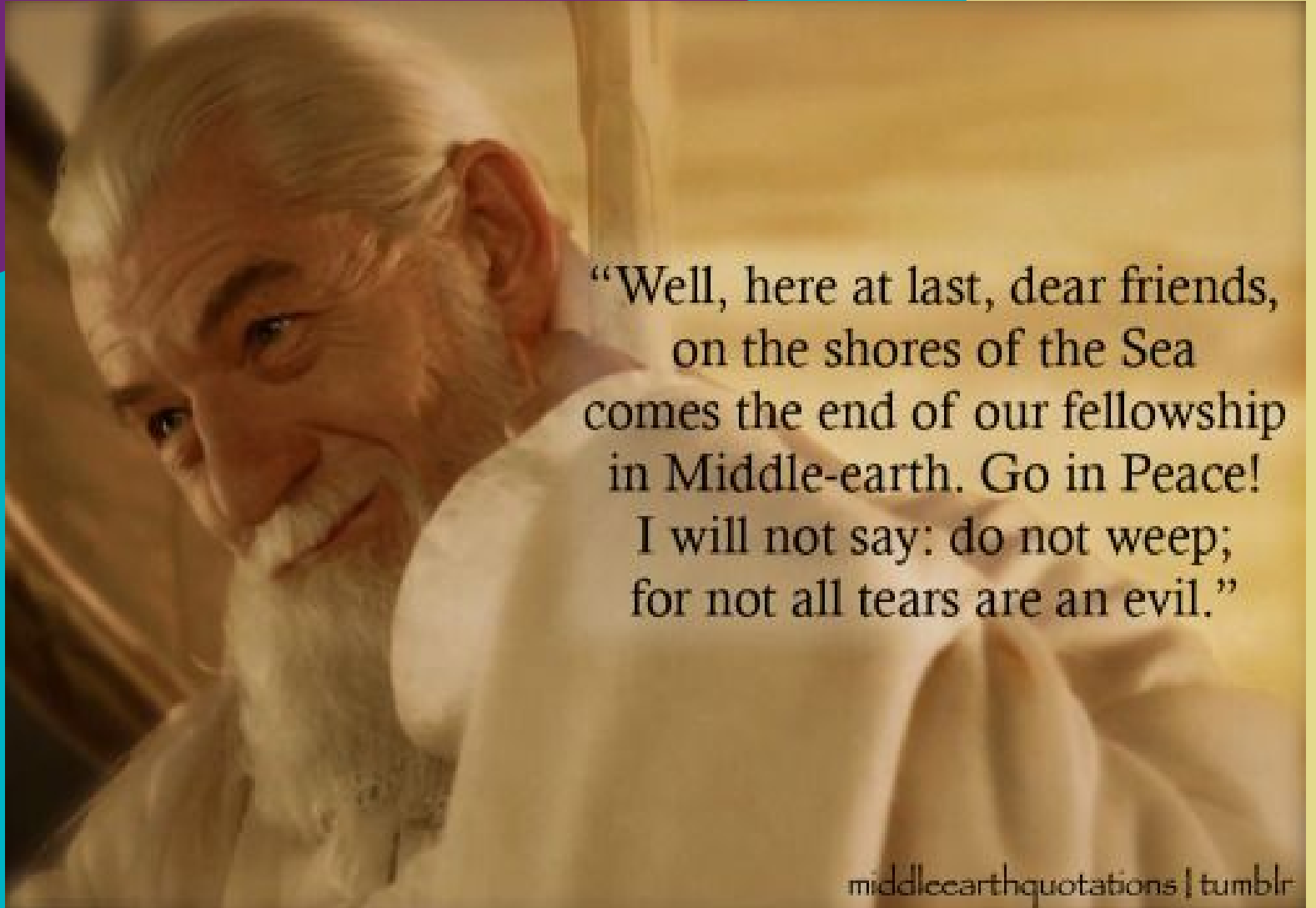


# GROUPMEETING

Rasa Muller

Nikhef



“Well, here at last, dear friends,  
on the shores of the Sea  
comes the end of our fellowship  
in Middle-earth. Go in Peace!  
I will not say: do not weep;  
for not all tears are an evil.”

