anyone (not just in the GW collaboration) who needs it

Three main work packages

- Cryogenic Technology (Koen)
- Seismic Attenuation System (Enrico)
- Optical setup (Enzo)

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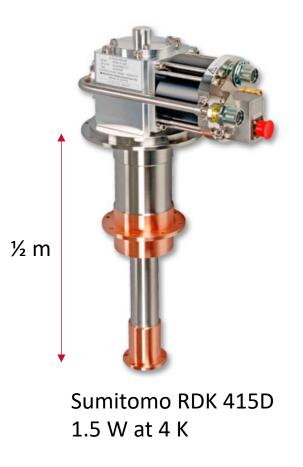
Cryogenic system requirements

| | ET, ETPF | Mirror Coating Test Facility |
|--------------------------|--------------|--|
| Operation of system | Continuous | Discontinuous |
| Vibration-free condition | At all times | Only during measurements (~ ½ hour) |
| System cooldown | Weeks | A few days |
| Operating temperature | 10 K, 123 K | 10 K (123 K*) |

Core concept

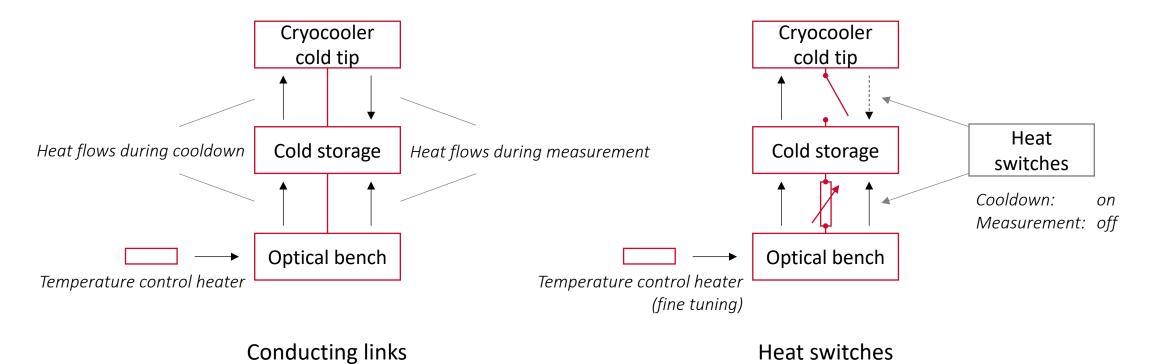
- Cold storage (battery)
- Mechanical cryocooler (charging battery)
- Cooler on, cool down cold storage
 Cooler off, measure using stored cold





Operation

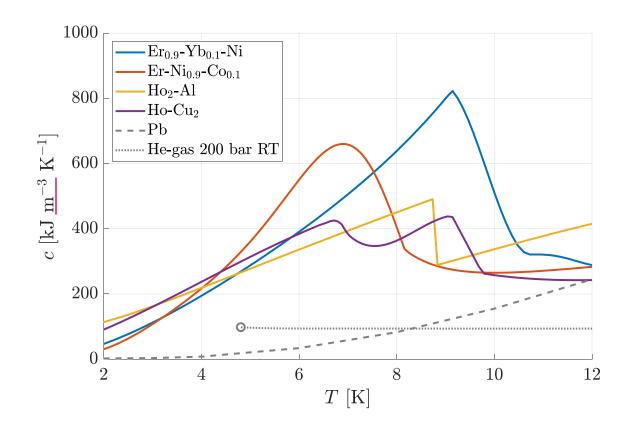
Other heat loads (radiative, supports/suspension) not drawn



Cold storage

High heat capacity at $T < 10 \ {\rm K}$

- Phase change material: latent heat
 - Helium (4.2 K, 1 bar)
 - Lanthanides (magnetic)
- Ordinary heat capacity
 - Lead
 - Helium gas

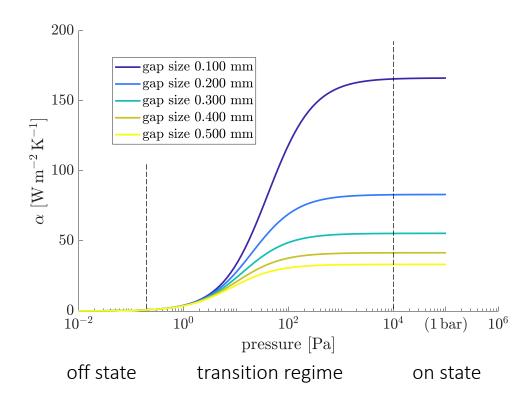


Heat switches

- Mechanical
- Gas-gap

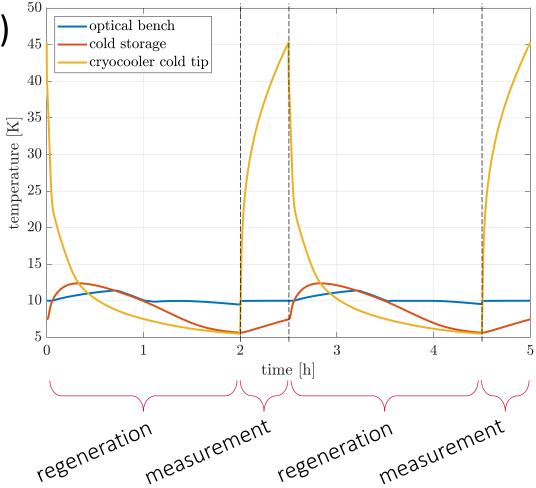


M.O. Kimball et al., NASA



Simulation: cyclic operation

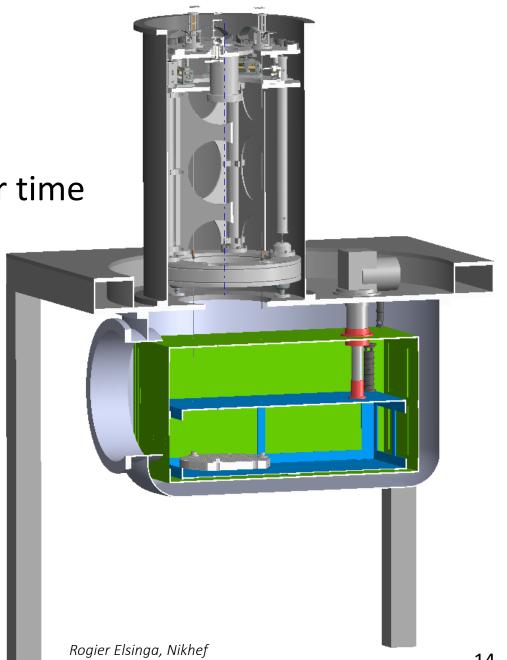
- 2 hours of regeneration (cooldown)
- ¹/₂ hour of measurement time



