

Gravitational wave data analysis

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Testing general relativity and the nature of compact objects

Testing the dynamics of binary black hole coalescence:





 Higher harmonics of the basic signal: Are the amplitudes consistent with GR?

Puecher et al., PRD 106 , 082003 (2022)

Looking for anomalous propagation of gravitational waves

• E.g. due to a non-zero graviton mass Current bound: $m_g \leq 1.76 \times 10^{-23} \, {\rm eV}/c^2$

LIGO, Virgo, KAGRA, arXiv:2112.06861

- Black holes, or black hole mimickers?
 - Search for gravitational wave echoes emitted by horizonless objects
 Tsang et al., PRD 101, 064012 (2020)
 LIGO, Virgo, KAGRA, arXiv:2112.06861



Lo et al., PRD **99**, 084052 (2019)

Lensing of gravitational waves

Just like light, gravitational waves can undergo lensing



- Gravitational wave images:
 - Copies of the original signal, but (de)magnified
 - Arriving at different times: minutes to months apart
 - Expected to be discovered in coming years

Difficult to search for:

- Developed an efficient way to check all pairs of detected signals for similarities Janquart et al., MNRAS 506, 5430 (2021)
- Use of lens models to keep false alarm probability under control Wierda et al., Astrophys. J. 921, 154 (2021) Janquart et al., MNRAS 519, 2046 (2022)

Waveform modeling

Effect of eccentric orbits on binary black hole waveforms



Ramos-Buades, Tiwari, Haney, et al., PRD **102**, 043005 (2020)

Phenomenological models for binary neutron star inspiral-merger-postmerger



igure 7. Gravitational waveforms (top) from large-scale numerical simulations of binary neutron-star mergers nottom) will give valuable input for the construction of phenomenological waveform models that can be used in ata analysis. Towards the end of the proposal duration, we will have access not only to tidal effects during spiral (bottom left), but also the signal from a hypermassive neutron-star resulting from the merger (bottom iddle and right). (Figure from Ref. [77].)

Puecher et al., arXiv:2210.09259

The equation of state of dense nuclear matter

Empirical channels:

- GW observations of binary neutron star coalescences
- Electromagnetic counterparts: jet and afterglow











Heavy ion collisions in accelerators on Earth



The equation of state of dense nuclear matter

Unified Bayesian data analysis framework to combine information from all these channels:



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Constraining neutron-star matter with microscopic and macroscopic collisions

<u>Sabrina Huth</u> ⊡, <u>Peter T. H. Pang</u> ⊡, <u>Ingo Tews, Tim Dietrich, Arnaud Le Fèvre, Achim Schwenk,</u> Wolfgang Trautmann, Kshitij Agarwal, Mattia Bulla, Michael W. Coughlin & Chris Van Den Broeck

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Huth et al., Nature 606, 274 (2022)

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Huth et al., Nature 606, 274 (2022)

Can we connect with LHC physics? (ALICE)



To be investigated through the grant "Probing the QCD phase diagram"

Towards Einstein Telescope

Einstein Telescope data analysis poses serious challenges:

 Loud signals, long signals, overlapping signals: "Traditional" search and parameter estimation techniques are inadequate



LIGO, Virgo, PRL 120, 091101 (2018)

Pursuing three avenues to address the parameter estimation problem:

(accepted by PR) (10000)

- Hierarchical subtraction
- Joint parameter estimat
- Machine learning

Will investigate how this cari



