

UnivEarthS

Cosmic rays in local clouds

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The Local Medium

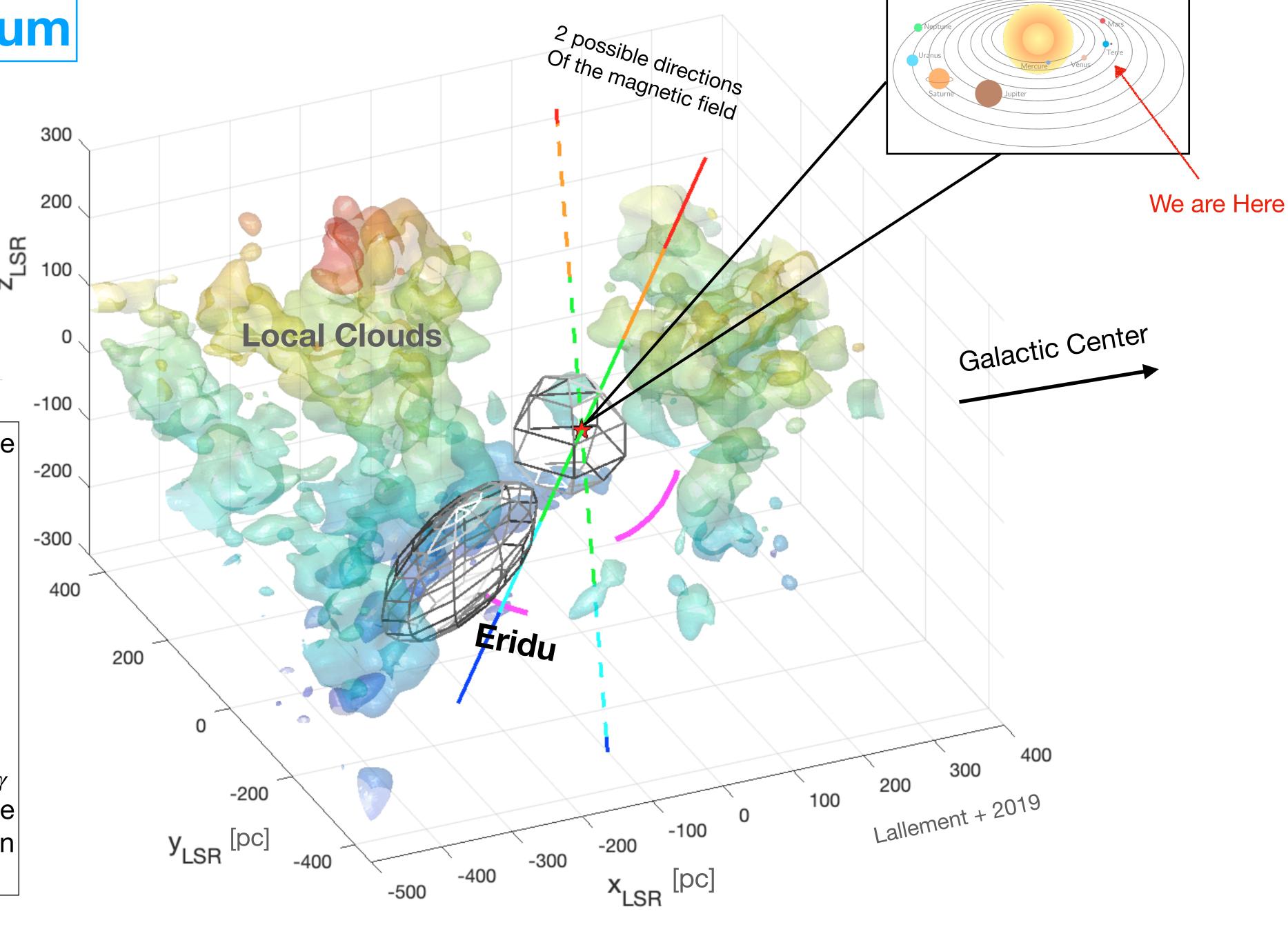