Phased approach

Modular system, replace components over time

- Lead cold storage, conducting links (no heat switches)
 - System testing
 - Will not meet ½ hour measurement criterium
 - Target cryogenic assembly: end of 2024
- 2. Replace links by heat switches
- 3. Replace cold storage by helium gas or magnetic phase change material



Three main work packages

- Cryogenic Technology (Koen)
- Seismic Attenuation System (Enrico)
- Optical setup (Enzo)

GAS filters Six degrees of freedom IP vibration isolator Pulse tube cryocooler Evans-Gras interferometer 80K heat shield Vacuum chamber нι Suspending Wires

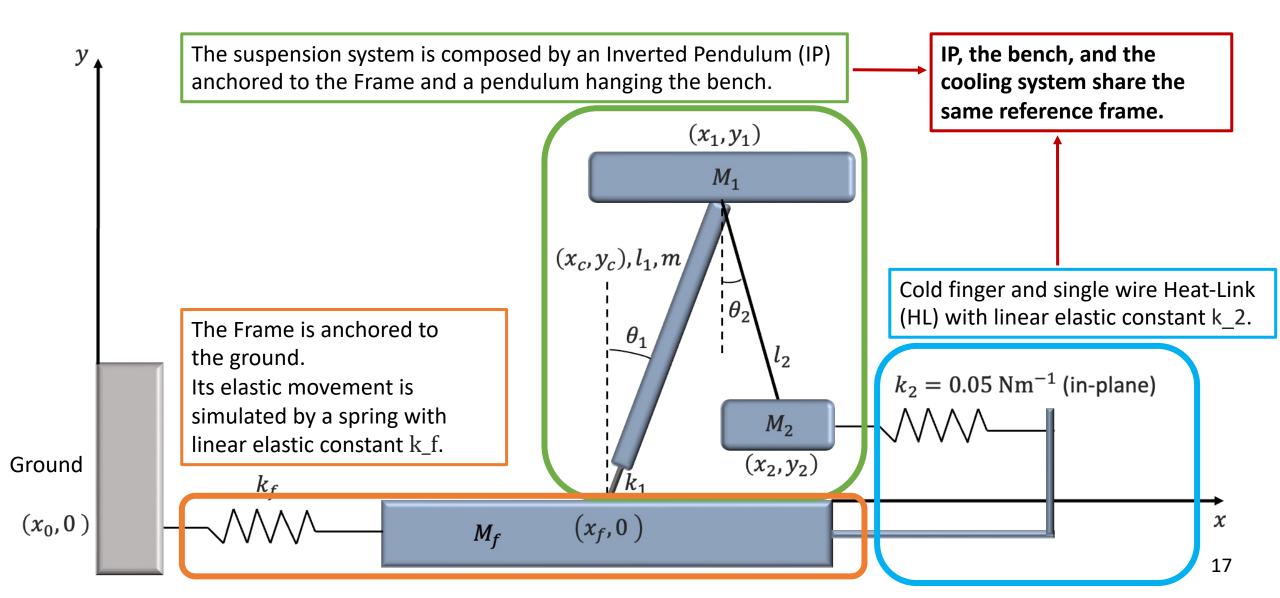
Why a vibration-isolation system?

- 1. Measuring the vibration injection due to heat-links by avoiding direct contribution from the ground.
- 2. Testing sensors/actuators at cryogenic temperature (≥10 K).

The whole system is supported by a Frame and is composed by:

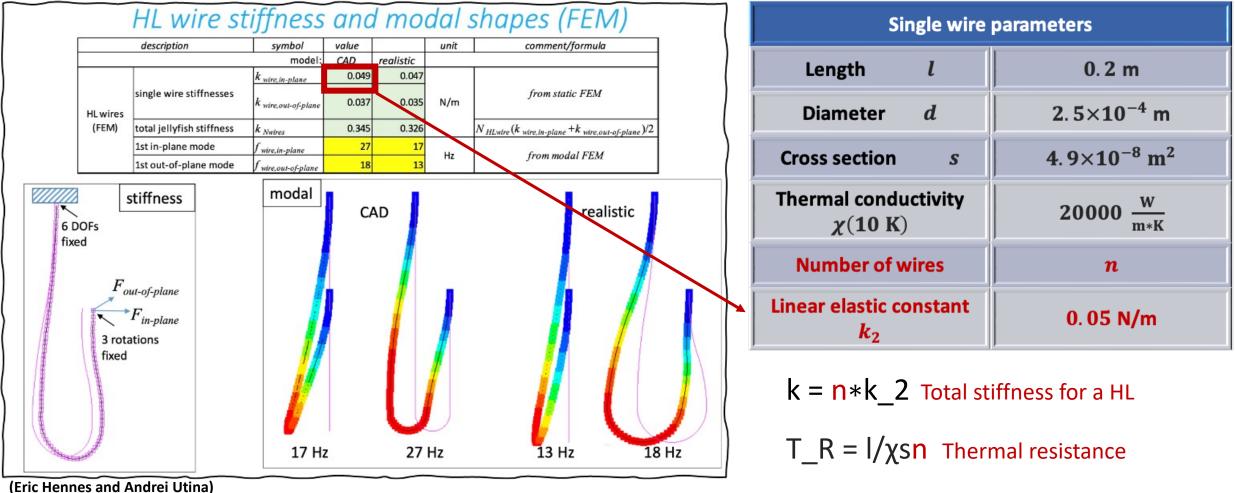
- 3 legs Inverted Pendulum (IP) Longitudinal vibration.
- 1 stage with 3 pairs of Geometrical Anti Springs (GAS) Vertical vibration.
- 3 Titanium alloy suspending wires hanging the optical bench.
- Heat-Links (HL) for thermal conduction from the Cryocooler to the bench.

