



UNIVERSITY OF AMSTERDAM



# Models of Particle Signatures in KM3NeT ORCA

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Photo by Blaque X from Pexels

- Introduction
- Procedure
- Models
- Reconstructions
- Conclusion

# Introduction

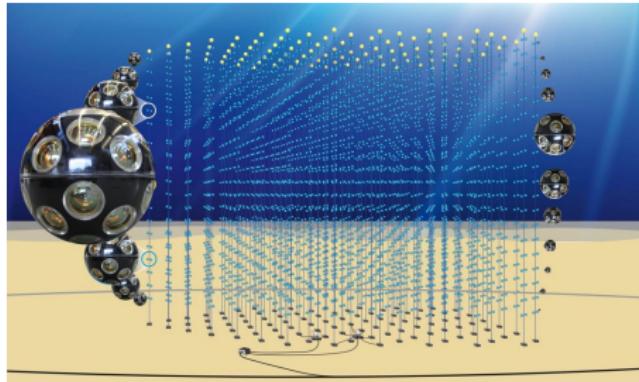
$\mathcal{O}(\text{GeV})$  detector, KM3Net  
**ORCA**

Dense compared to ARCA

68 000 PMTs.

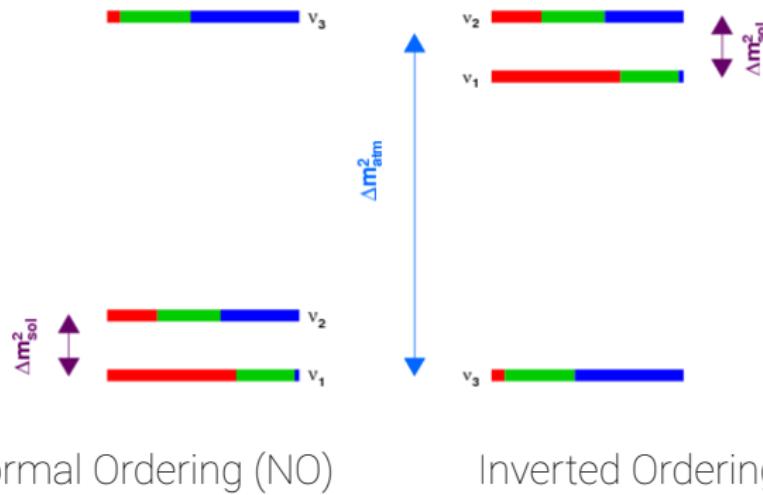
$0.004 \text{ km}^3 = 4 \text{ MegaTonnes}$   
=  $80 \times$  Super-Kamiokande.

**very** cool multi-directional detection modules.



# Introduction

KM3NeT ORCA's goal: **What is the Neutrino Mass Ordering (NMO)?**



<sup>1</sup>[Gomez-Cadenas et al., 2012]

# Introduction

Here are the parameters necessary to predict the oscillation probability of a neutrino through matter:

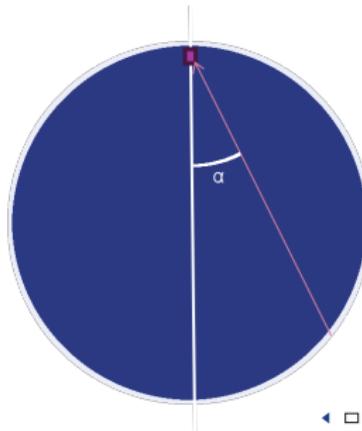
- **Oscillation parameters**
- **The number of electrons in the neutrino's path**
- **Energy of the neutrino**
- **Flavor of the neutrino**
- **Neutrino Mass Ordering (NMO)**

$$P_{3\nu} m(\nu_\mu \rightarrow \nu_\mu) \simeq 1 - \sin^2 2\theta_{23} \cos^2 \theta_{13}^m \sin^2 \left( \frac{AL}{4} + \frac{\Delta m_{31}^2 + \Delta^m m^2}{8E_\nu} L \right)$$

\*

+some other terms

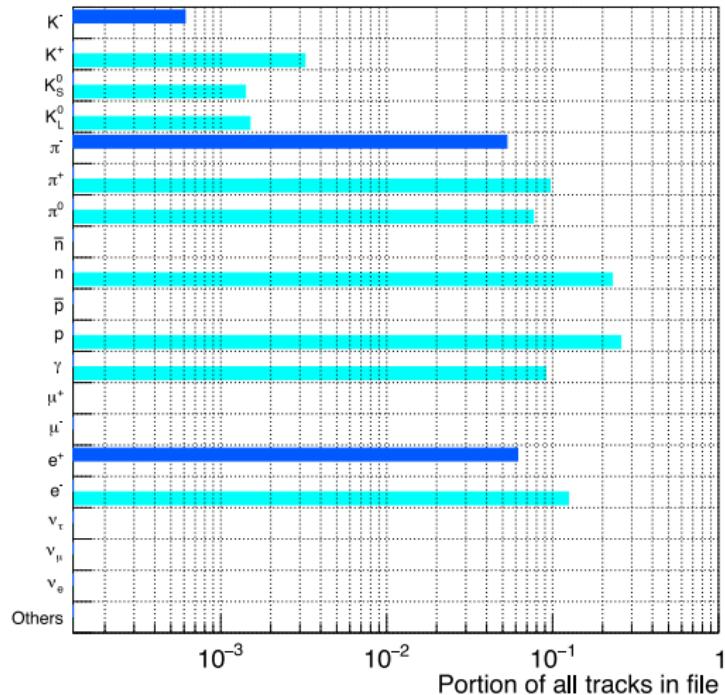
(1)



# Introduction

Atmospheric neutrinos can interact in water and create *product particles*.

3-5 GeV  $\nu_e$  Charged Current interaction products

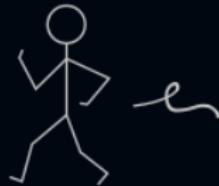


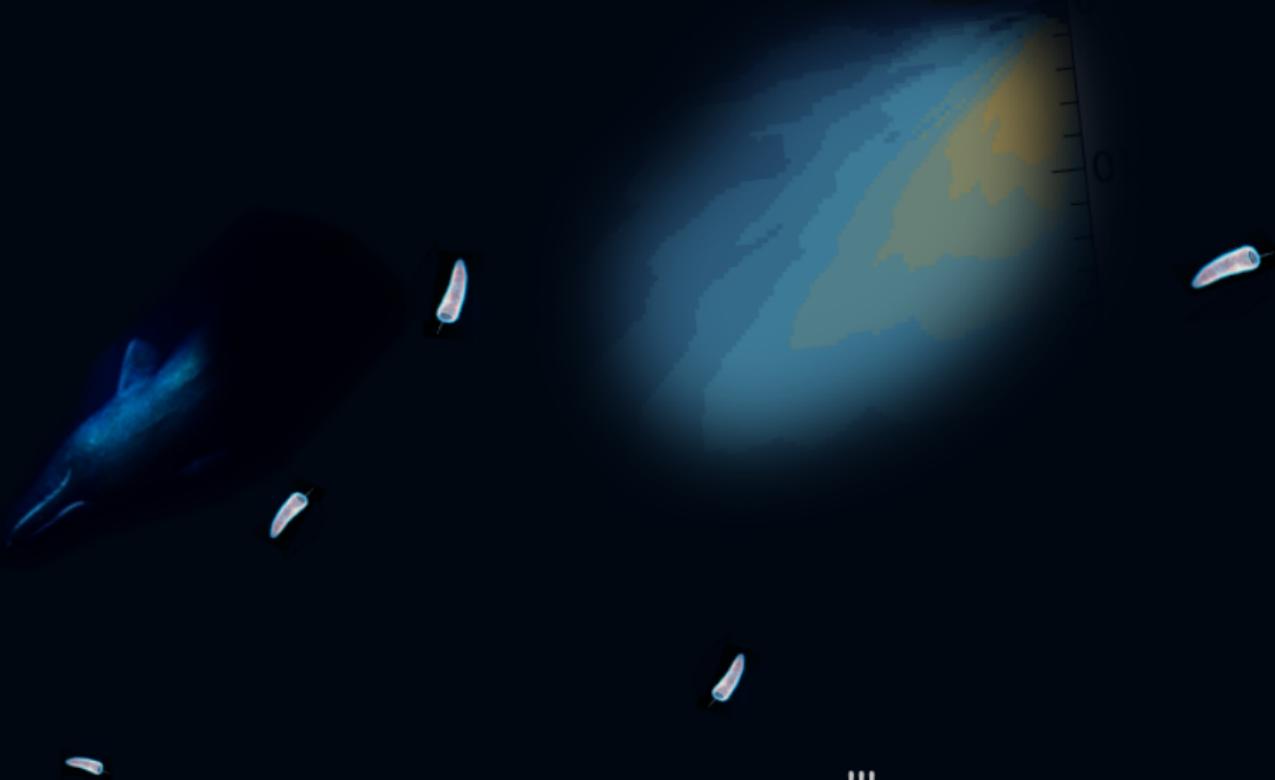
Neutral current →  
**hadronic particles.**