Galactic Plane

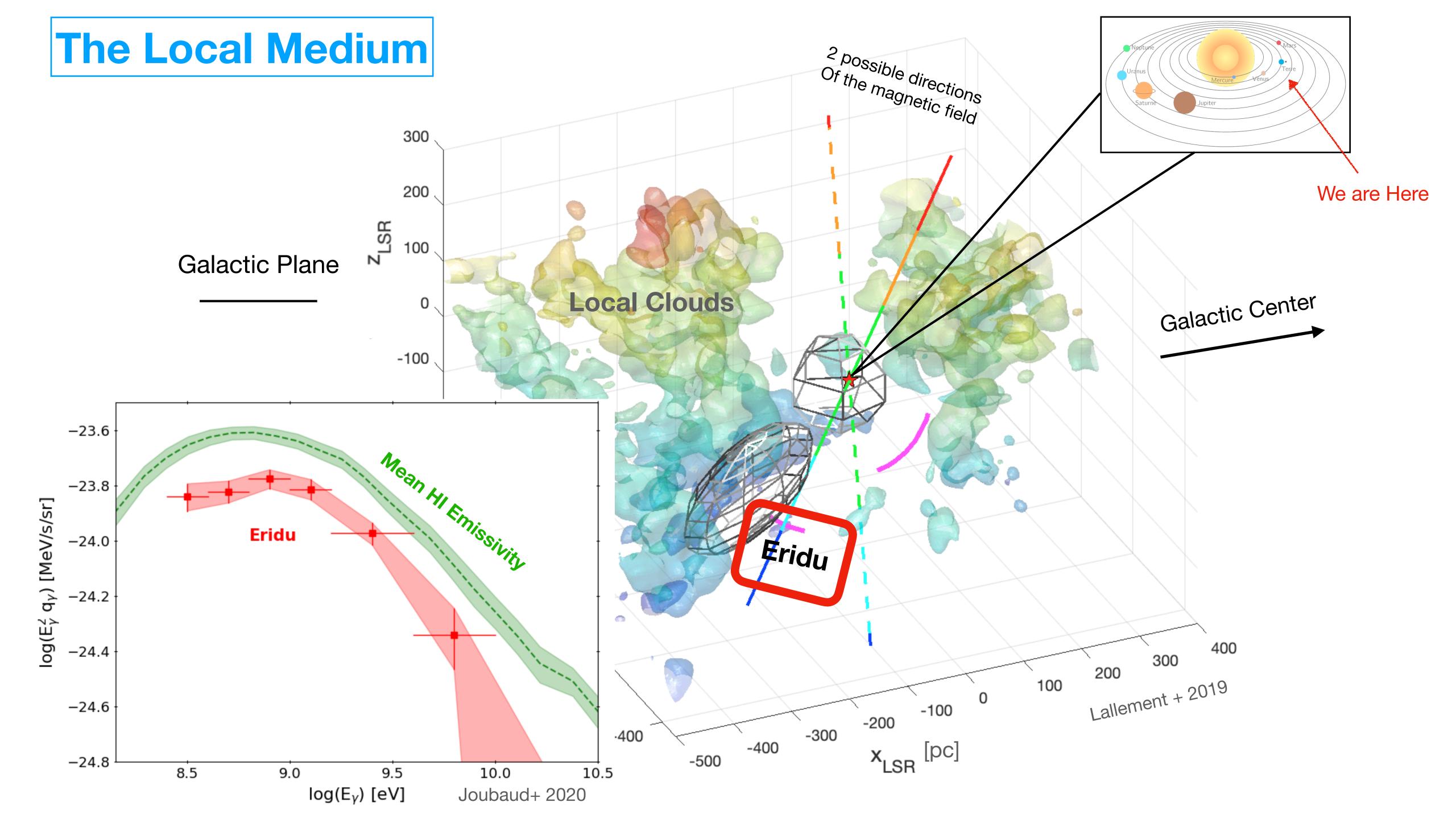
- 3D gas distribution around the Solar System
- Red : Above Galactic plane
- Bleu: Below Galactic plane