Lagrangian model of the mechanics in a longitudinal DoF



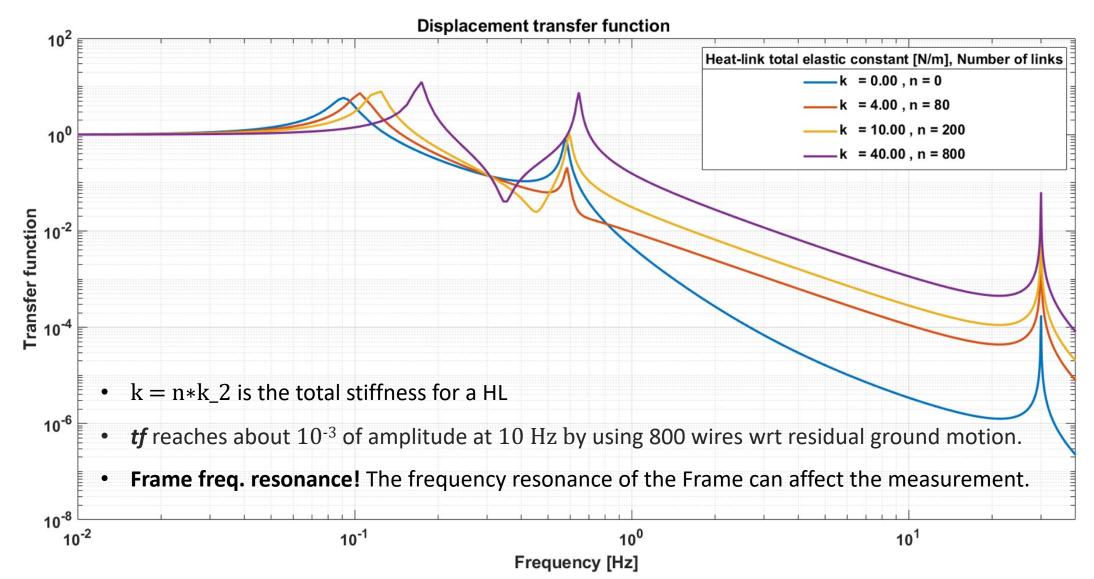
Stiffness of a single wire HL

For a fixed geometrical model

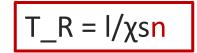


How many wires do we need in order to have a good cooling down without affecting the seismic isolation?

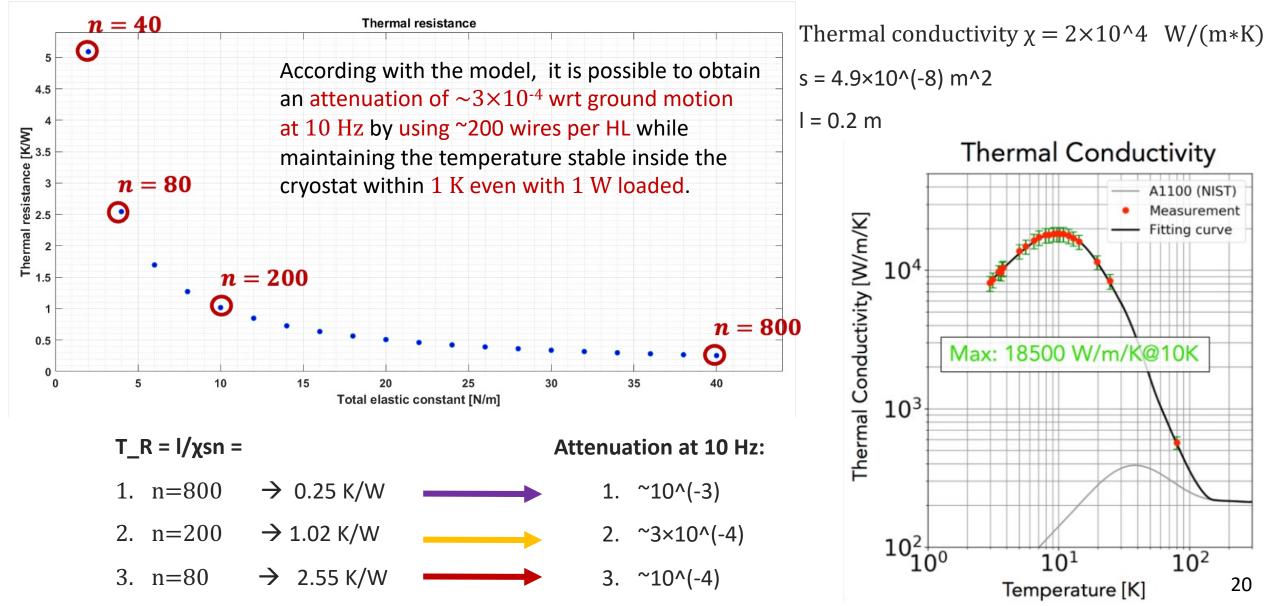
Displacement transfer function (tf) of the mechanics at the bench level