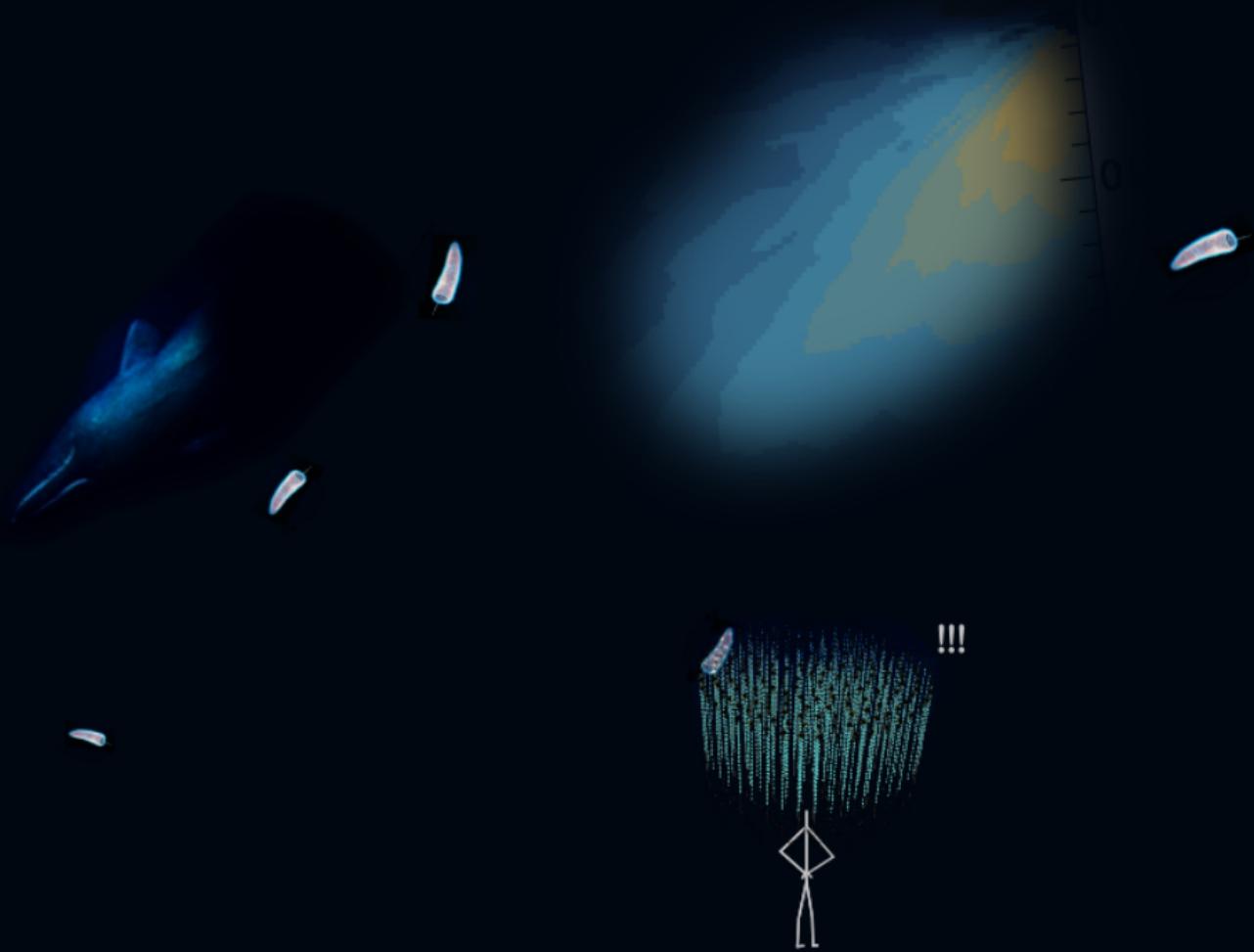
Charged current →  
**lepton + hadronic  
particles**



Image of Pyrosome courtesy of Jeff Millisen



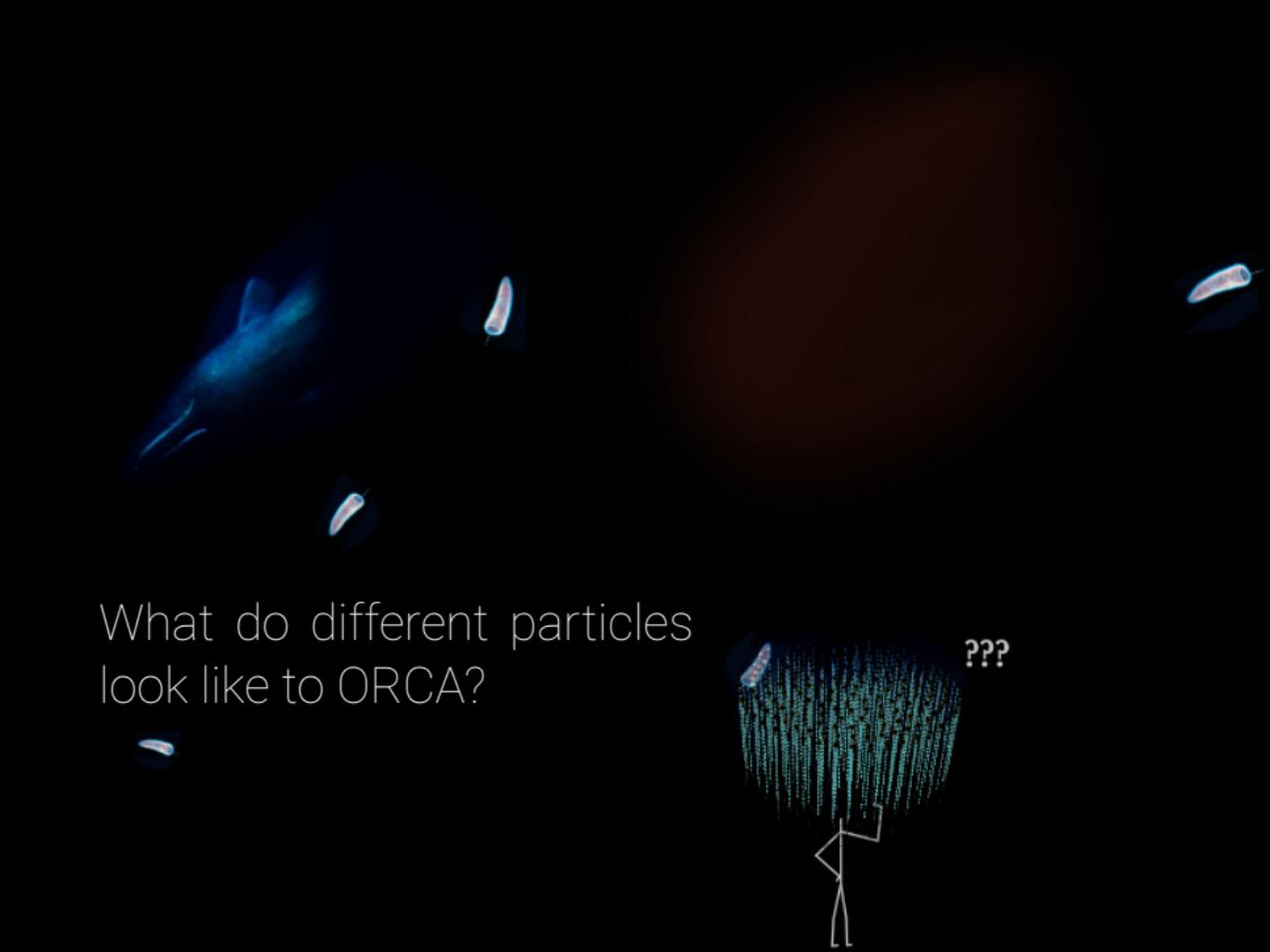






???





What do different particles  
look like to ORCA?



# The plan

1. Create models of **light signatures** in ORCA from  $\nu$  interaction **product particles**.
2. Use models for reconstruction of particle showers and  $\nu$  events.

# Procedure

Monte Carlo sample of signal from energetic particles in ORCA.

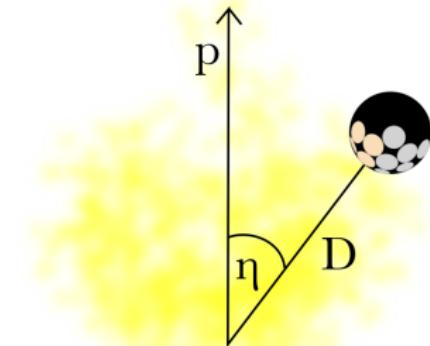
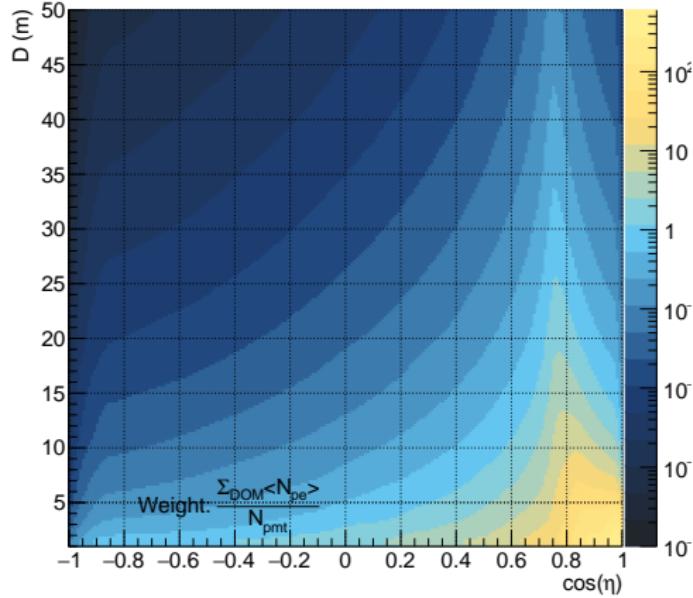


Create **photo-electron Pattern PDFs (PEPPs)** from sample for each particle type.

The PEPP tells you **the probability of obtaining a p.e. given particle type and position in phase space**

*"It's basically an interpolation of our particle interaction modelling in water." – me*

# PEPPs Geometry

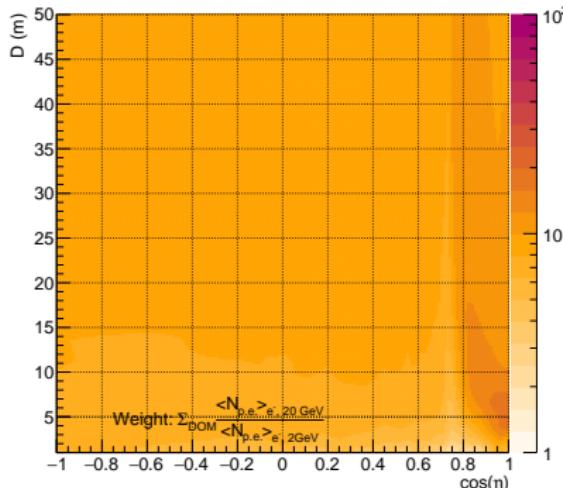


- In water,  $\cos \theta_{\text{Cher.}} \simeq 0.75$  for  $\beta \simeq 1$ , explaining the peak.
- For 30 GeV electron, expect an EM shower of  $\sim 6$  m.