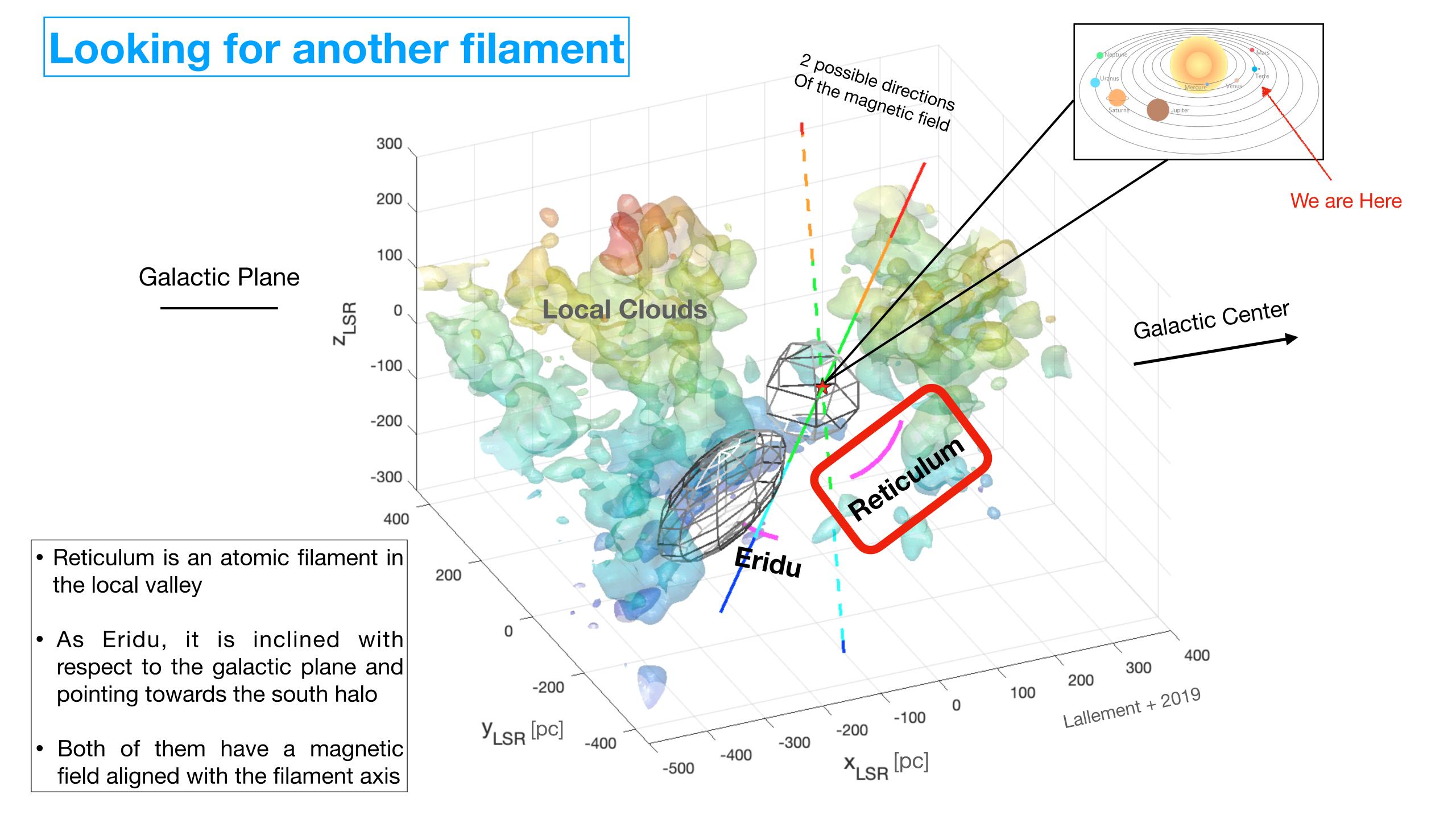
$$I_{\gamma} \propto \int n_{gas} n_{CR} dl$$

$$I_{\gamma} \approx q_{\gamma} \int n_{gas} dl$$

• Several measurements of q_{γ} showed that local clouds are consistent with the mean emissivity in the local medium