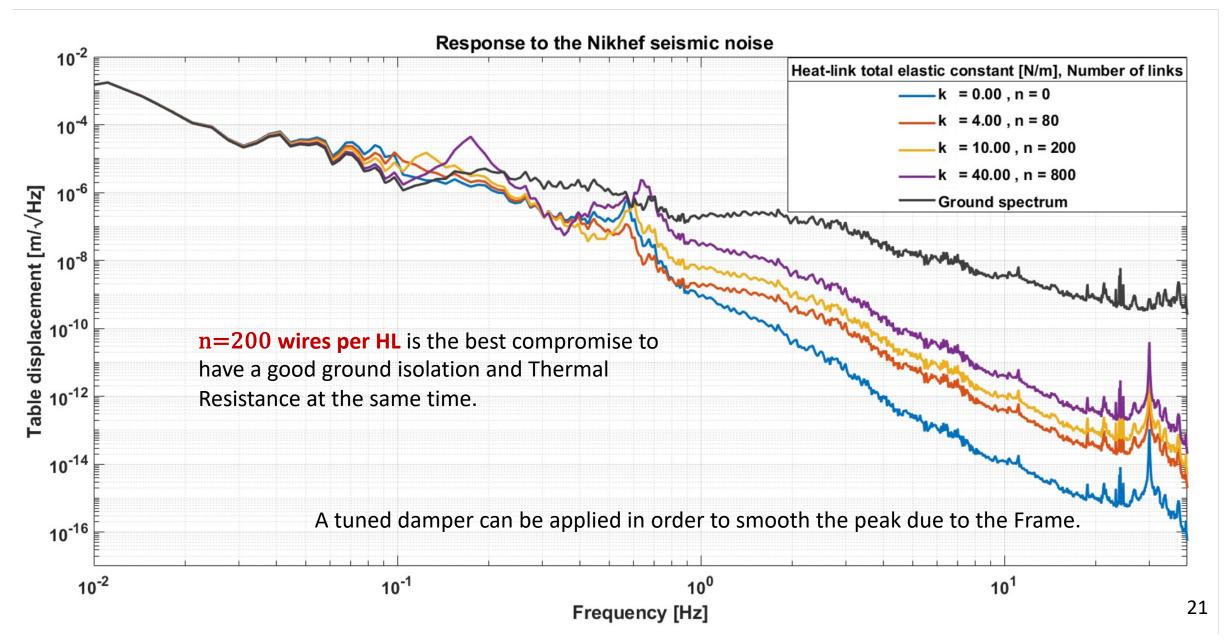
HL Thermal Resistance



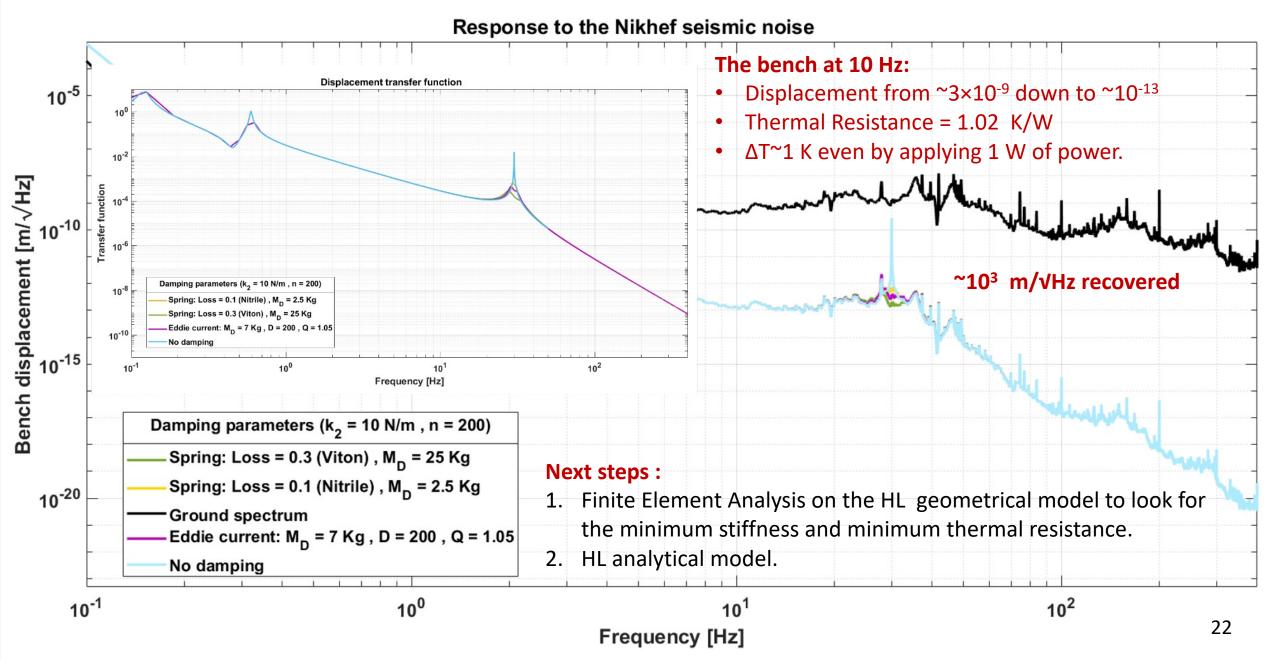
Taking advantage from KAGRA experience



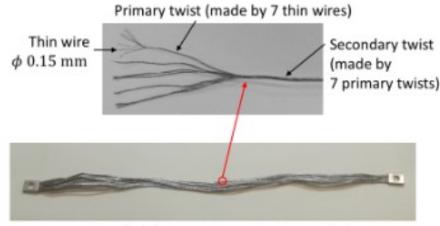
Expected noise level of the bench



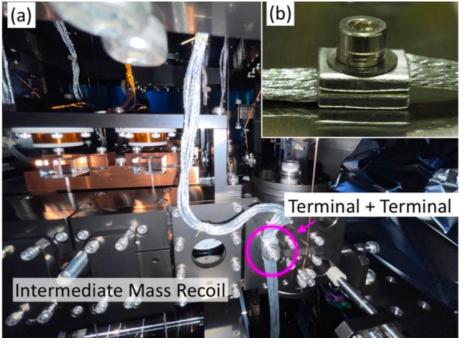
Bench displacement after applying a tuned damper



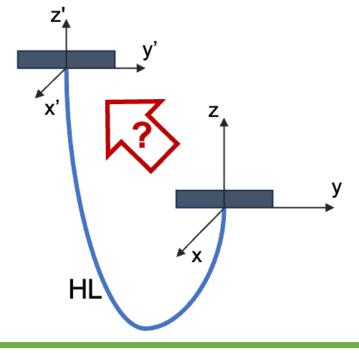
Heat Link analytical study

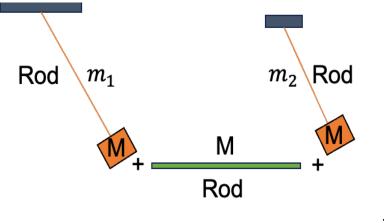


Heat link (7 secondary twists in parallel)



In both pictures: Heat-Links used at KAGRA





 $-Ely(z,t)'''+Ty(z,t)''=\rho\ddot{y}(z,t)$

Where:

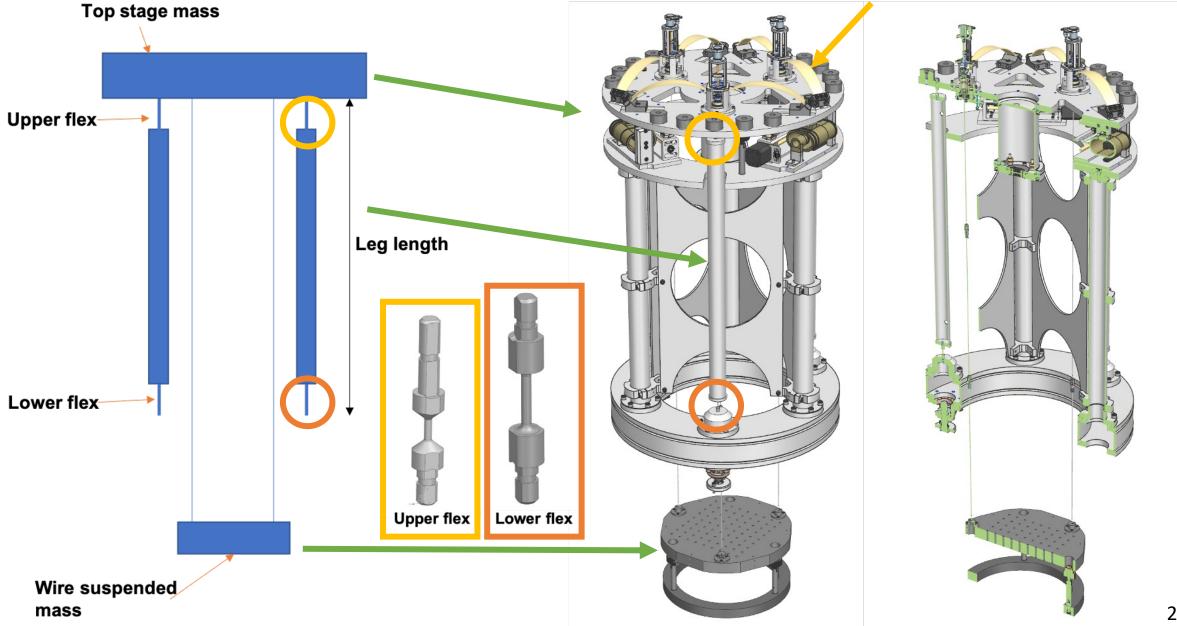
- $E \rightarrow$ Young Modulus
- I → Second moment of area
- T \rightarrow Tension
- $\rho \rightarrow$ Mass per unit length

To solve in frequency domain for:

- Rod+mass pendulum
- Simply supported rod

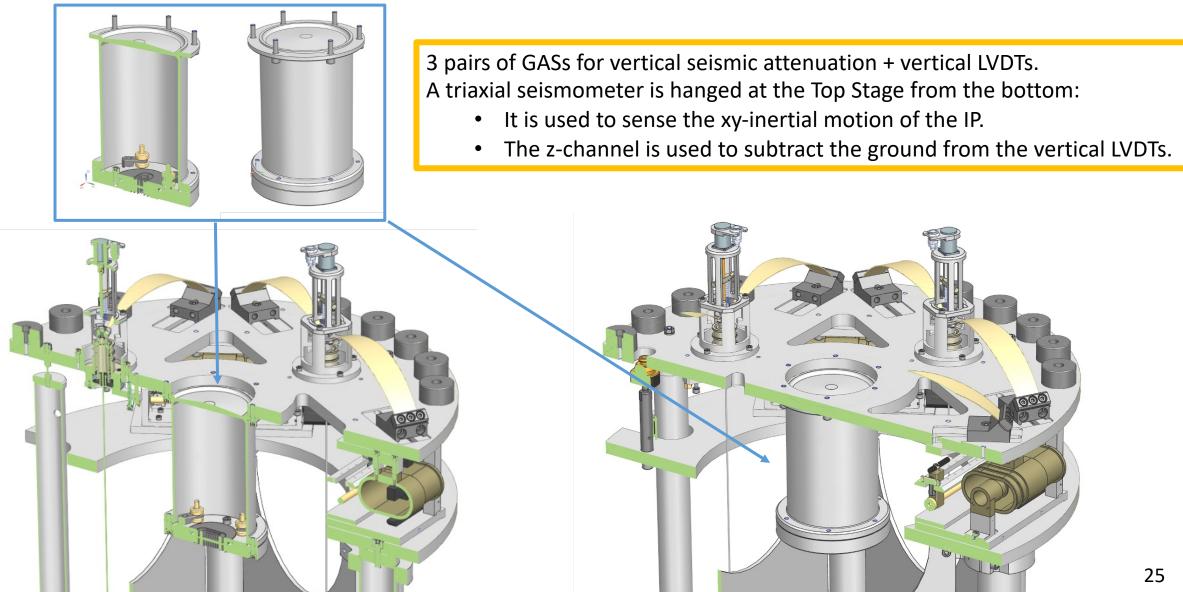
Vibration Isolation System Overview

Geometrical Anti-Spring (GAS)

