



University of Amsterdam

Nikhef

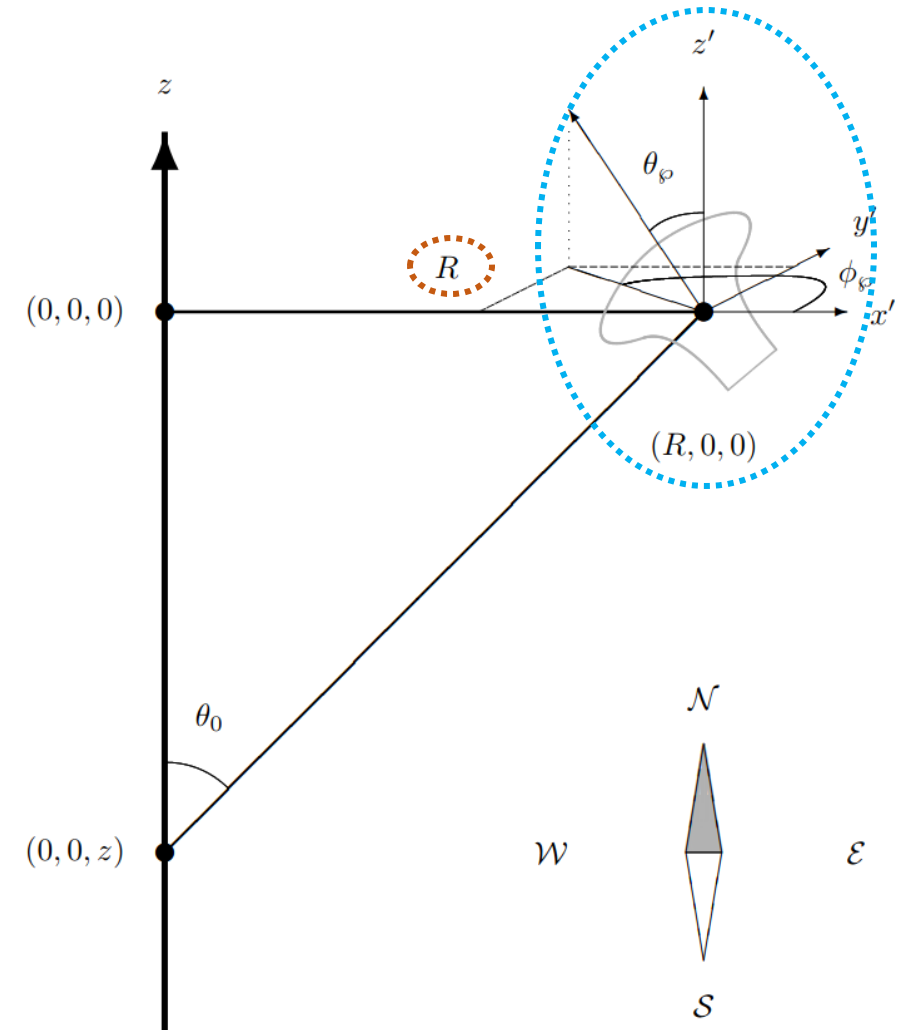


# No-hit inclusion

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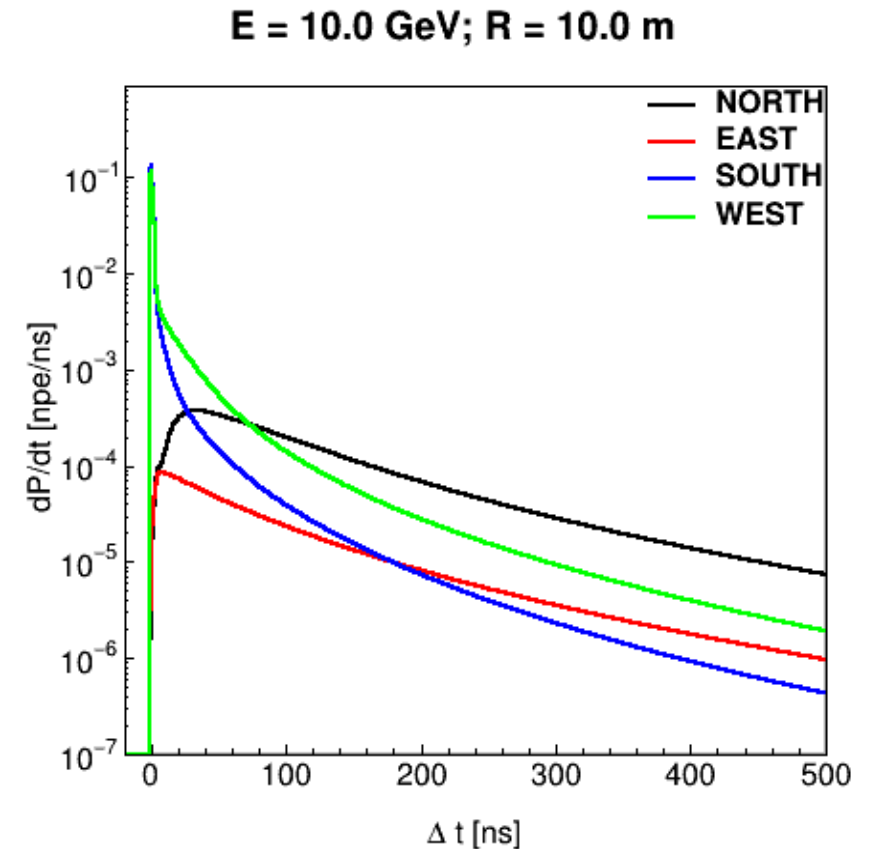
# Track reconstruction