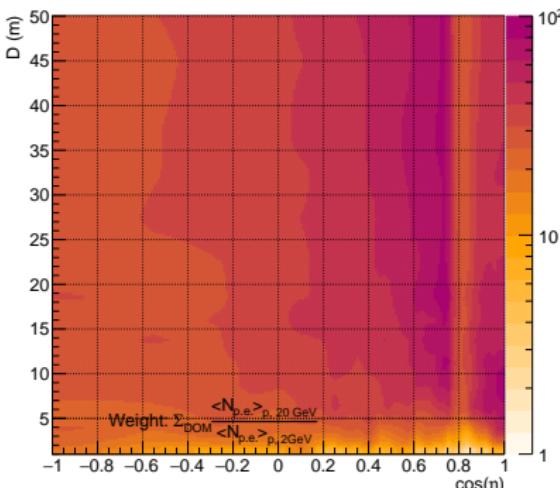
**PEPPs agree with theoretical expectations**

# PEPPs Energy Dependence

Left (right):  $\frac{20 \text{ GeV}}{2 \text{ GeV}}$  electron (proton)



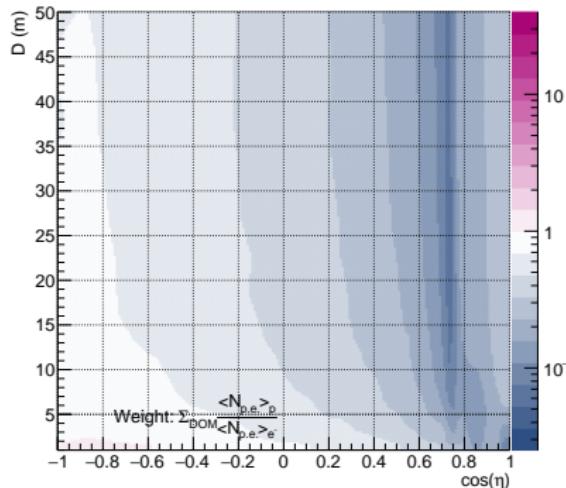
Uniform increase of  
 $10 \times N_{\text{particles}}$



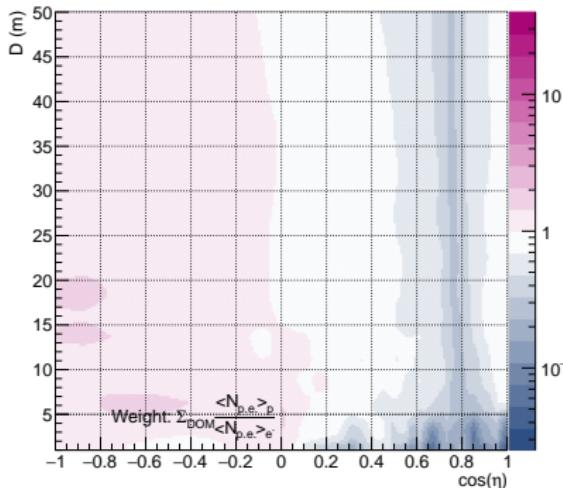
Features near Cherenkov angle  
(note: log scale)

# PEPPs EM/Hadronic Comparison

$$\frac{\text{protonPEPP}}{\text{electronPEPP}}$$



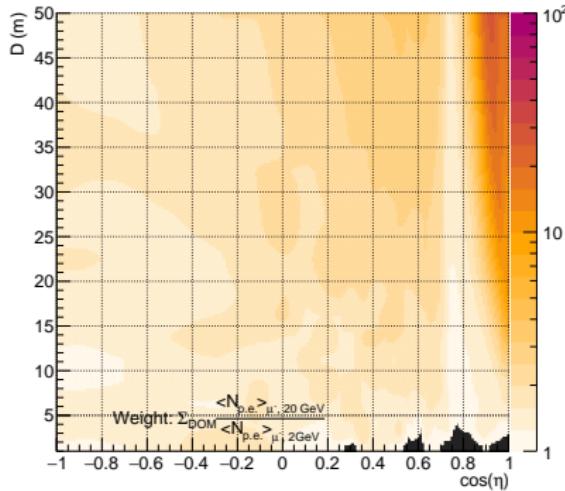
3 GeV: very different



30 GeV: similar  $\rightarrow$  larger portion of EM particles in hadronic shower.

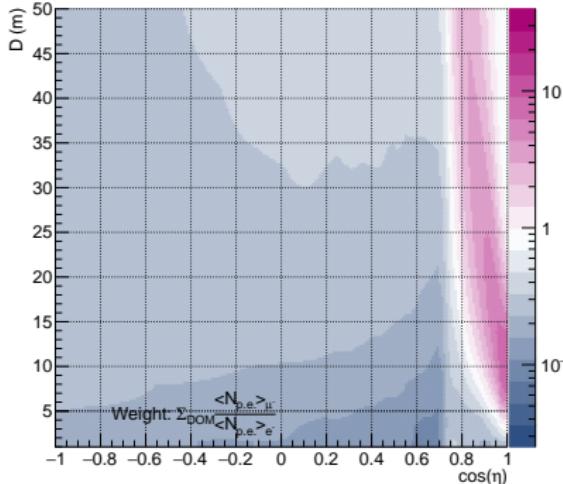
# PEPPs Geometry

$\frac{20 \text{ GeV}}{2 \text{ GeV}}$  muon



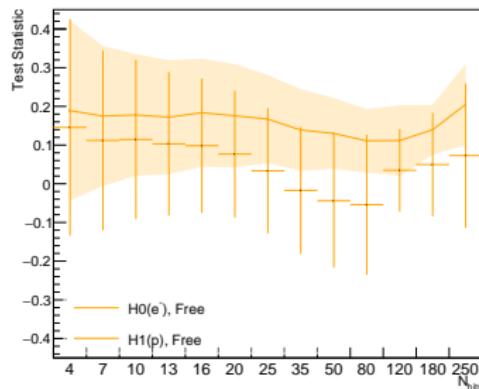
Bulk of increase in very forward direction.

$\frac{\text{muonPEPP}}{\text{electronPEPP}}$  3GeV



Strictly brighter than electron in front of Cherenkov angle.

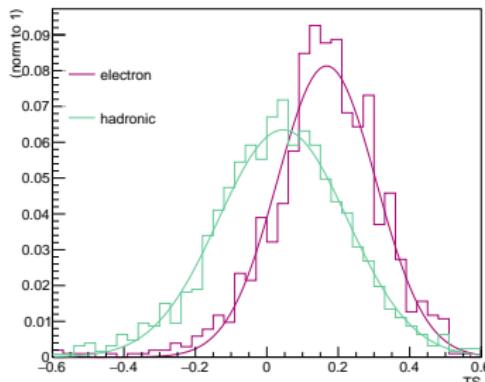
# Preliminary Reconstruction of EM and Hadronic Showers



Test statistic for **single electron (EM shower)** and **hadronic shower**.

Top: 25 - 80 hit region shows promise for  $1\sigma$  separation.

Bottom: Projection along X for 25 - 35 hits.

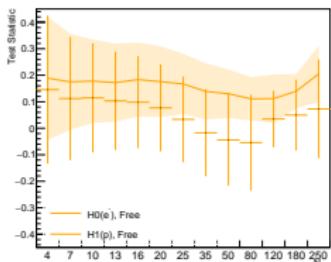
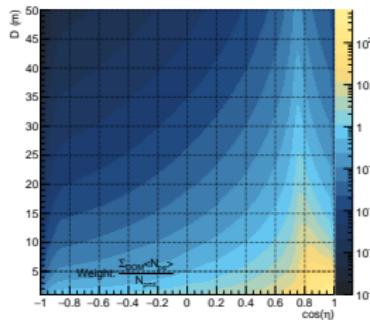


**Distinguish hadronic shower from electron shower possible?**

At higher energies,  
EM shower  $\simeq$  hadronic shower  
 $\rightarrow$  loss of distinguishing power.

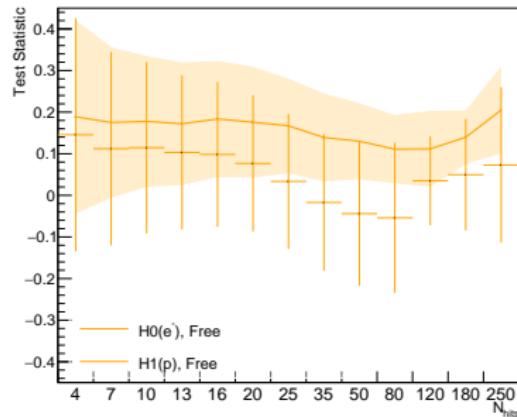
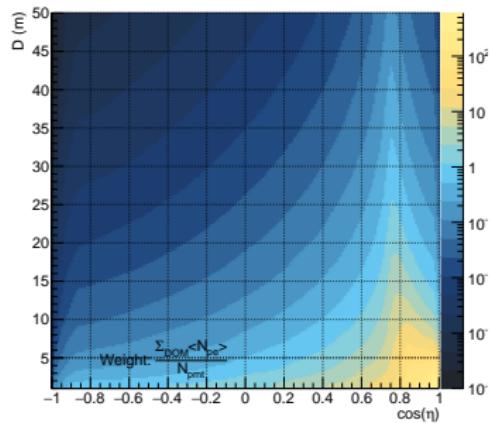
# Conclusion

- KM3NeT **ORCA** aims at measuring the *Neutrino Mass Ordering*.
- Identifying the type of neutrinos and secondary particles assists this goal.
- Produced models of various particles which show their expected signals.
- The models can be put to use for reconstruction of events in the detector, including particle type.



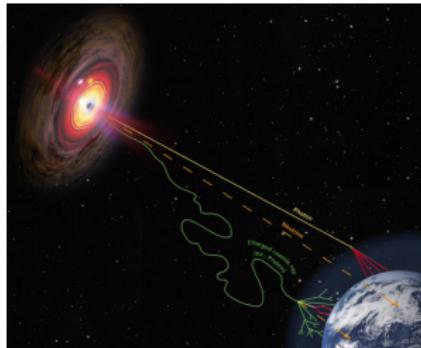
Thank you for your attention!

Questions please.

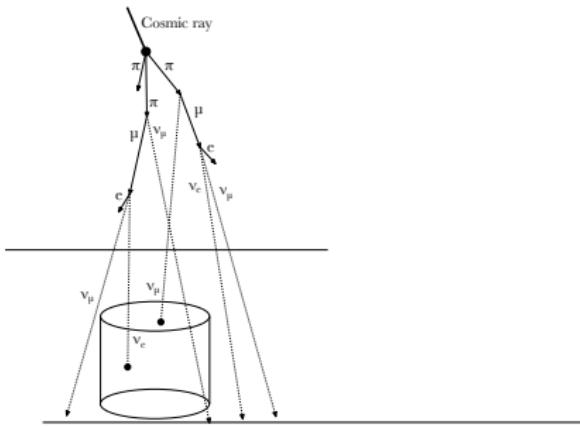


# Leftover slides

# Introduction



3



<sup>3</sup> Artist's impression by HAP / A. Chantelauze

<sup>4</sup> [Mollerach and Roulet, 2018][Honda et al., 1995]

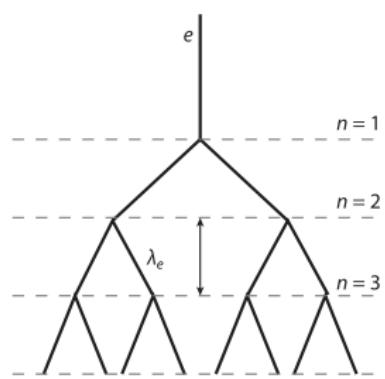
Certain extrasolar objects accelerate particles to high energies, which are called *cosmic rays*.