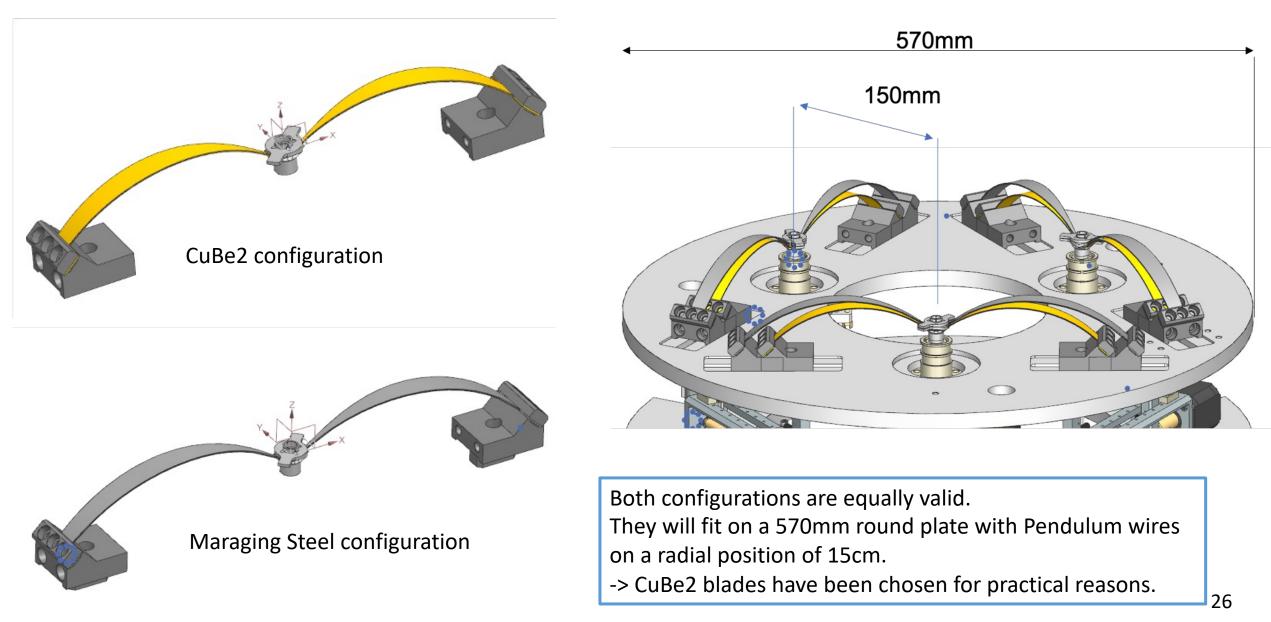


Top Stage: GAS blades+Seismometer

Nanometrics Trillium 120

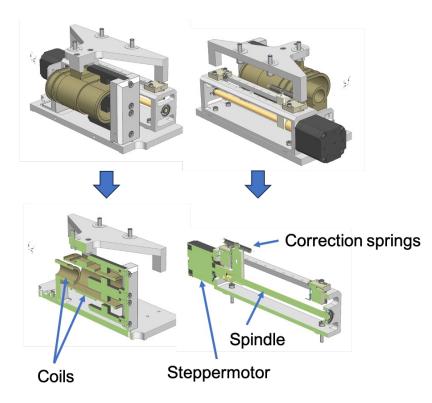


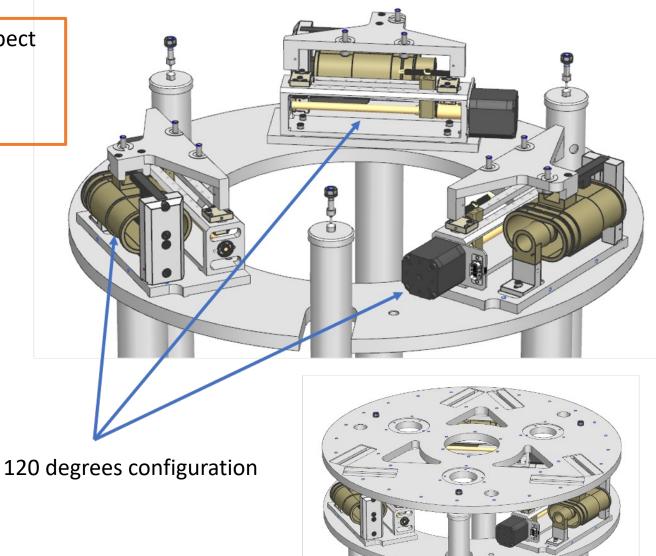
GAS filters



Top Ring: IP Correction Spring Assembly

- LVDTs are used to sense the position of the IP with respect to the Frame/Ground.
- Voice-coils are used for correction.
- StepMotors are used for static positioning.





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Next Steps

Short term:

- Finalization of the analytical model of the heat-links
- Procurement of the mechanical components of the vibration-isolation system in progress;

Mid term:

• Implementation and checks of the vibration-isolation system in stand-alone;

Long term:

• Integration of the vibration-isolation system with the optical bread-board and the cryogenic system in the vacuum tank;

Three main parts

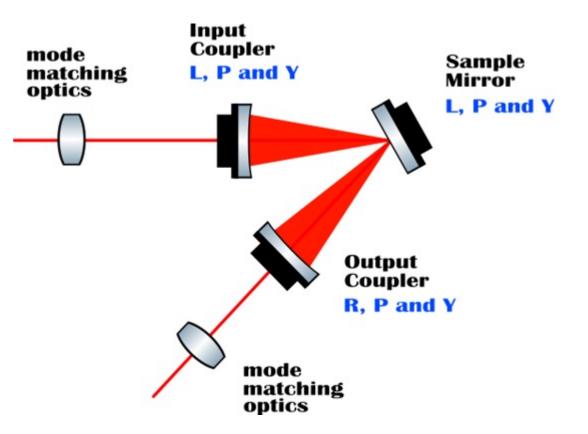
- Cryogenic Technology (Koen)
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Working Principle

Fluctuations of the longitudinal motion of the optical axis can be caused by the internal thermal noise of a mirror in an optical cavity. Fluctuations are produced by the internal damping inside the test mass.

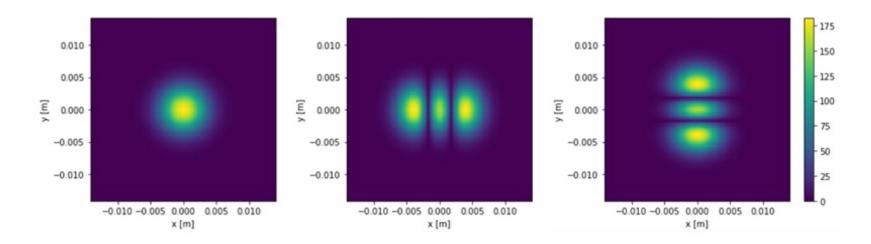
Measurement of the longitudinal displacement of the sample mirror by direct application of the fluctuation-dissipation theorem to the readout of the setup.

First implementation at MIT: S. Gras and M. Evans Phys. Rev. D **98**, 122001



Working Principle

Folded cavity must allow co-resonance for the fundamental mode and two higher order modes (HoMs). Thermal noise that is sensed by the TEM02 mode is different from the thermal noise sensed by the TEM20 mode, since they do not explore the same area of the sample mirror.



In the readout, for the HOMs, common noises such as length noise, frequency noise and thermal noise of the mirror can be decoupled from the Coating Thermal Noise (CTN). CTN is obtained from the beat note of the resonant frequencies of TEM02 and TEM20.

Optical Setup: Overview

1550 nm laser beam split into three beams: one is used to lock the cavity with the TEM00 mode, the two of others are shifted in frequency with Acousto-optic Modulators (AOMs) to have TEM02 and TEM20 resonant in the cavity.

latching Ontic

Polarizing B

DC locking of TEM02 and TEM20

Polarizing R

olarizing BS

to ISS for TEM20 and TEM02

PLL

CTN

In-vacuum high Finesse multi resonant folded cavity.