[middleearthquotations](#) | [tumblr](#)

THANKS BRIAN!

# TRY NEW SETUP



Week 1: regular groupmeeting

Week 2: **paper meeting**

Week 3: regular groupmeeting

Week 4: **junior meeting**

# TRY NEW SETUP



Week 1: regular groupmeeting

Week 2: **paper meeting** ←

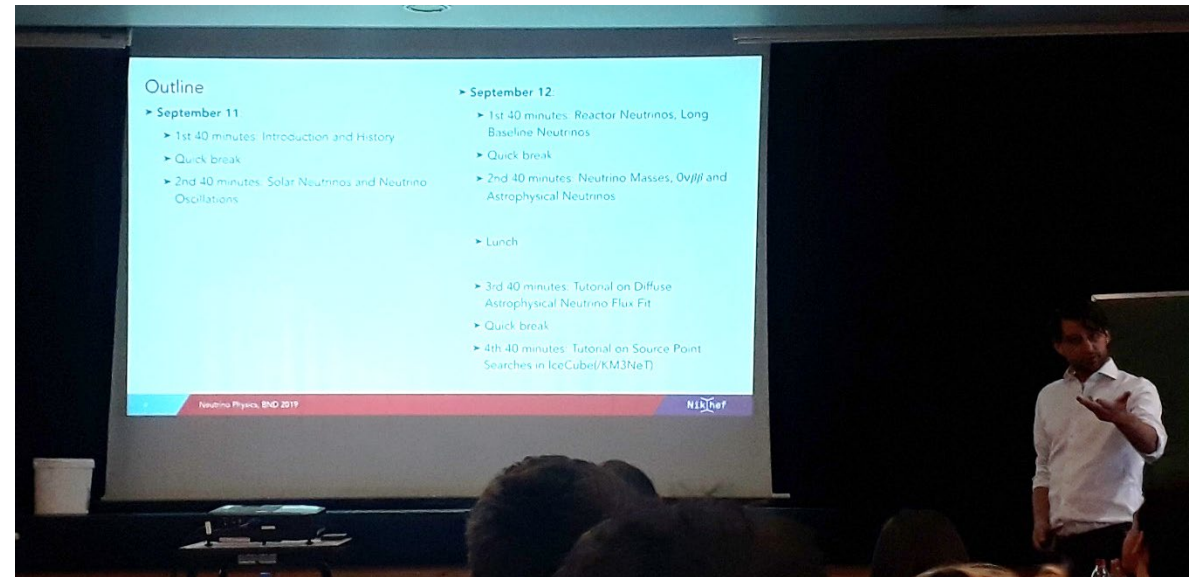
Week 3: regular groupmeeting

Week 4: **junior meeting**

# BND

- Quantum field theory
- Electroweak theory
- Gravitational waves
- Electroweak experiments
- Tracking
- Flavour physics
- Long lived particles
- Cosmology
- Neutrino physics

Daan



# CATALOG SEARCHES

- 1) What sources emit High Energy neutrinos?  
What other signals do they emit?
- 2) Which experiments/telescopes detect all these signals?
- 3) How is this info stored in catalogs?
- 4) What are the most relevant catalogs? How complete are they (upto which redshift)?