Cosmic rays collide with the Earth's atmosphere and produce neutrinos.<sup>4</sup>

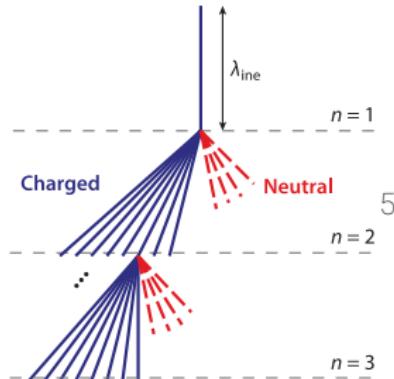
Atmospheric neutrinos can travel through the entire Earth virtually unaffected, causing a ubiquitous flux.

# Introduction

Product particles from neutrino interactions produce more particles in *showers*



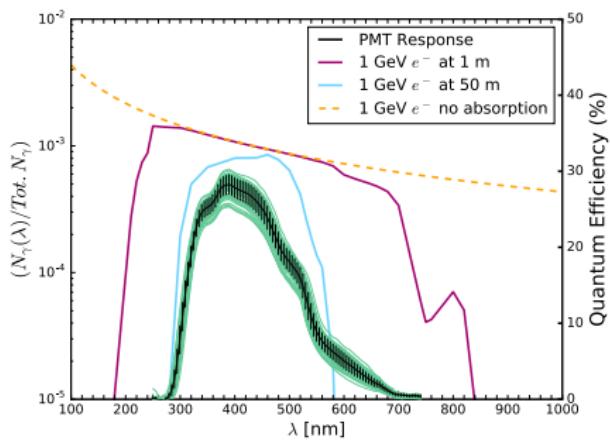
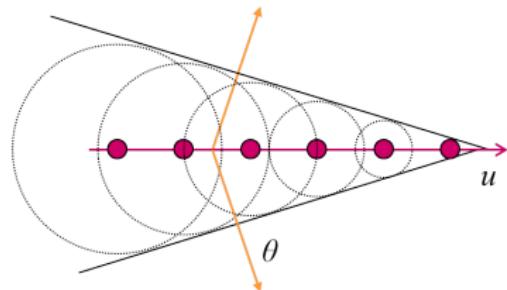
$$\begin{aligned} & \vdots \\ & \text{EM shower} \\ & N_{\text{particles}} \propto E \\ & X \propto \log E \end{aligned}$$



**Hadronic shower**  
**Complicated!**

<sup>5</sup>Diagram from [Engel, 2011]

# Introduction

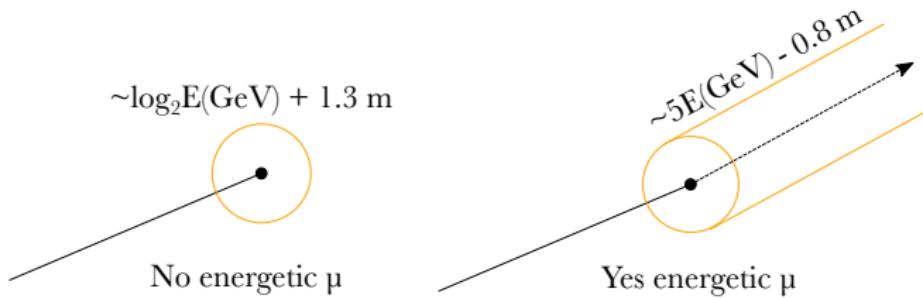


Charged particles emit *Cherenkov light* in water.

This light is emitted at an angle  $\theta$ , and can be seen by *Photomultiplier tubes* (PMT)<sup>6</sup>

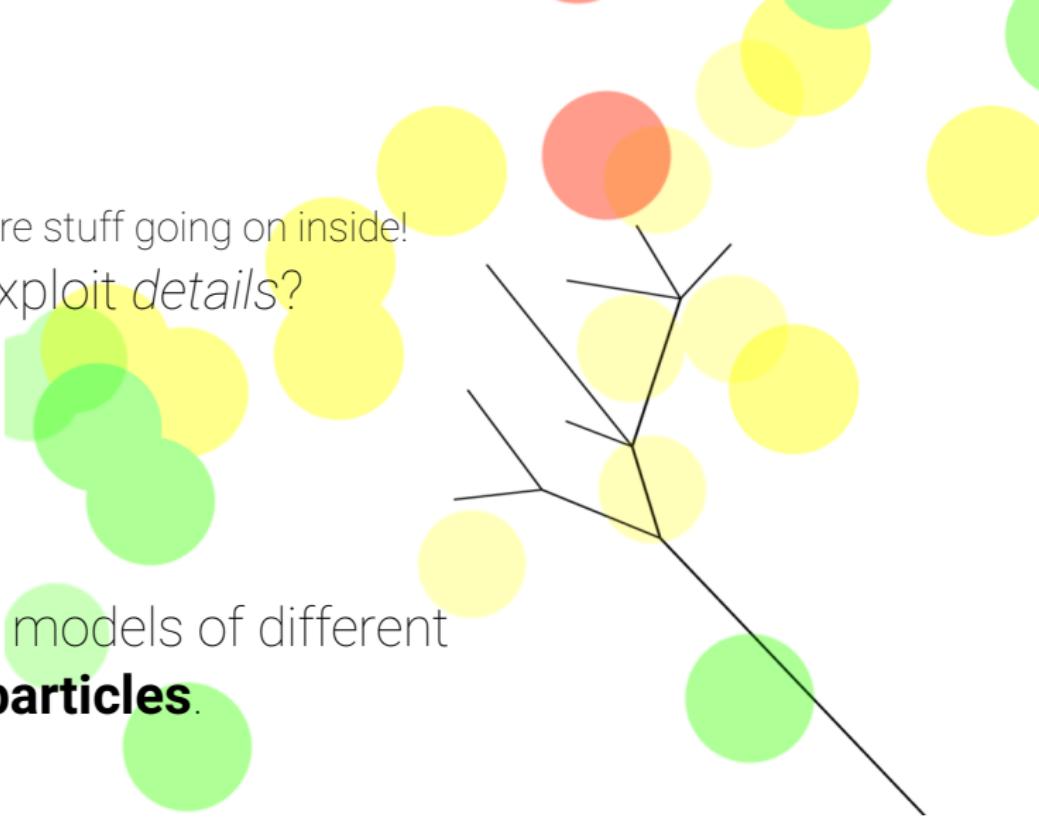
<sup>6</sup> Diagram from [Alaeian, 2014]

# Motivation



**Current models lump all events into showers or tracks.**

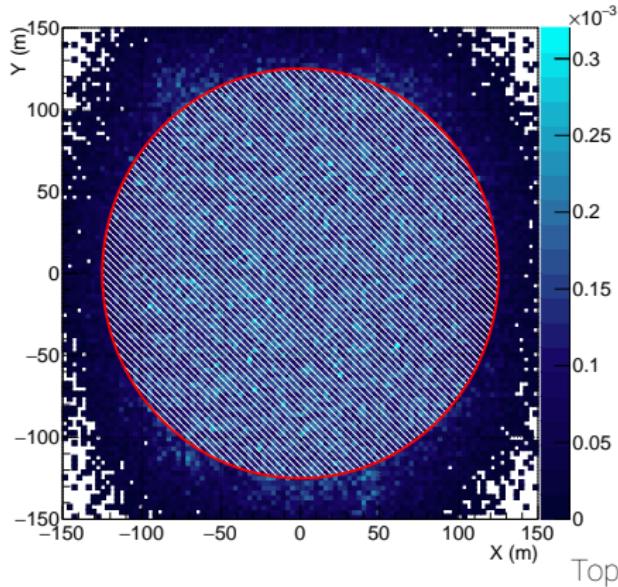
- Only large scale differences used
- Simplified energy scaling



There is more stuff going on inside!  
Can we exploit *details*?

→ Create models of different  
**product particles**.

# Procedure



view of simulated interaction vertices and  
ORCA outline.

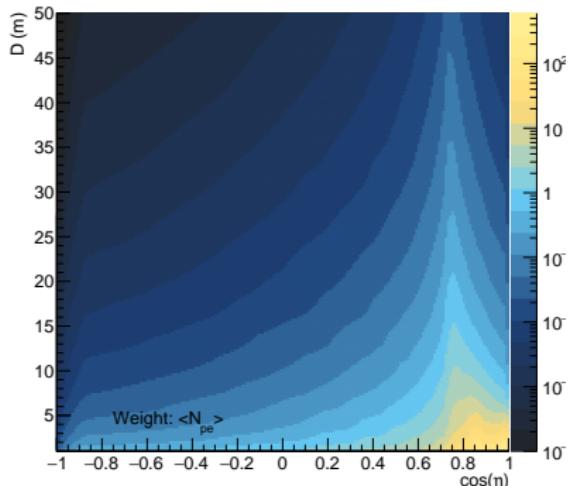
Start off with simulated sample  
of  $\nu$ -events in ORCA