From VCO

From VCO

AOM for TEM20

AOM for TEM02

from PDH TEMO(

SM Coupler 1x2 50/50

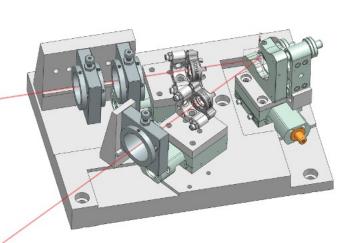
Laser

1550nm

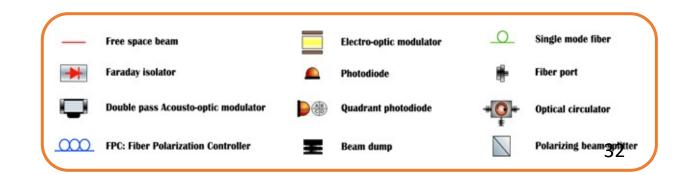
SM Coupler

1x2 10/90

PDH TEMOO

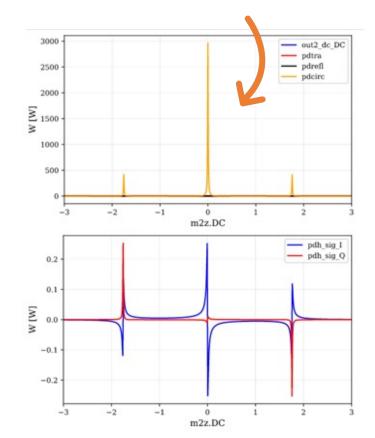


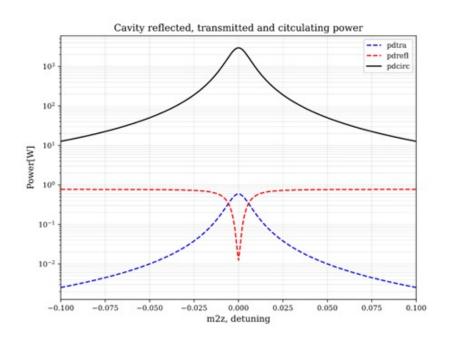
*Cavity design based on a similar AEI cavity design

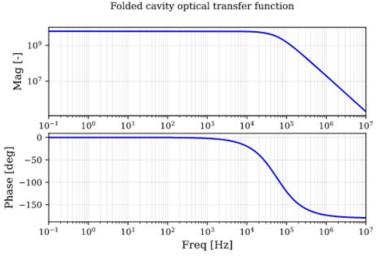


Optical Setup: Simulations

Sample mirror inside a folded high-Finesse optical cavity. RoC Input mirror: 50 mm RoC Output mirror: 50 mm Fundamental (TEM00) is locked using a Pound-Drever-Hall (PDH) loop in reflection.



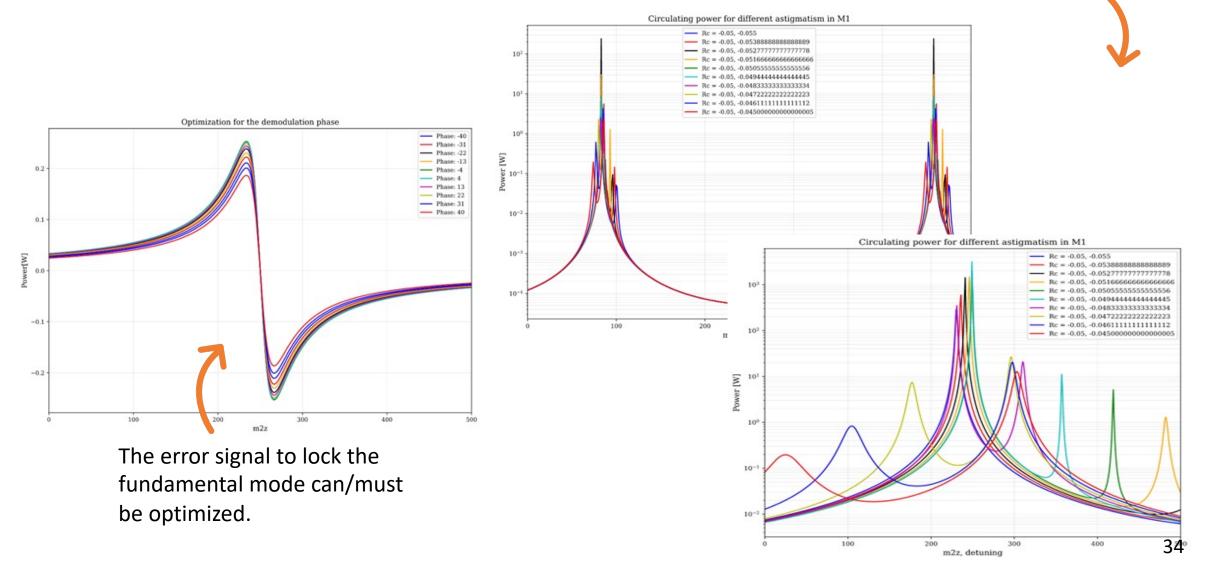




Pole of the cavity @ 56kHz Cavity Finesse: ~13000

Optical Setup: Simulations

TEM00, TEM02 and TEM20 can co-resonate inside the folded cavity by tuning the orientation of the input mirror.



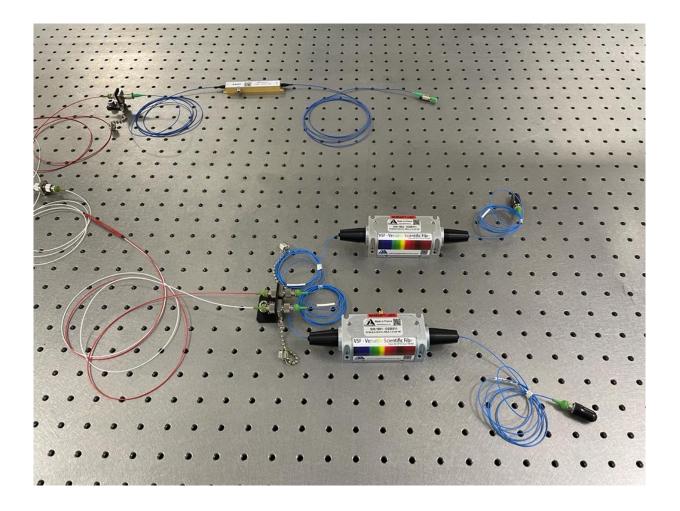
Optical Setup at Nikhef Lab

Main Laser and Faraday Isolator: 1550nm Main laser with 0.1-2W Output Power.