# CATALOG SEARCHES

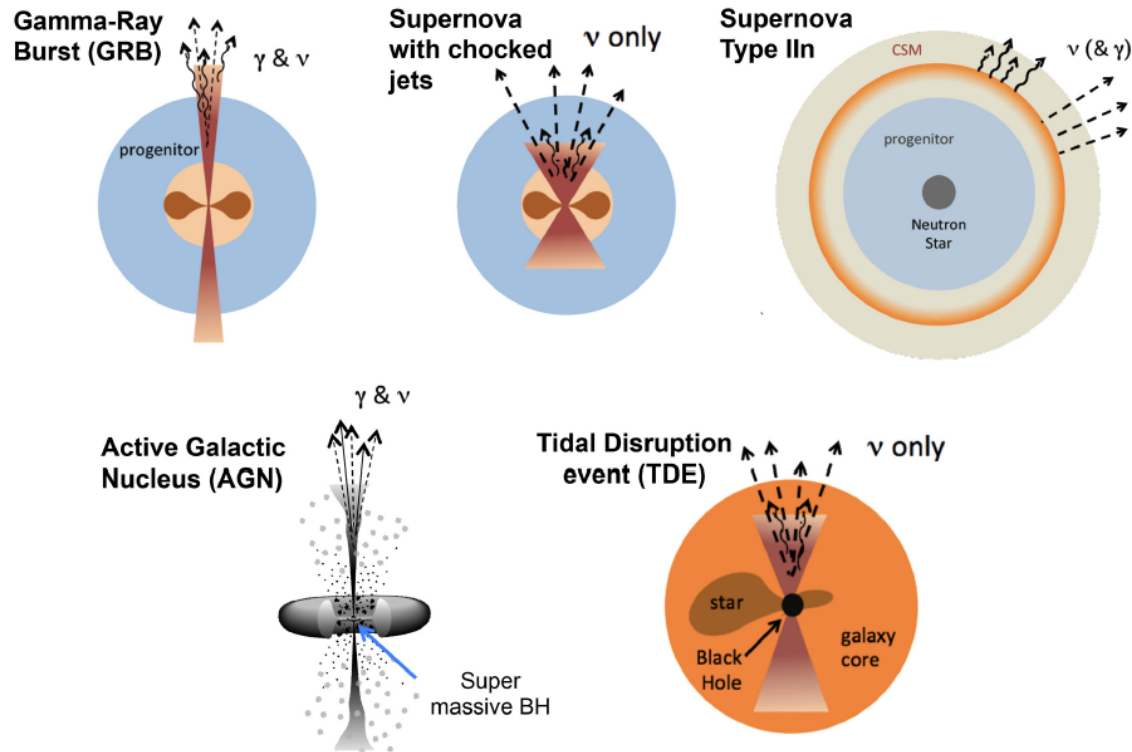


Figure 2. Scenarios for sources of neutrinos, with varying degrees of jet formation.

## Multimessenger astronomy:

The exploration of the Universe through combining information from a multitude of cosmic messengers

- I. cosmic rays
- II. gravitational waves
- III. electromagnetic radiation
- IV. neutrinos

1) What sources emit High Energy neutrinos?  
What other signals do they emit?

# CATALOG SEARCHES

## I. Cosmic rays

→ Detector for atm particle showers + accompanying fluorescence and Cherenkov radiation

PierreAugier

## II. Gravitational waves

LIGO/VIRGO and others

## III. Electromagnetic radiation

→  $\gamma$ -ray, X-ray unable to cross Earth's atmosphere thus primarily observed using satellites

- Fermi LAT
- Imaging Atmospheric Cherenkov Telescope (IACT) systems
  - i. H.E.S.S.
  - ii. FACT
  - iii. VERI-TAS
  - iv. GFU



2) Which experiments/telescopes detect all these signals?



# CATALOG SEARCHES

- 1) What sources emit High Energy neutrinos?  
What other signals do they emit?
- 2) Which experiments/telescopes detect all these signals?
- 3) **How is this info stored in catalogs?**
- 4) **What are the most relevant catalogs?**  
**How complete are they (upto which redshift)?**

- ➔ Generate pseudodatasets
- ➔ Analyse pseudo datasets

# CATALOG SEARCHES WITH KM3NET

PHYSICAL REVIEW D **96**, 023003 (2017)

## Prospects of establishing the origin of cosmic neutrinos using source catalogs

I. Bartos,<sup>1,\*</sup> M. Ahrens,<sup>2</sup> C. Finley,<sup>2</sup> and S. Márka<sup>1</sup>

<sup>1</sup>Department of Physics, Columbia University, New York, New York 10027, USA

<sup>2</sup>Oskar Klein Centre & Dept. of Physics, Stockholm University, SE-10691 Stockholm, Sweden

(Received 26 November 2016; published 14 July 2017)