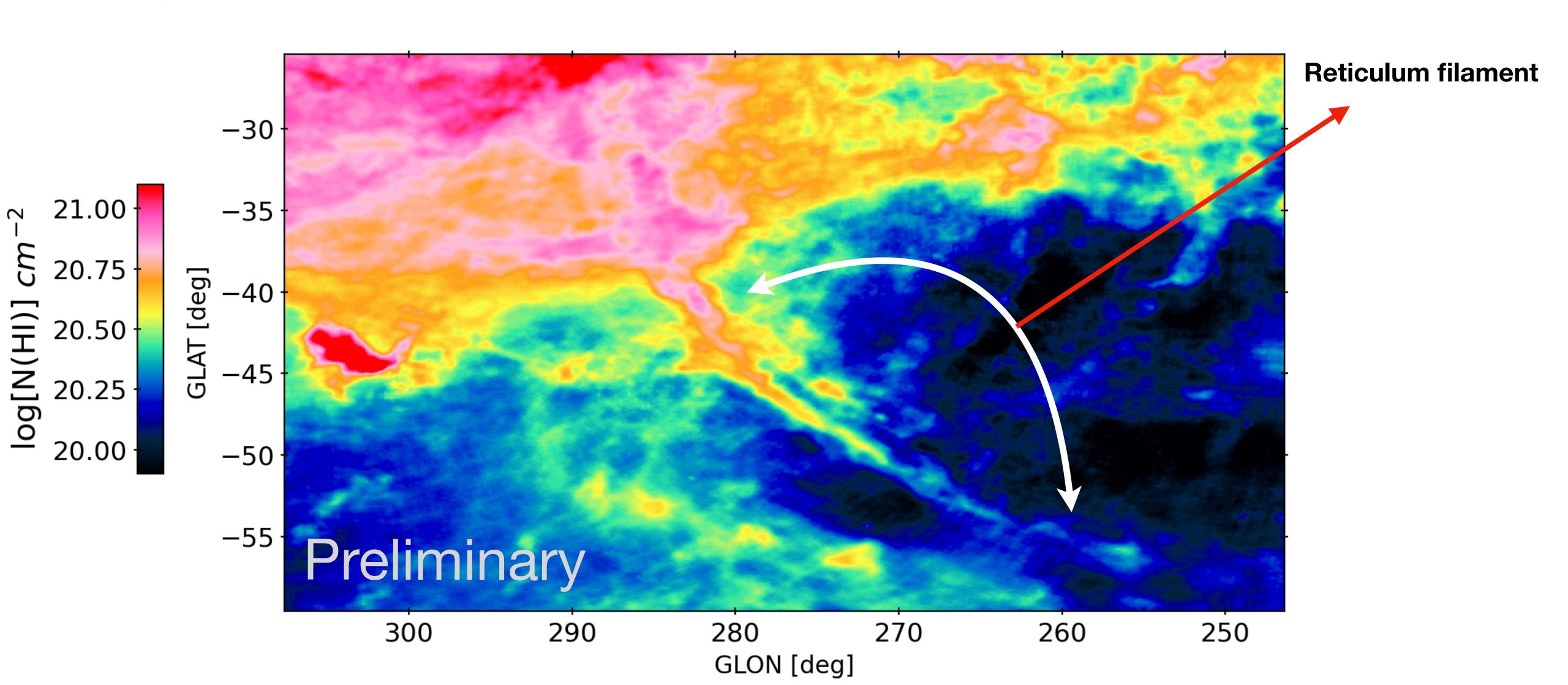




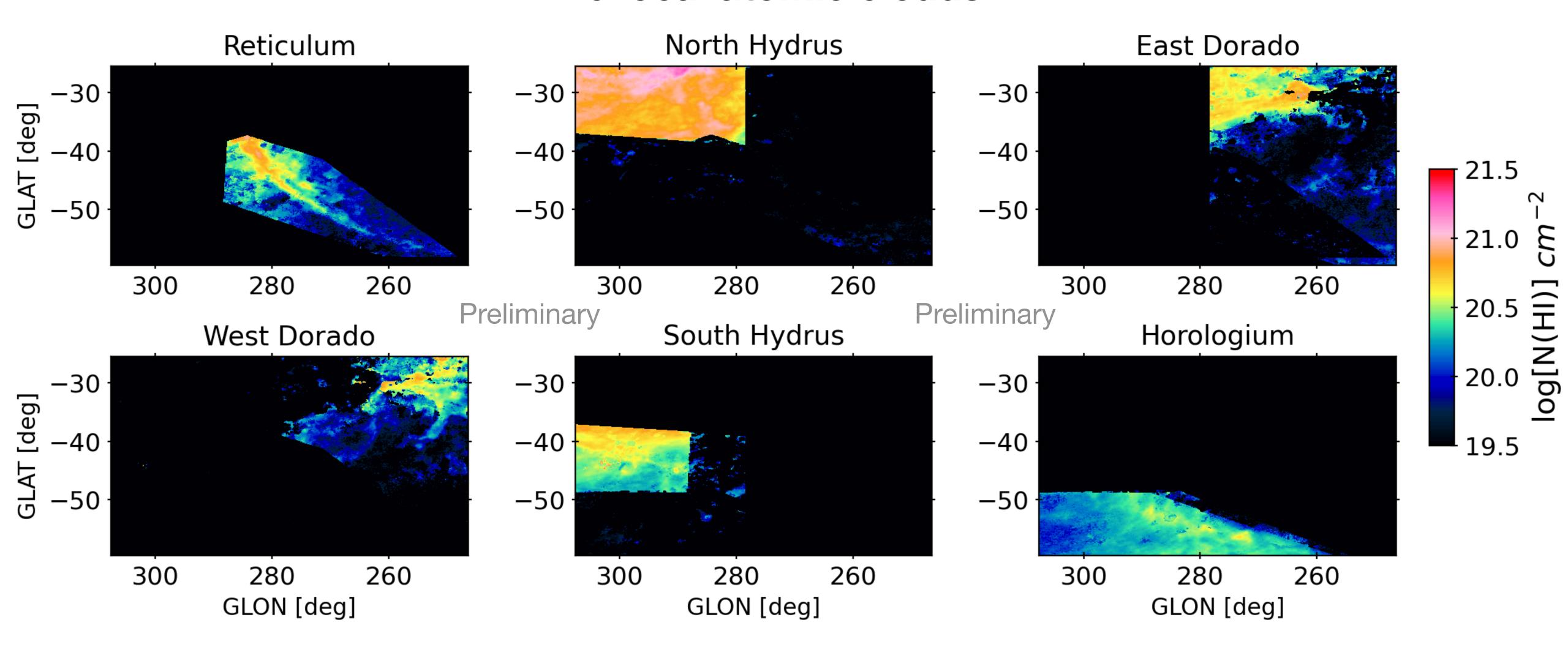


First, we need to measure the quantity of gas