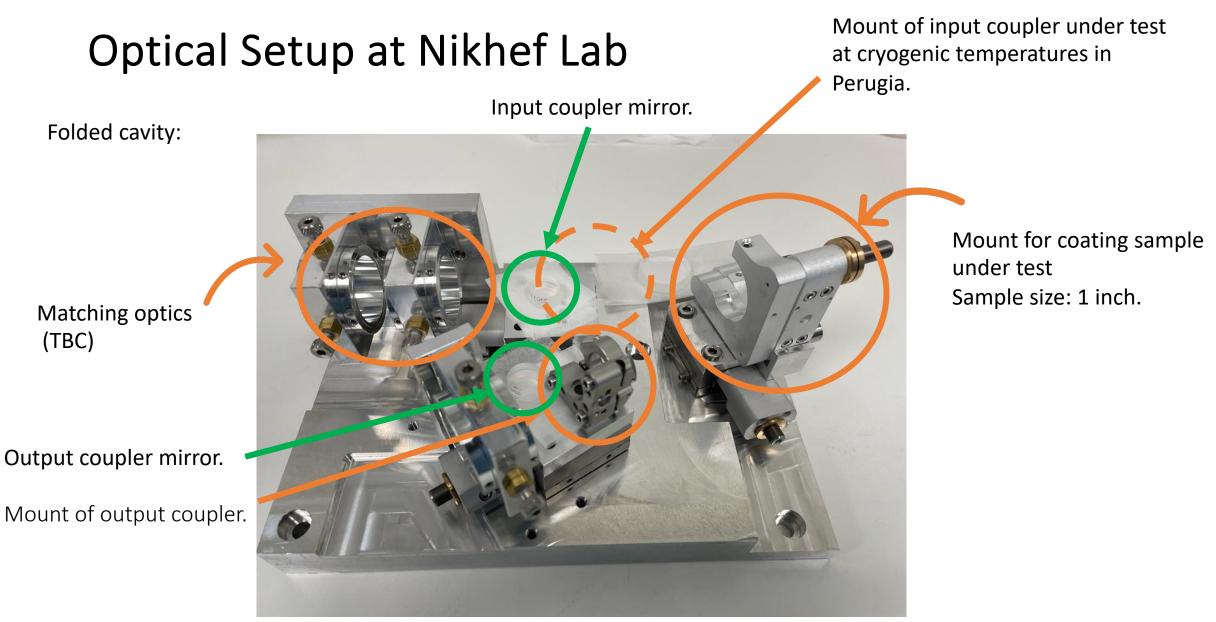
Optical Setup at Nikhef Lab

Electro-optic Modulator (EOM) and Acousto-optic Modulators (AOMs) installed in lab:



EOM modulation frequency ~ 15 MHz AOM modulation frequency ~280MHz

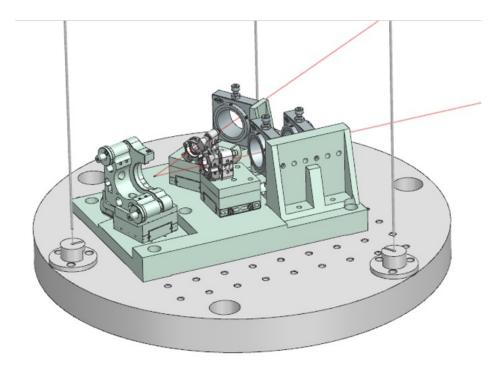
From here towards and in the folded cavity the laser is in-air. Collimator and matching optics to be installed.

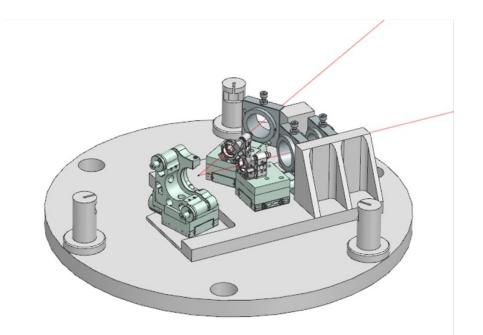


Integration of the optical breadboard

Two possible solutions:

- Removable breadboard to allow the cavity pre-alignment outside the vacuum tank.
- Monolithic cavity to reduce the residual motions.





Next Steps

Short term:

- Characterization of the optical beams in the lab;
- Design and procurement of the mode-matching telescope;
- Procurement of some electronics in progress;

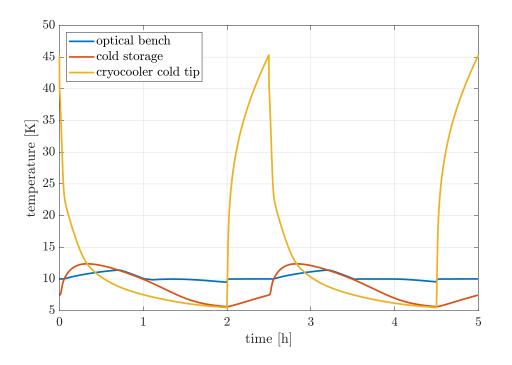
Mid term:

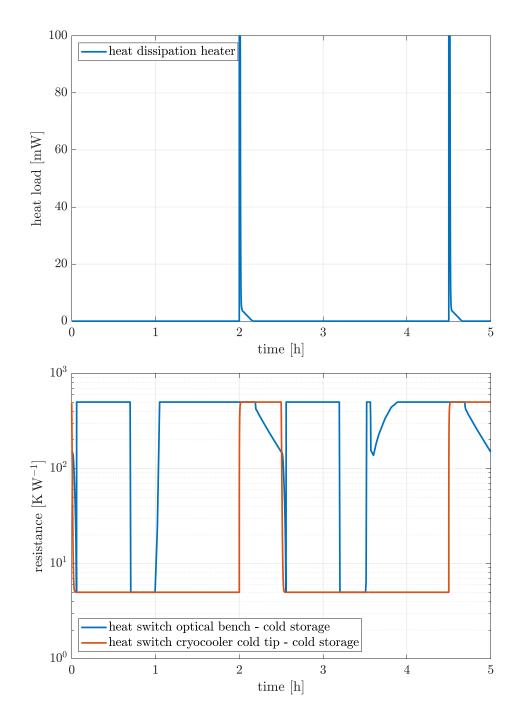
- Have the 3 beams co-resonant in the cavity in lab;
- Lock the cavity in-air;

Long term:

- Integration of the optics in the full experiment;
- Measurements and characterization of the full system at room temperature;
- Measurements and characterization of the full system at cryogenic temperature

Backup: simulation





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