The cosmic neutrino flux recently discovered by IceCube will be instrumental in probing the highest-energy astrophysical processes. Nevertheless, the origin of these neutrinos is still unknown. While it would be more straightforward to identify a transient, or galactic source, class, finding a population of distant, continuous sources is challenging. We introduce a source-type classification technique that incorporates all available information from catalogs of source candidates. We establish the origin of cosmic neutrinos, even for the most challenging AGN, or galaxy clusters—if neutrino track directions can be reconstructed. We show that the source catalog out to  $\sim 100$  Mpc can be sufficient for identifying sources, allowing for more straightforward source surveys. We also characterize the angular resolution, size, and veto power in order to understand the

DOI: 10.1103/PhysRevD.96.023003

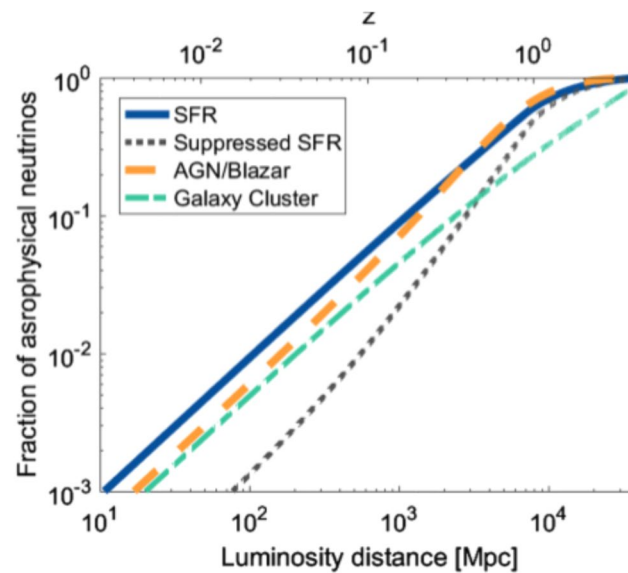


FIG. 1. Fraction of astrophysical neutrinos detected from sources within a luminosity distance, as a function of the luminosity distance, for different cosmic evolution models (see legend and Sec. II E). For comparison, the top axis shows the corresponding redshift.

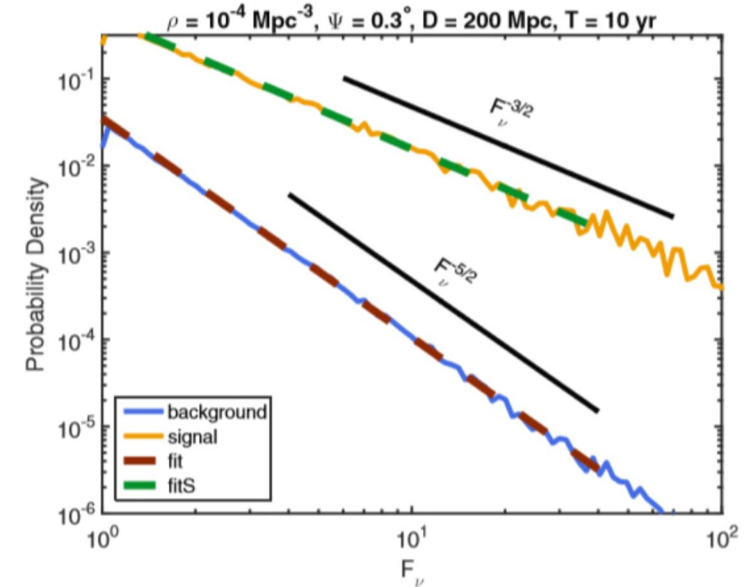


FIG. 2. Simulated distribution of neutrino flux  $\mathcal{F}_\nu$  corresponding to signal and background neutrinos. This example shows the densities for  $\psi = 0.3^\circ$ ,  $d_{\text{th}} = 200$  Mpc, for starburst galaxies. The dashed red line shows a power-law fit on the background density's tail. The black solid lines show power-law slopes with  $\mathcal{F}_\nu^{-3/2}$  and  $\mathcal{F}_\nu^{-5/2}$ , the theoretical expectations for the signal and neutrino models, respectively, for the case in which the expected number of sources within  $d_{\text{th}}$  coincident with a neutrino is  $\ll 1$ .