Using GASS data at 21cm, we obtained the HI column density map



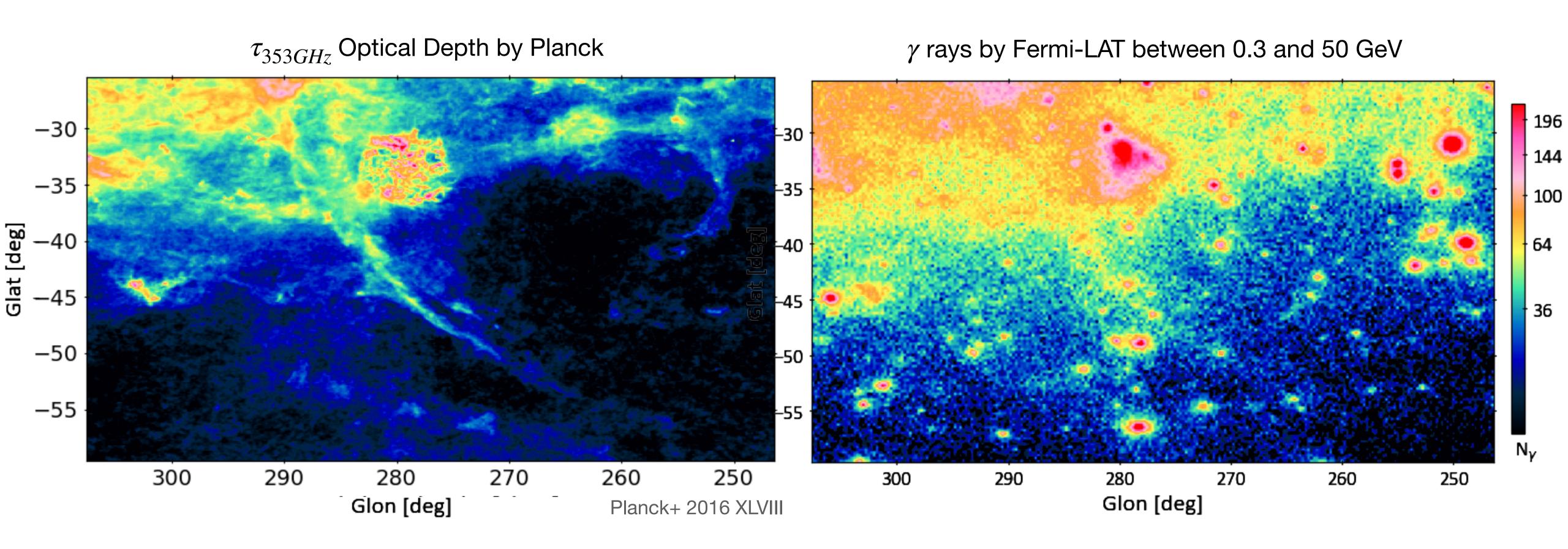
6 local atomic clouds



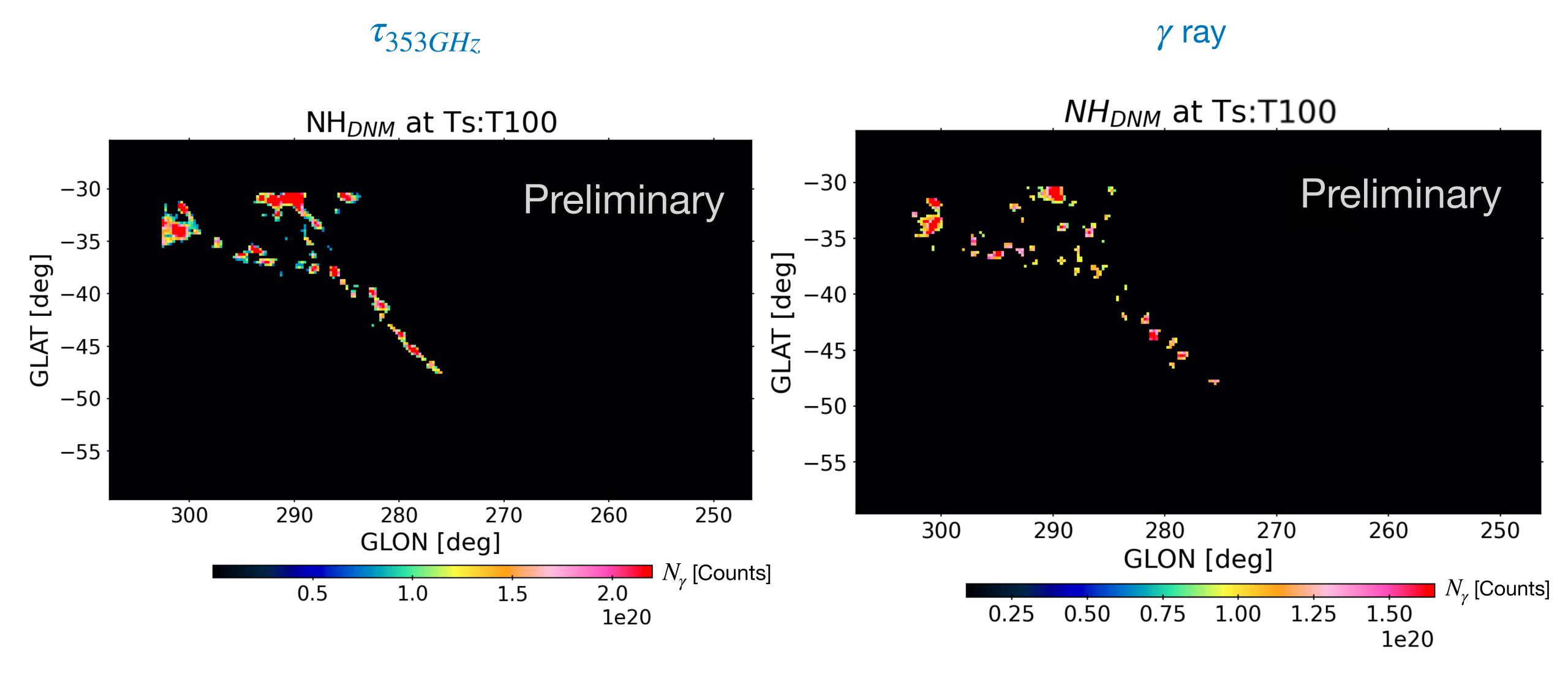
- Decomposition in space and velocity
- N(HI) derivation for optically thin HI or for self-absorption corrections with : $100K \le T_{spin} \le 800K$