Propagator and hit simula-  
tor is GEANT4 based **KM3Sim**

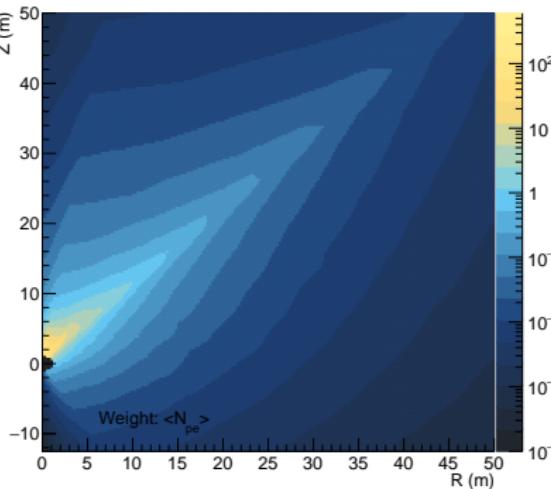
~ 250 M events

# PEPPs Geometry

3 GeV electrons



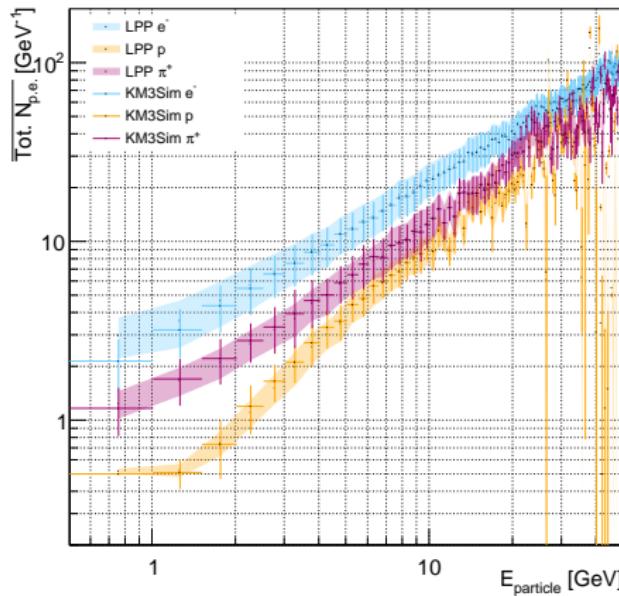
Native coordinates



Cartesian transformation

Remember this because only native coordinates will be shown from now on.

# PEPP Monte Carlo Simulations

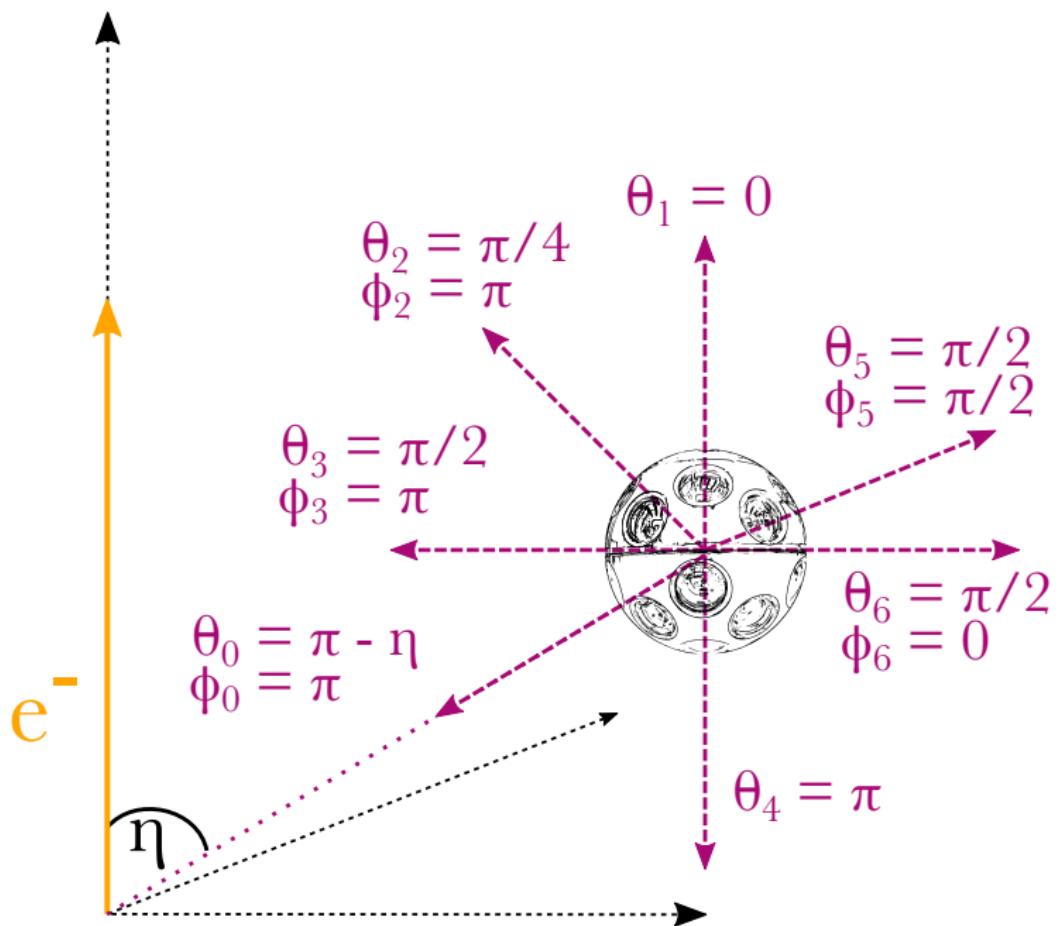


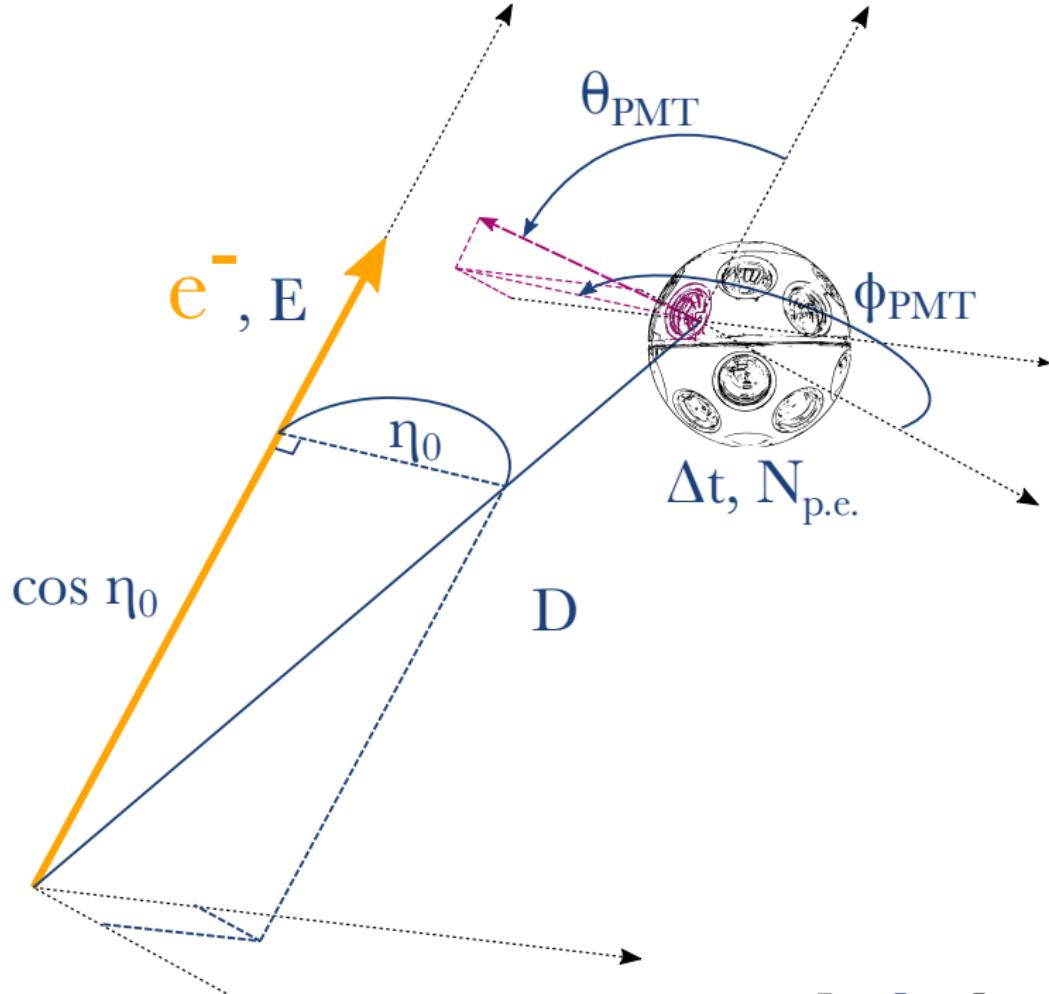
Comparison of KM3Sim and PEPP MC, error bars and bands are for  $\frac{1}{10}\sigma$ .

Excellent match for

$$\overline{TotN_{p.e.}} \equiv \sum^{N_{\text{events}}} N_{p.e.}$$

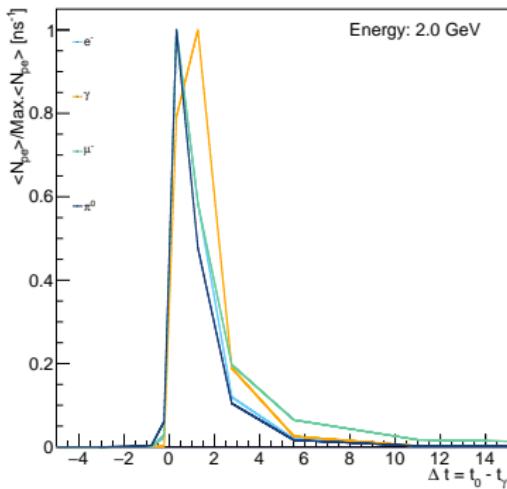
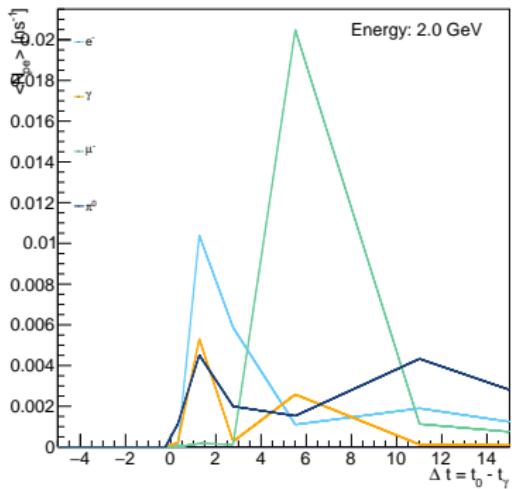
**PEPPs agree with original Monte Carlo simulation**





