KM3Net Outing 01/06/2018



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- High energy neutrinos come from two different sources.
 - O Cosmic rays interactions in the Earth -> BACKGROUND
 - O Extraterrestrial sources -> SIGNAL





- Background should be isotropic.
- Signal should cluster in a particular region of the sky.
 - O Reconstruct the direction of the incoming neutrino.





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- Need other variables to discriminate between SIGNAL & BACKGROUND.
 - O Reconstruct the energy of the incoming neutrino.
 - O Very handy for SHOWERS because they are contained in the detector.





• Reconstructing energy and direction we obtain some discrimination power.







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Where is cross section playing a role??

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 v_{μ} cross section Aeff $\mathcal{C}\mathcal{O}(\mathbf{E})$

Differential Cross section.

- It tells us the energy of the outgoing lepton and shower.
- It affects to the energy reconstruction.
 - NN trained with simulated neutrino interactions -> very model dependent!!!
 - Tracks -> How much energy is given to the muon?
 - Shower (NC) -> How much energy is given to the neutrino?

$E_{rec} \propto d\sigma/dy$





Differential Cross section.

$E_{rec} \propto d\sigma/dy$





Quark content:

- It affects to the energy reconstruction.
 - O Neural network trained with simulated neutrino interactions -> very model dependent!!!
- It tells us the ID of the outgoing hadrons.
 - O Current simulations only take into account light quark mesons (K, π) in the final state.
 - O At high energies, more exotic mesons (B,D) can contribute.
 - They will immediately decay into leptons? other hadrons?



Quark content:

$E_{rec} \propto \sigma(q)$

$\mathbf{E} = \mathbf{10^6} \; \mathbf{GeV}$

Quark	I c e C u b e	КМЗЛЕТ
u	14 %	15 %
d	43 %	48 %
С	8 %	10 %
S	23 %	25 %
b	10 %	1 %