- Five(-and-a-half) stages:
  1. JMuonPrefit  $\longrightarrow$  Directional grid scan
  2. JMuonSimplex  
(JMuonStart) } Directional MLE
  3. JMuonGandalf }
    - Based on arrival time PDF written by Maarten
    - Function of:
      1. E: the muon energy
      2. **R**: radial distance to PMT
      3.  **$\theta$  &  $\phi$** : PMT orientation
      4.  $\Delta t$ : arrival time difference w.r.t. direct Cherenkov
  4. JMuonStart  $\longrightarrow$  Finds track origin
    - Back-project hits onto track
    - Determine first emission point above background level
  5. JMuonEnergy  $\longrightarrow$  Energy MLE
    - Likelihood based on hit probabilities
    - Use hits / no-hits within cylinder surrounding muon path

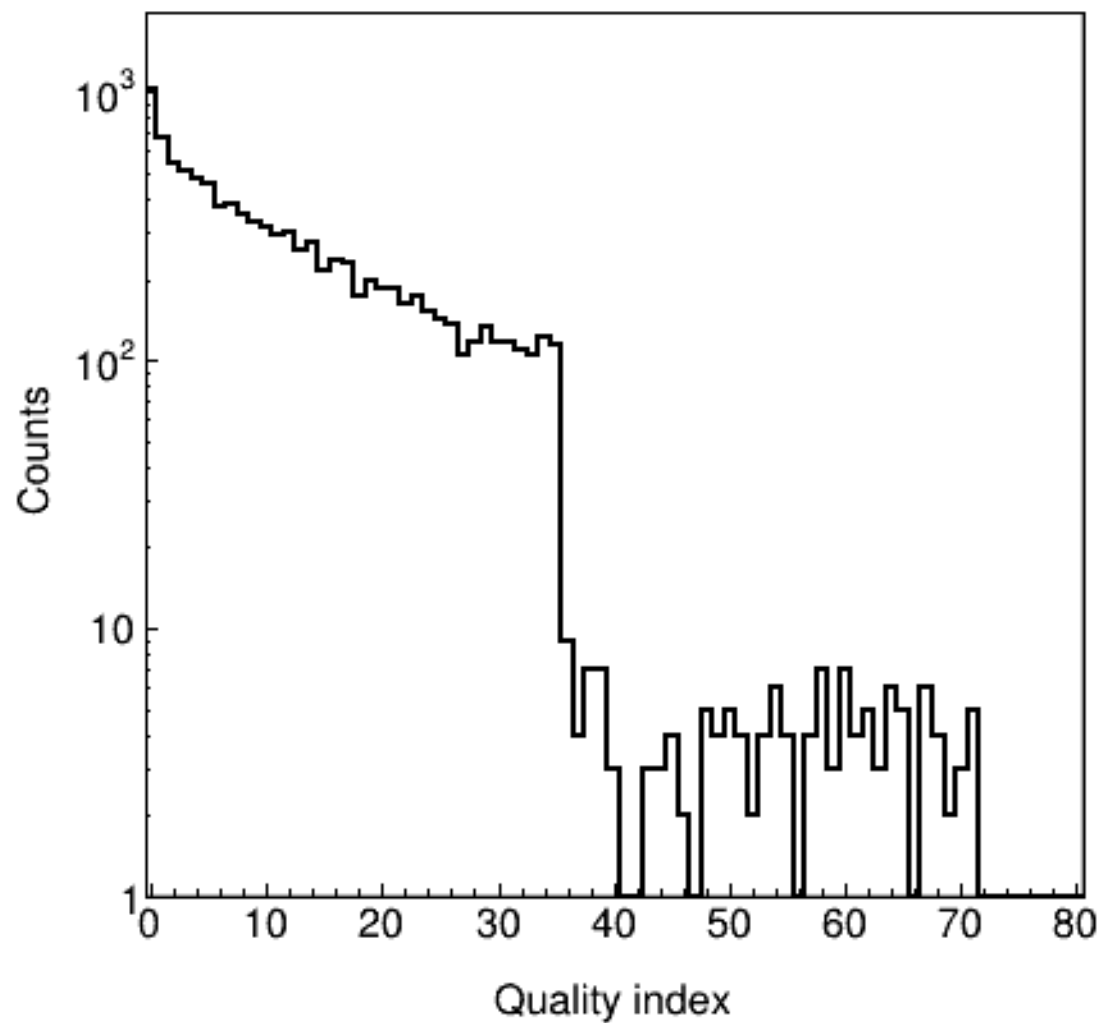
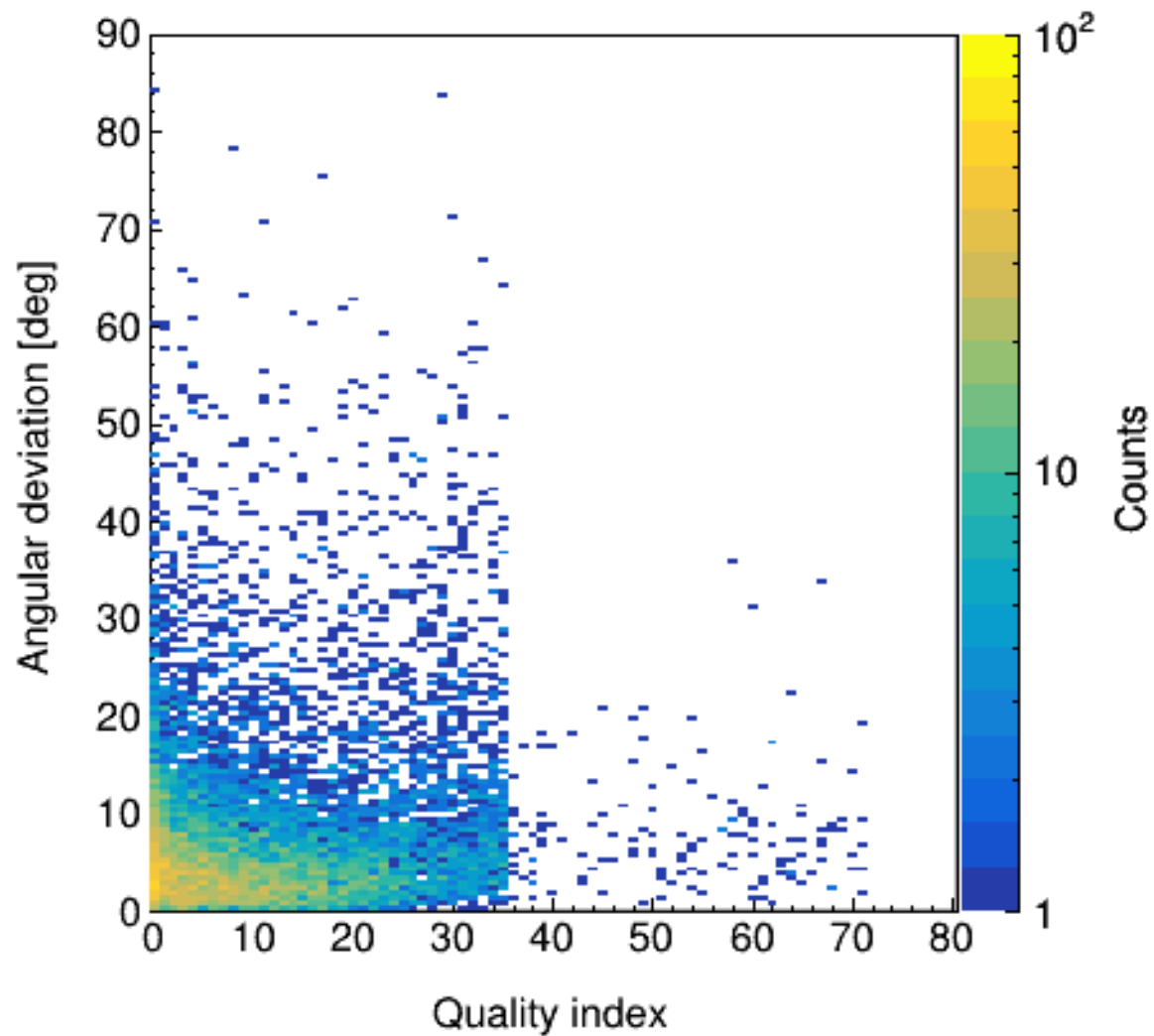