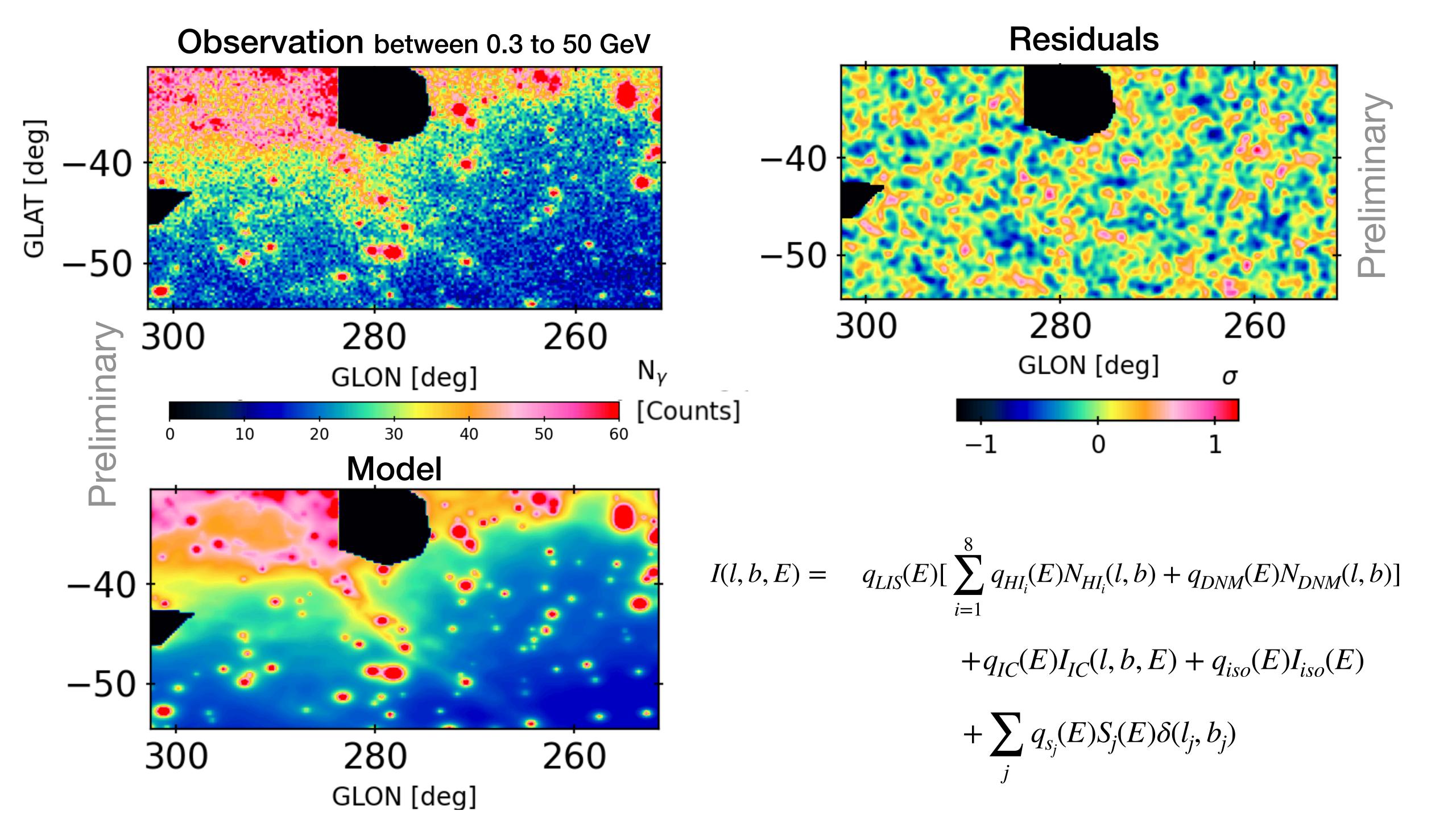
Is there additional gas?