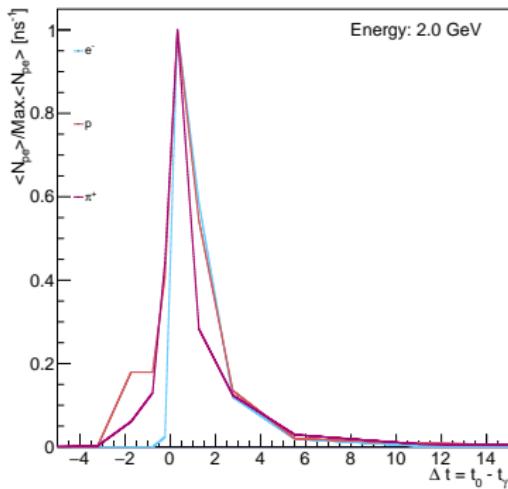
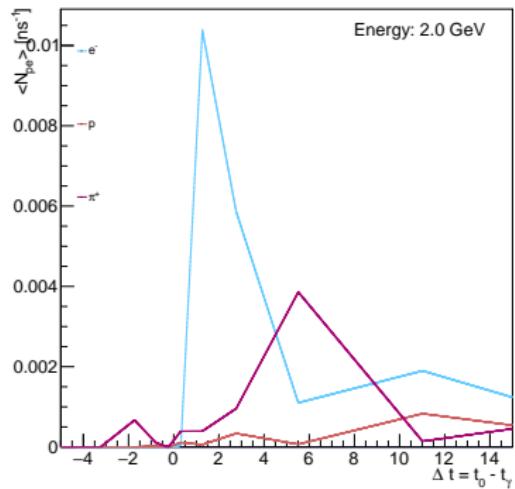
# PEPPs Time Arrival

Discerning power in time dependence



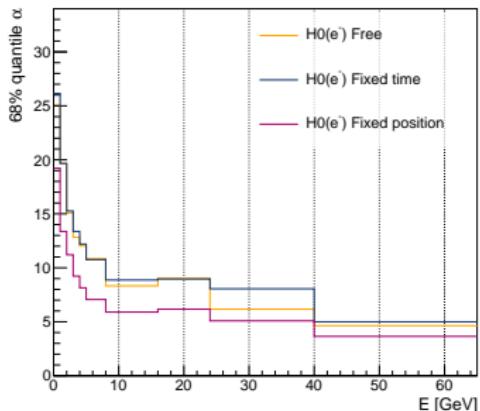
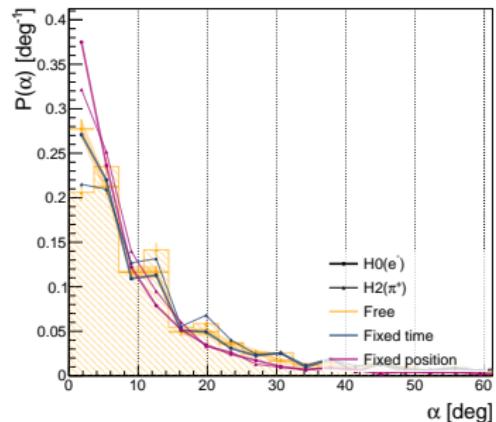
# PEPPs Time Arrival

Discerning power in time dependence



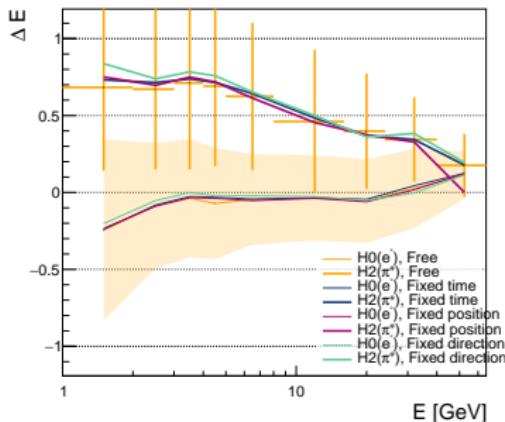
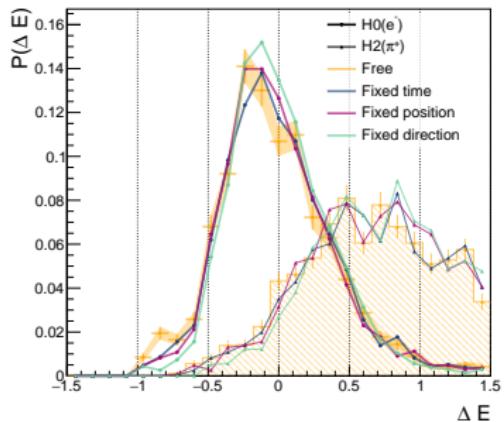
# Reconstruction Single Shower

Direction electron



# Reconstruction Single Shower

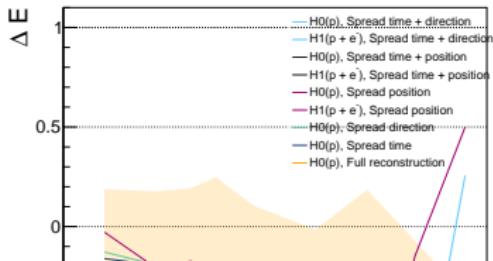
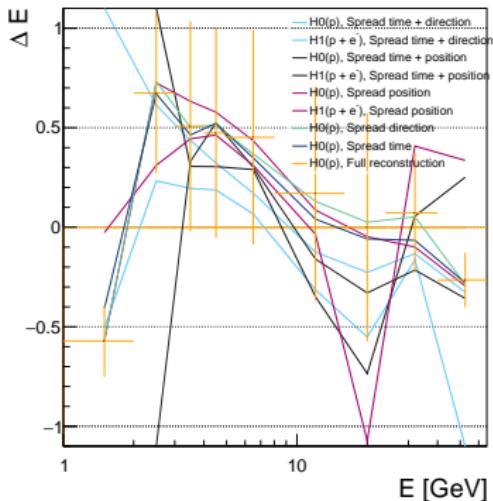
Energy electron



# Reconstruction Neutrino Event

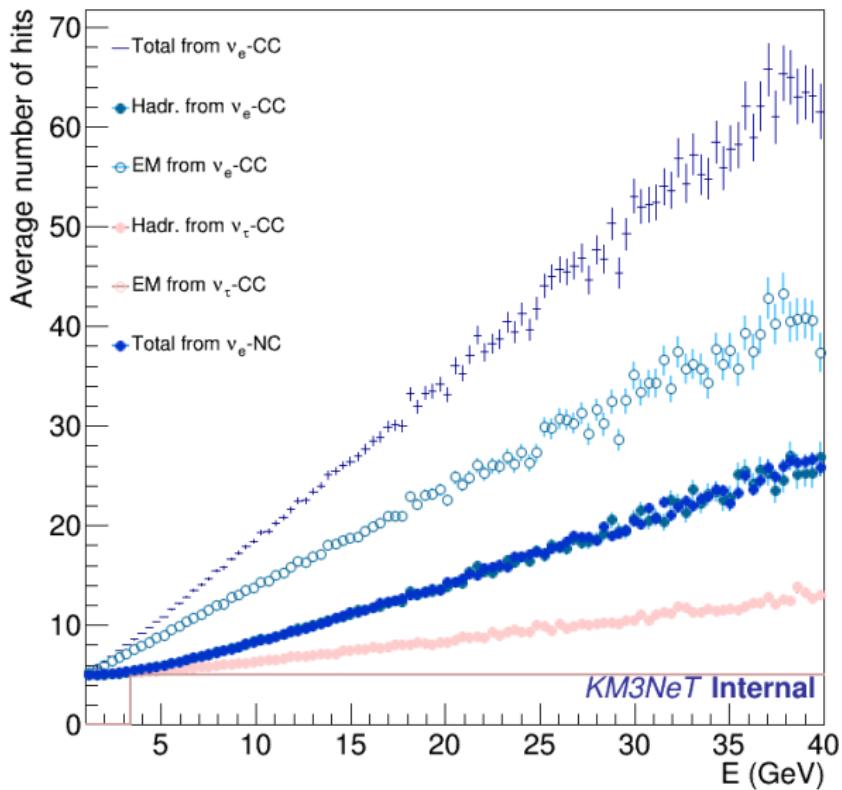
# Reconstruction Neutrino Event

$\Delta E$  for  $\nu_e$ -CC ( $\nu_e$ -NC) above (below)