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# Directionality performance



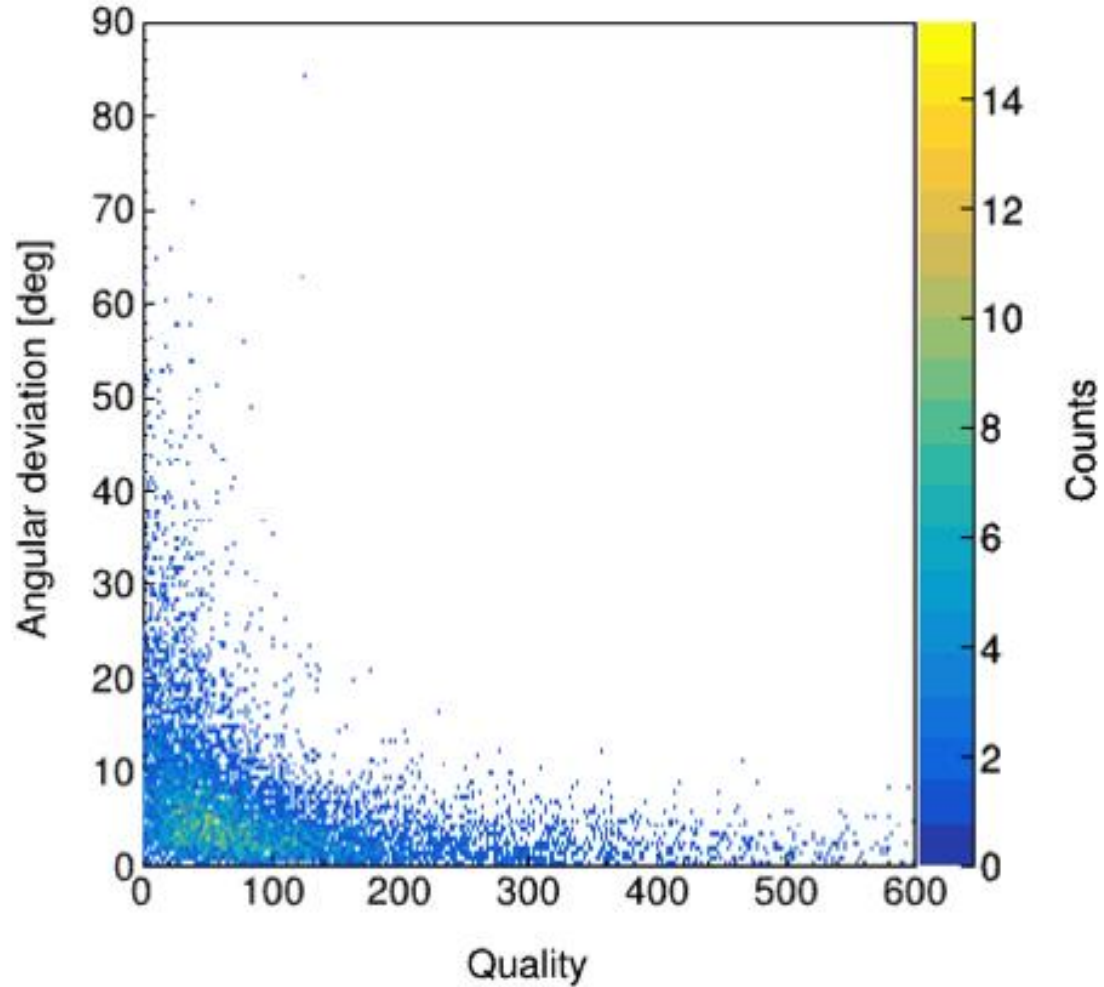
Decreasing quality



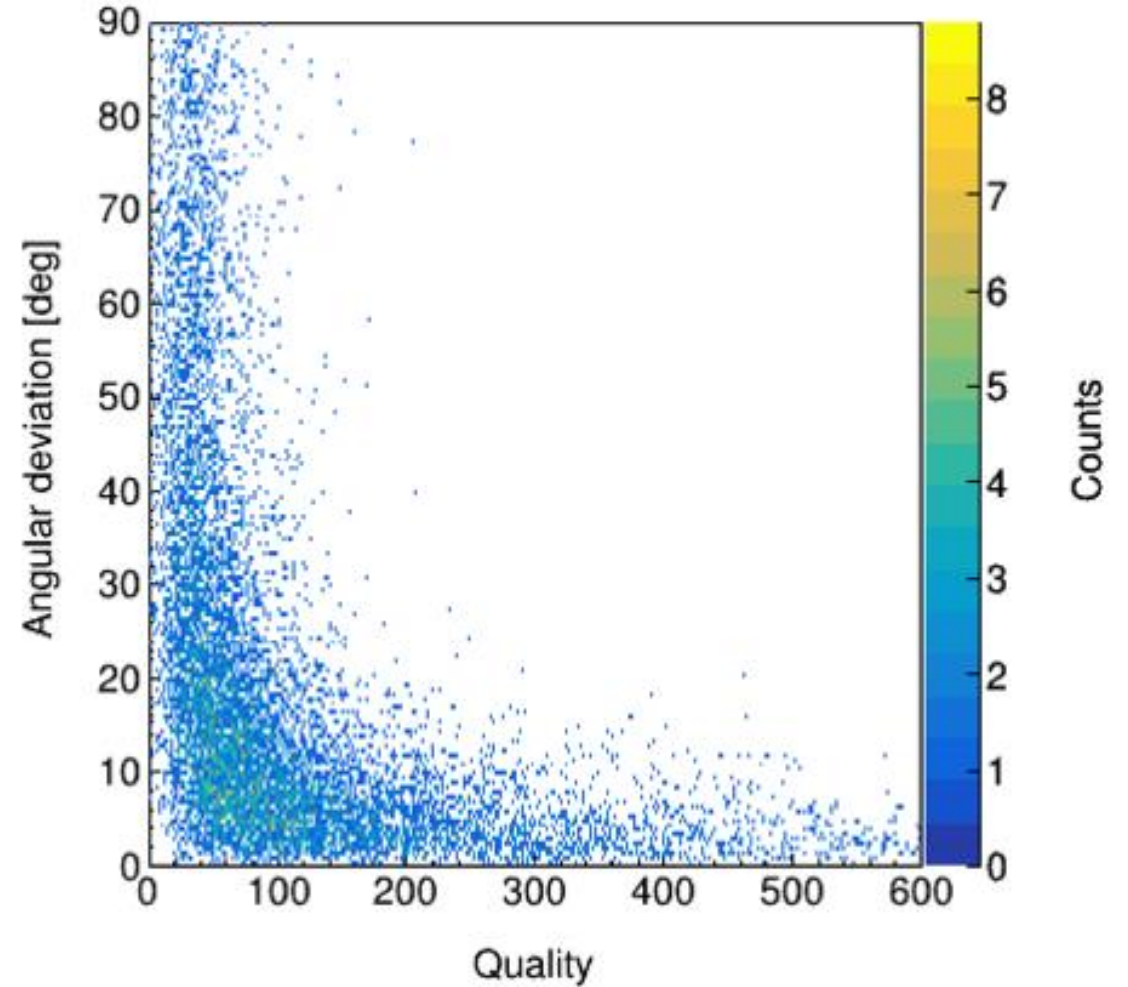
$$Q = -\ln \left( \prod_{i=0}^N \frac{S_i}{B_i} \right)$$

# Directionality performance

## Best angular deviation




## Best quality



## No-hit inclusion

- Best quality fits do not always yield smallest angular deviation
- But we are not exploiting all available topological information
  - No-hits also contain information!
- Not implemented previously for 2 reasons:
  1. Muon track length not yet accounted for previously
  2. For high energies in ARCA, no-hits are negligible

# Plan

- Focus on ORCA muon neutrino MC-data first
- Use JMuonGandalf and JEnergy as stepping stones
  1. Loop over hits  loop over PMTs
  2. Include computation expected number of hits (using Jpp muon PDFs)
  3. Determine no-hit likelihood
    - Including threshold inefficiency
  4. Multiply hit and no-hit likelihoods
- Compare fit results with and without no-hit information