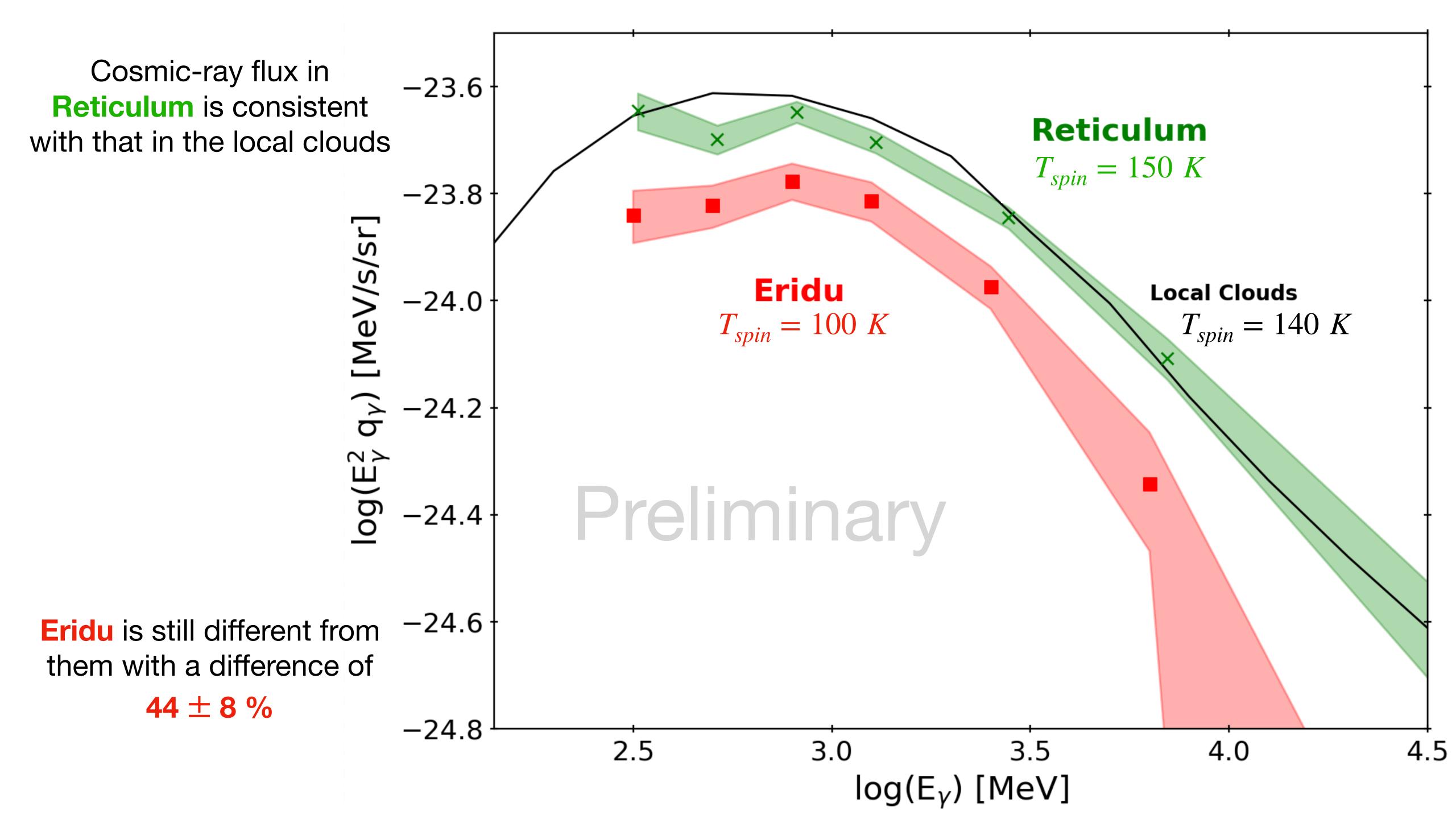
- No detected CO emission
- Dark Neutral Gas : obtained by coupling the analyses of dust and γ rays as tracers of the total gas quantity



Dark Neutral Medium extracted from dust and γ rays







Why 44 \pm 8 % more CRs in Reticulum compared to Eridu ?