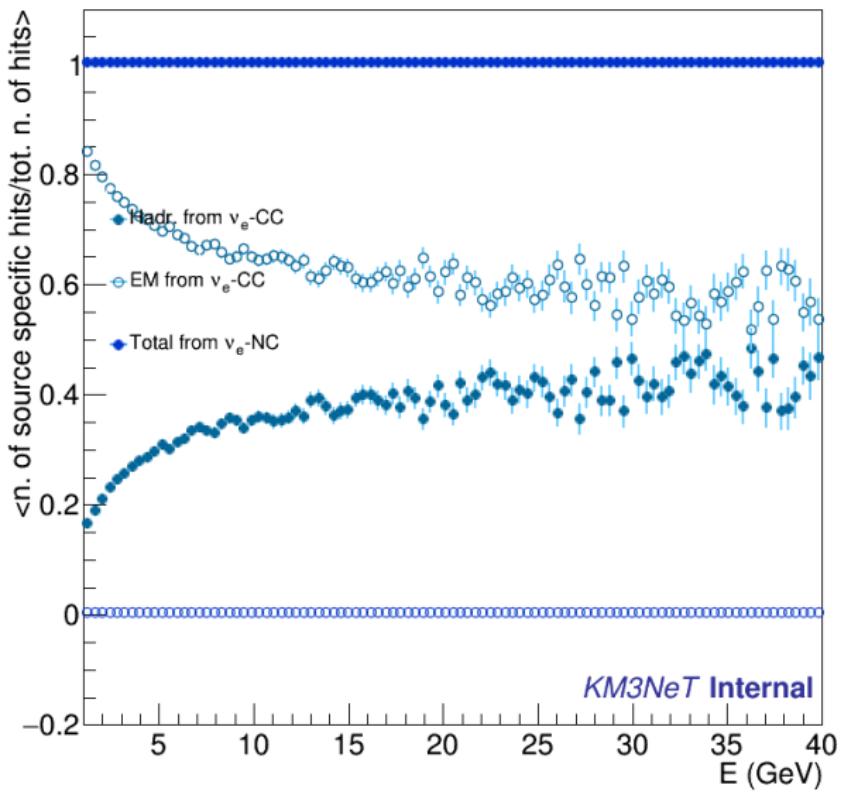
# Secondaries

Number of EM and Hadronic related hits

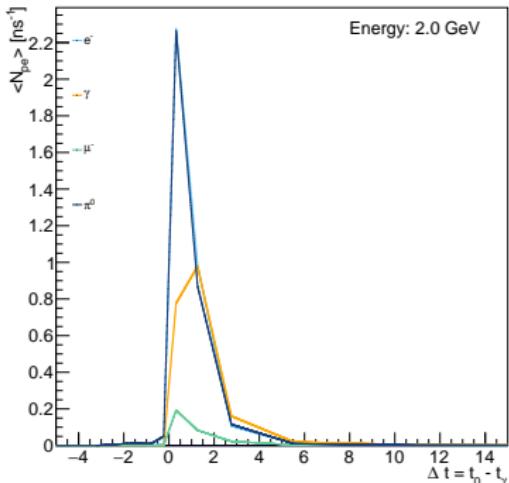


# Secondaries

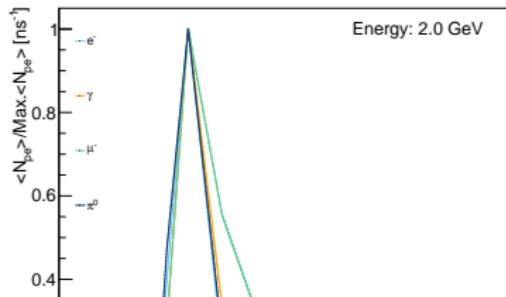
Event dependent hit yield



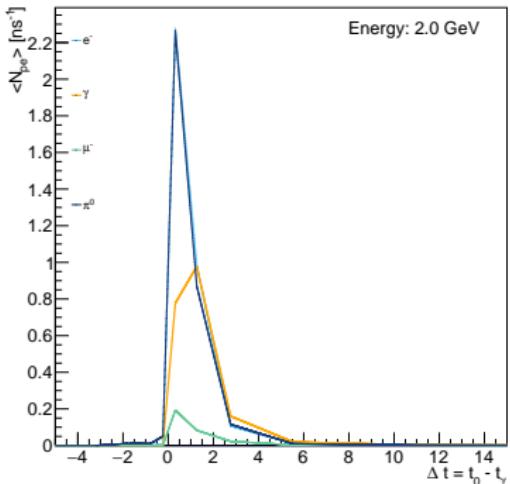
# PDFs Time Arrival



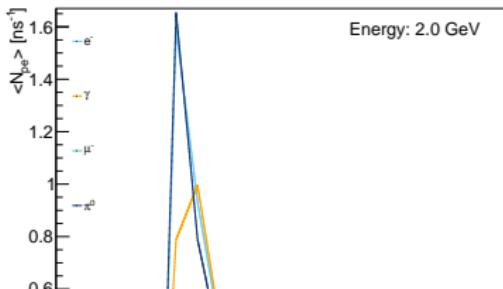
Normalised view



# PDFs Time Arrival

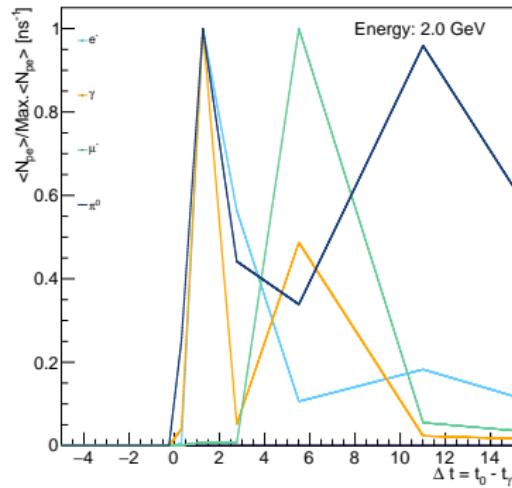
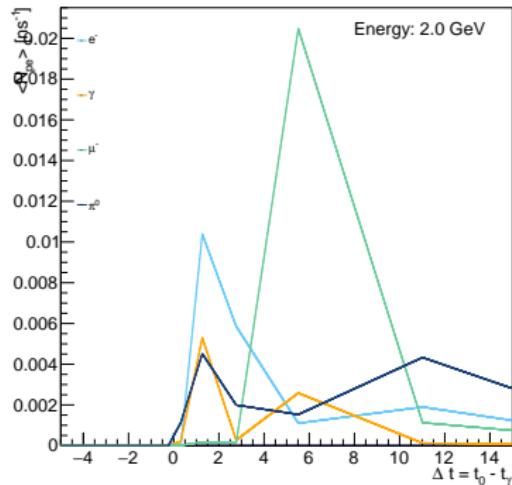


Angle dependence



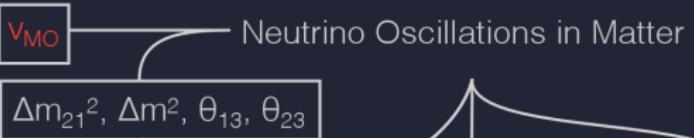
# PDFs Time Arrival

Discerning power in time dependence





Event level  
(simulation level)



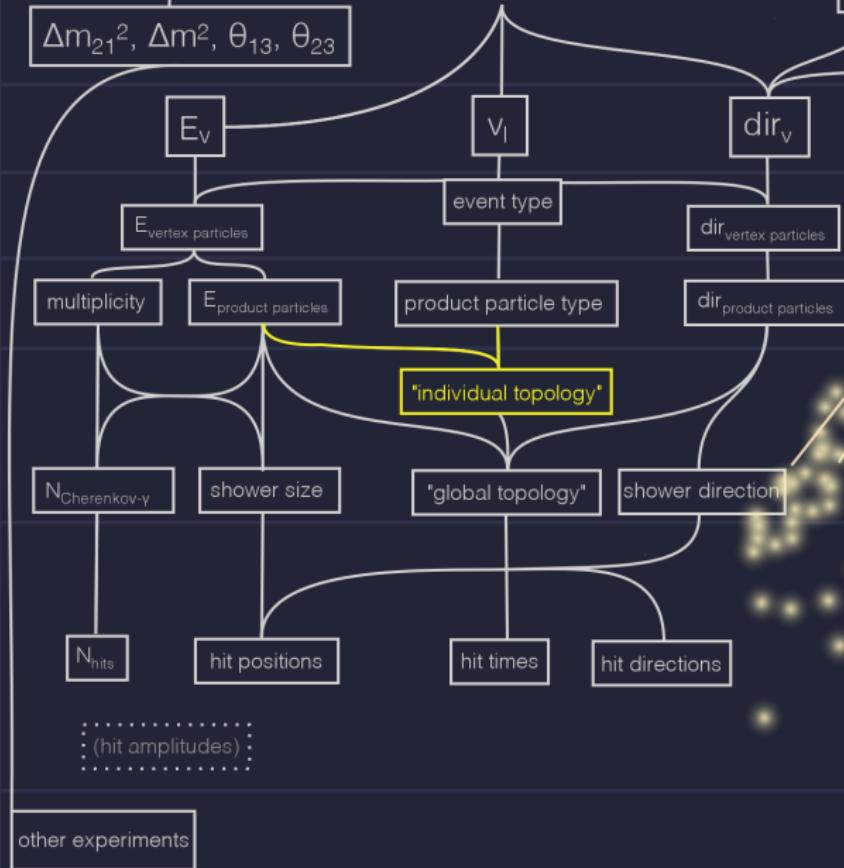
neutrino  
(input)

interaction  
(generator)

product  
(generator)

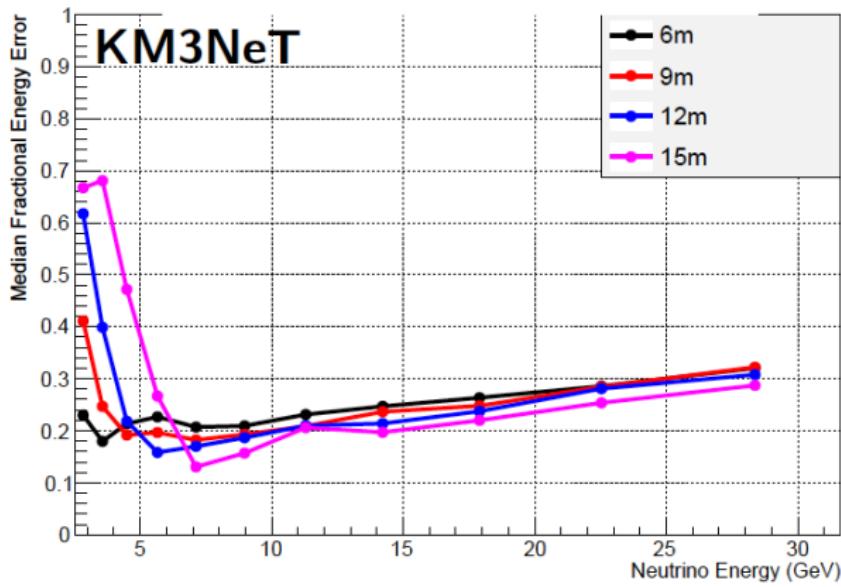
propagation  
(KM3Sim)

detector  
(JTE)



# Motivation

## Orca Energy resolution

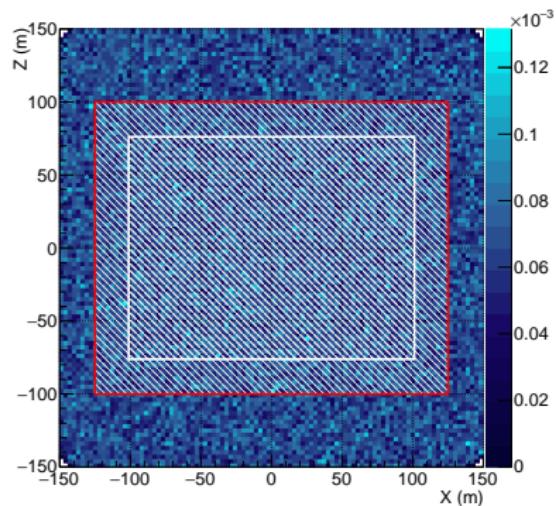


7

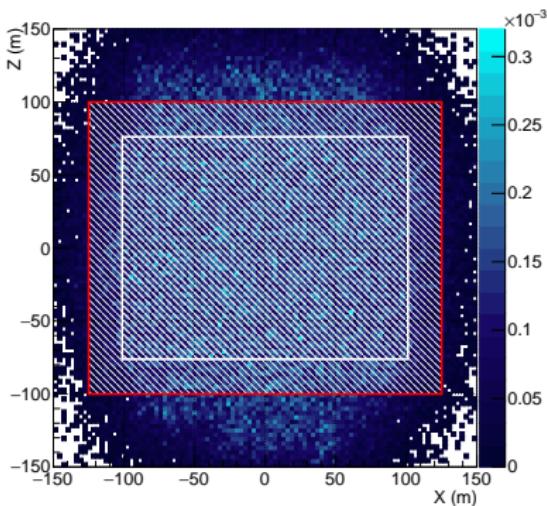
# Reconstructions

## Data cuts

1. > 4 hits
2. Within inner half volume of ORCA

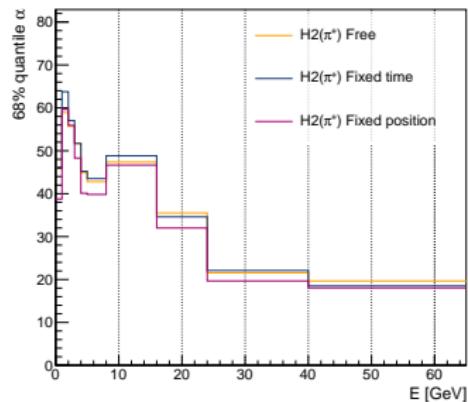


Before hit number cut



After hit number cut

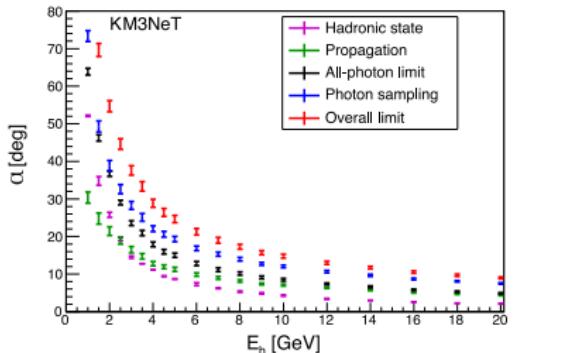
# Reconstruction Single Shower



Single hadronic shower  $\equiv$  all  $\nu$ -interaction secondaries minus leading lepton.

Angle difference

$$\alpha \equiv \cos^{-1} \hat{p}_{\text{true}} \cdot \hat{p}_{\text{reco.}}, \text{ where}$$
$$\hat{p} \equiv \frac{1}{E_{\text{tot}}} \sum_i^N \hat{p}_i E_i.$$

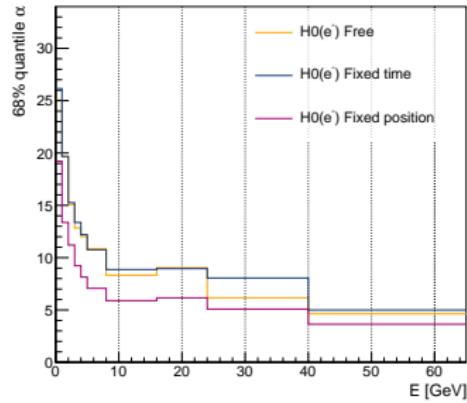


Intrinsic limit paper<sup>8</sup> shows **best possible resolution** of ORCA.

Not directly comparable due to hit cut, but gives an idea.

<sup>8</sup>[Adrian-Martinez et al., 2017]

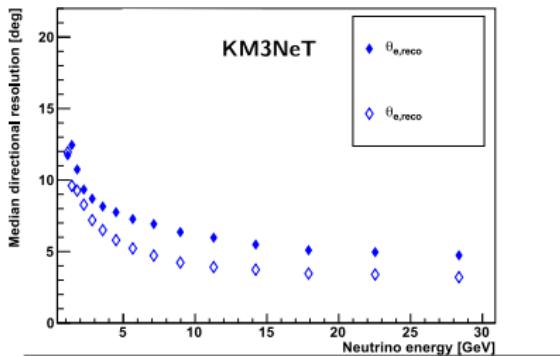
# Reconstruction Single Shower



Single electron.

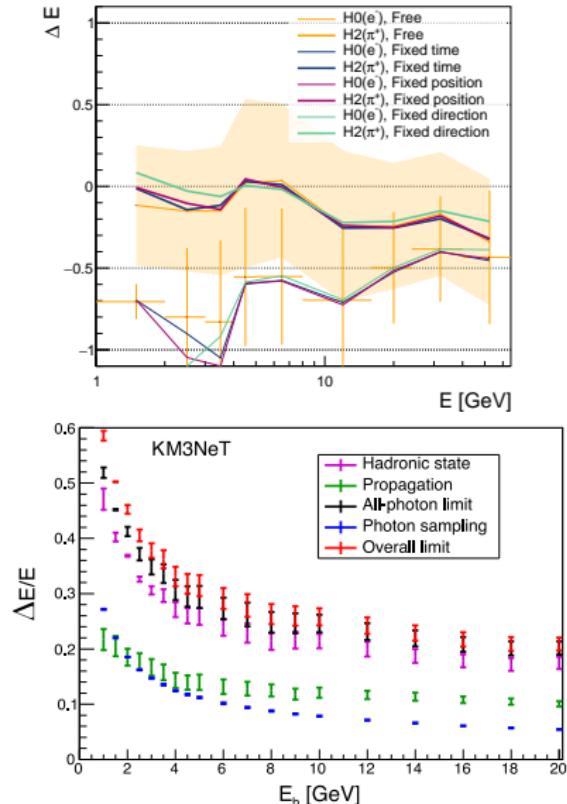
Angle difference

$$\alpha \equiv \cos^{-1} \hat{p}_{\text{true}} \cdot \hat{p}_{\text{reco}}, \text{ where } \hat{p} \equiv \frac{1}{E_{\text{tot}}} \sum_i^N \hat{p}_i E_i.$$



Resolution reproduces that of LOI

# Reconstruction Single Shower

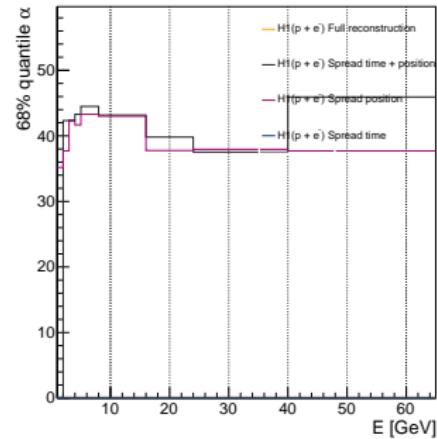


Single hadronic shower energy difference  $\Delta E \equiv \frac{E_{h,\text{true}} - E_{h,\text{reco}}}{E_{h,\text{true}}}$ ,  
where  $E_h \equiv E_\nu - E_{\text{lep.}}$ .

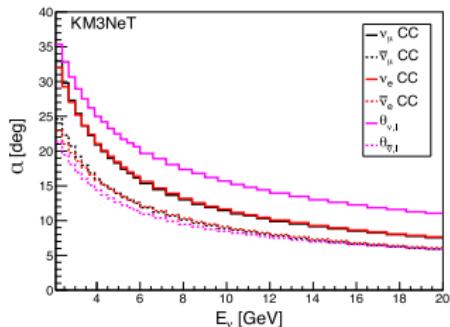
Energy difference resolution for low energies, close to intrinsic resolution for >4 hits.

$\pi_+$  best at reconstructing hadronic showers, supposedly due to high presence of  $\pi_{+/-}$  in hadronic showers.

# Reconstruction Neutrino Event



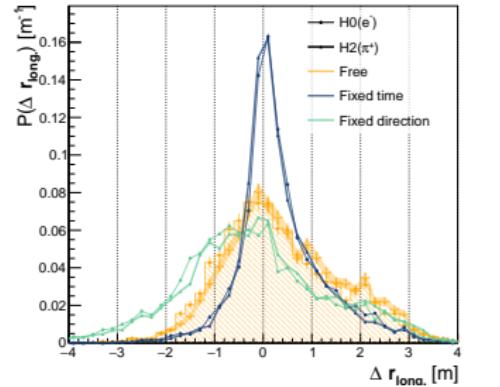
$\nu_e$ -charged current angle difference.



Assuming 3 m position resolution and 5 ns timing resolution

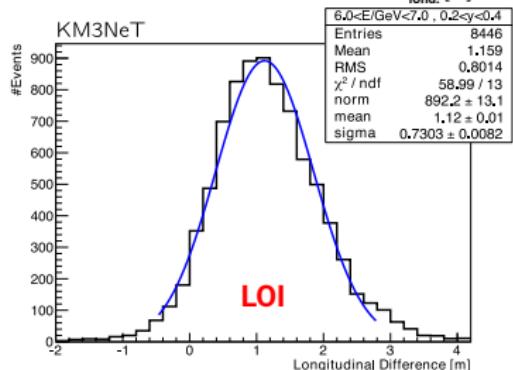
Close to intrinsic limits at low energies.

# Reconstruction Neutrino Event



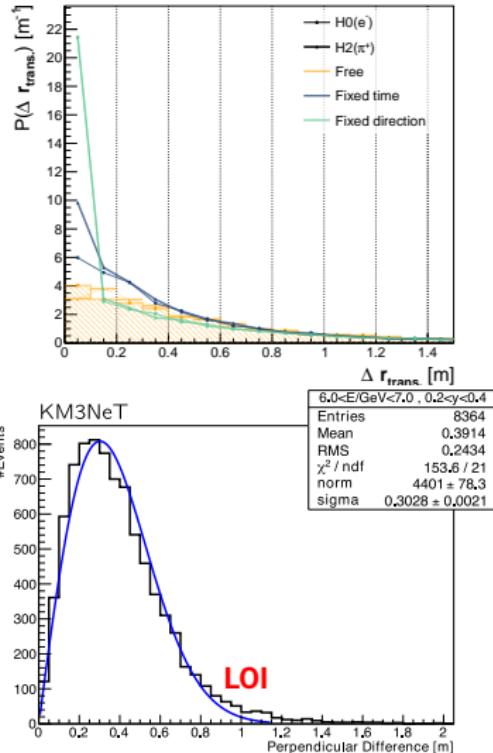
Position of single hadronic shower, identical to  $\nu$ -NC.

Reproduces resolutions for  $\nu$ -CC as reported in LOI, but better resolution in other parameters accentuates this resolution!



Distance between shower maximum and vertex folded into model  
→ naturally centres at zero.

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